

Quantum Entanglement at all distances

The Racah Memorial Lecture
The Racah Institute of Physics
June 21, 2021



האוניברסיטה העברית בירושלים
THE HEBREW UNIVERSITY OF JERUSALEM



Subir Sachdev

Talk online: sachdev.physics.harvard.edu



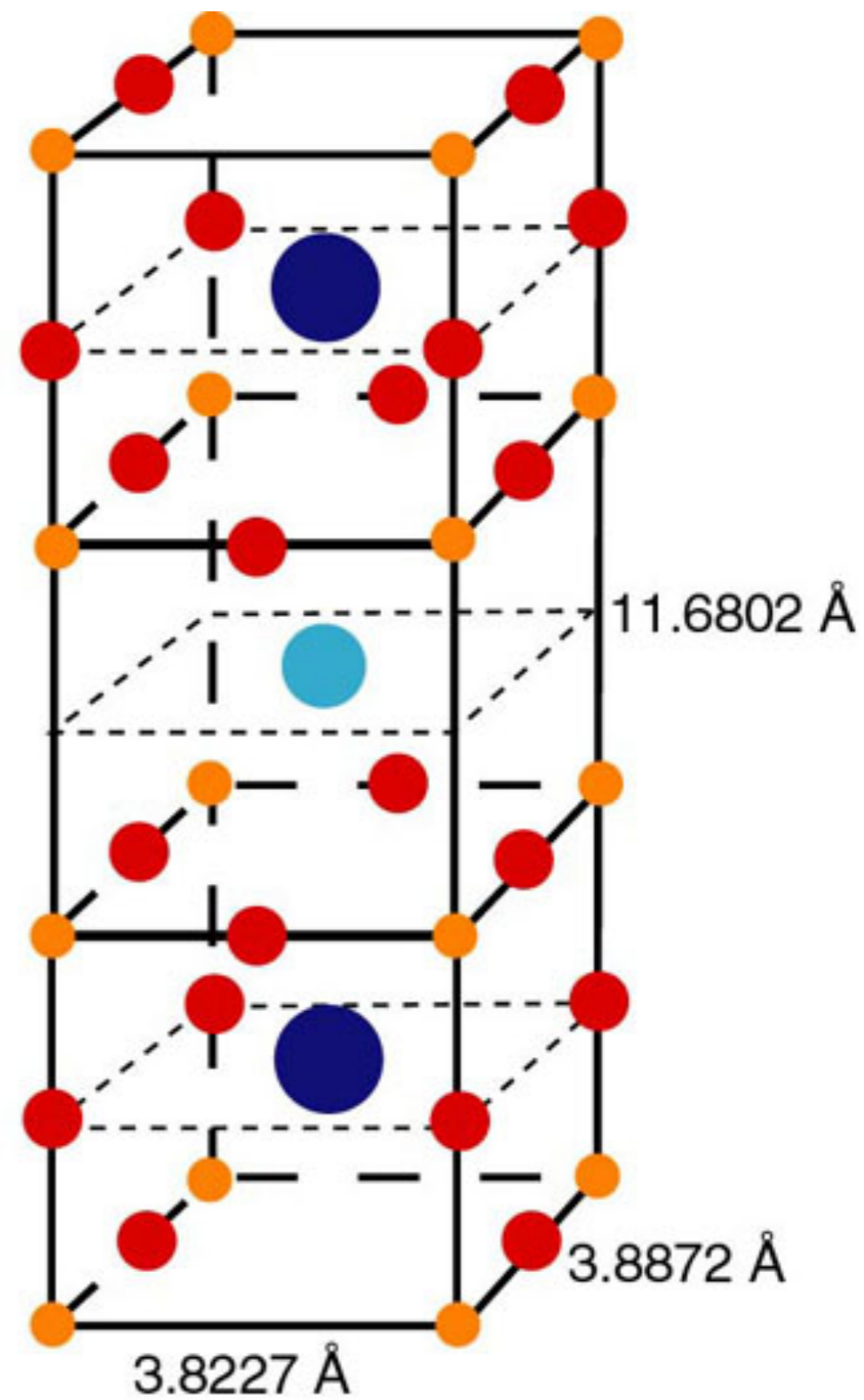
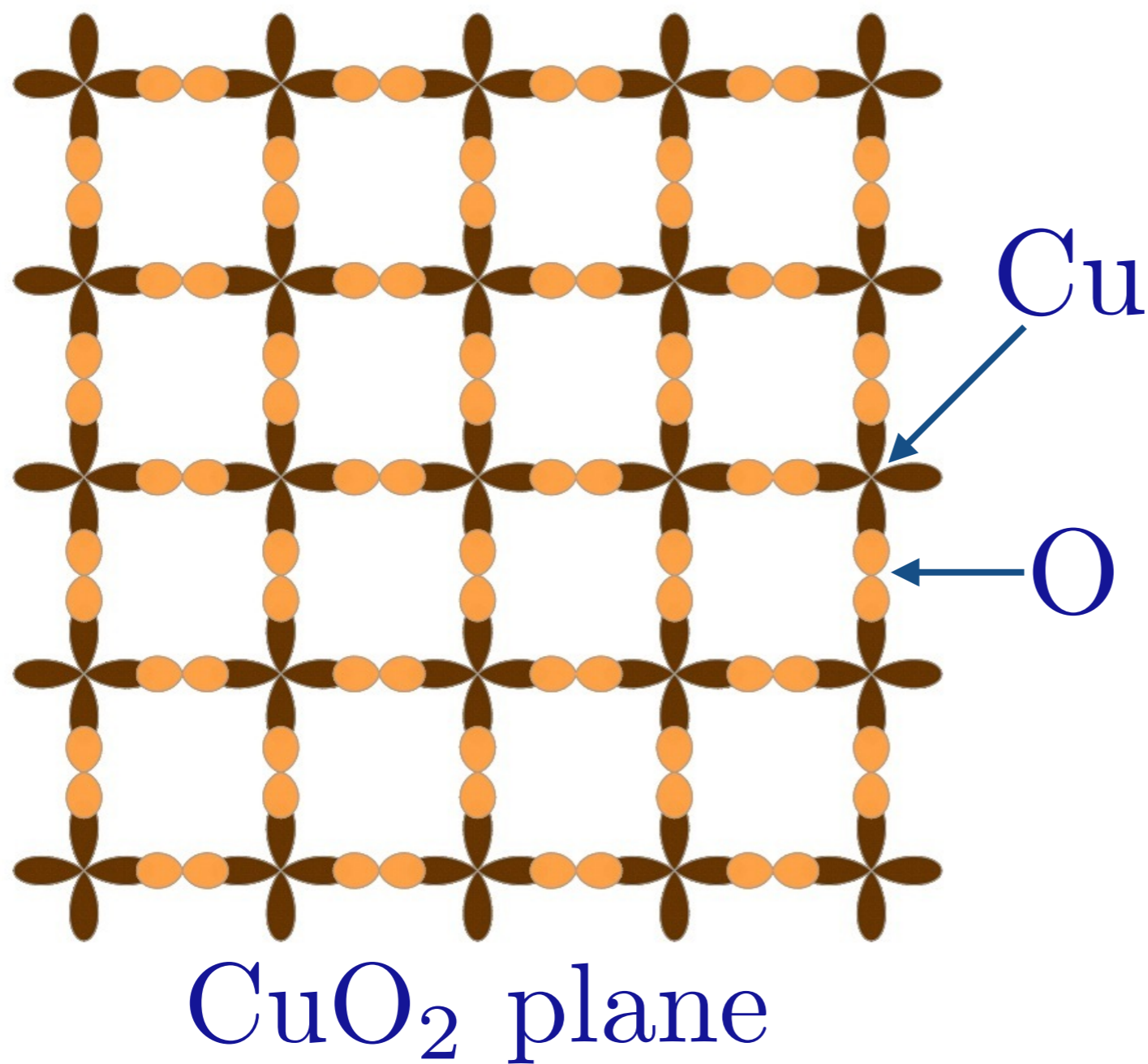
1. Introduction to the cuprates

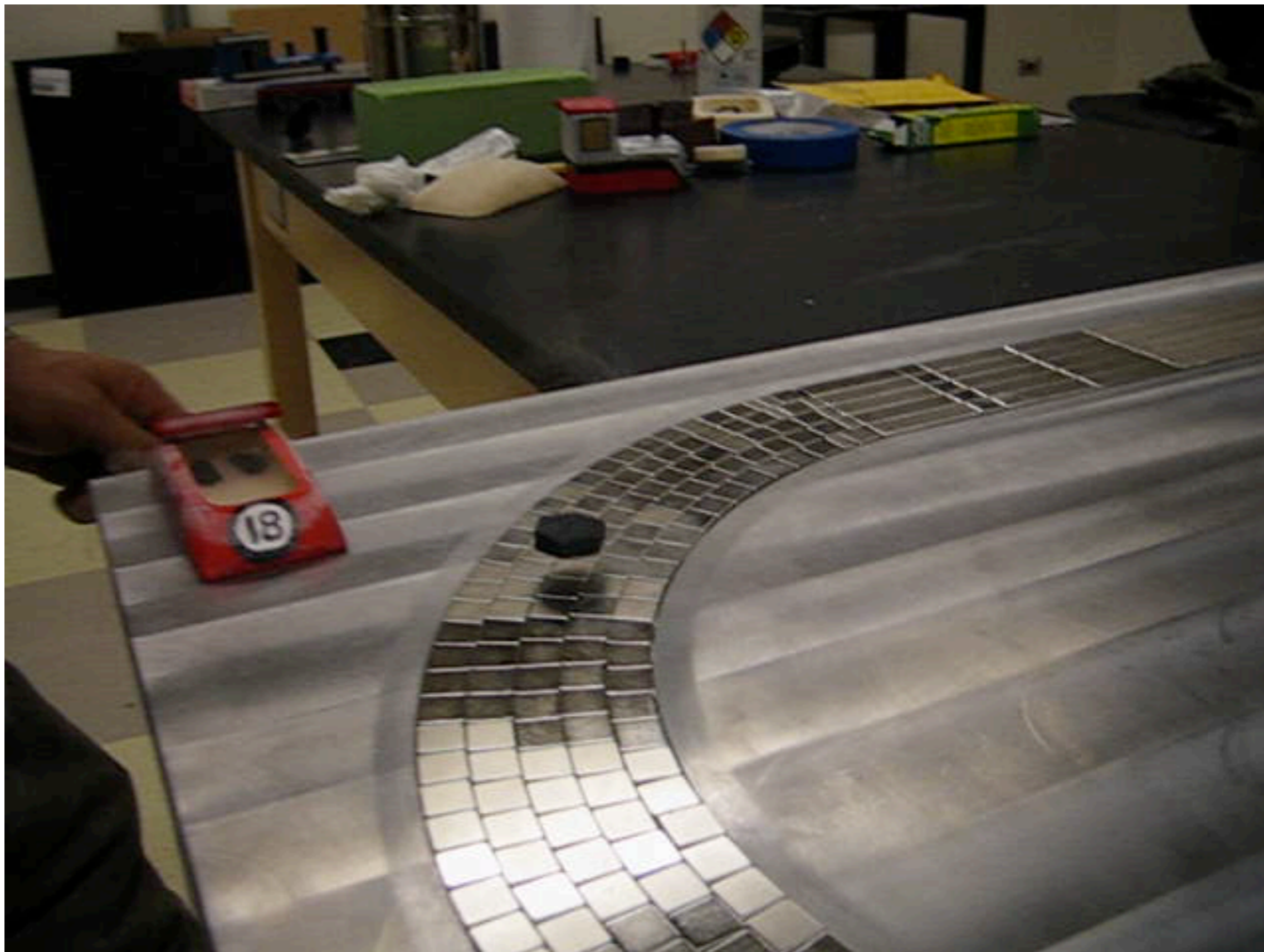
2. Random matrix and SYK models

3. Black holes

4. Back to the cuprates

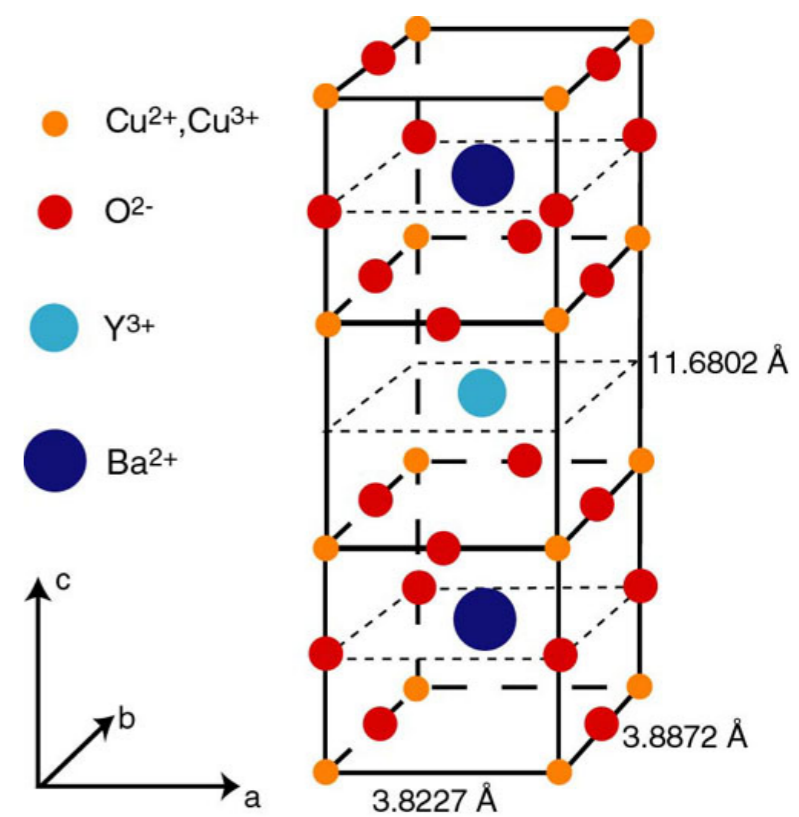
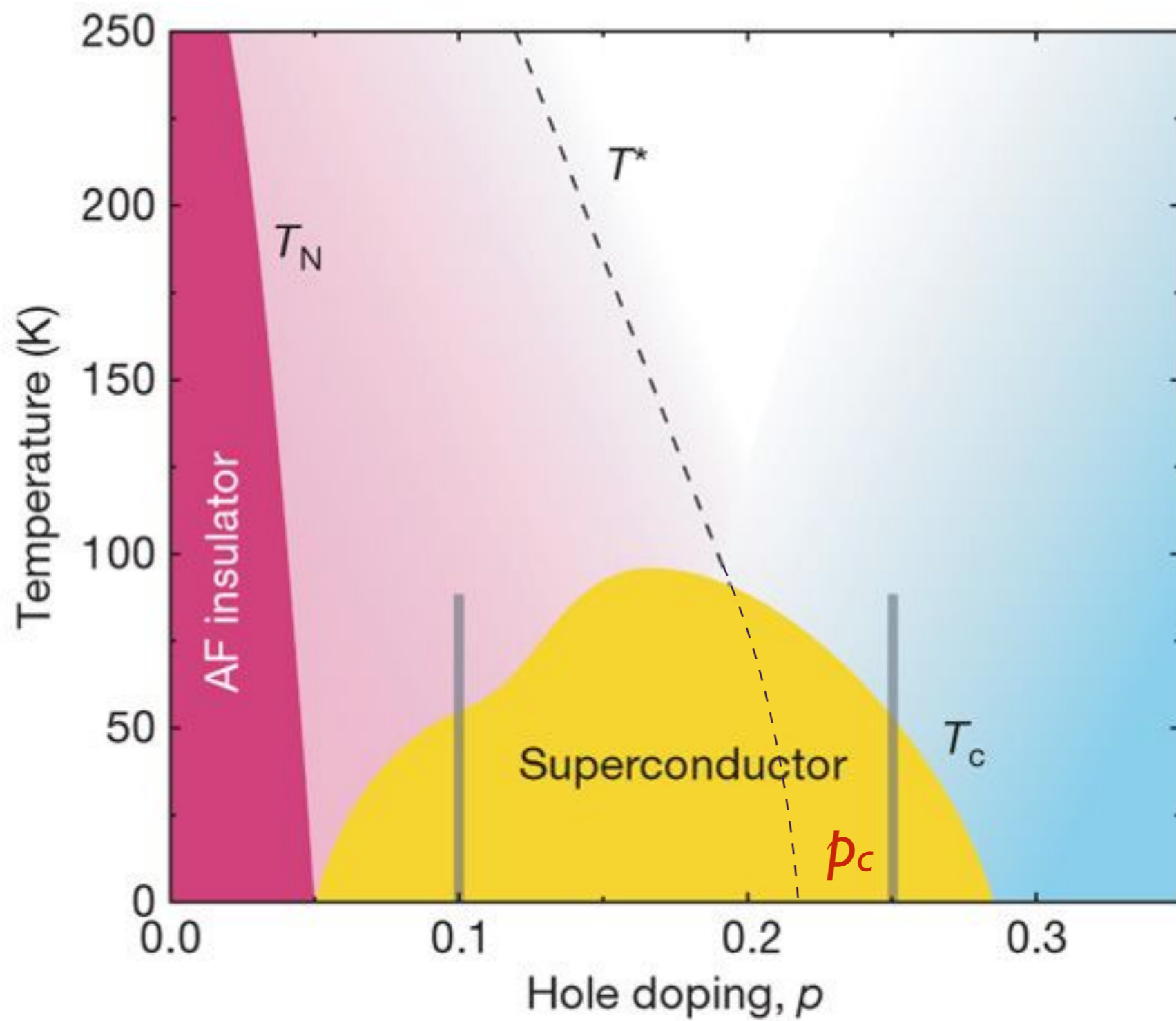
High temperature superconductors

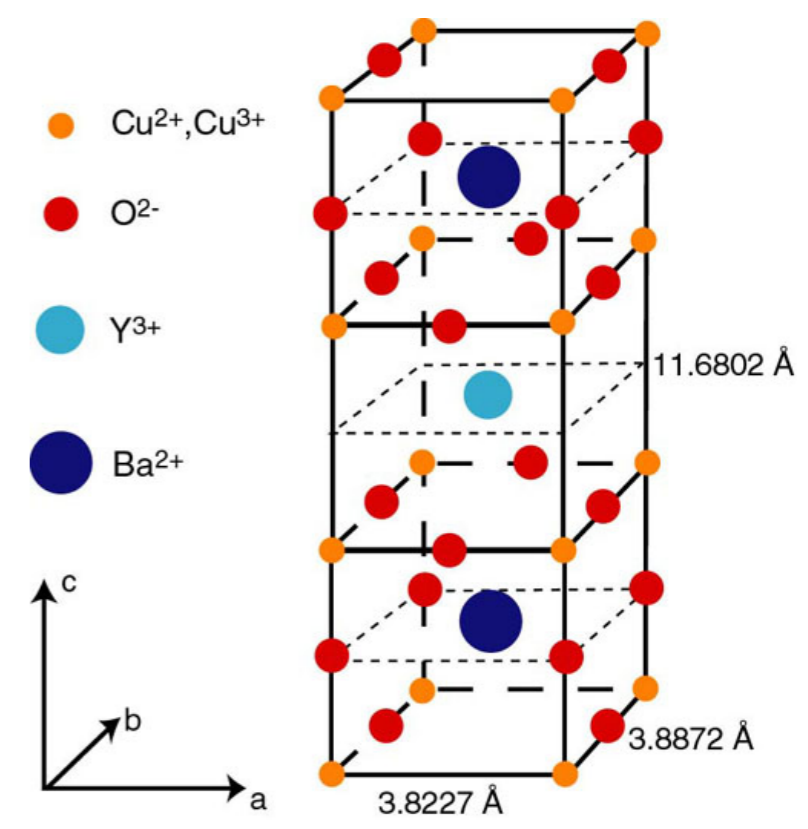
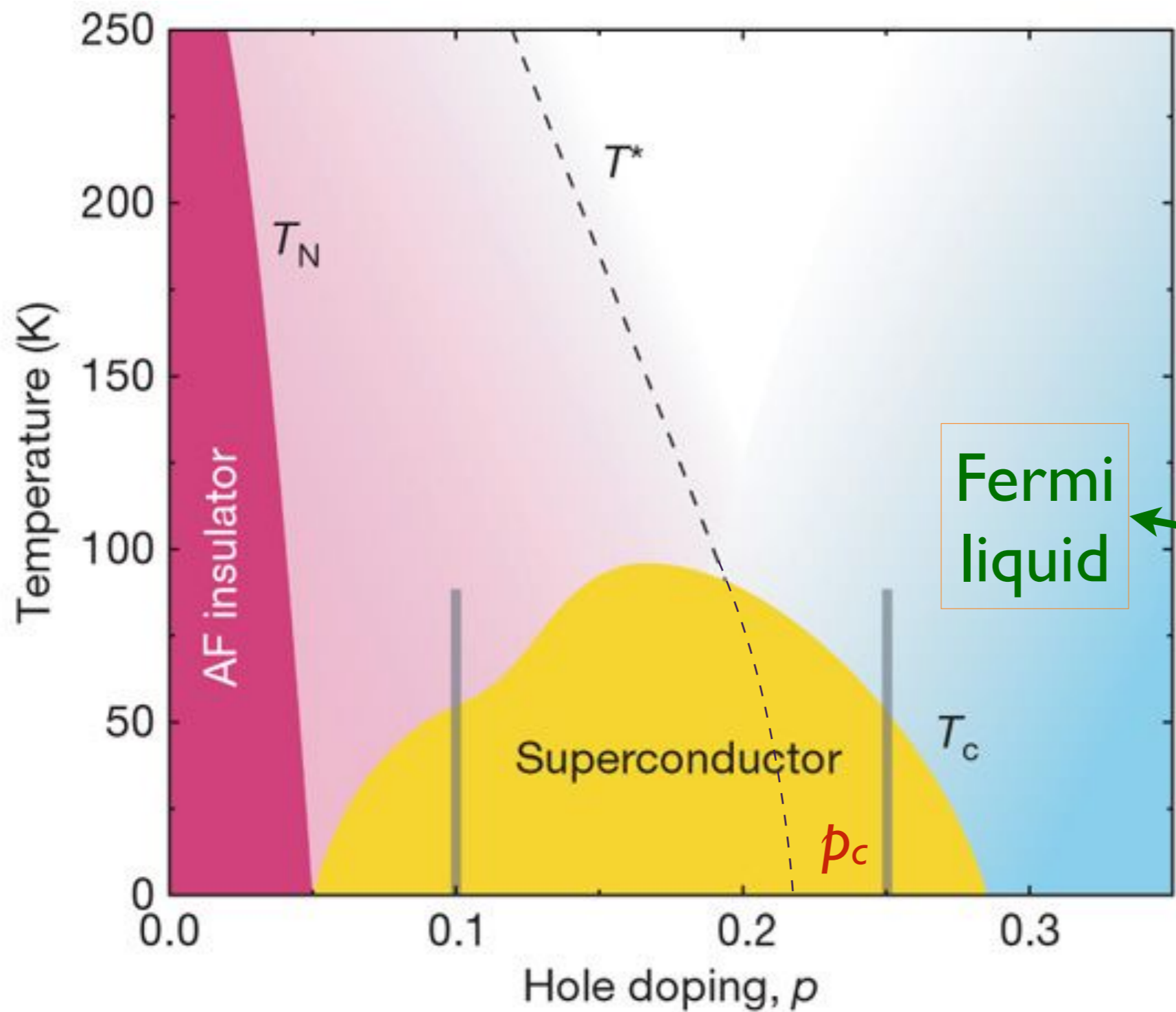




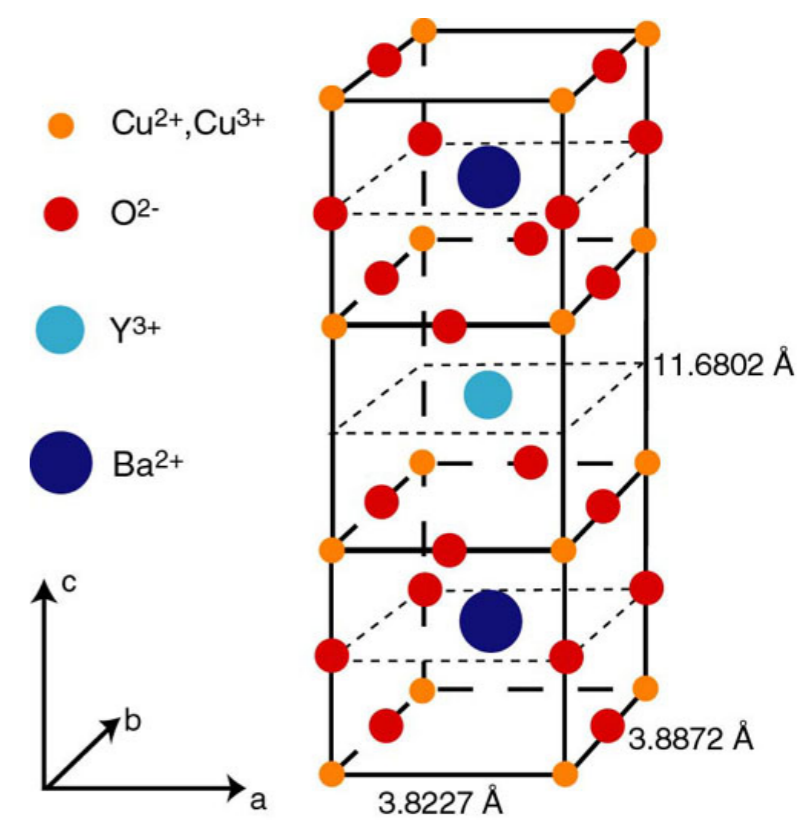
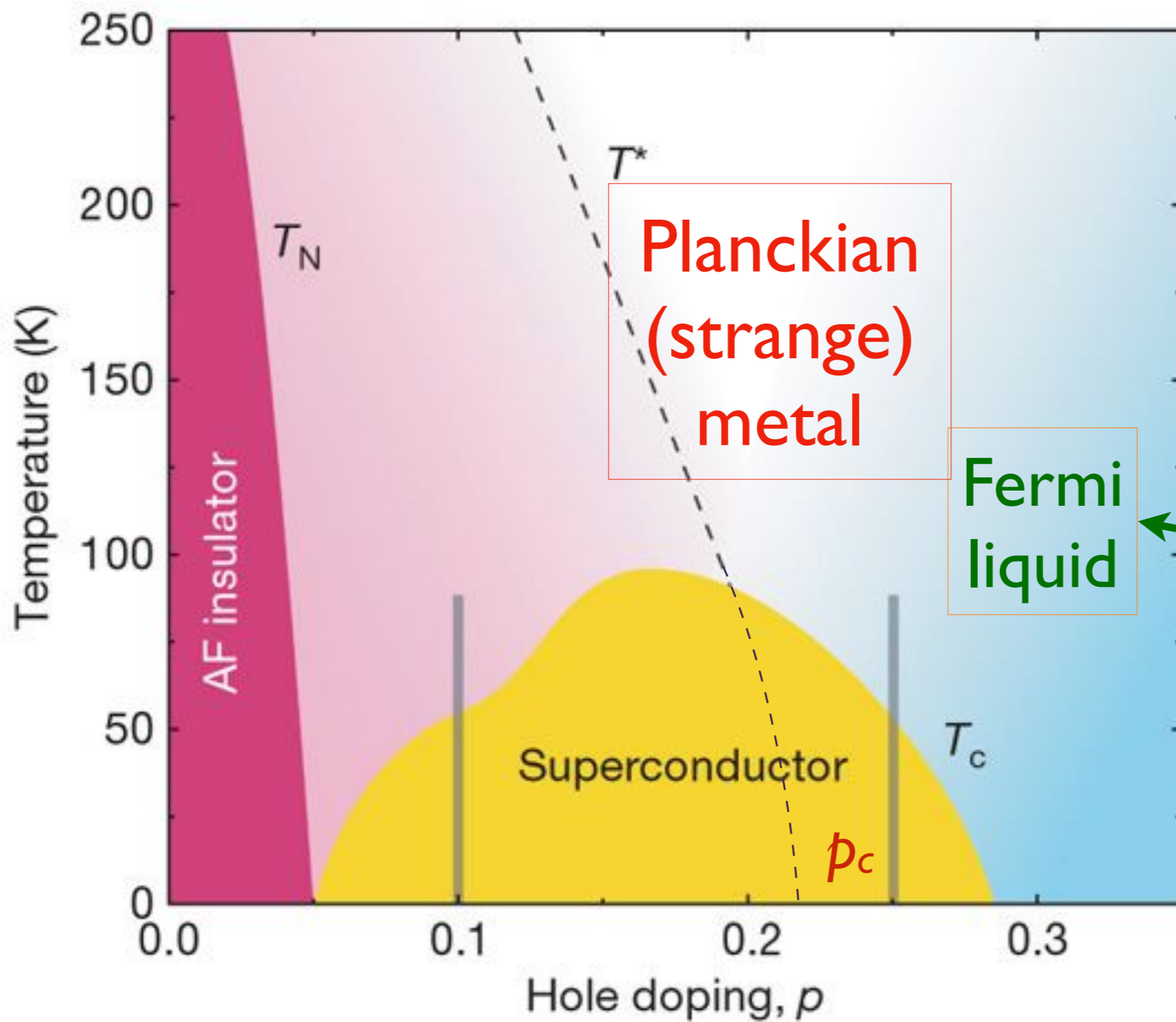
Nd-Fe-B magnets, YBaCuO superconductor

Julian Hetel and Nandini Trivedi, Ohio State University

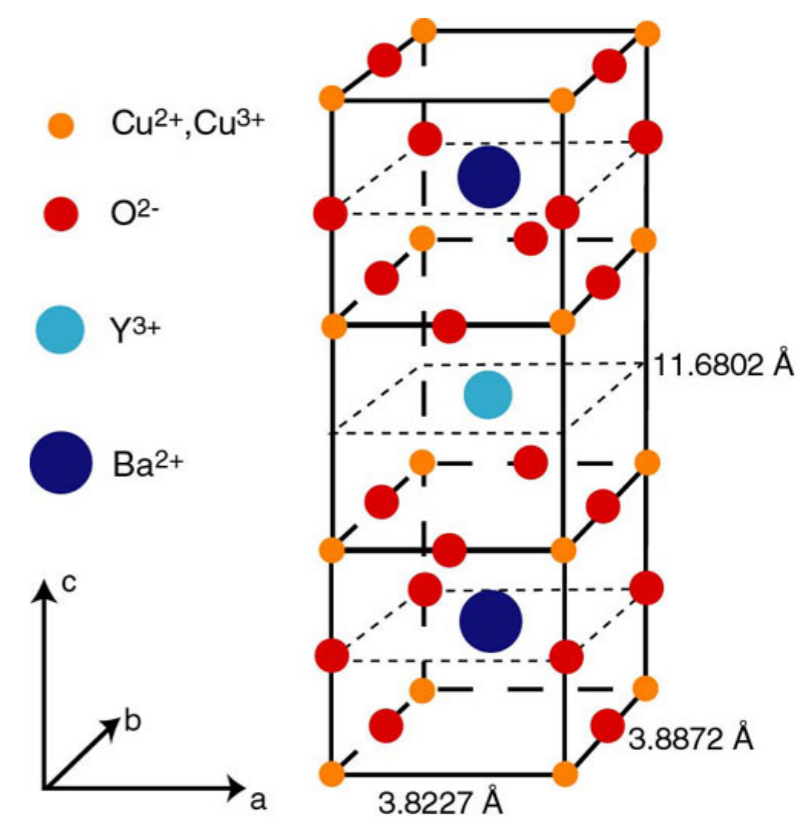
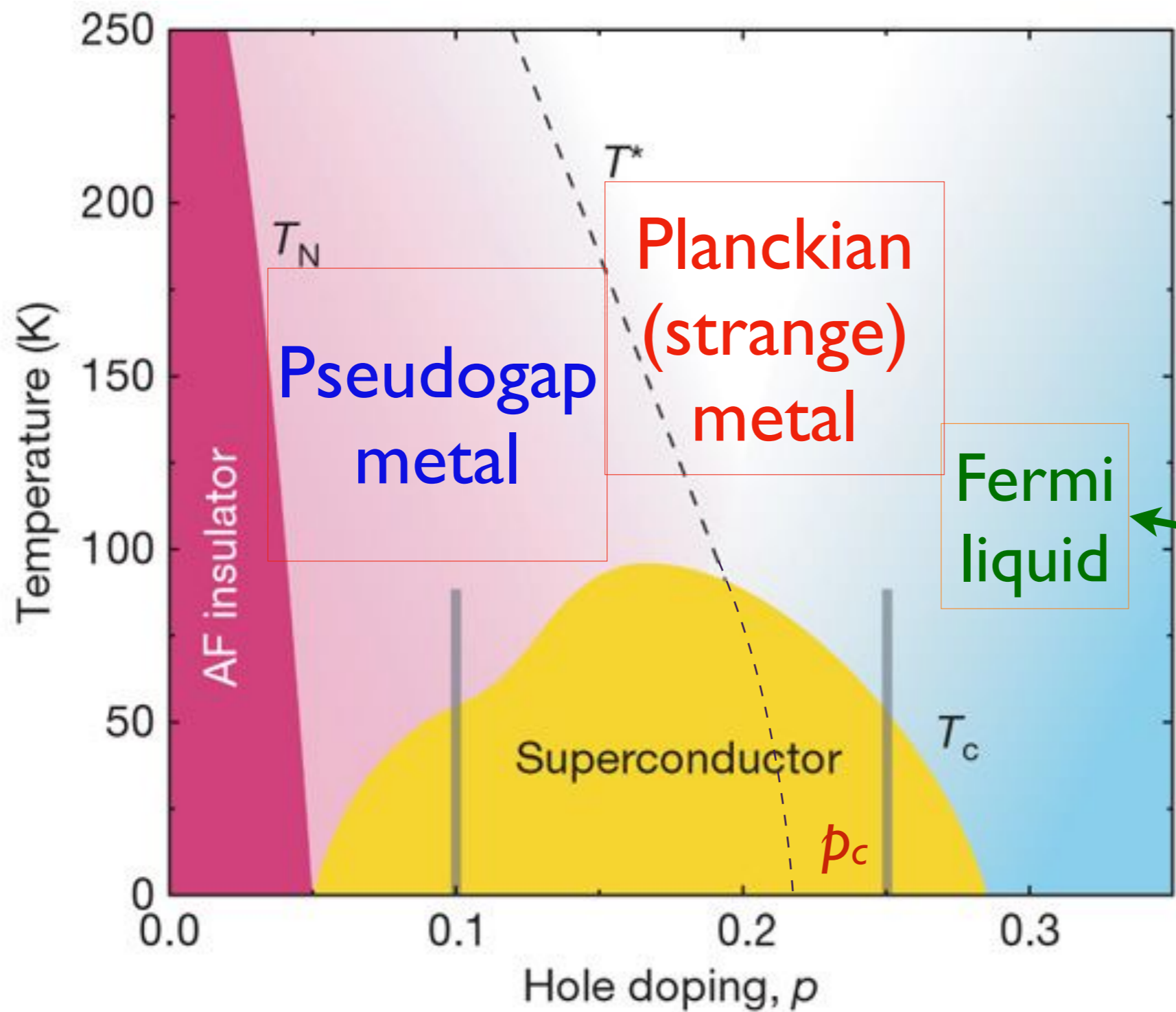




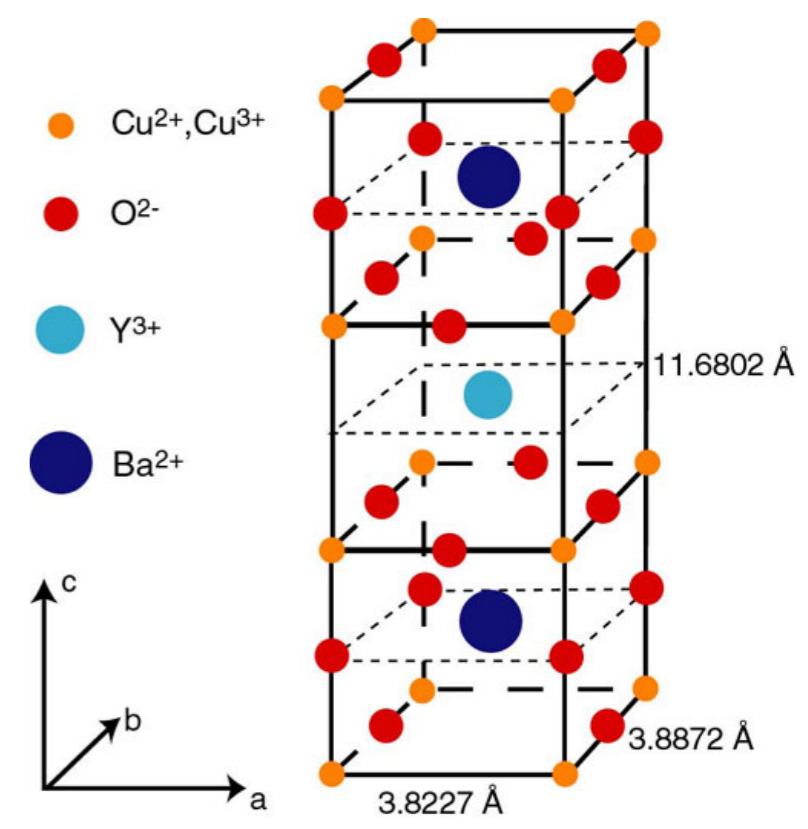
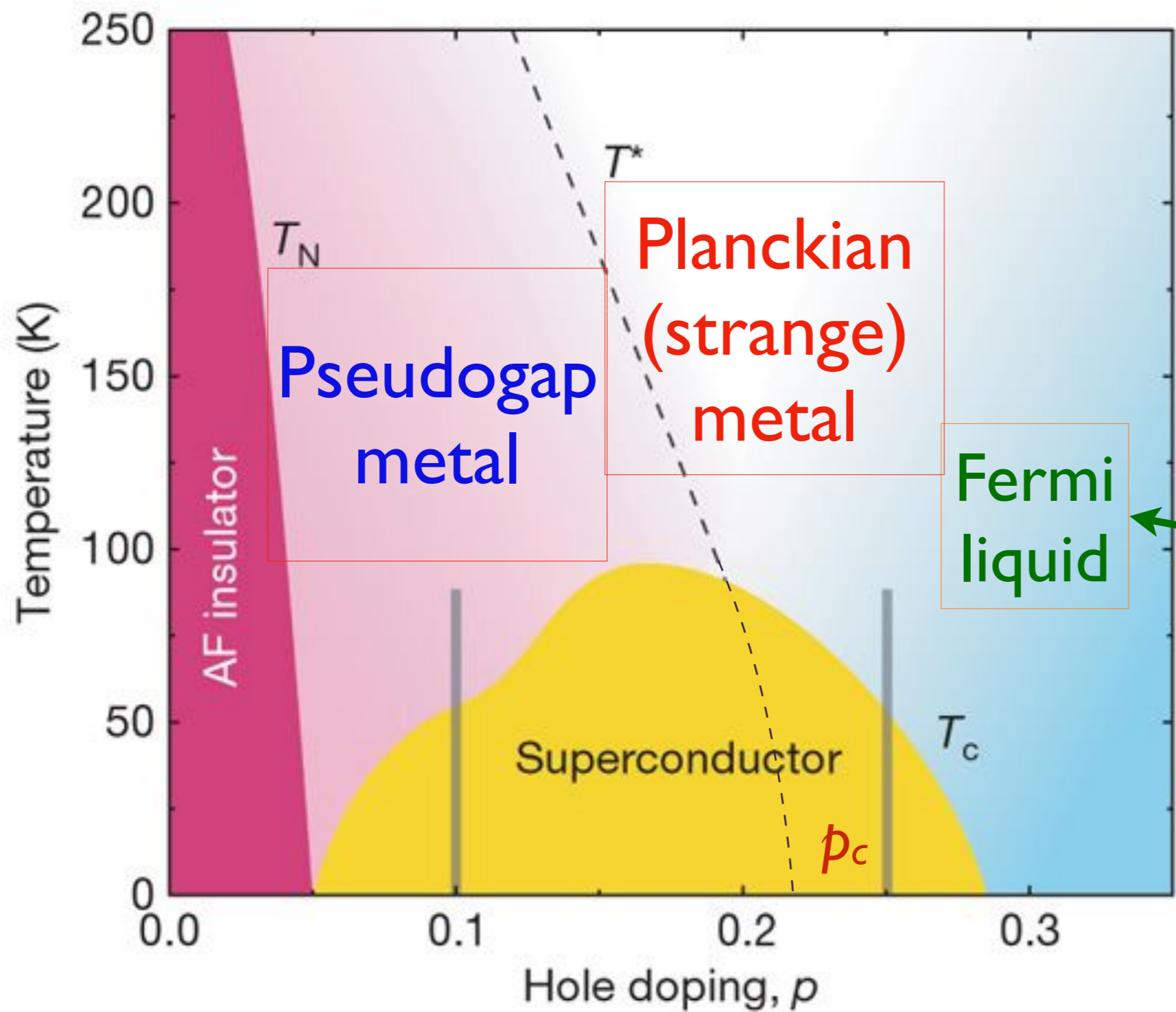
Like ordinary metals



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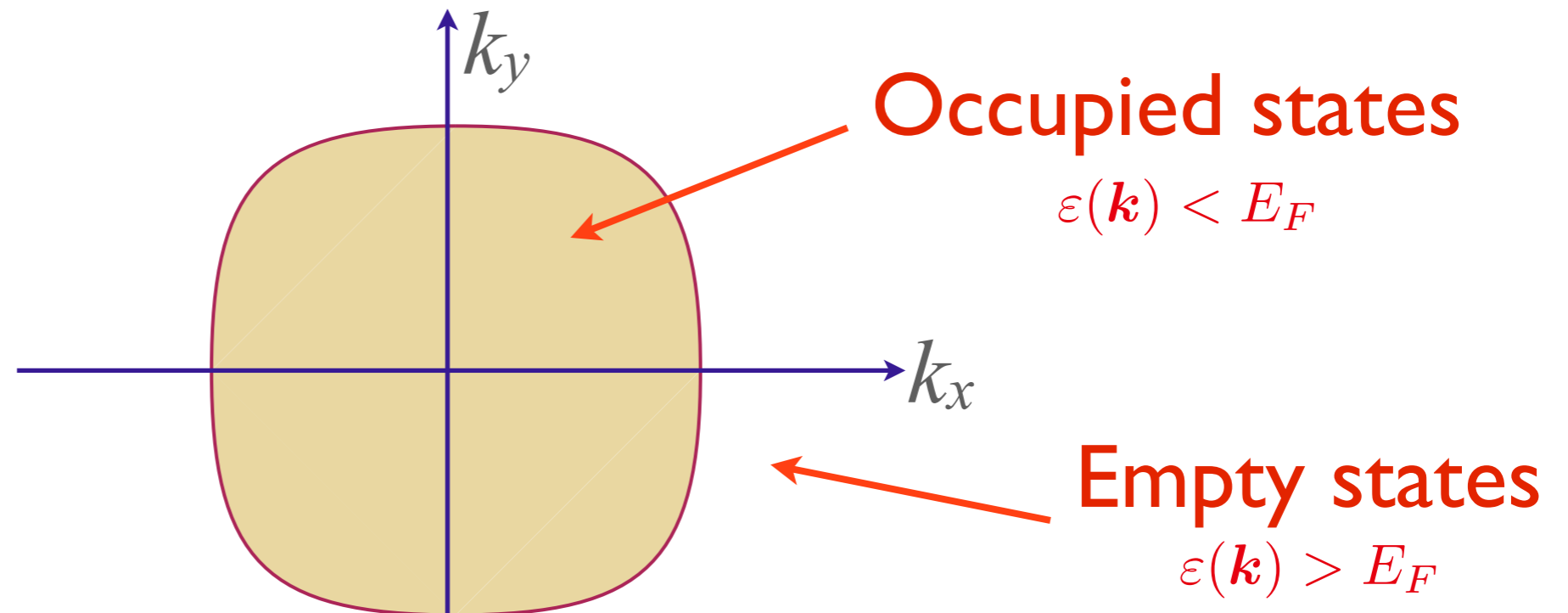
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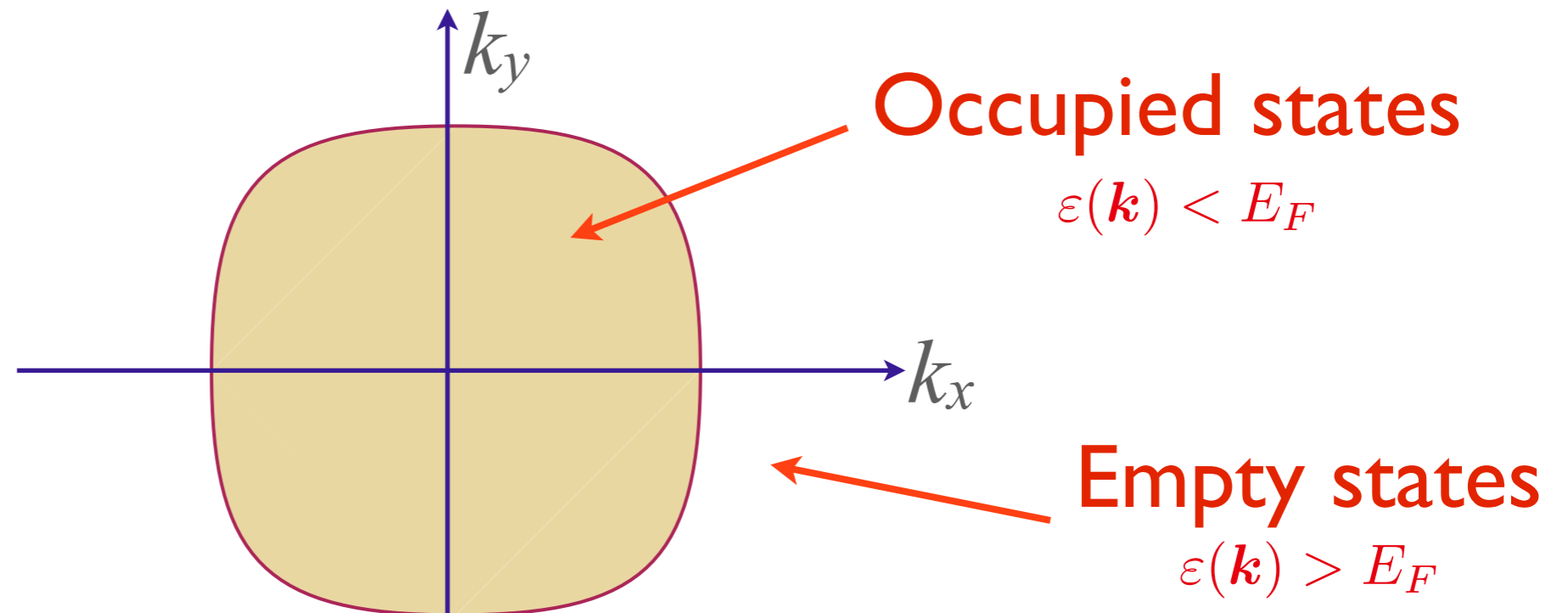
Theory of ordinary metals

Electrons move with momentum \mathbf{k} through the lattice with dispersion $\varepsilon(\mathbf{k})$



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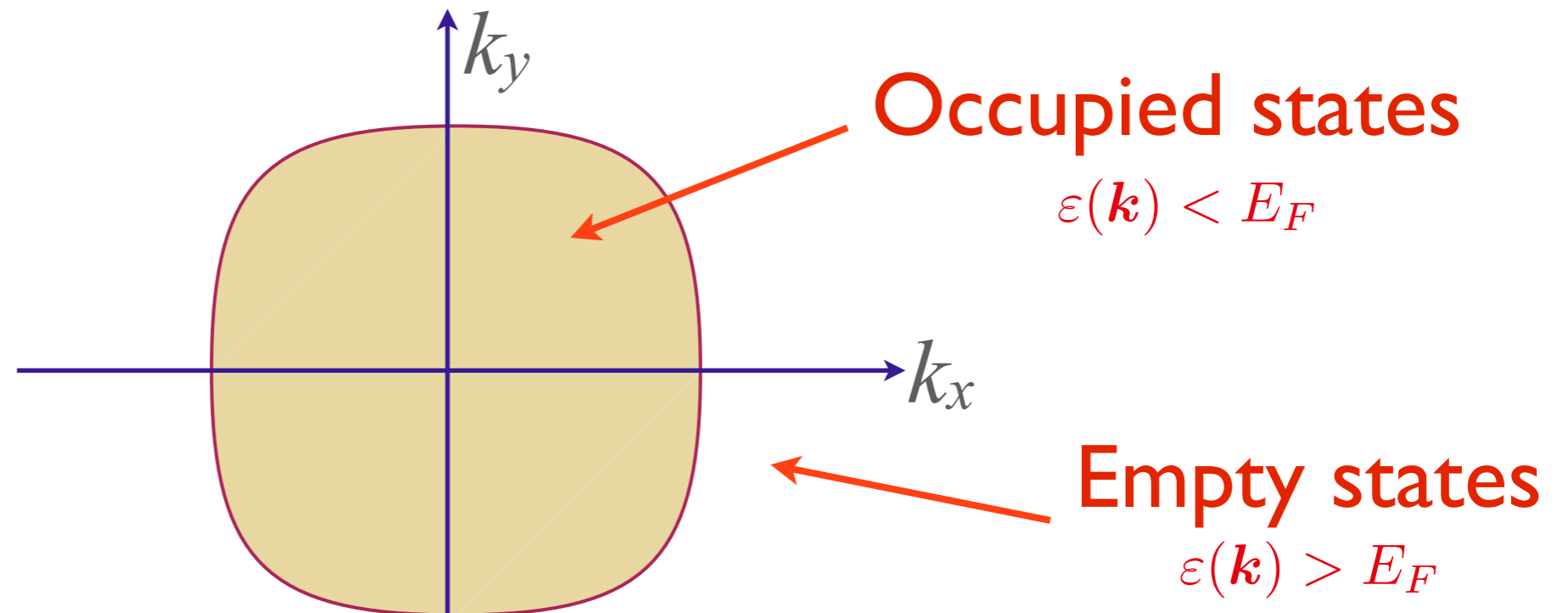


Fundamental principles of Fermi liquid theory:

- Low energy excitations are nearly-free “quasielectrons” near the Fermi surface.

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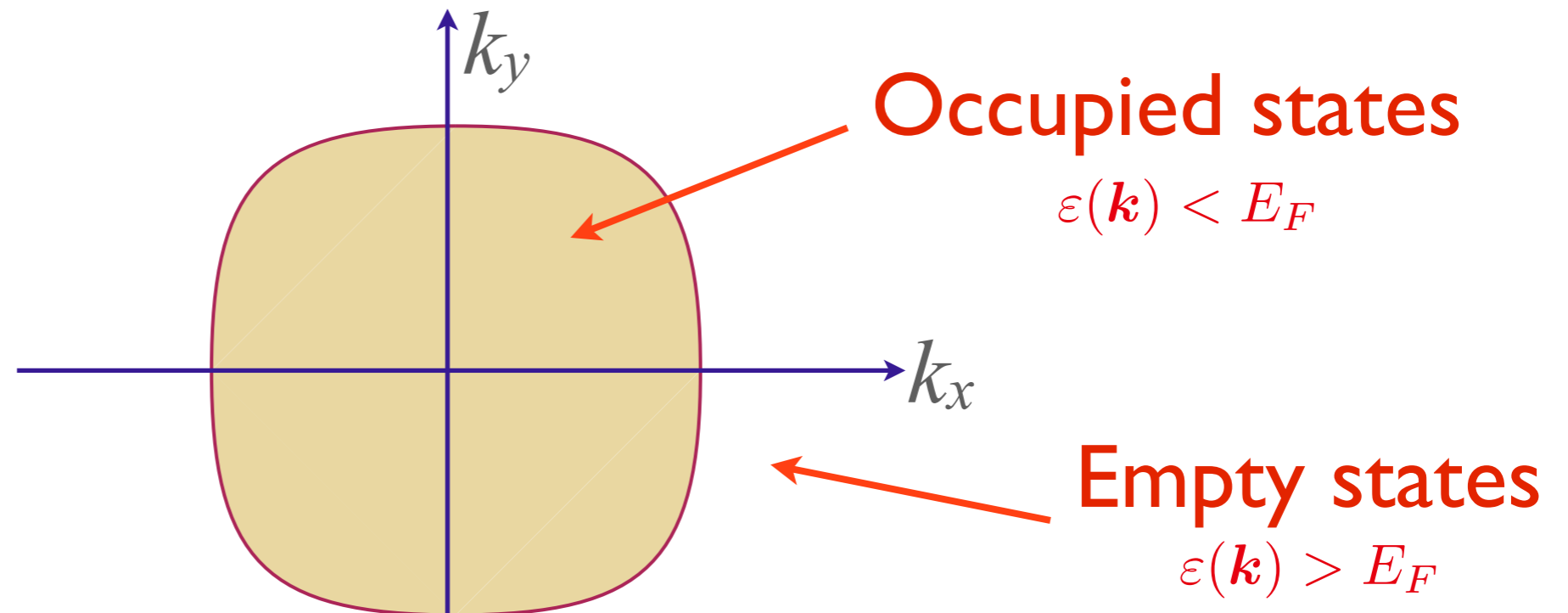
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S. Sachdev, *Quantum Phase Transitions* (1999)

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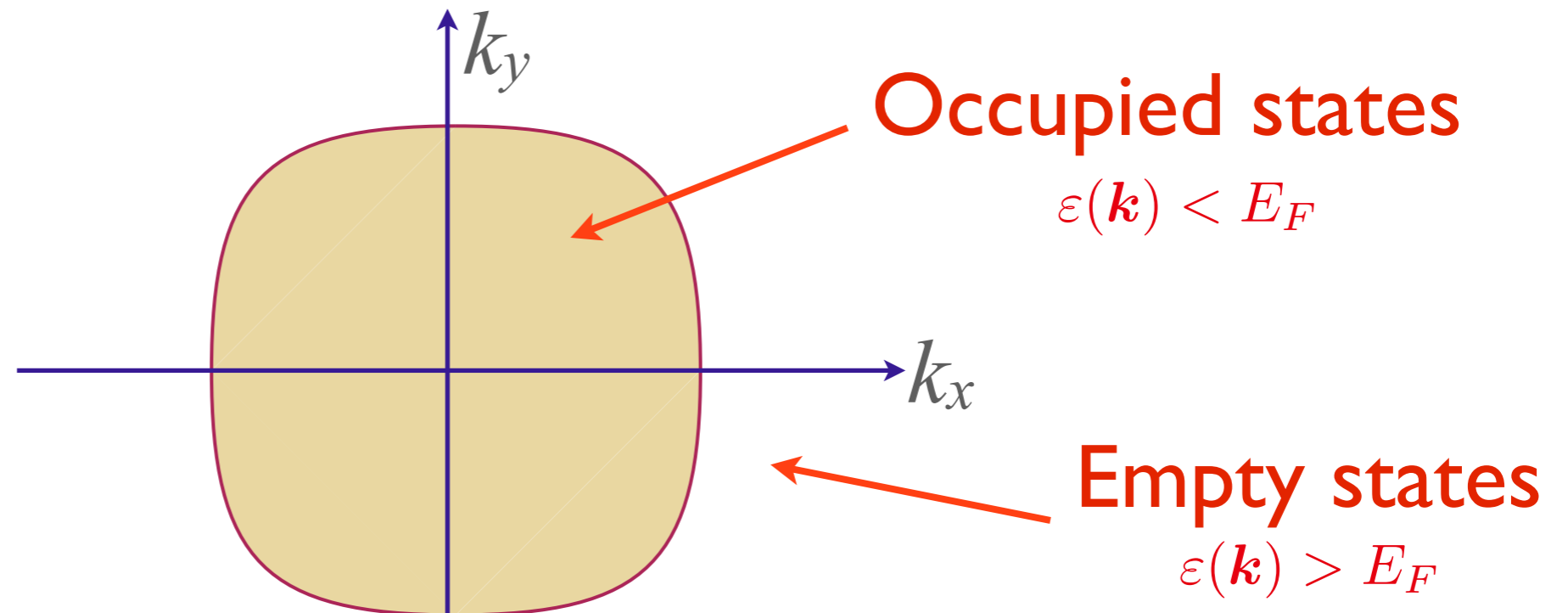


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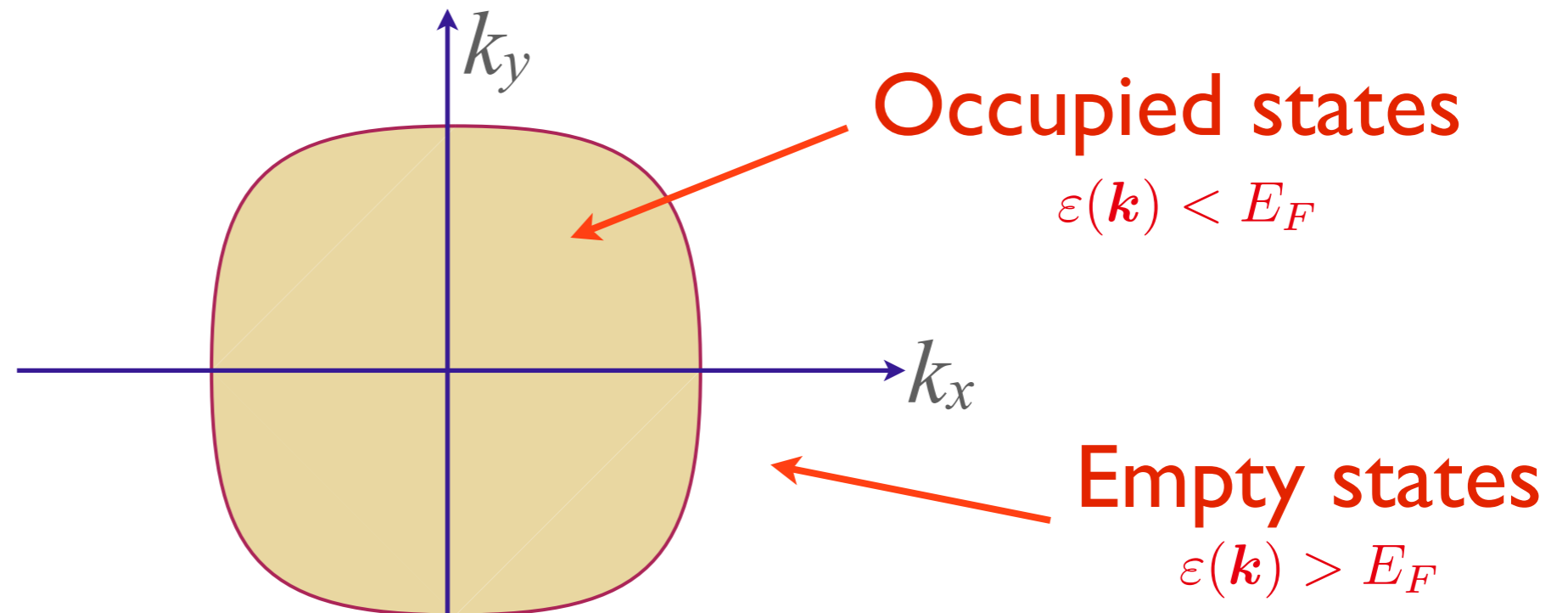
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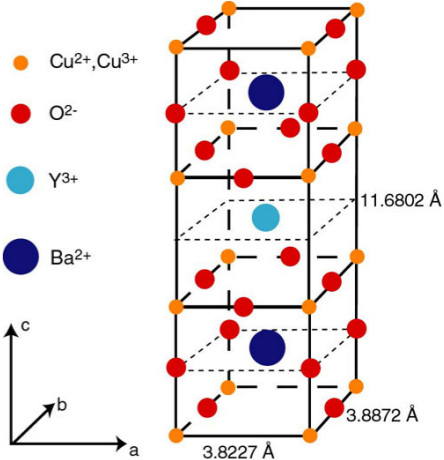
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Violated
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metal

Material		n (10^{27} m^{-3})	m^* (m_0)	A_1 / d (Ω / K)	$h / (2e^2 T_F)$ (Ω / K)	α
Bi2212	$p = 0.23$	6.8	8.4 ± 1.6	8.0 ± 0.9	7.4 ± 1.4	1.1 ± 0.3
Bi2201	$p \sim 0.4$	3.5	7 ± 1.5	8 ± 2	8 ± 2	1.0 ± 0.4
LSCO	$p = 0.26$	7.8	9.8 ± 1.7	8.2 ± 1.0	8.9 ± 1.8	0.9 ± 0.3
Nd-LSCO	$p = 0.24$	7.9	12 ± 4	7.4 ± 0.8	10.6 ± 3.7	0.7 ± 0.4
PCCO	$x = 0.17$	8.8	2.4 ± 0.1	1.7 ± 0.3	2.1 ± 0.1	0.8 ± 0.2
LCCO	$x = 0.15$	9.0	3.0 ± 0.3	3.0 ± 0.45	2.6 ± 0.3	1.2 ± 0.3
TMTSF	$P = 11 \text{ kbar}$	1.4	1.15 ± 0.2	2.8 ± 0.3	2.8 ± 0.4	1.0 ± 0.3

Planckian electron scattering time τ in 6 different cuprates

Linear- T Resistivity $\rho = \frac{m^*}{ne^2\tau} \Rightarrow \frac{1}{\tau} = \alpha \frac{k_B T}{\hbar}$

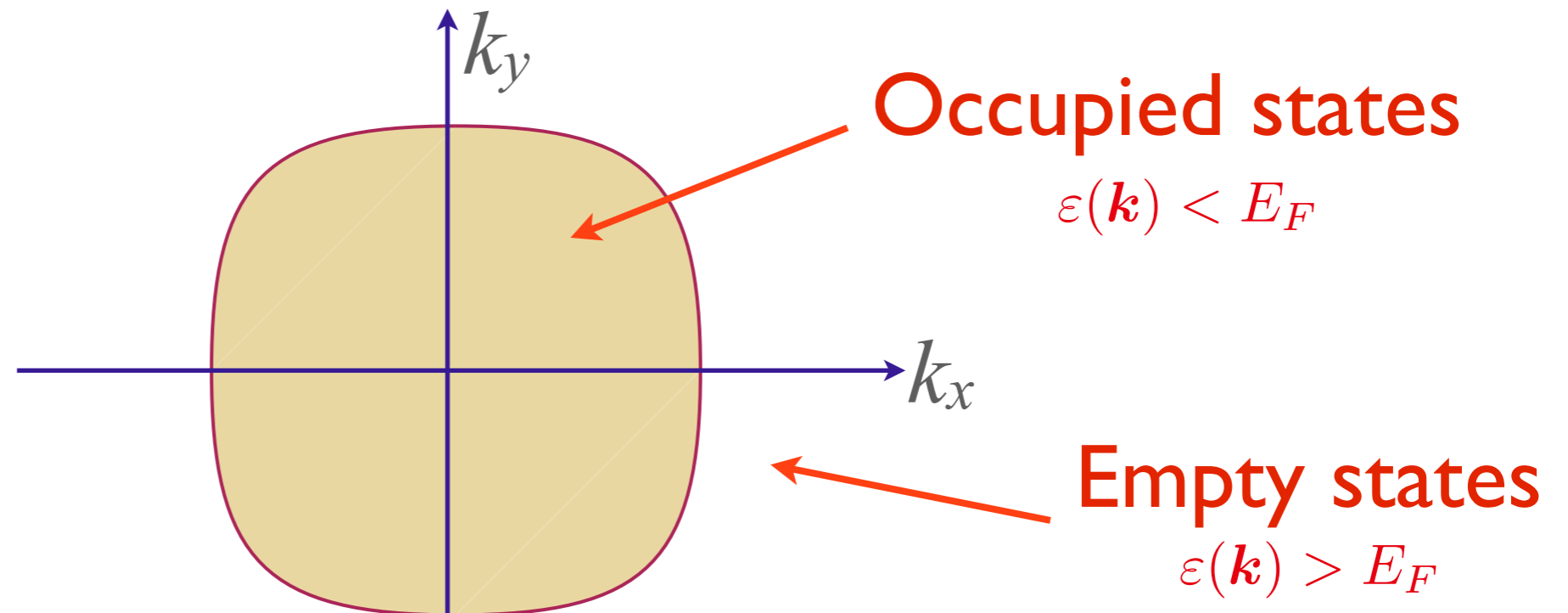


Current flow without quasiparticles

See also: G. Grissonnanche et al., arXiv:2011.13054

Theory of metals

Electrons move with momentum \mathbf{k} through the lattice with dispersion $\varepsilon(\mathbf{k})$



- Needed: a theory of a Planckian metal with $1/\tau \sim k_B T/\hbar$

Missing ingredient:
multi-particle, long-range
quantum entanglement

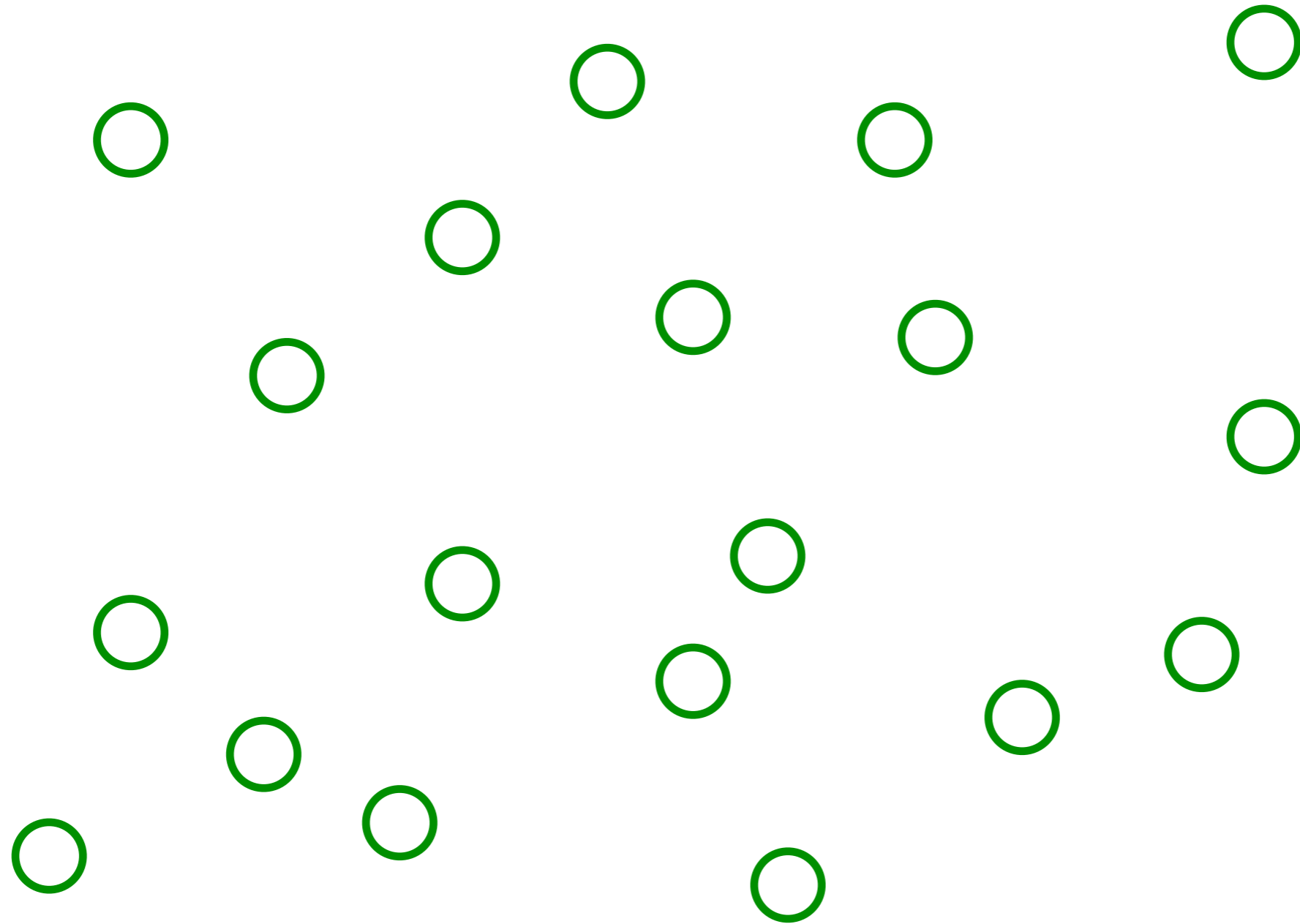
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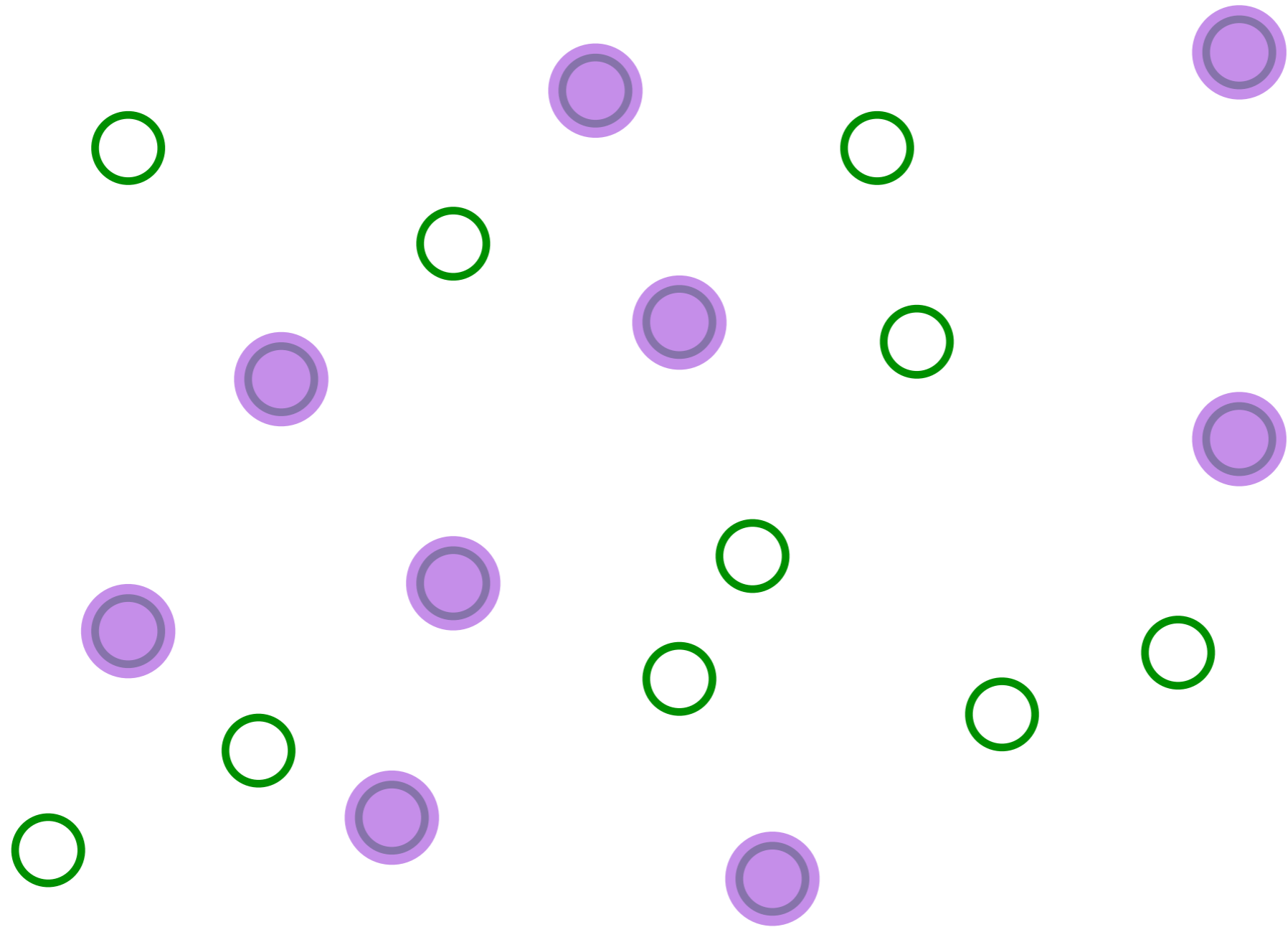
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The random matrix model



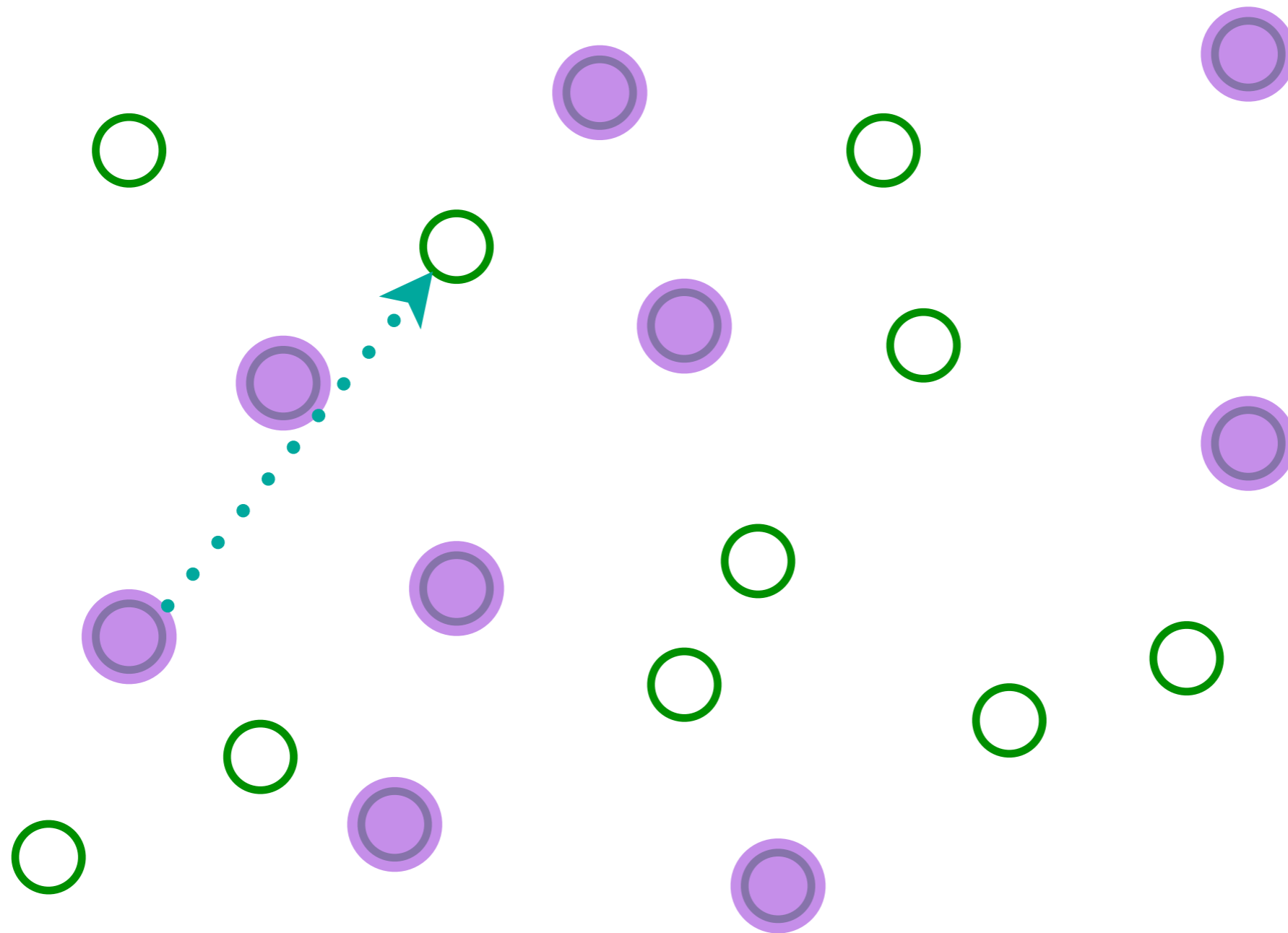
Pick a set of random positions

The random matrix model



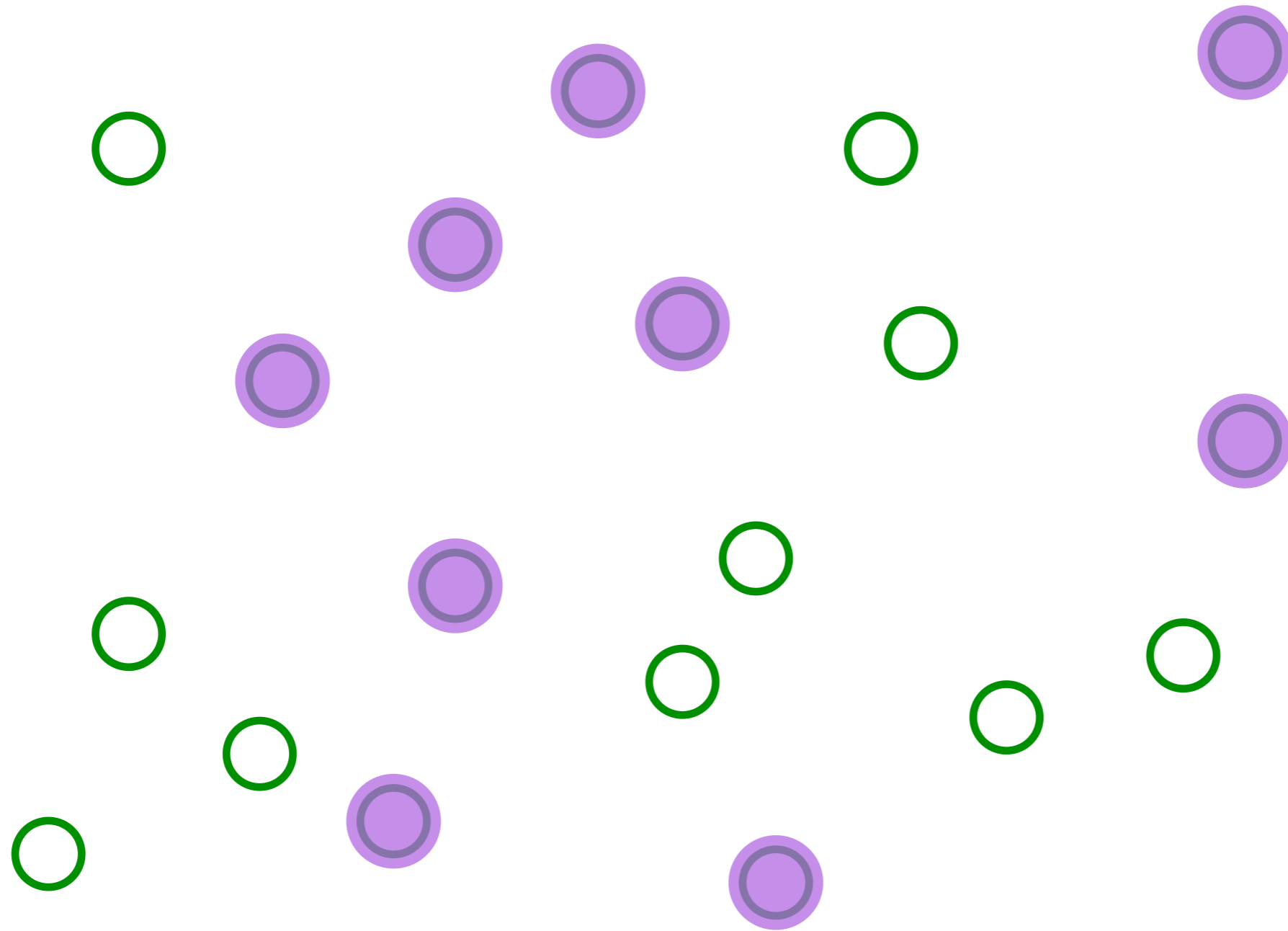
Place electrons randomly on some sites

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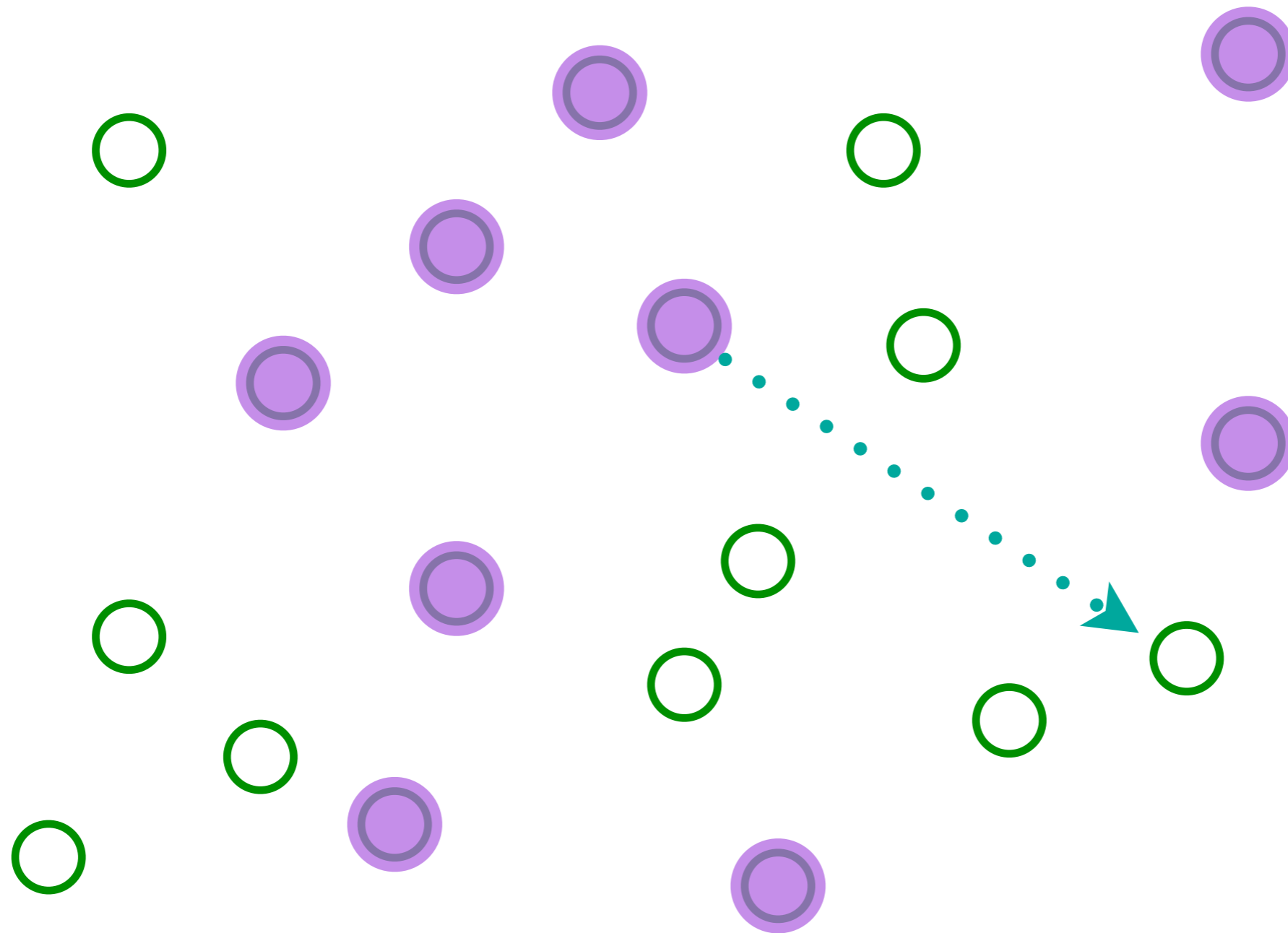
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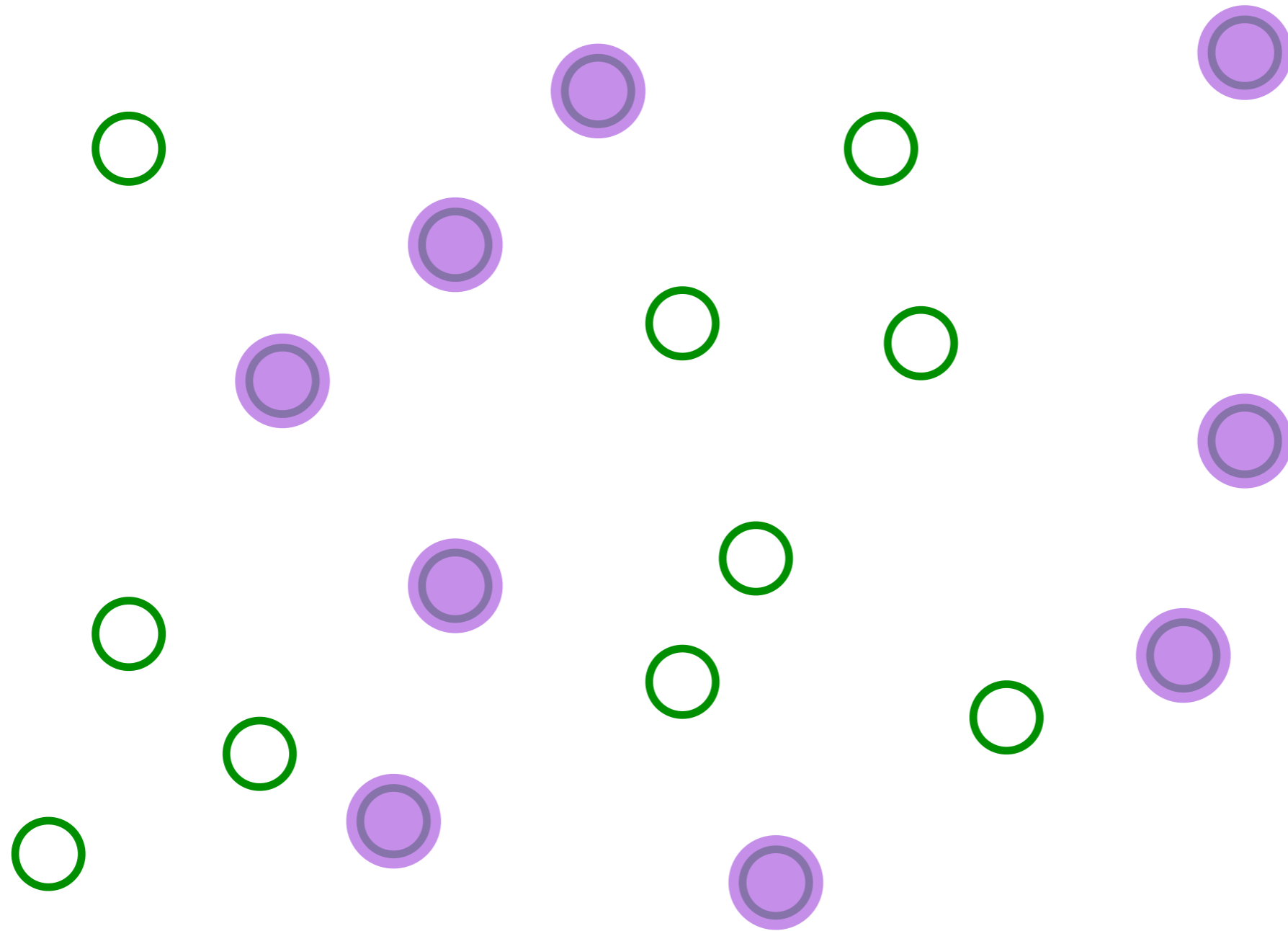
Electrons can hop anywhere with a random amplitude

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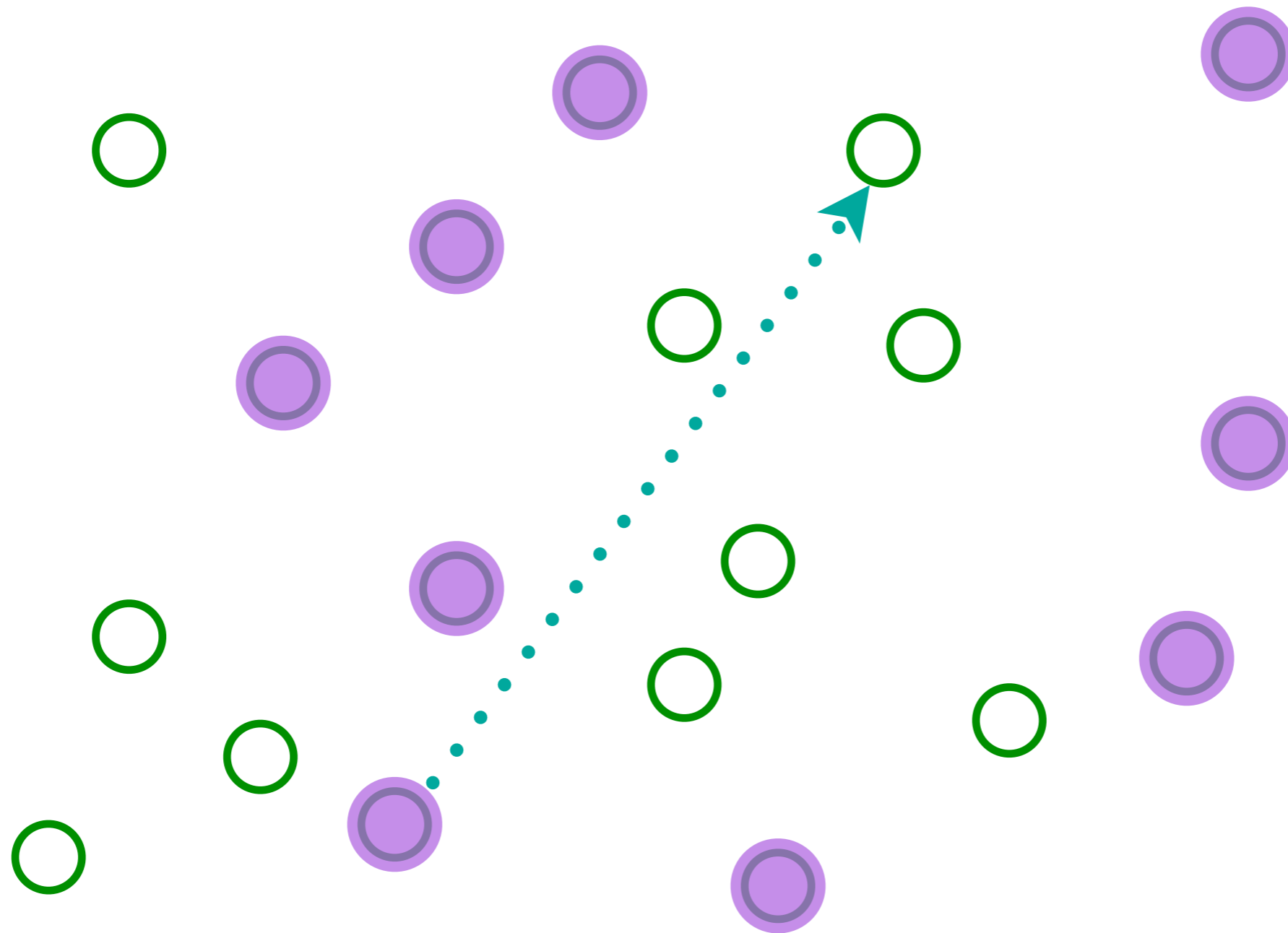
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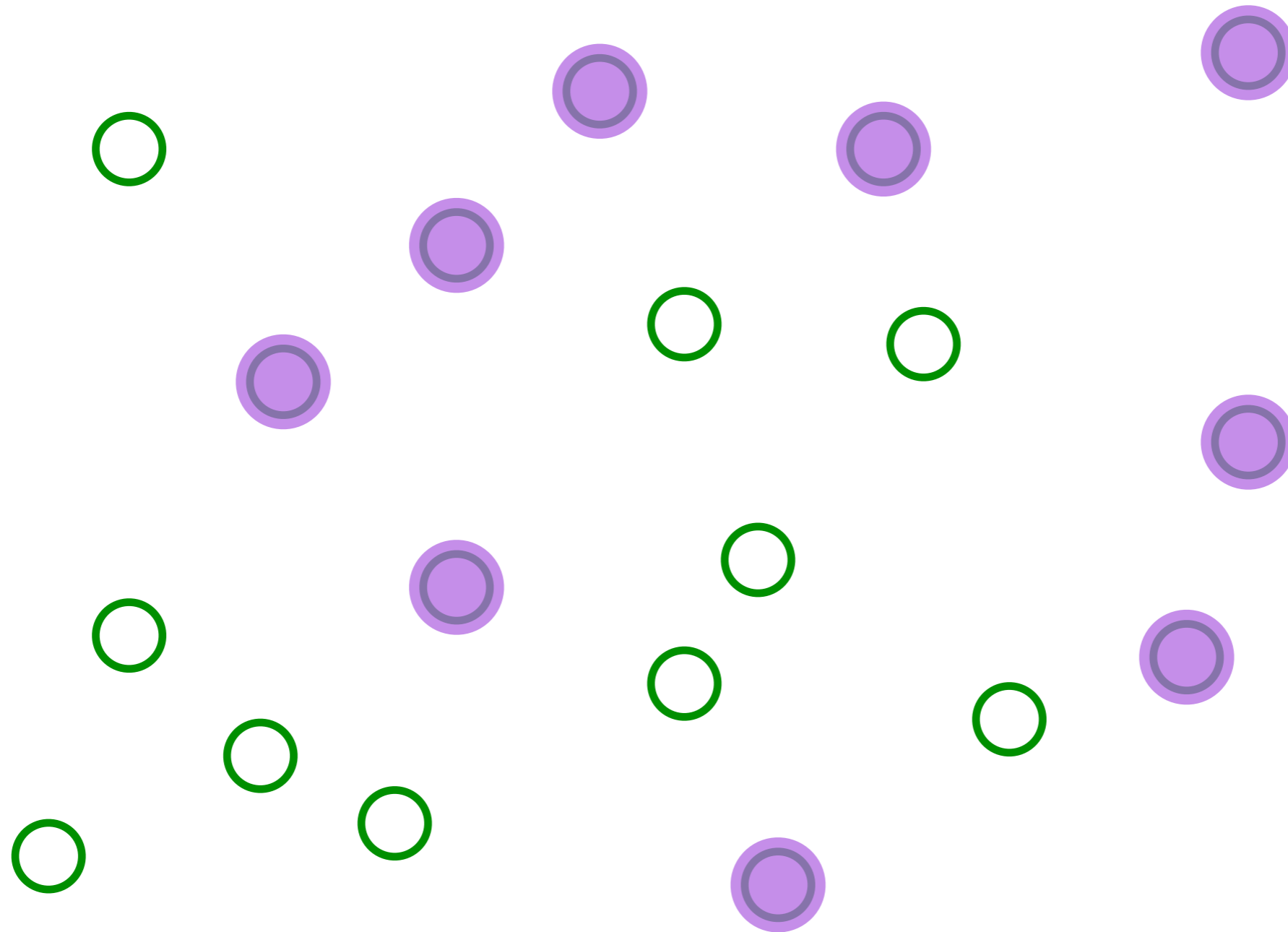
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A simple model of a metal with quasiparticles

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i$$

$$\frac{1}{N} \sum_i c_i^\dagger c_i = \mathcal{Q}$$

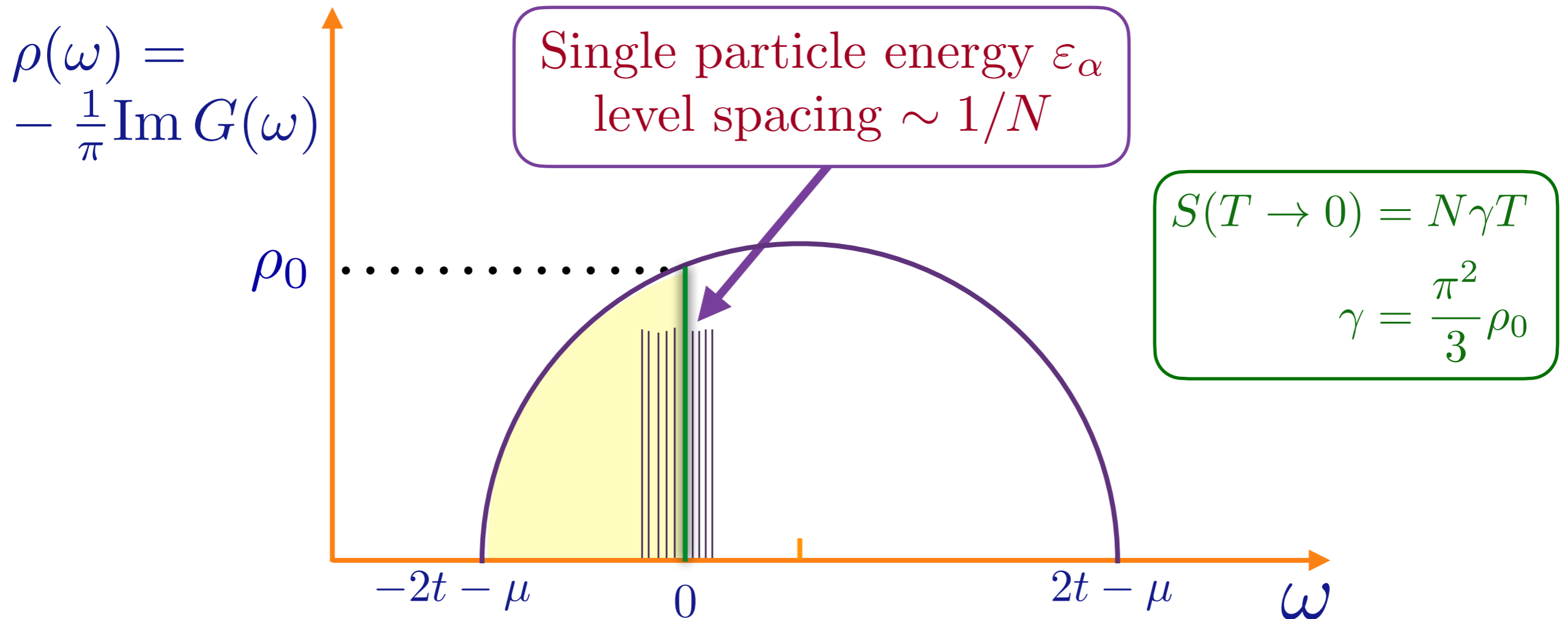
t_{ij} are independent random variables with $\overline{t_{ij}} = 0$ and $\overline{|t_{ij}|^2} = t^2$

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Many-body density of states

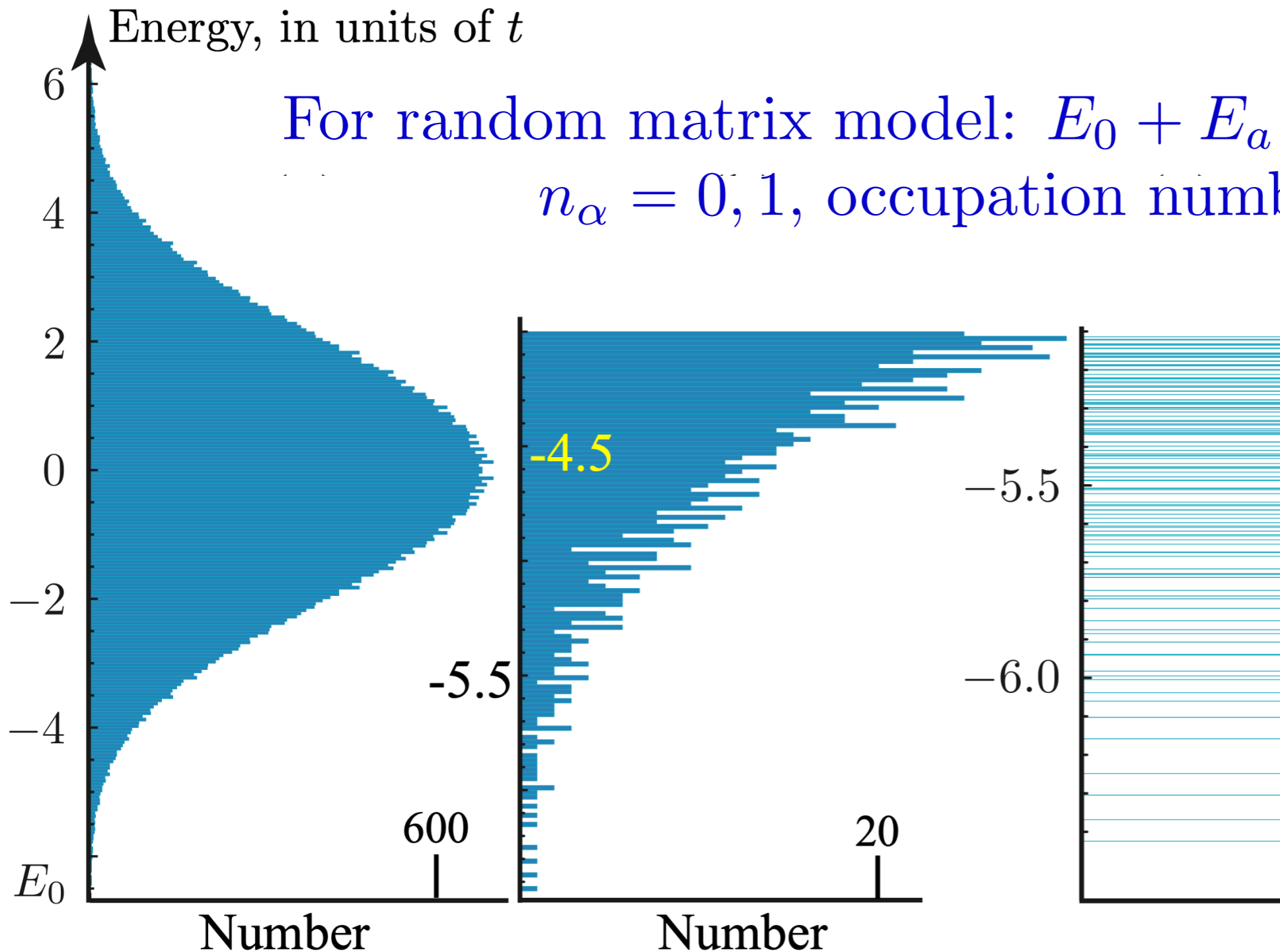
$$D(E) = \sum_a \delta(E - E_a); \quad E_0 + E_a \Rightarrow \text{Many body eigenvalue}$$

For random matrix model: $E_0 + E_a = \sum_{\alpha} n_{\alpha} \varepsilon_{\alpha}$
 $n_{\alpha} = 0, 1$, occupation number

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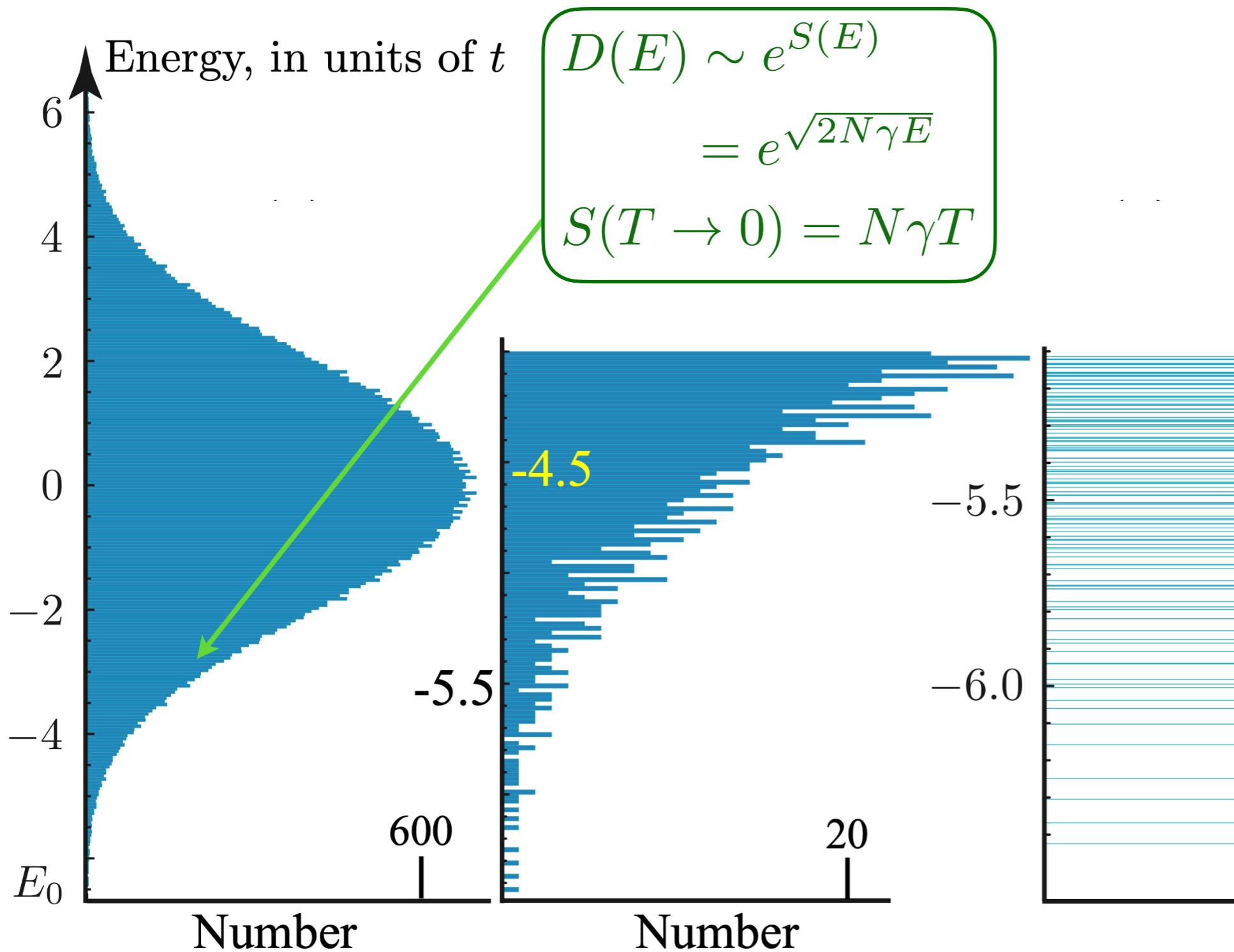


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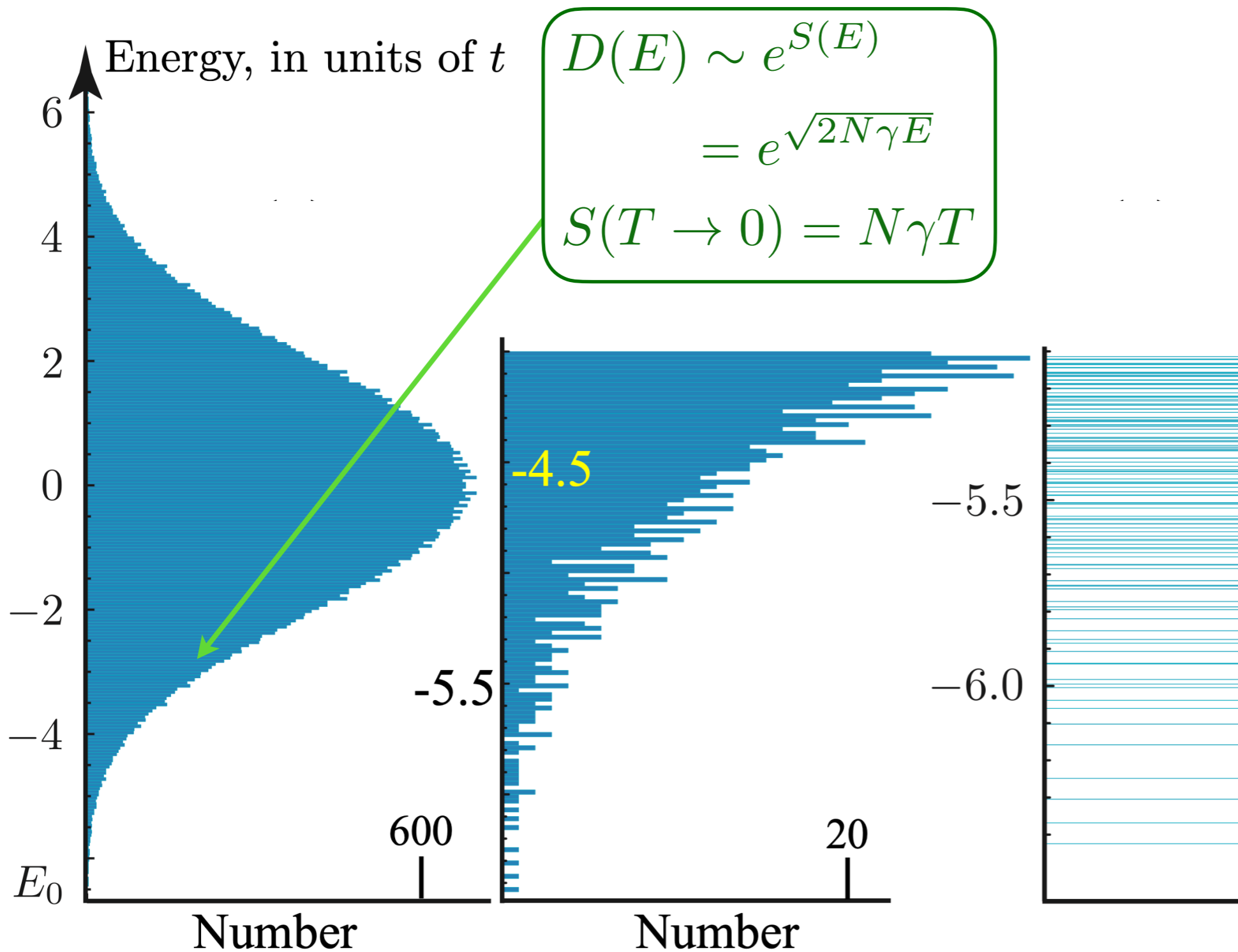


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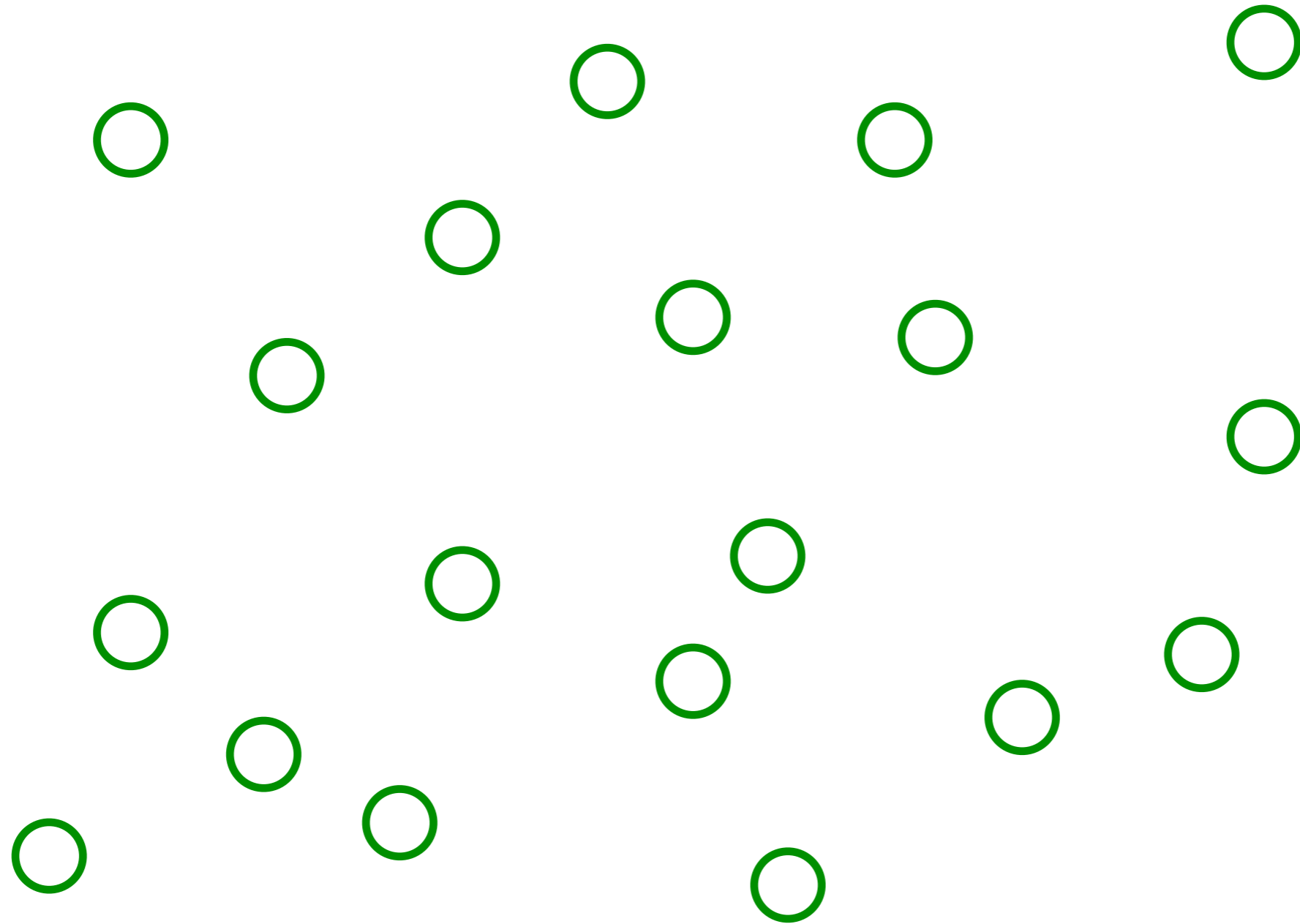
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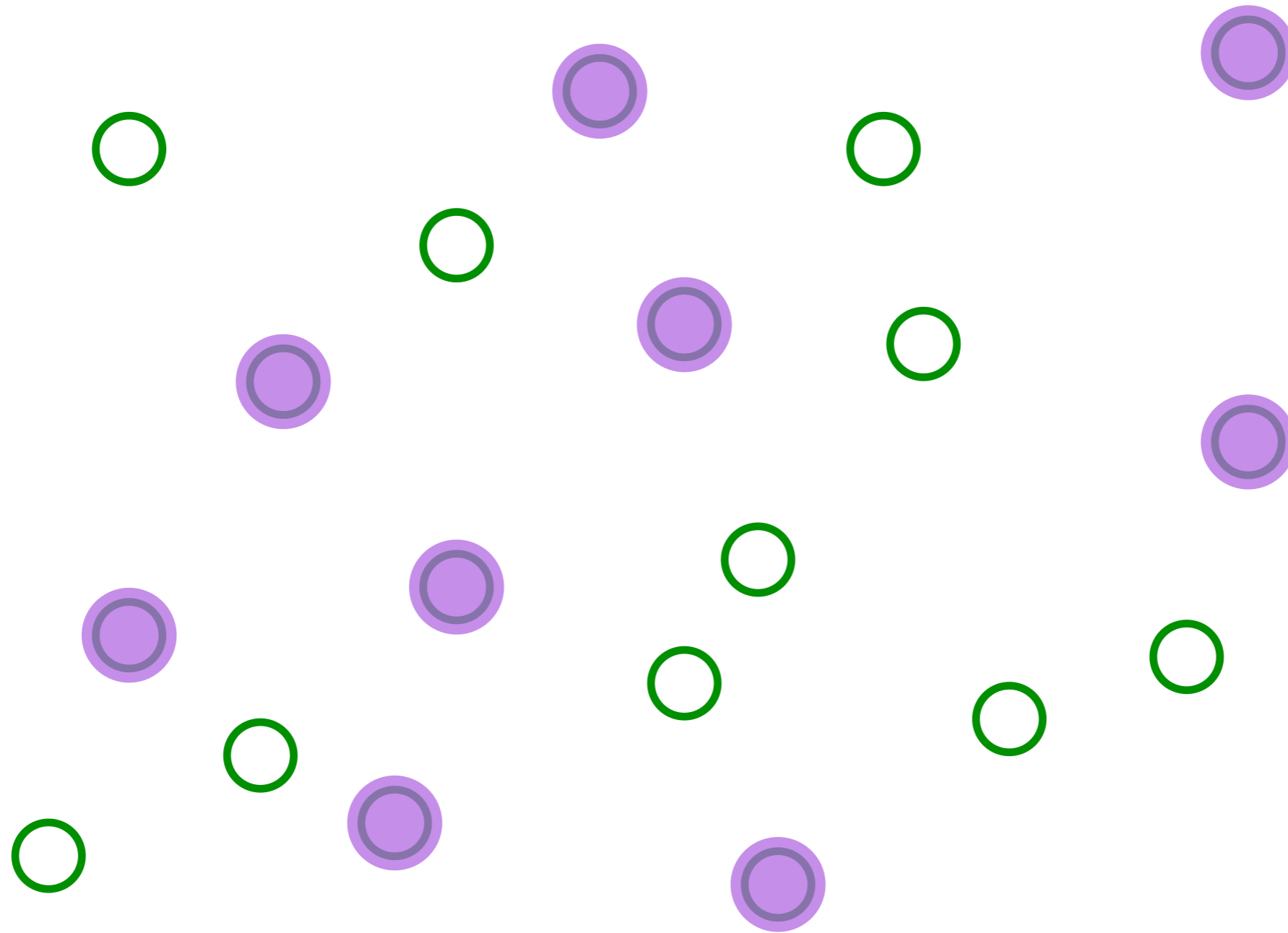
Random matrix model

The SYK model



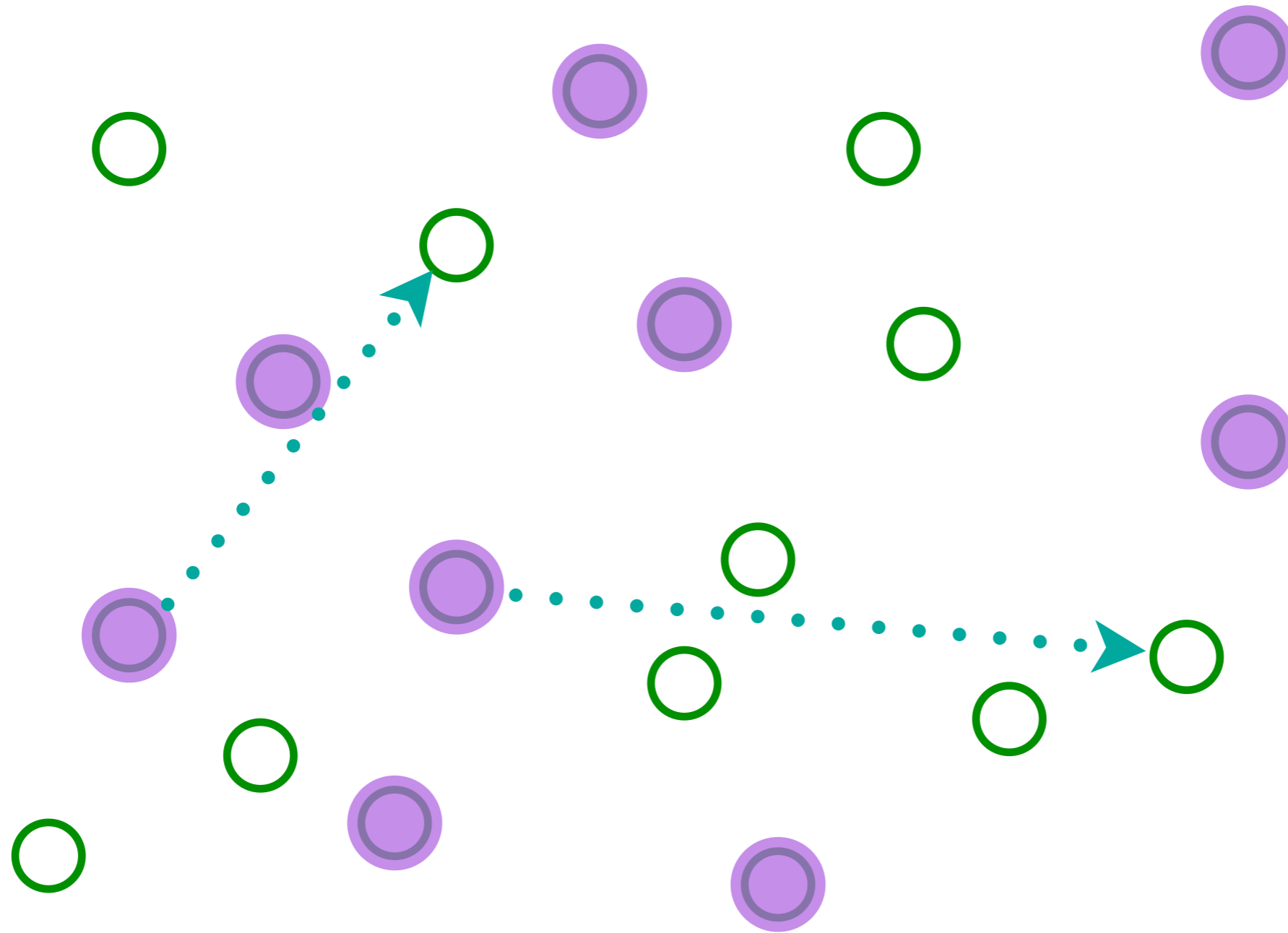
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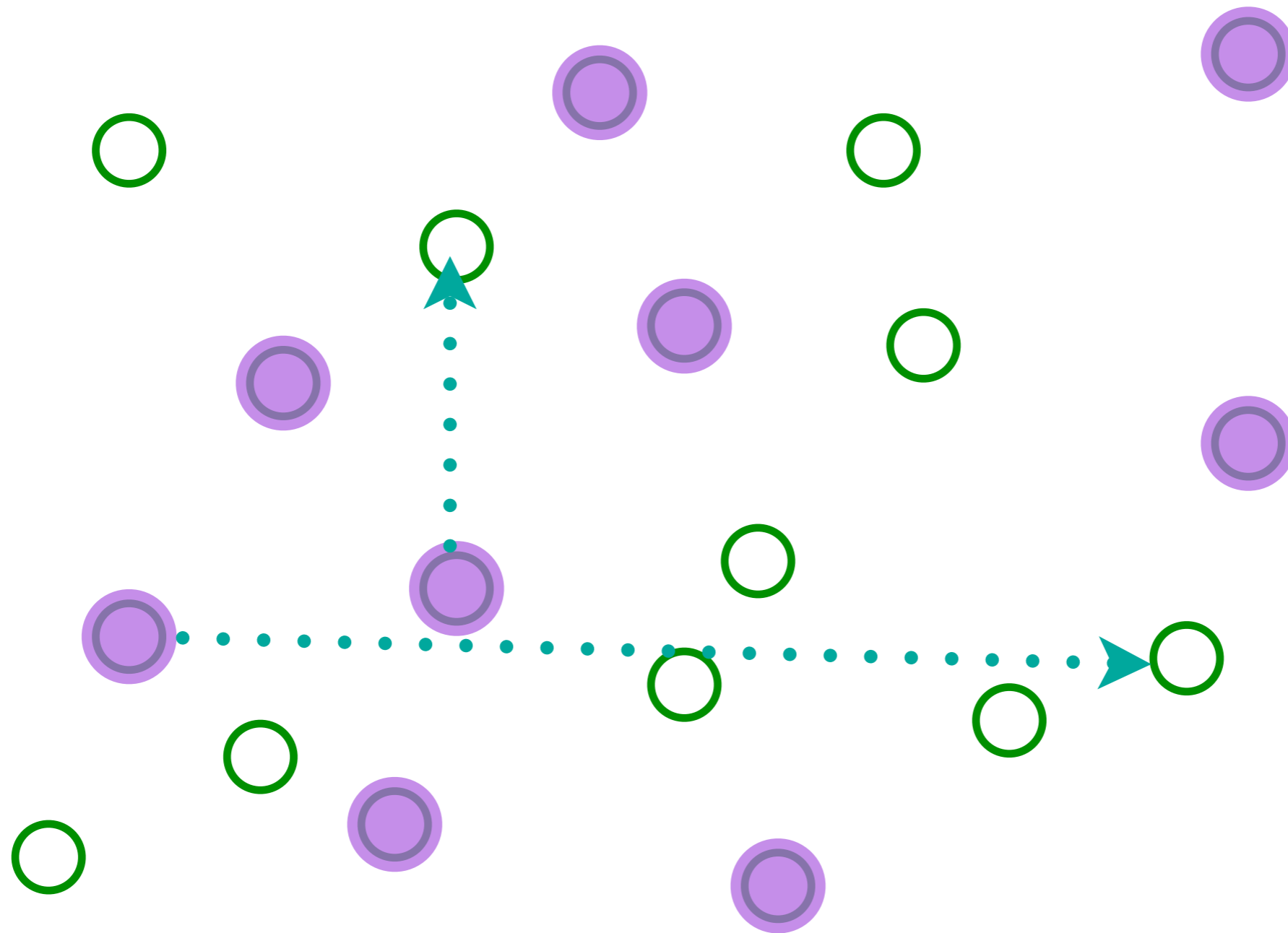
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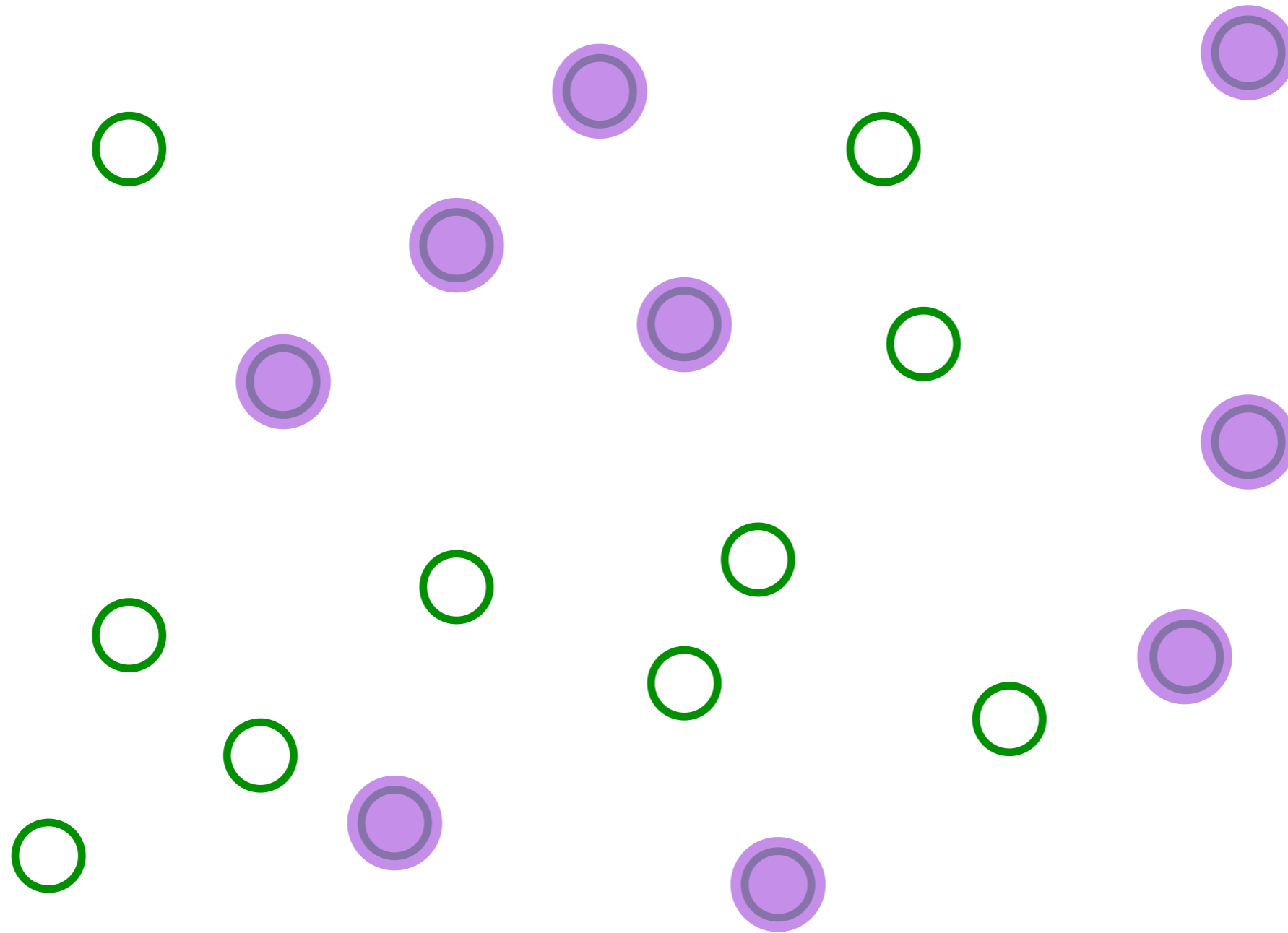
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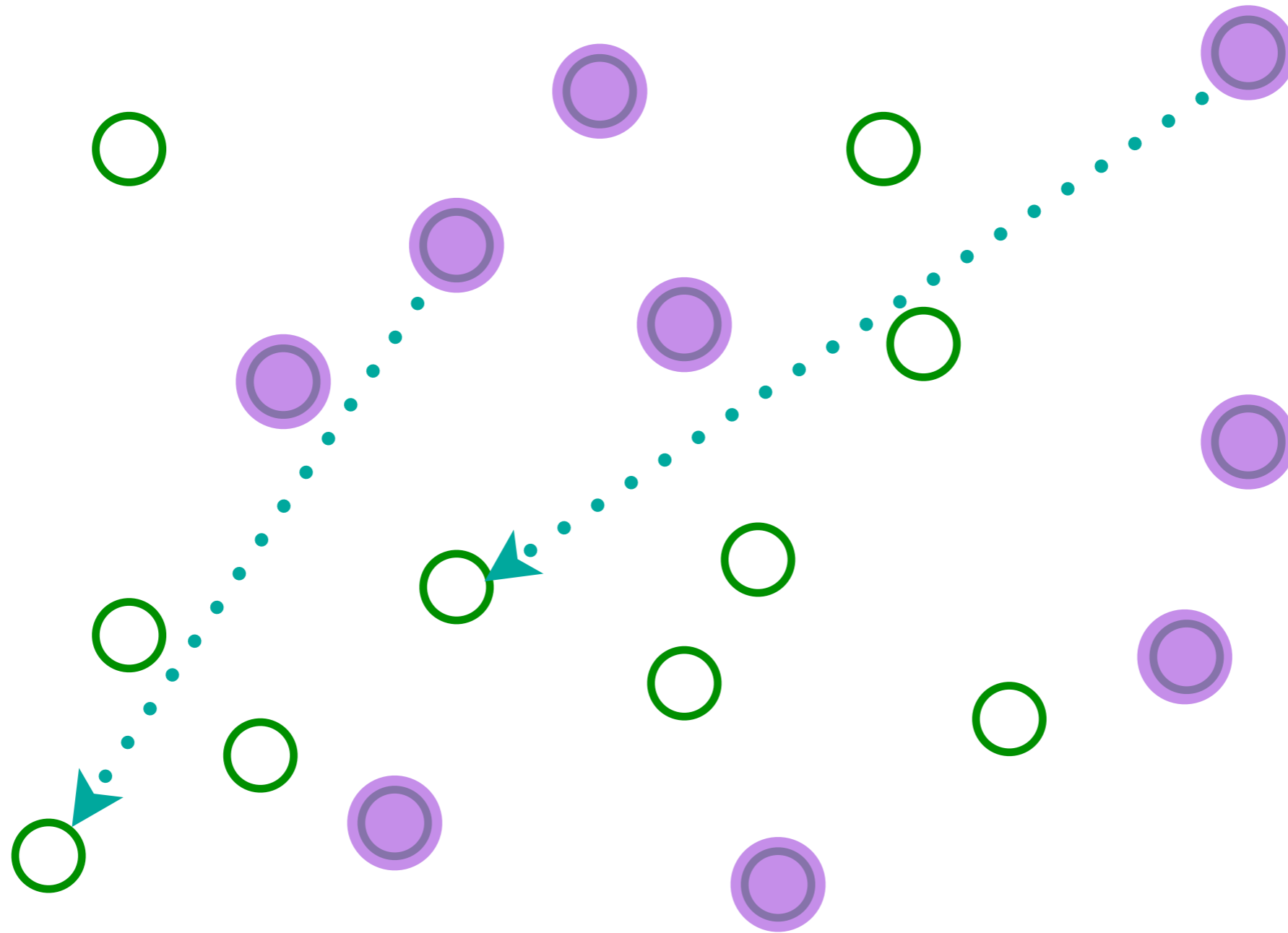
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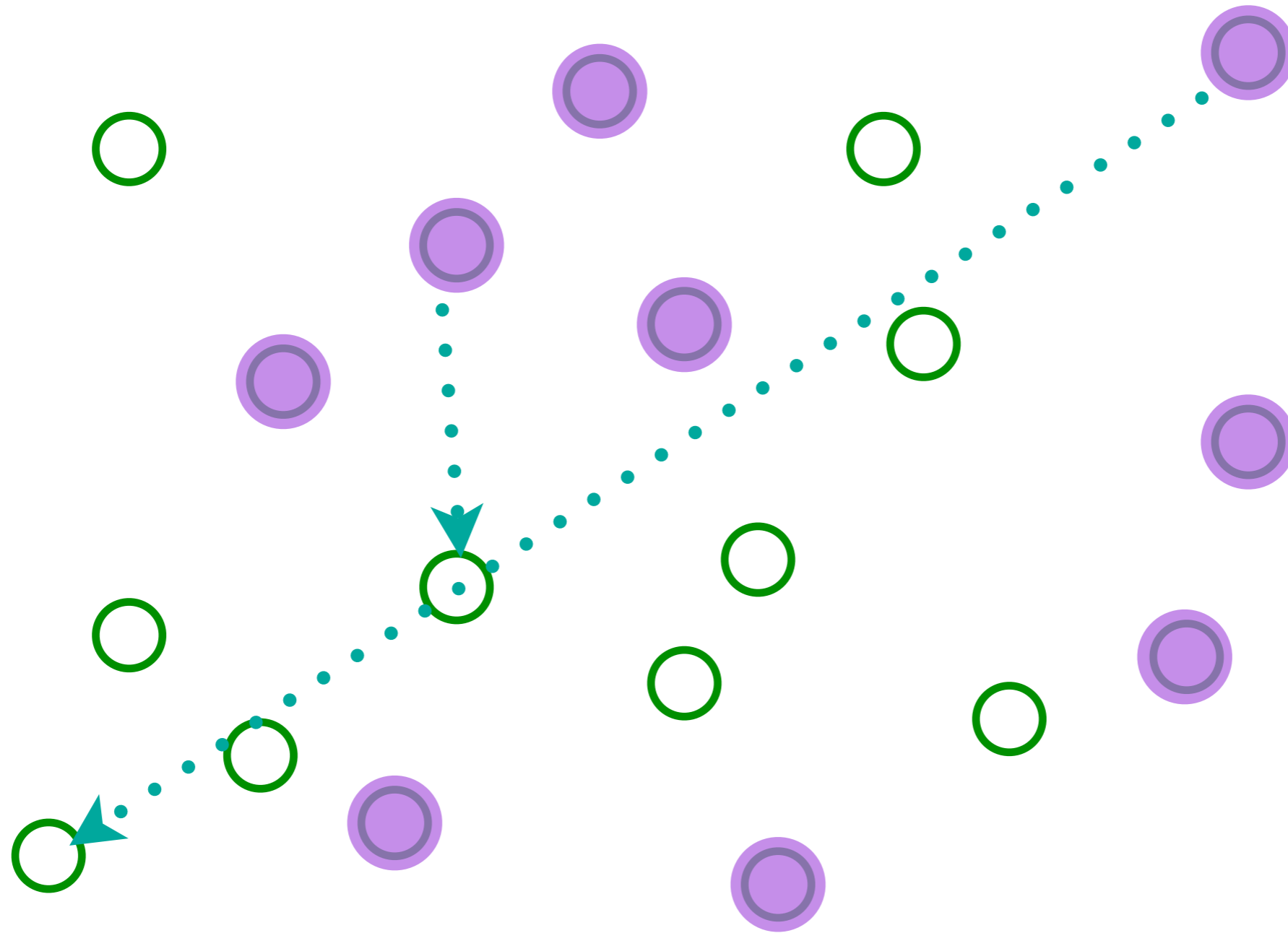
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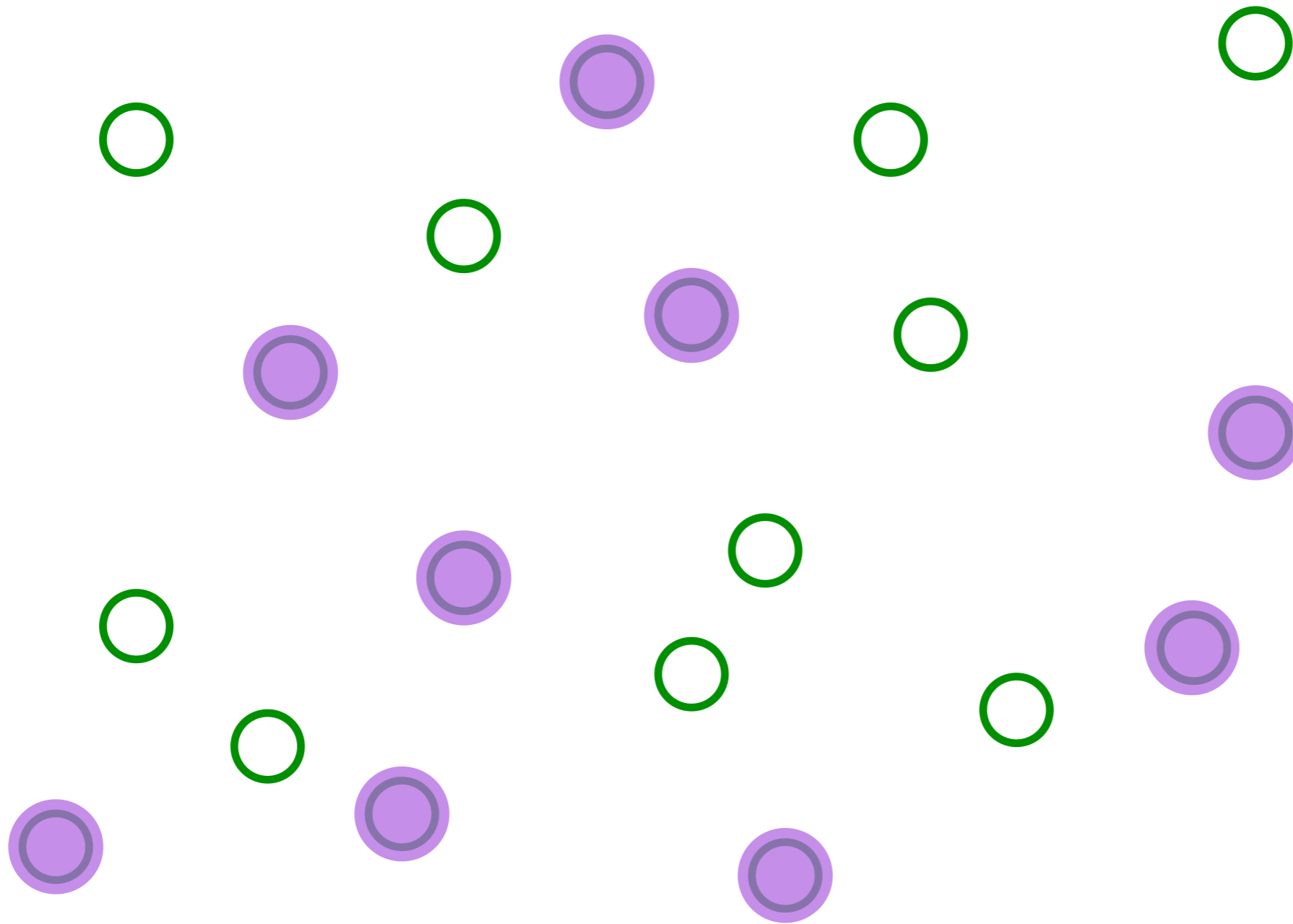
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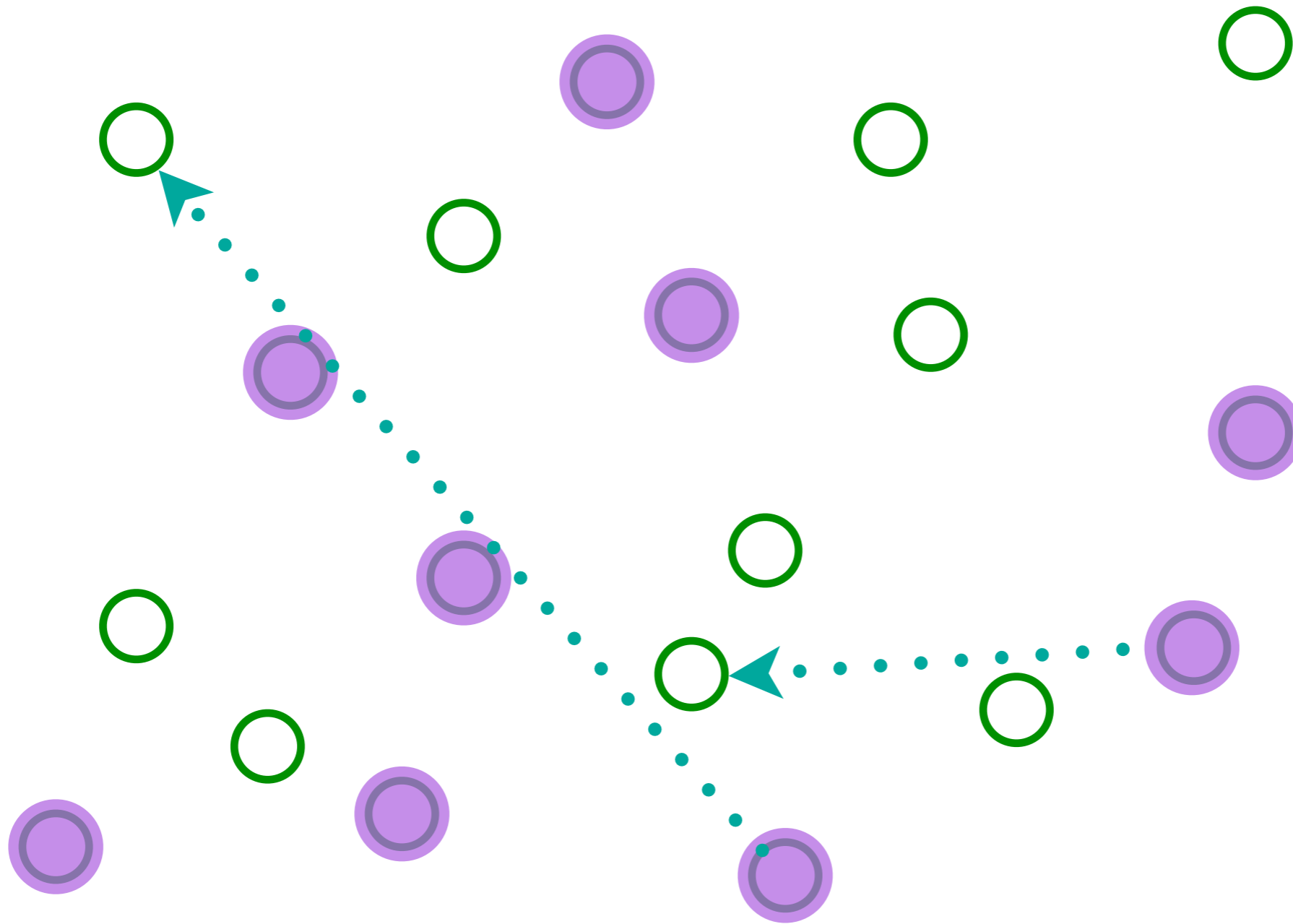
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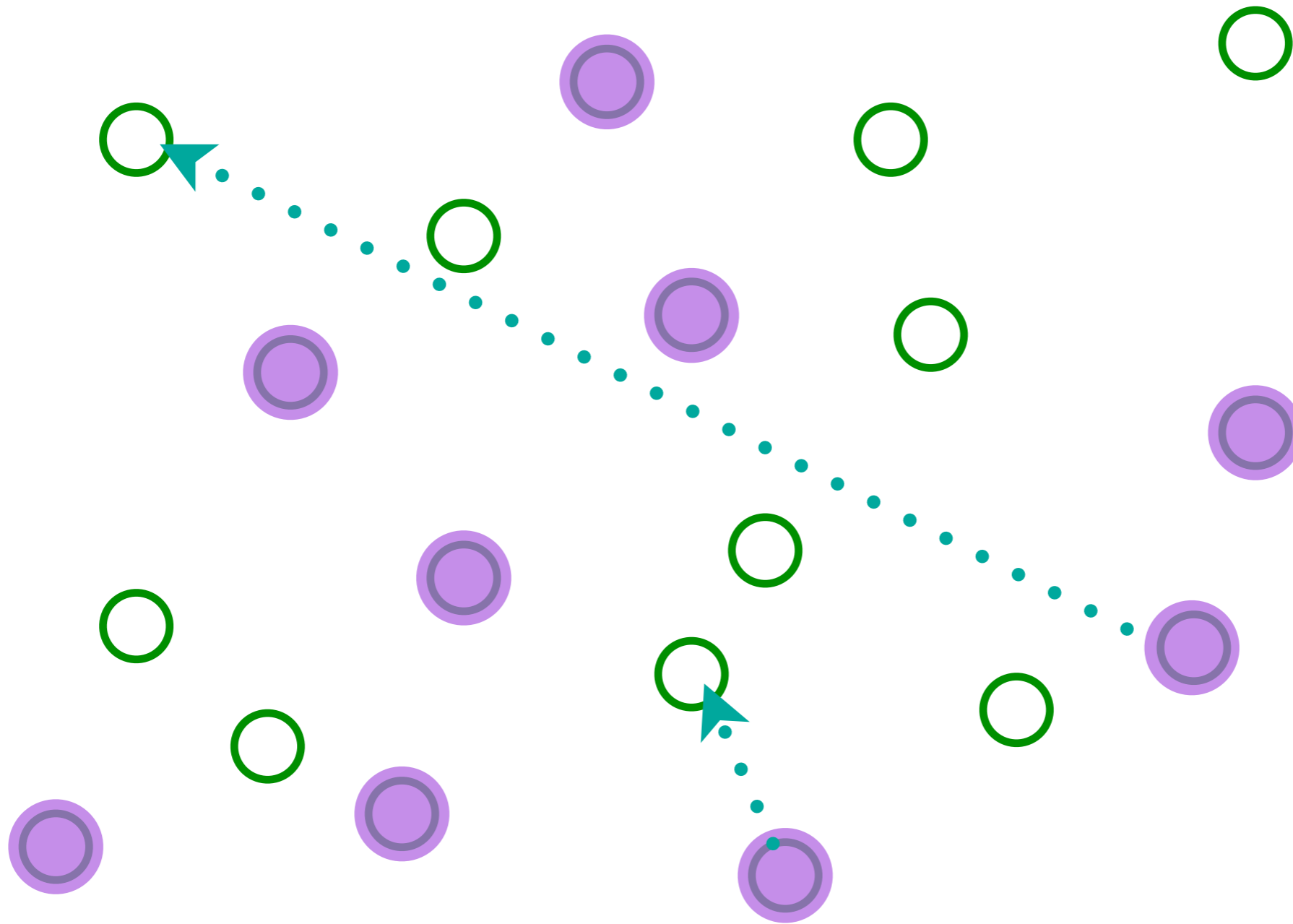
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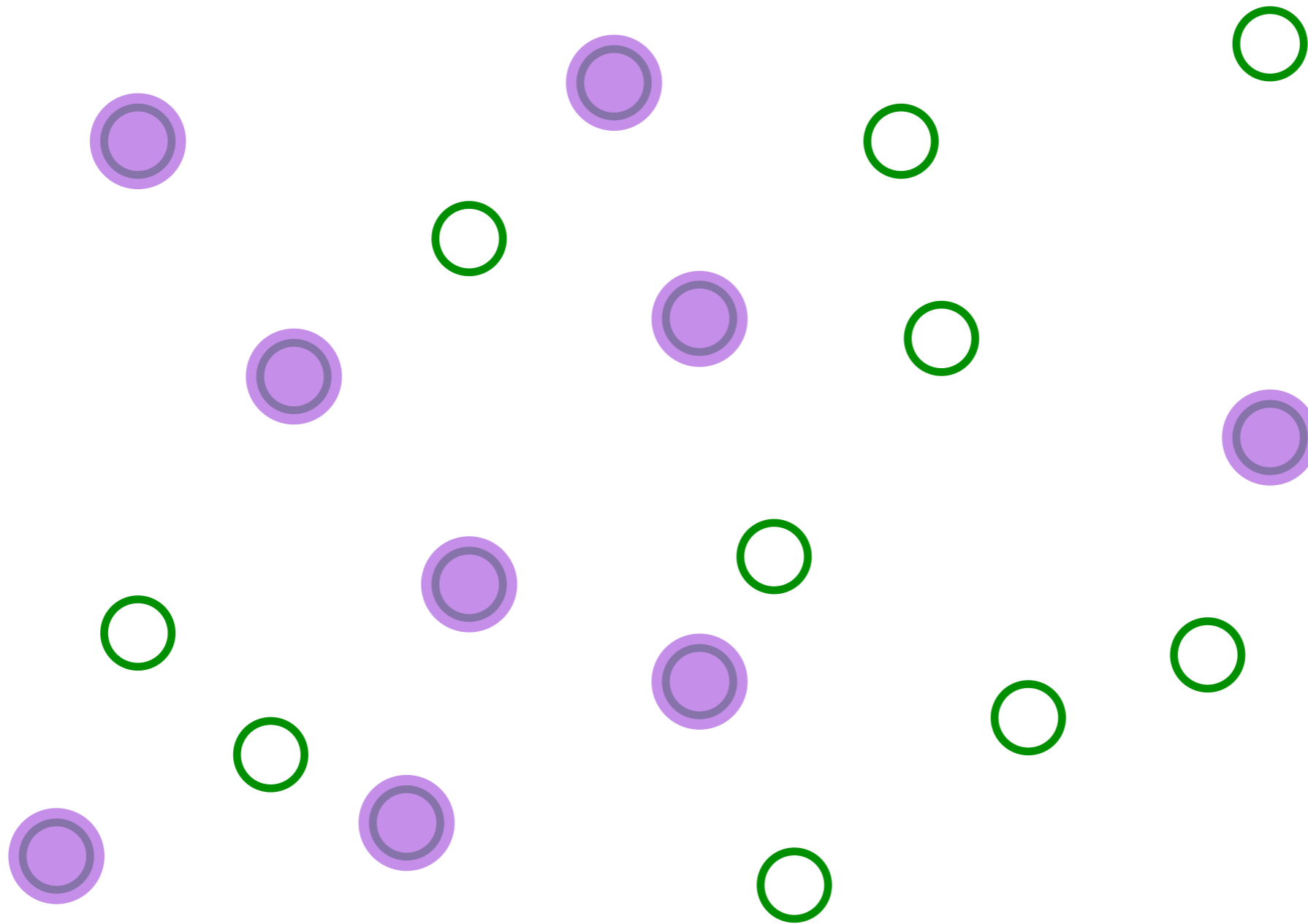
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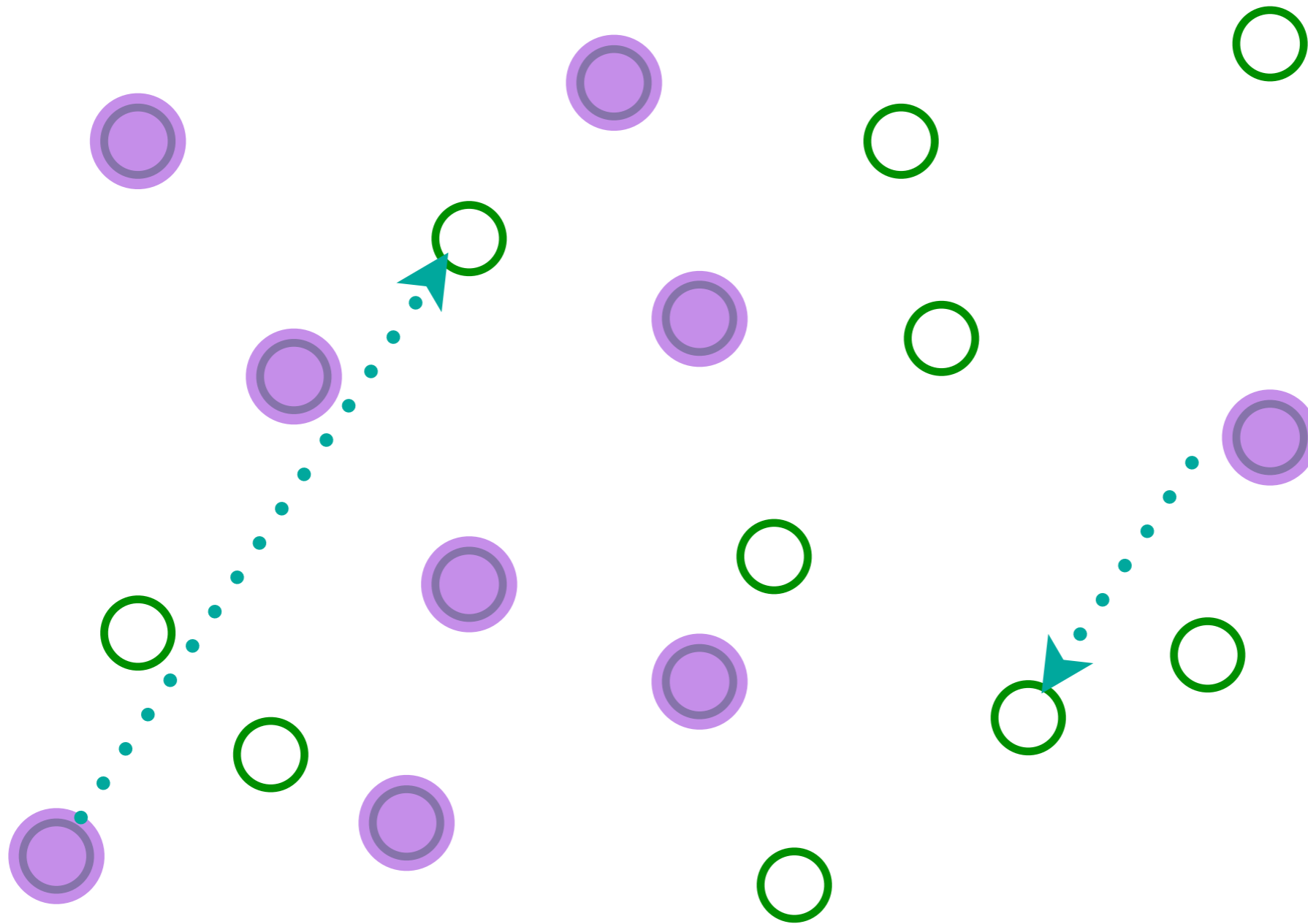
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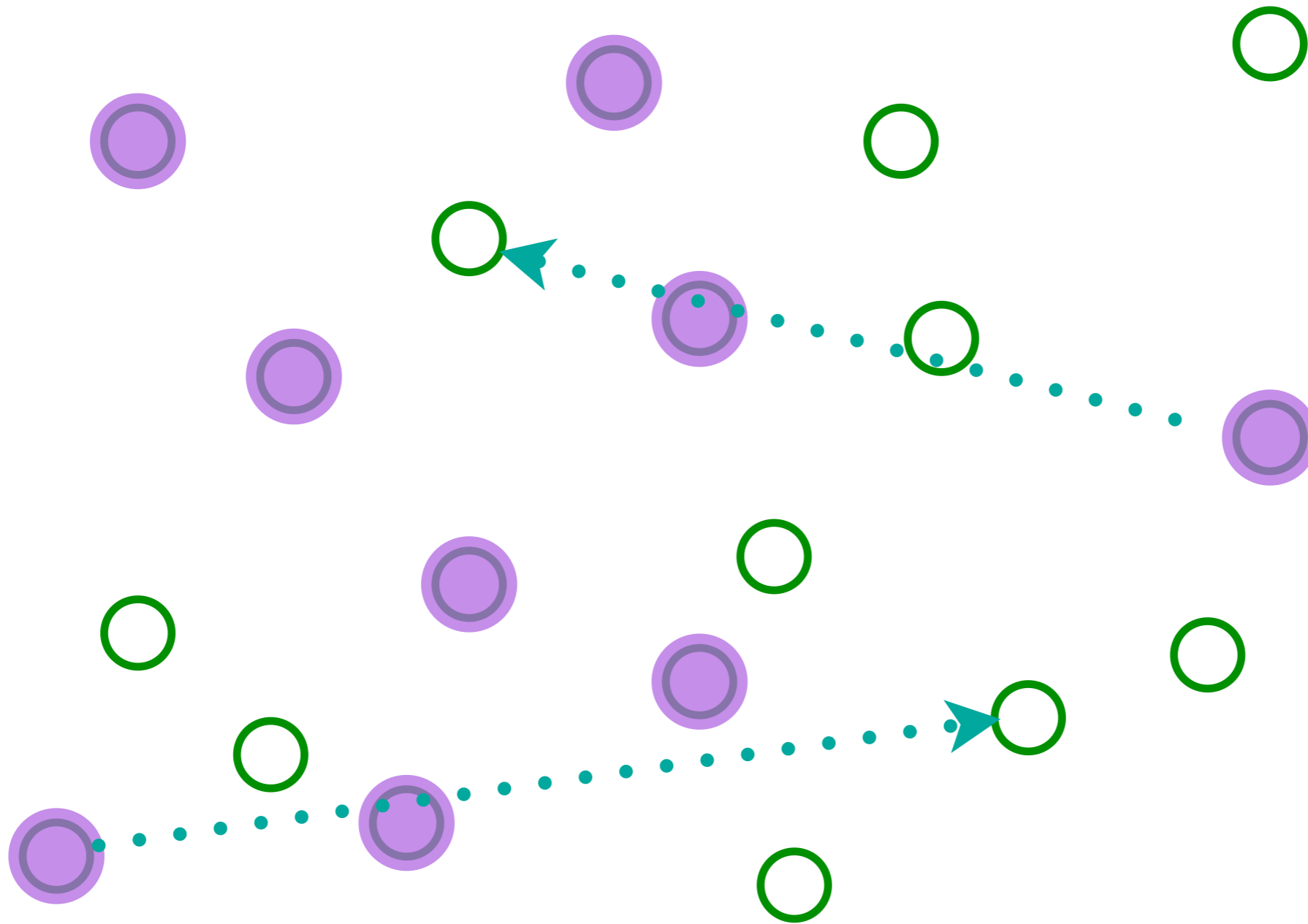
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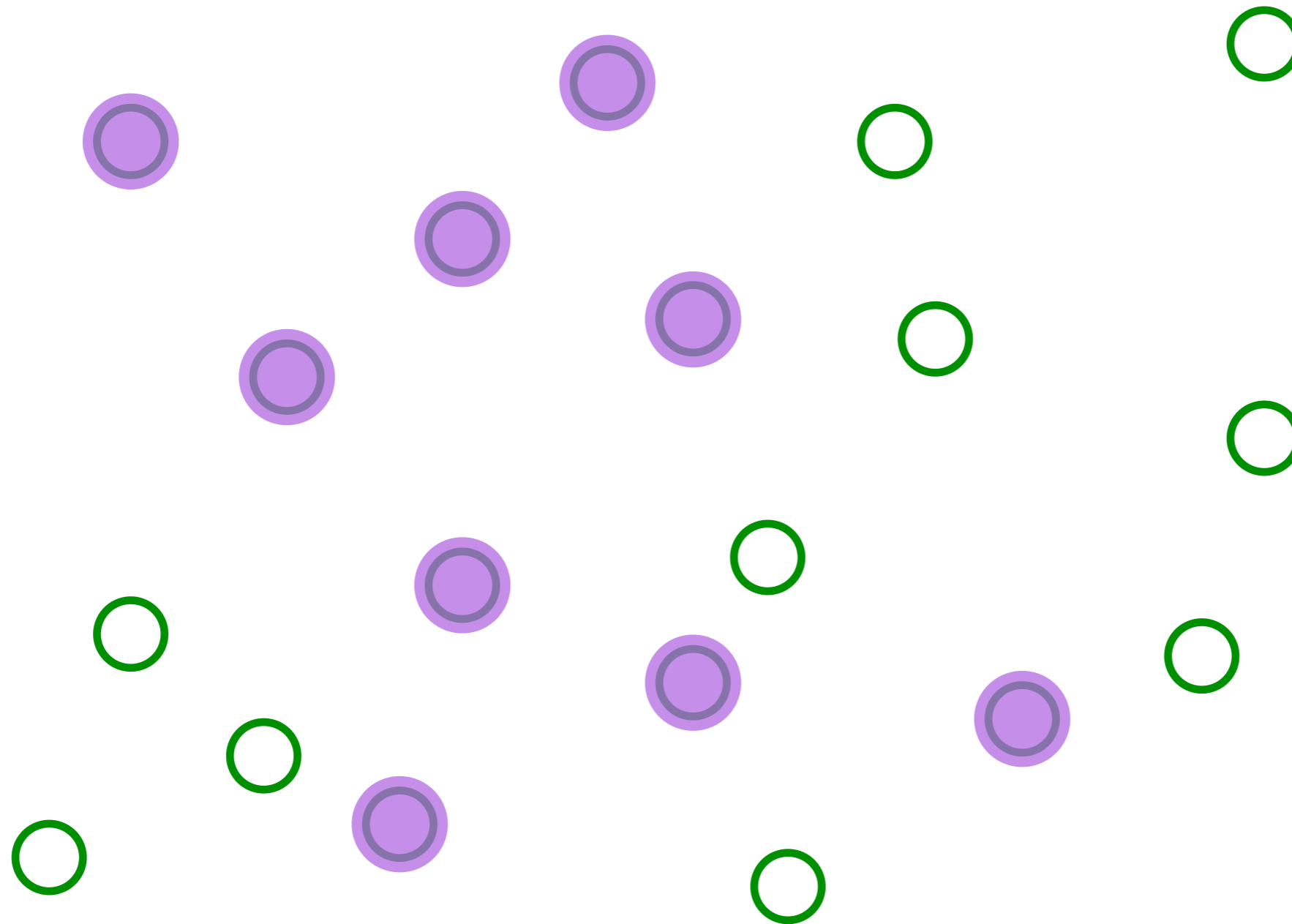
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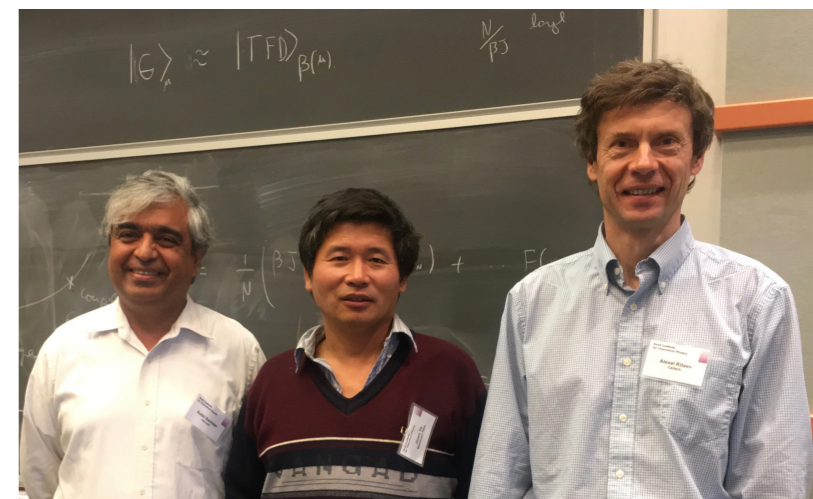
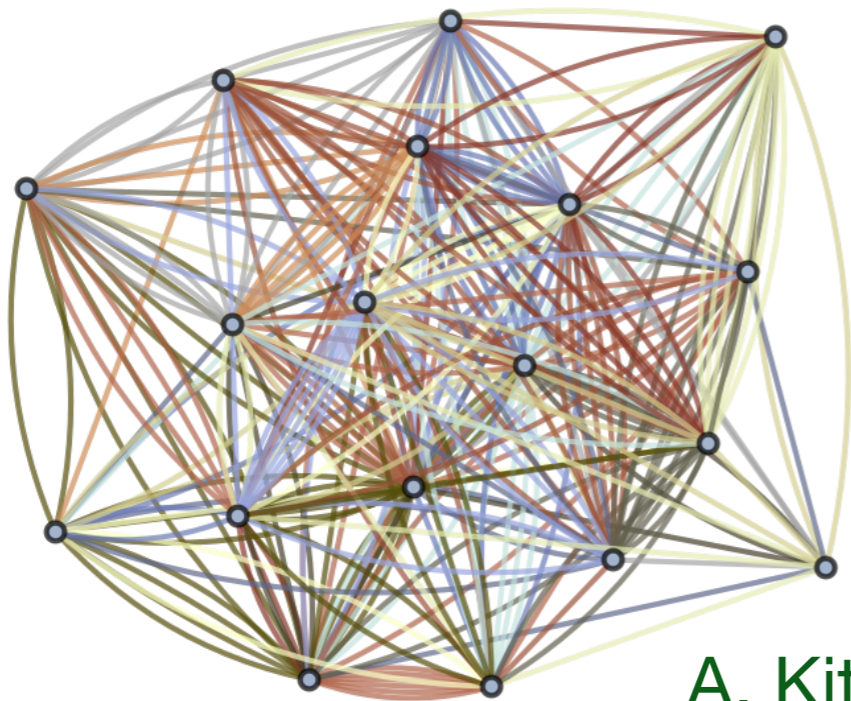
The Sachdev-Ye-Kitaev (SYK) model

(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large N limit;
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,\gamma,\delta=1}^N U_{ij;k\ell} c_i^\dagger c_j^\dagger c_k c_\ell - \mu \sum_i c_i^\dagger c_i$$

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$U_{ij;k\ell}$ are independent random variables with $\overline{U_{ij;k\ell}} = 0$ and $\overline{|U_{ij;k\ell}|^2} = U^2$
 $N \rightarrow \infty$ yields Planckian metal.

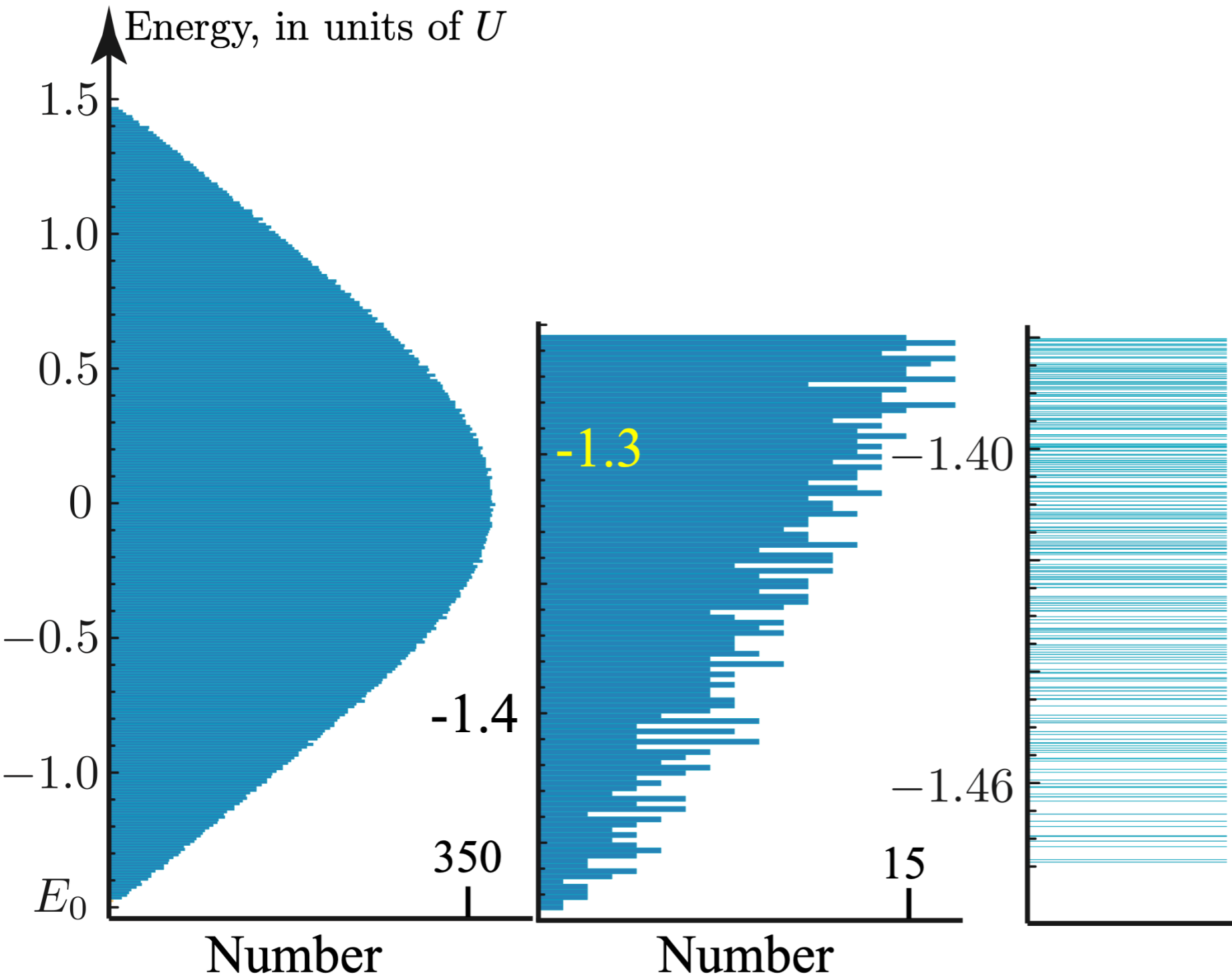


S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

Many-body density of states

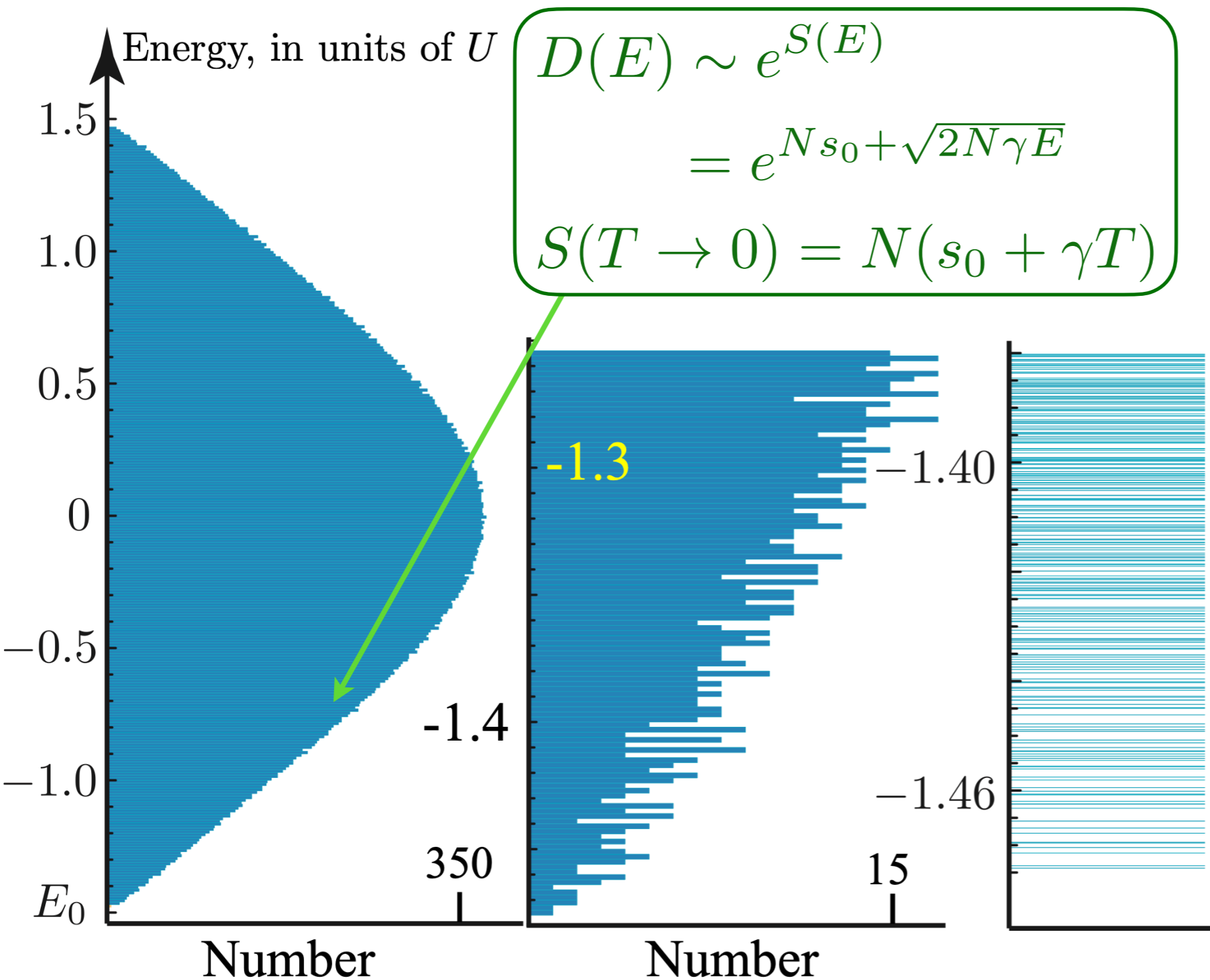
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Complex SYK model

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$$D(E) \sim e^{S(E)}$$
$$= e^{Ns_0 + \sqrt{2N\gamma E}}$$
$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

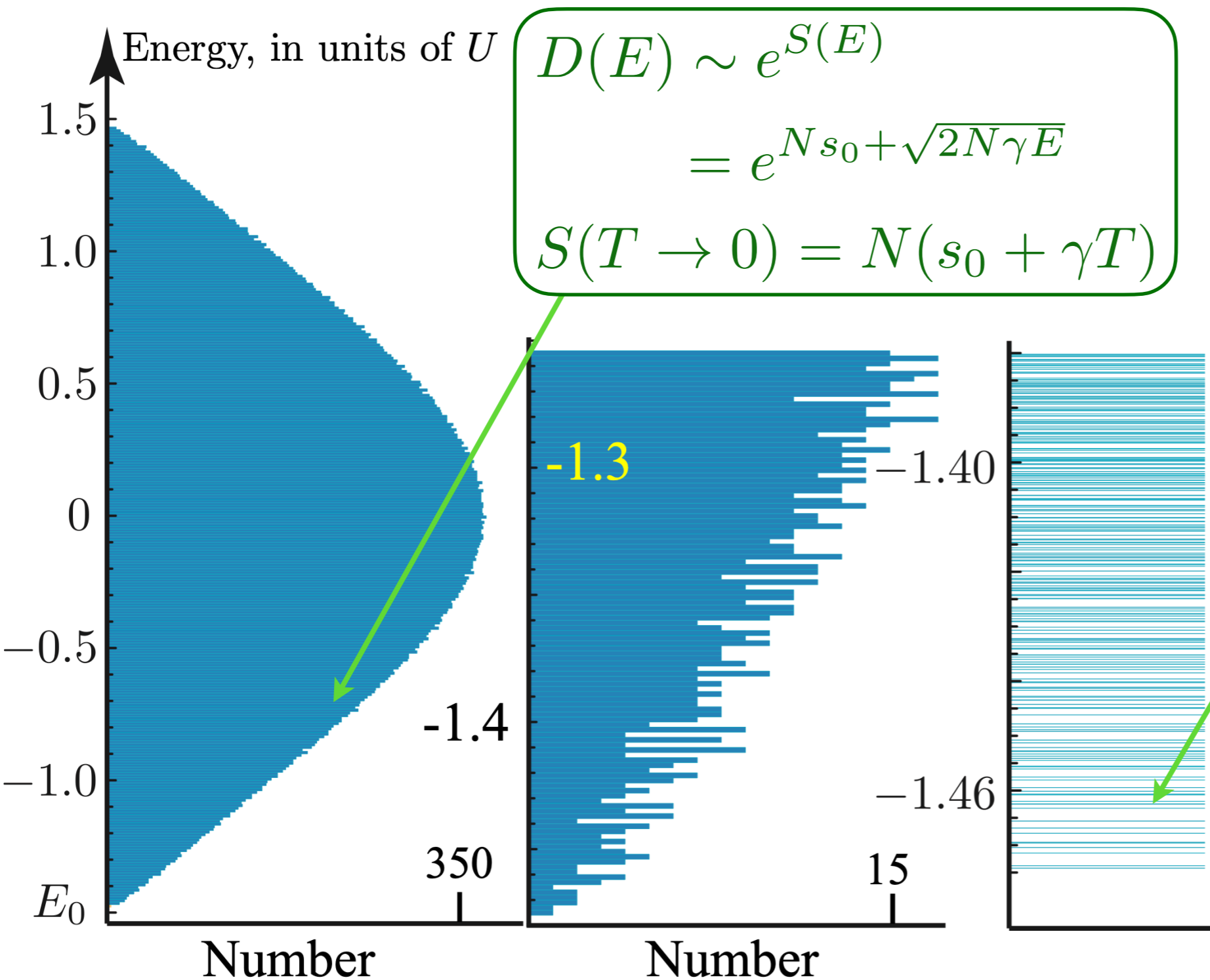
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A. Georges, O. Parcollet, and
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PRB **63**, 134406 (2001)

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$$D(E) \sim$$

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No quasiparticle
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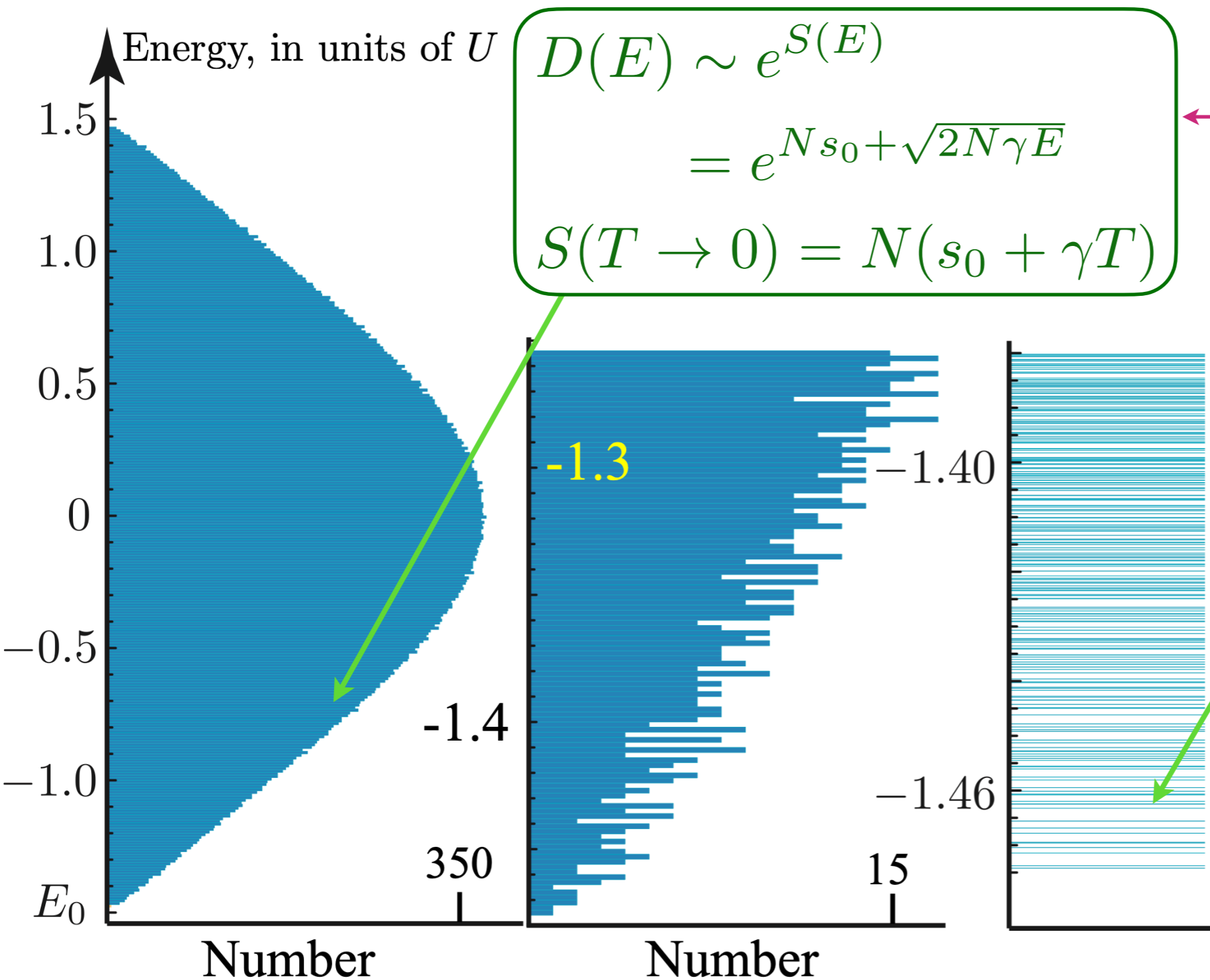
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$$D(E) \sim 2 e^{N s_0} \sinh(\sqrt{2N\gamma E})$$

$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln \left(\frac{U}{T} \right)$$

$$D(E) \sim 2 e^{N s_0} \sqrt{2N\gamma E}$$

No quasiparticle decomposition of many-body states

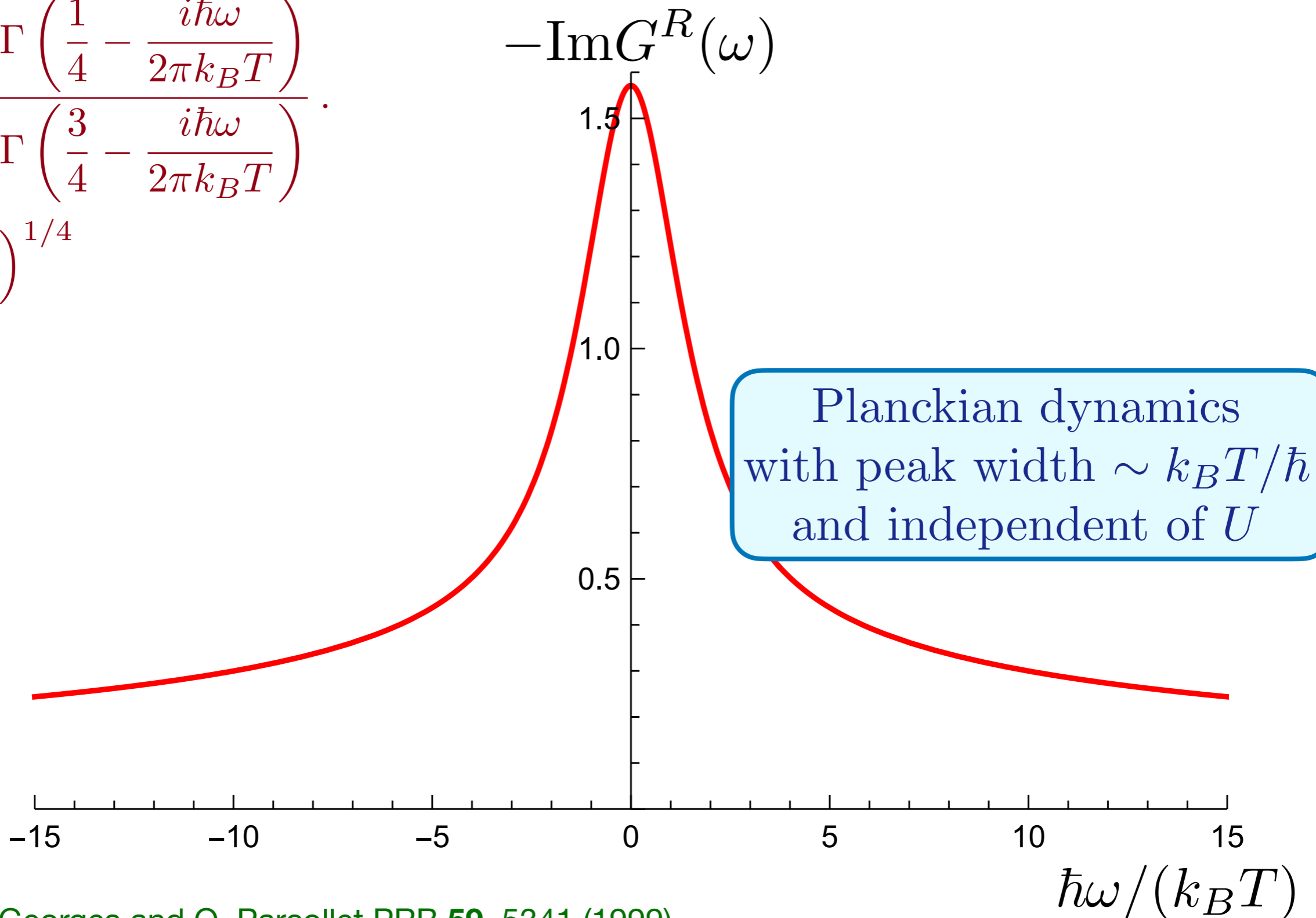
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Complex SYK model

The SYK model

$$G_{\text{SYK}}^R(\omega) = \frac{-iC}{(2\pi T)^{1/2}} \frac{\Gamma\left(\frac{1}{4} - \frac{i\hbar\omega}{2\pi k_B T}\right)}{\Gamma\left(\frac{3}{4} - \frac{i\hbar\omega}{2\pi k_B T}\right)}$$
$$C = \left(\frac{\pi}{U^2}\right)^{1/4}$$



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The SYK model

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- Low energy theory is invariant under time reparameterizations $\tau \rightarrow f(\tau)$.

The SYK model

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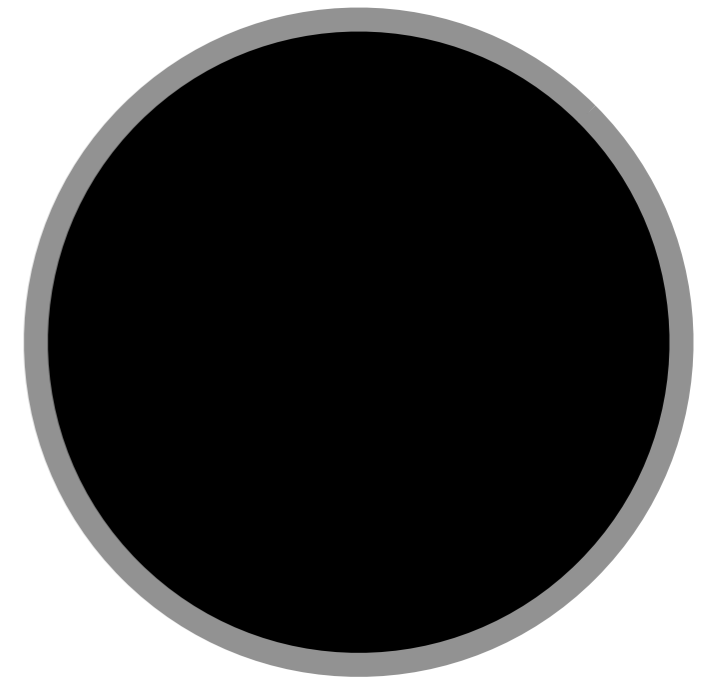
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 - Linear-in- T resistivity in the random t - J model
(after including *fractionalization*)

1. Introduction to the cuprates
2. Random matrix and SYK models
3. Black holes
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Black Holes

Objects so dense that light is gravitationally bound to them.

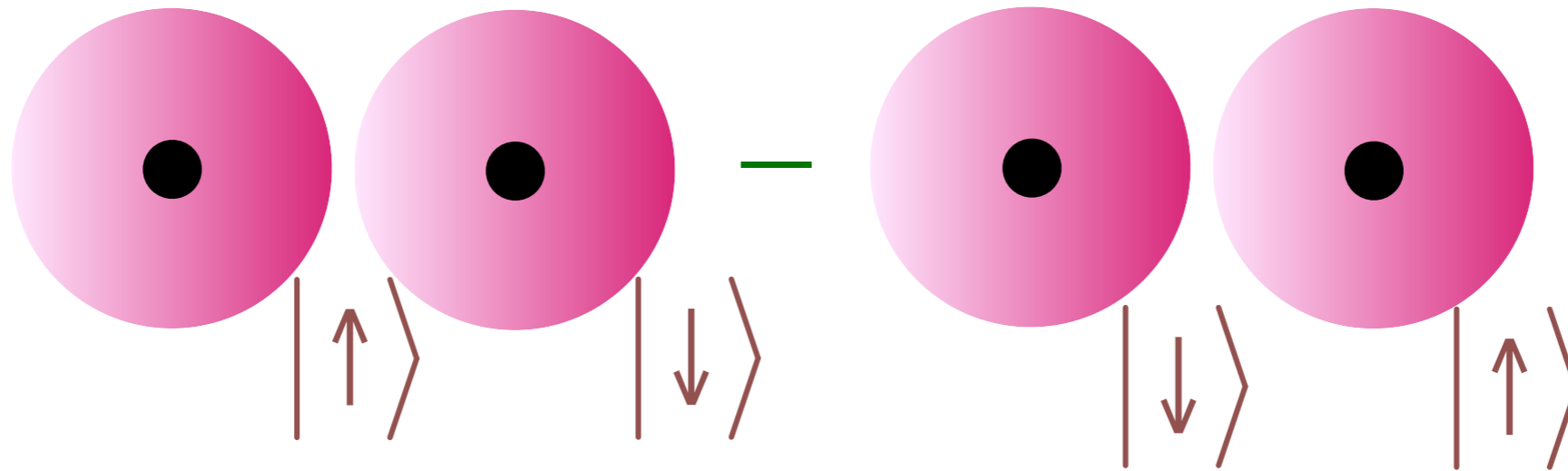
In Einstein's theory, the region inside the black hole **horizon** is disconnected from the rest of the universe.



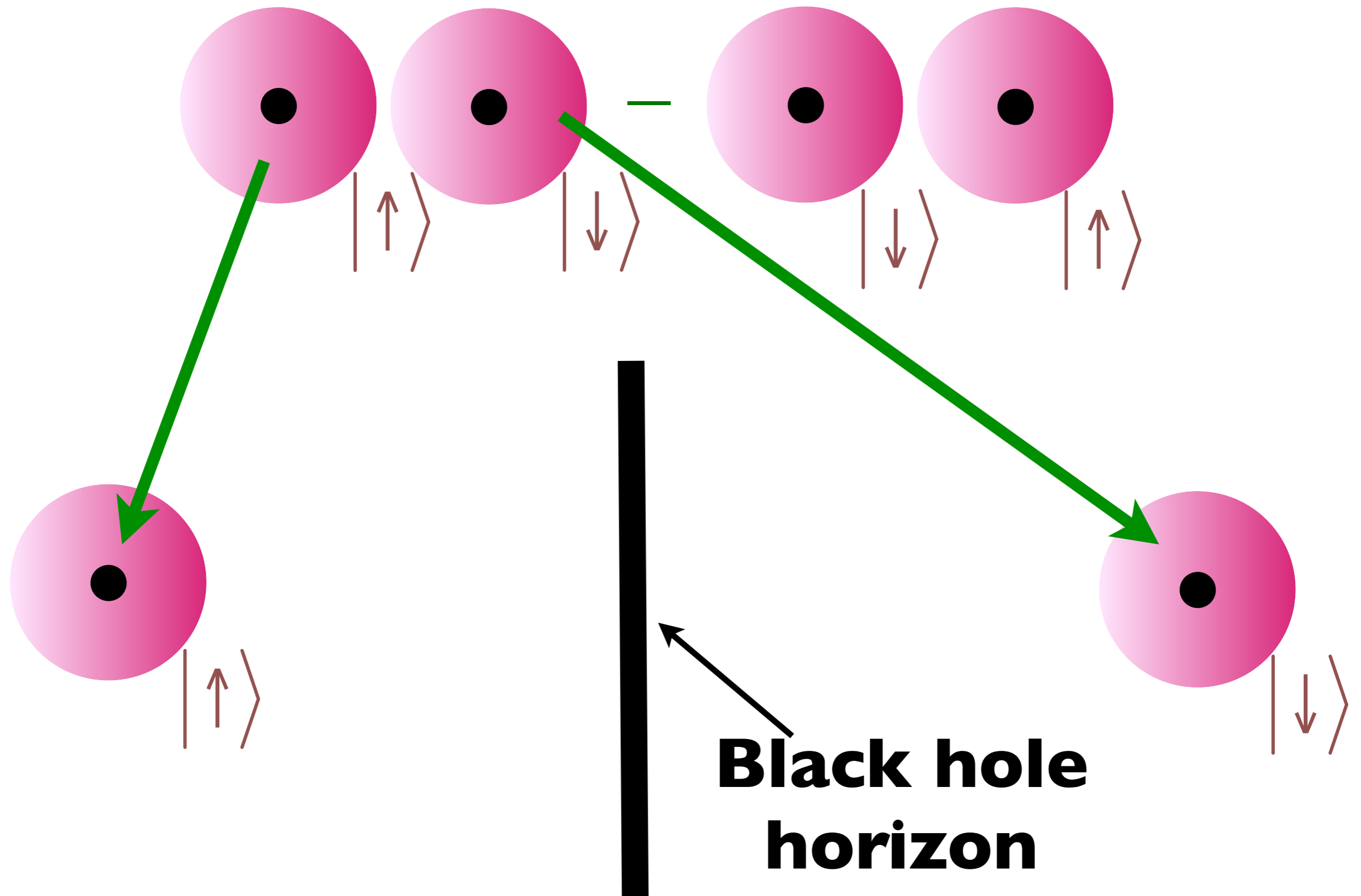
$$\text{Horizon radius } R = \frac{2GM}{c^2}$$

G Newton's constant, c velocity of light, M mass of black hole
For $M = \text{earth's mass}$, $R \approx 9 \text{ mm!}$

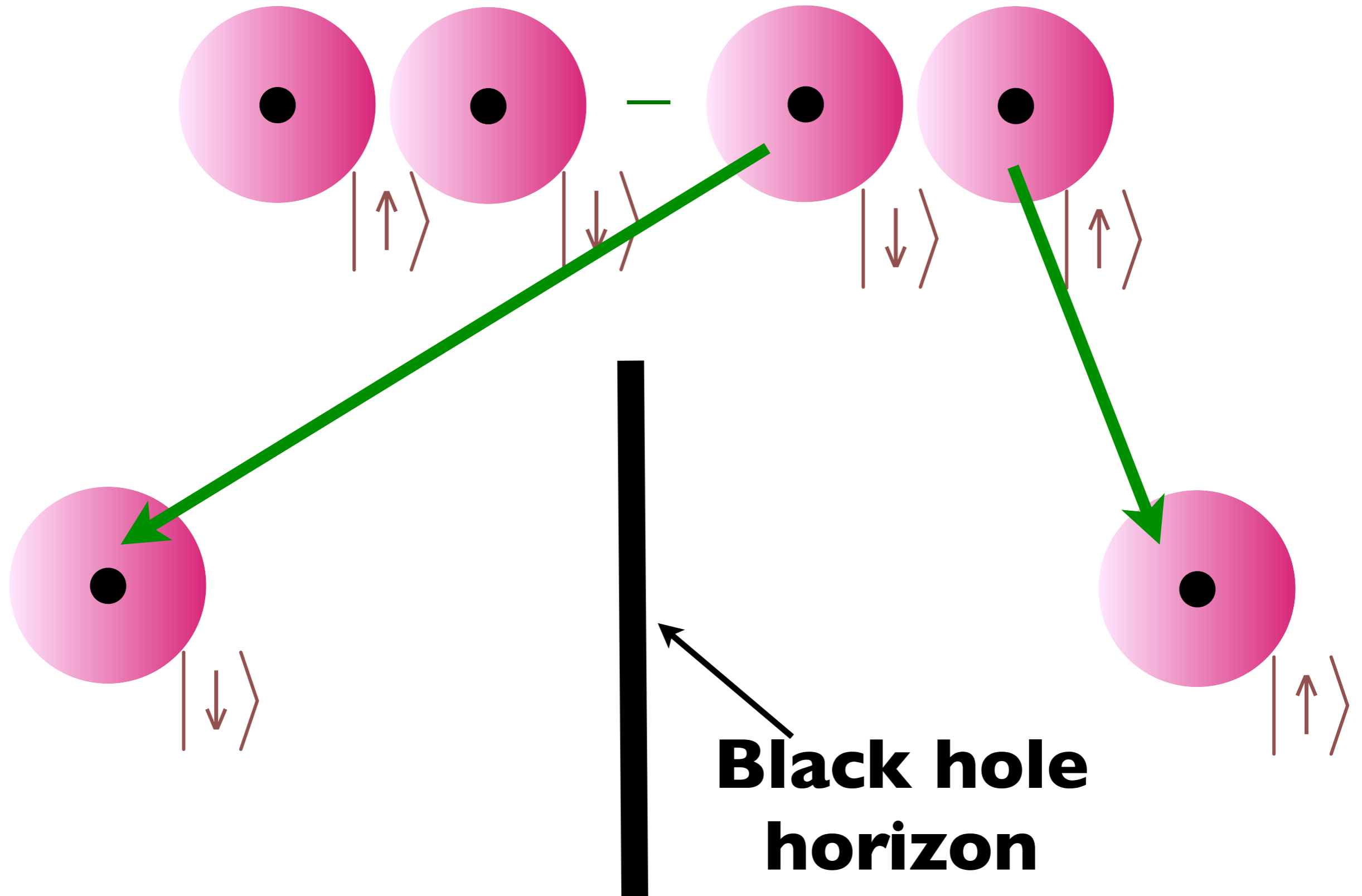
Quantum Entanglement across a black hole horizon



Quantum Entanglement across a black hole horizon

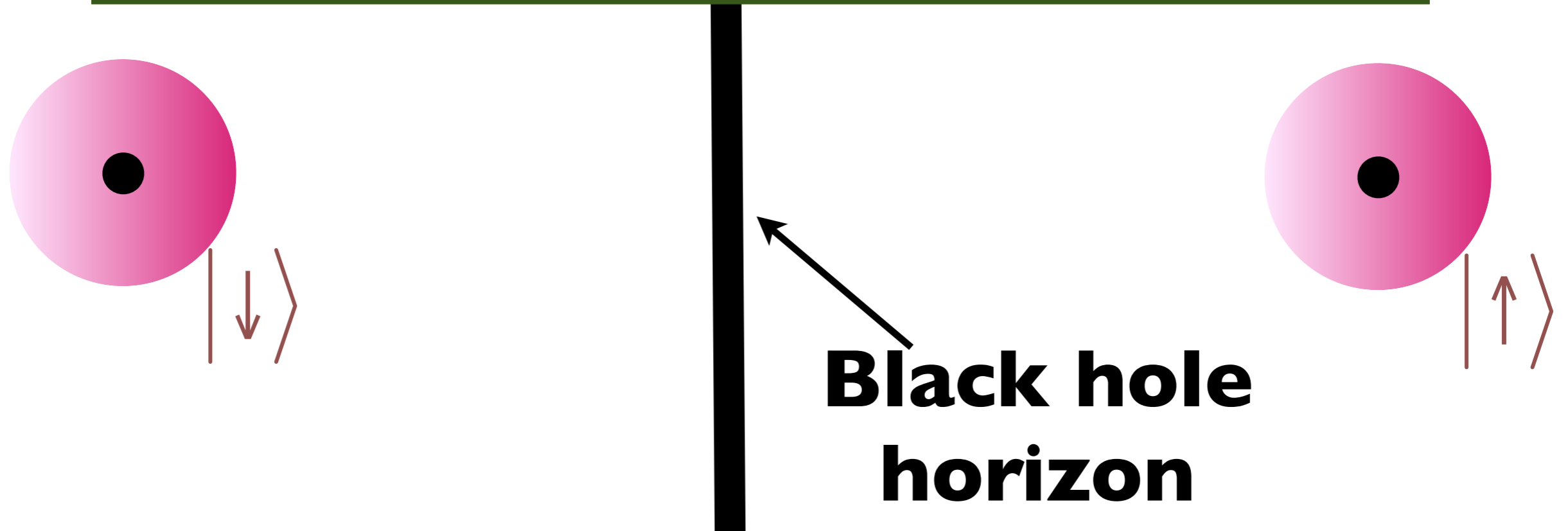


Quantum Entanglement across a black hole horizon



Quantum Entanglement across a black hole horizon

To an outside observer, the state of the electron inside the black hole is an unknown:
black holes have an entropy and a temperature



Quantum Black holes

- Black holes have an entropy and a temperature, T_H .
- The entropy, S_{BH} is proportional to their surface area.

Bekenstein (1973)
Hawking (1974)

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- What are the microscopic degrees of freedom that lead to black hole entropy (strings ?)

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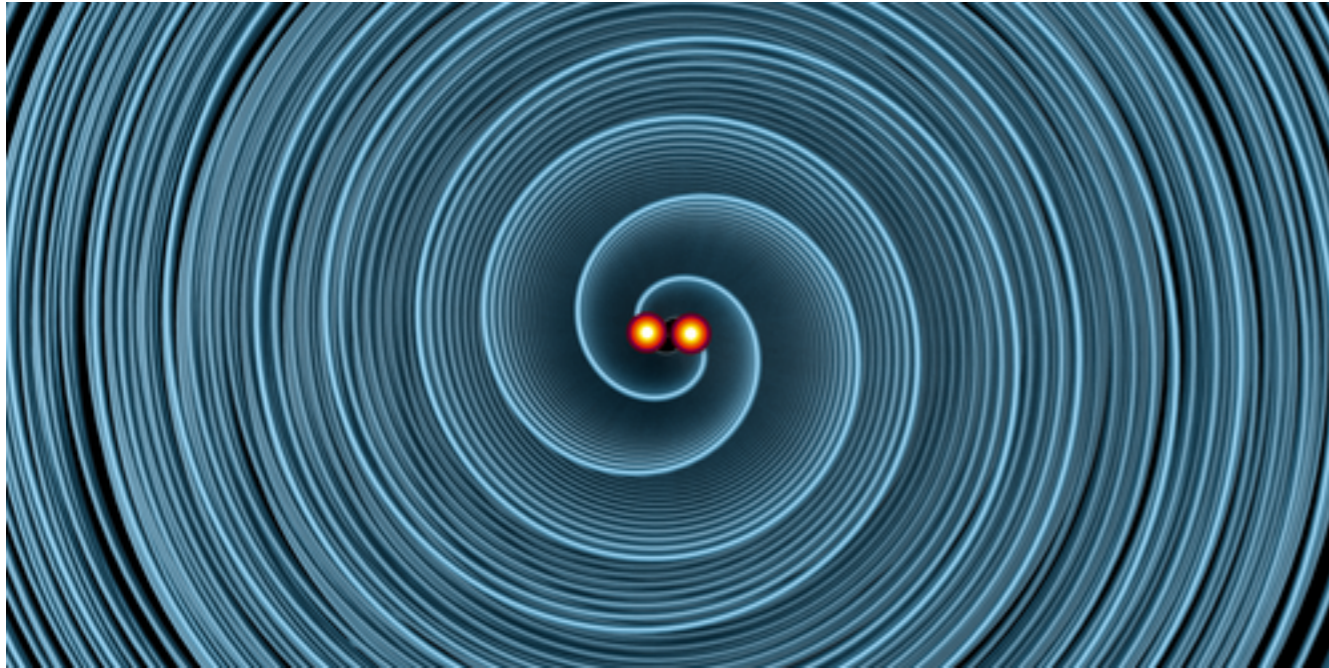
- What are the microscopic degrees of freedom that lead to black hole entropy (strings ?)
- How do we keep track of the quantum information carried by additional matter added to the black hole, and what happens to it when the black hole evaporates ?

Quantum Black holes

- Black holes have an entropy and a temperature, T_H .
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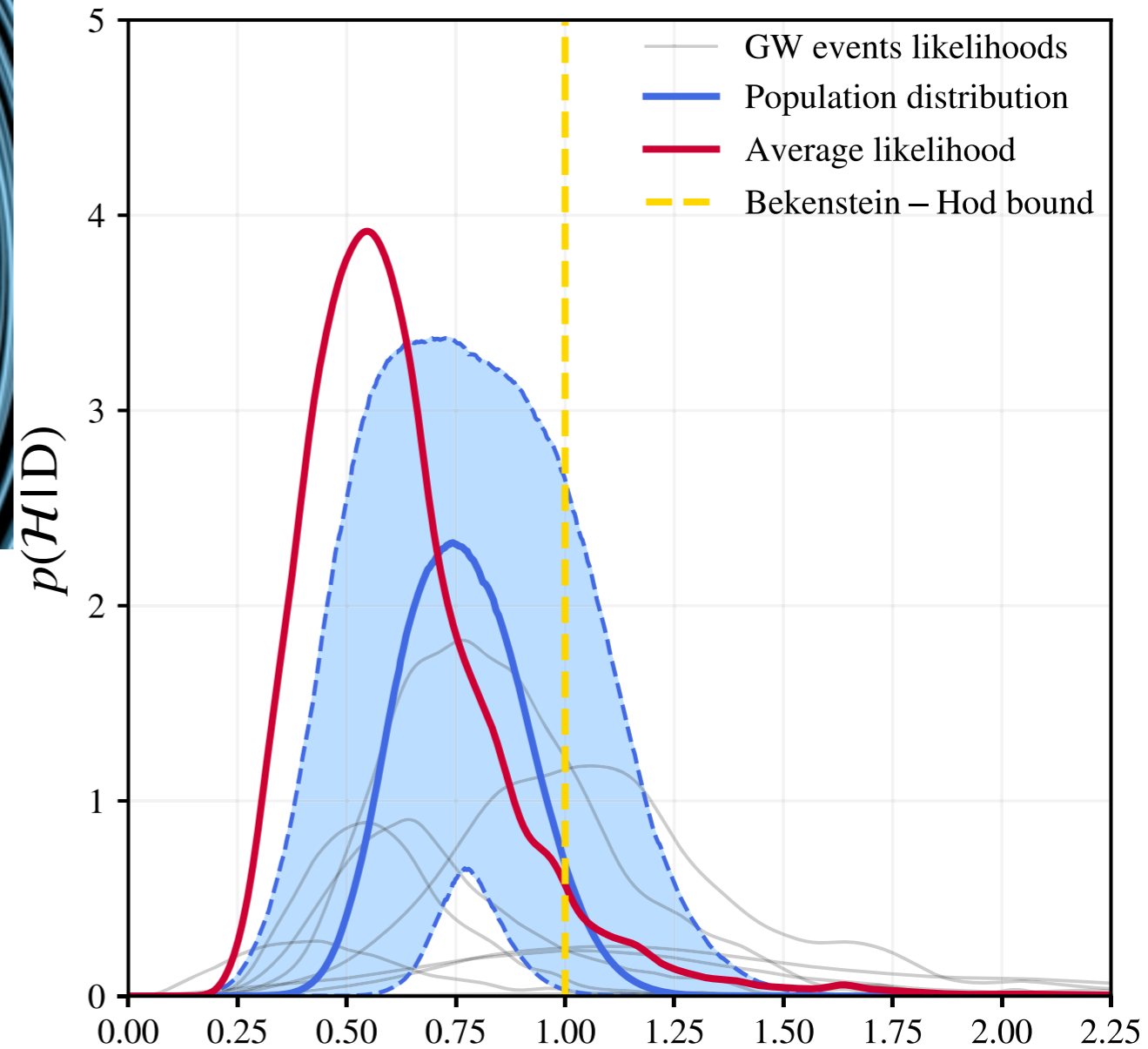
Recent developments: Many aspects of black hole entropy and quantum information are *independent* of microscopic (*i.e.* UV) details, and can be understood in a semiclassical theory of Einstein gravity.

G. Carullo, D. Laghi, J. Veitch, W. Del Pozzo, PRL **126**, 161102 (2021);
S. Hod, PRD **75**, 064013 (2007)



Gravity wave observations of 8 different black holes show a relaxation time

$$\tau \sim \frac{\hbar}{k_B T_H}$$



$$\mathcal{H} = \frac{1}{\pi} \frac{\hbar/\tau}{k_B T_H}$$

Quantum Black holes

- Black holes have an entropy and a temperature, T_H .
- The entropy, S_{BH} is proportional to their surface area.
- They relax to thermal equilibrium with a Planckian rate $k_B T_H / \hbar$, implying connection to non-quasiparticle dynamics.

Recent developments: Many aspects of black hole entropy and quantum information are *independent* of microscopic (*i.e.* UV) details, and can be understood in a semiclassical theory of Einstein gravity.



Holographic Metals and the Fractionalized Fermi Liquid

Subir Sachdev

Department of Physics, Harvard University, Cambridge, Massachusetts 02138, USA

(Received 23 June 2010; published 4 October 2010)

We show that there is a close correspondence between the physical properties of holographic metals near charged black holes in anti-de Sitter (AdS) space, and the fractionalized Fermi liquid phase of the lattice Anderson model. The latter phase has a “small” Fermi surface of conduction electrons, along with a spin liquid of local moments. This correspondence implies that certain mean-field gapless spin liquids are states of matter at nonzero density realizing the near-horizon, $\text{AdS}_2 \times \text{R}^2$ physics of Reissner-Nordström black holes.



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String theorists (2010): Models with random couplings ????



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String theorists (2010): Models with random couplings ????

Current understanding: Many key properties of suitable random Hamiltonians self-average, and allow easier access to the physics of strongly entangled and chaotic quantum matter

Thermodynamics of quantum black holes:

$$\mathcal{Z} = \int \mathcal{D}g_{\mu\nu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity}}^{(3+1)} [g_{\mu\nu}] \right)$$



Metric of
spacetime

In general, this summation is not well defined, because to the uncontrollably large number of spacetime configurations.

Thermodynamics of quantum black holes:

$$\mathcal{Z} = \int \mathcal{D}g_{\mu\nu} \exp\left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity}}^{(3+1)}[g_{\mu\nu}]\right)$$
$$= \exp(S_{BH}) \times \left(\dots????\dots \right)$$

Metric of spacetime

$$S_{BH} = \frac{A(T_H)c^3}{4G\hbar}$$

Gibbons, Hawking (1977)

With $\mathcal{Z} = \text{Tr}e^{-\mathcal{H}/(k_B T_H)}$,

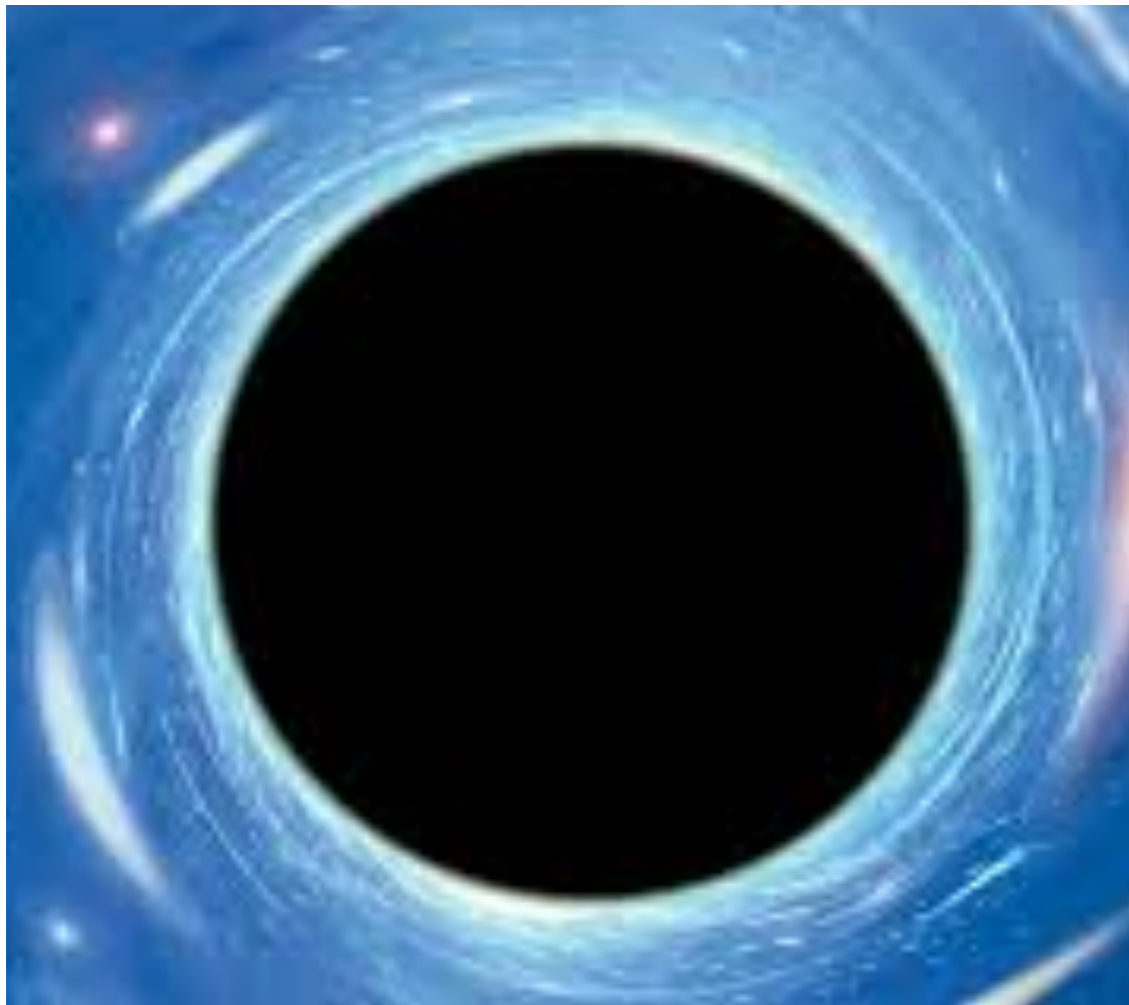
$\hbar/(k_B T_H)$ is the length of the Euclidean time circle.

$A(T_H)$ is the area of the black hole horizon at a temperature T_H .

Interpretation: Black hole entropy is entanglement entropy across the horizon.

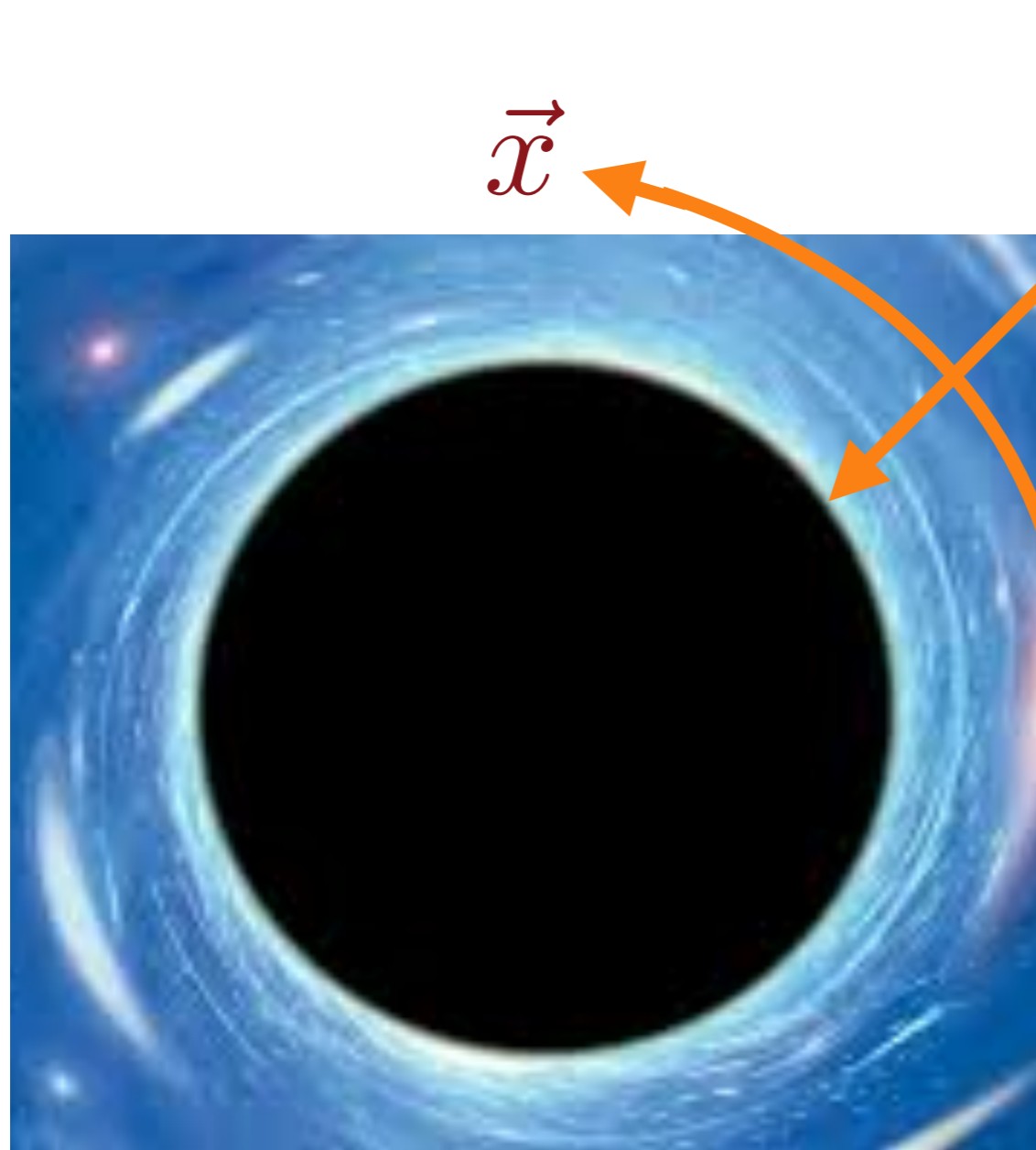


Maxwell's electromagnetism
and Einstein's general relativity
allow black hole solutions with a net charge





Maxwell's electromagnetism
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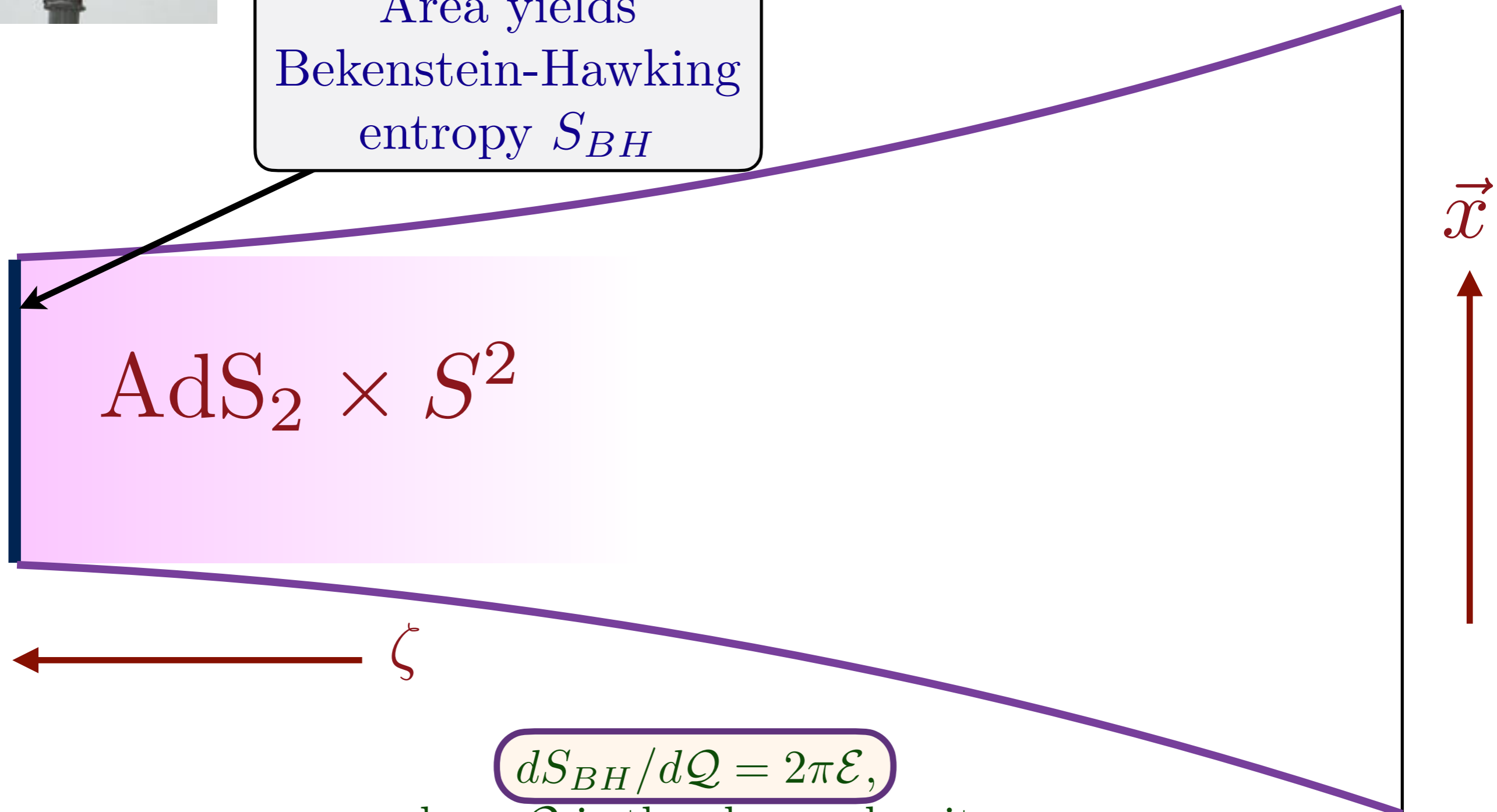


Zooming into the near-horizon region of a charged black hole at low temperature, yields a gravitational theory in one space (ζ) and one time dimension

SYK model and charged black holes



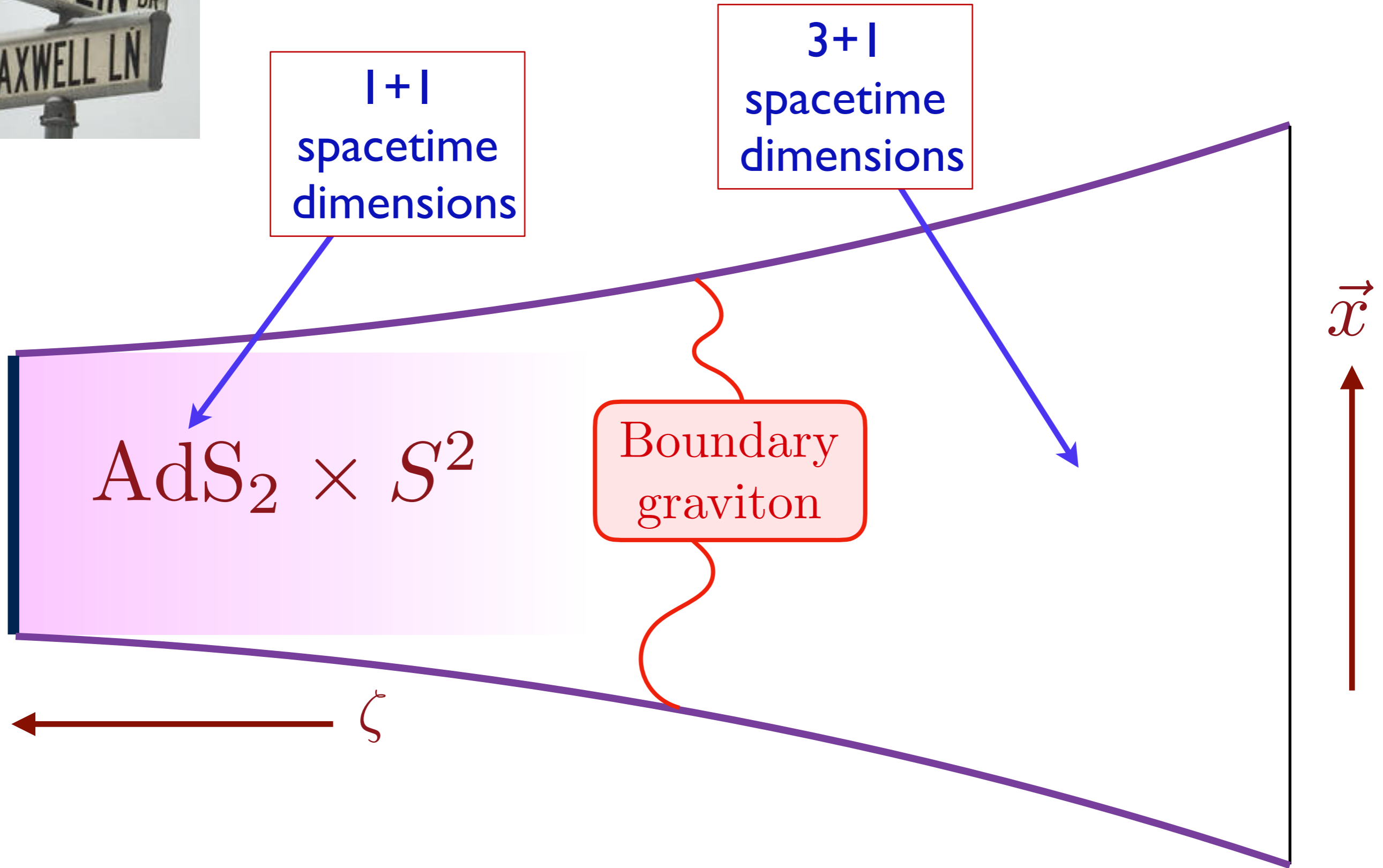
Horizon
Area yields
Bekenstein-Hawking
entropy S_{BH}



$$dS_{BH}/dQ = 2\pi\mathcal{E},$$

where Q is the charge density,
and \mathcal{E} is the electric field on the horizon.

SYK model and charged black holes



SYK model and charged black holes



1+1
spacetime
dimensions

3+1
spacetime
dimensions

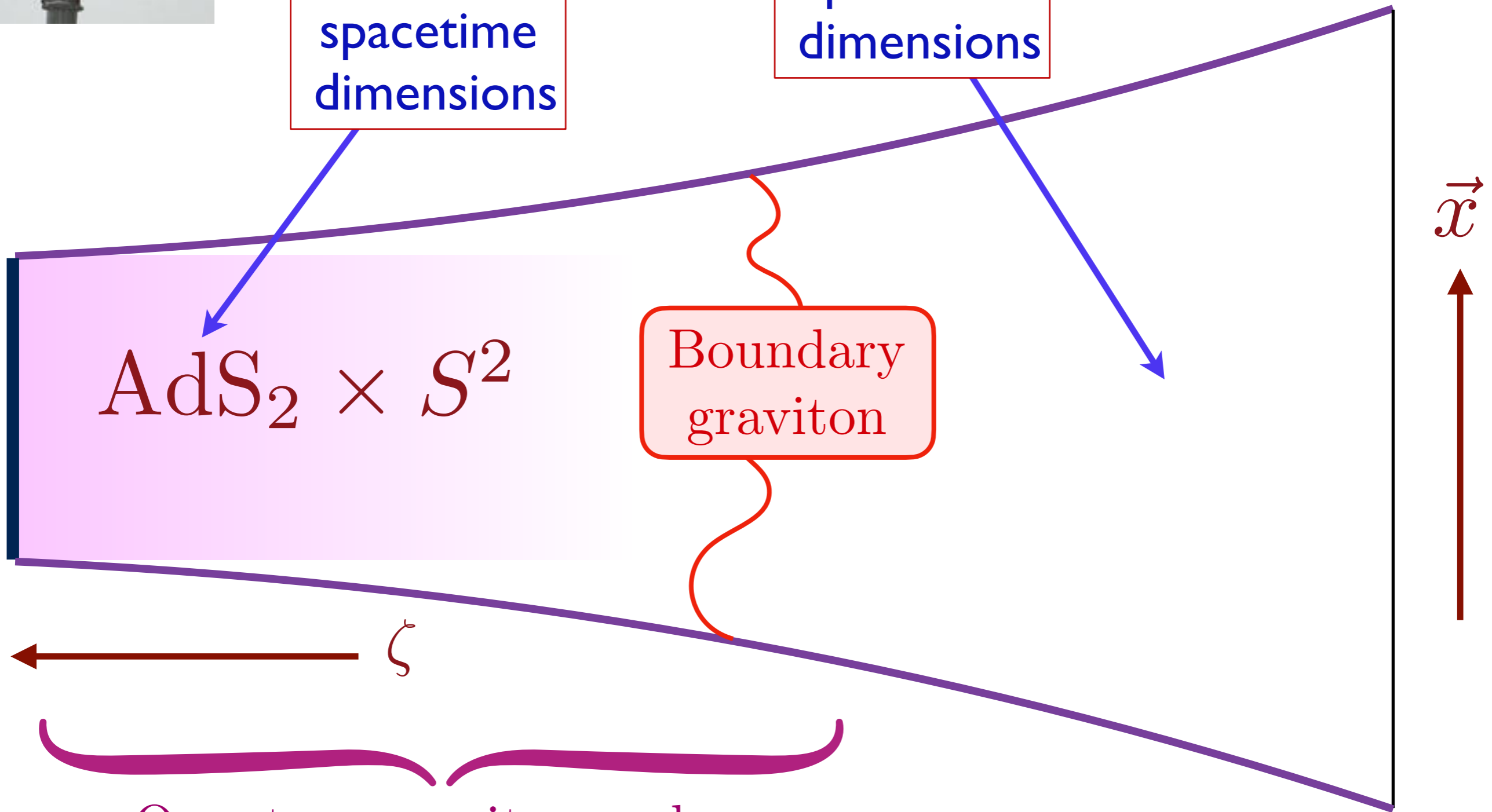
$\text{AdS}_2 \times S^2$

Boundary
graviton

\vec{x}

ζ

Quantum gravity can be
exactly solved in this region!



Thermodynamics of quantum black holes with charge Q :

$$\mathcal{Z} = \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right)$$
$$= \exp(S_{BH}) \times \left(\dots????\dots \right)$$

Gibbons, Hawking (1977)

$$S_{BH}(T_H \rightarrow 0, Q) = \frac{A(T_H)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2}T_H}{\hbar c} \right)$$

A_0 is the area of the charged black hole horizon at $T_H = 0$.

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$$\mathcal{Z} = \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right) \\ \approx e^{S_{BH}} \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{JT gravity of AdS}_2 \text{ and boundary}}^{(1+1)}[g_{\mu\nu}, A_{\mu}] \right)$$

$$S_{BH}(T_H \rightarrow 0, Q) = \frac{A(T_H)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2}T_H}{\hbar c} \right)$$

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Thermodynamics of quantum black holes with charge Q :

$$\begin{aligned} \mathcal{Z} &= \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_\mu] \right) \\ &= e^{S_{BH}} \int \mathcal{D}f(\tau) \exp \left(\begin{aligned} &-\text{Boundary graviton action } [f] \\ &= \text{SYK low energy action}[f]! \end{aligned} \right) \end{aligned}$$

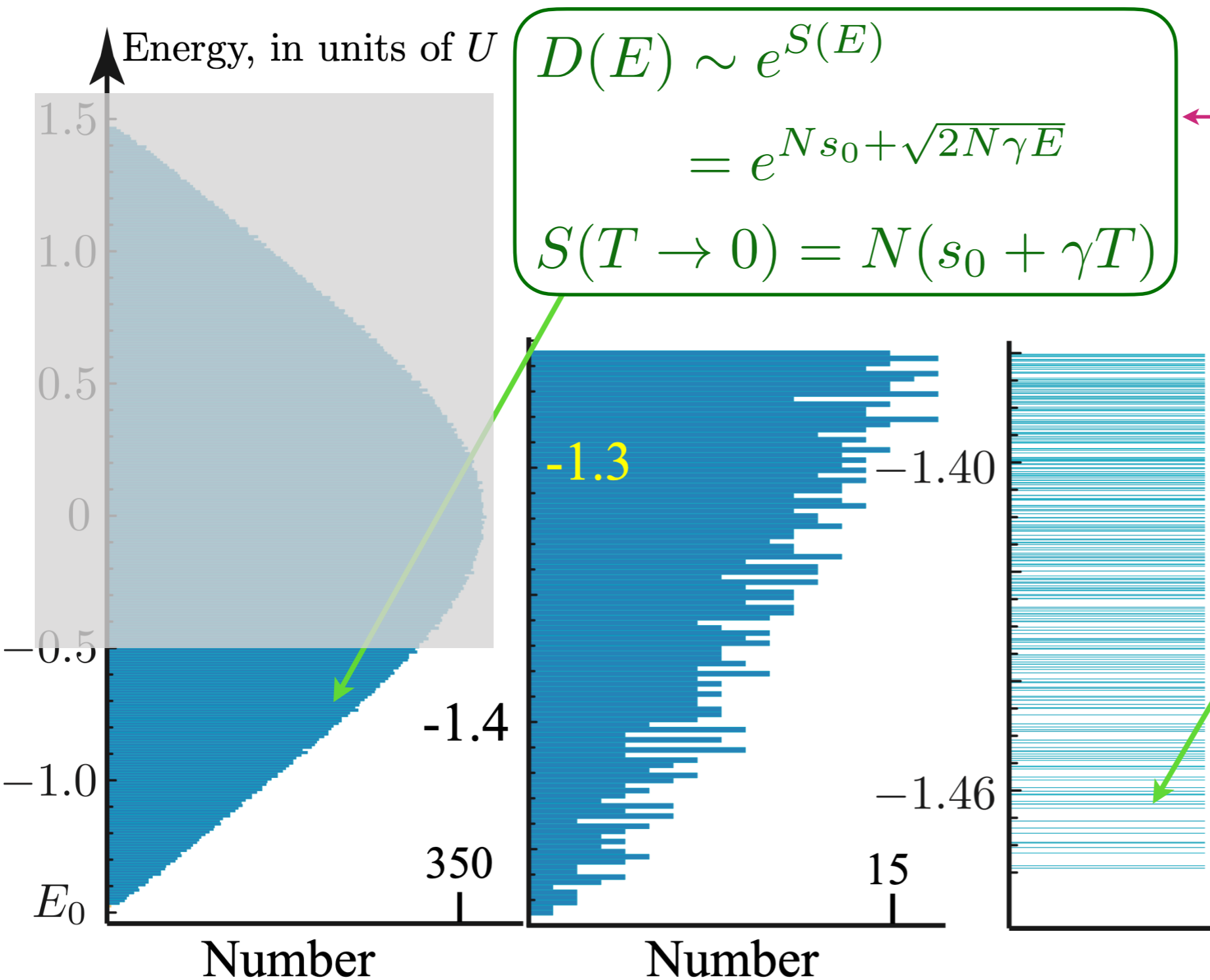
$$S(T_H \rightarrow 0, Q) = S_{BH} - \frac{3}{4} \ln \left(\frac{\hbar c^5}{GT_H^2} \right)$$

$$S_{BH} = \frac{A(T_H)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2}T_H}{\hbar c} \right)$$

A_0 is the area of the charged black hole horizon at $T_H = 0$. The $\ln(1/T_H)$ term is the contribution of the boundary graviton, and is also independent of UV details!

Many-body density of states

$$D(E) = \sum_a \delta(E - E_a); \quad E_0 + E_a \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim e^{S(E)}$$

$$= e^{N s_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{N s_0} \sinh(\sqrt{2N\gamma E})$$

$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln\left(\frac{U}{T}\right)$$

$$D(E) \sim 2 e^{N s_0} \sqrt{2N\gamma E}$$

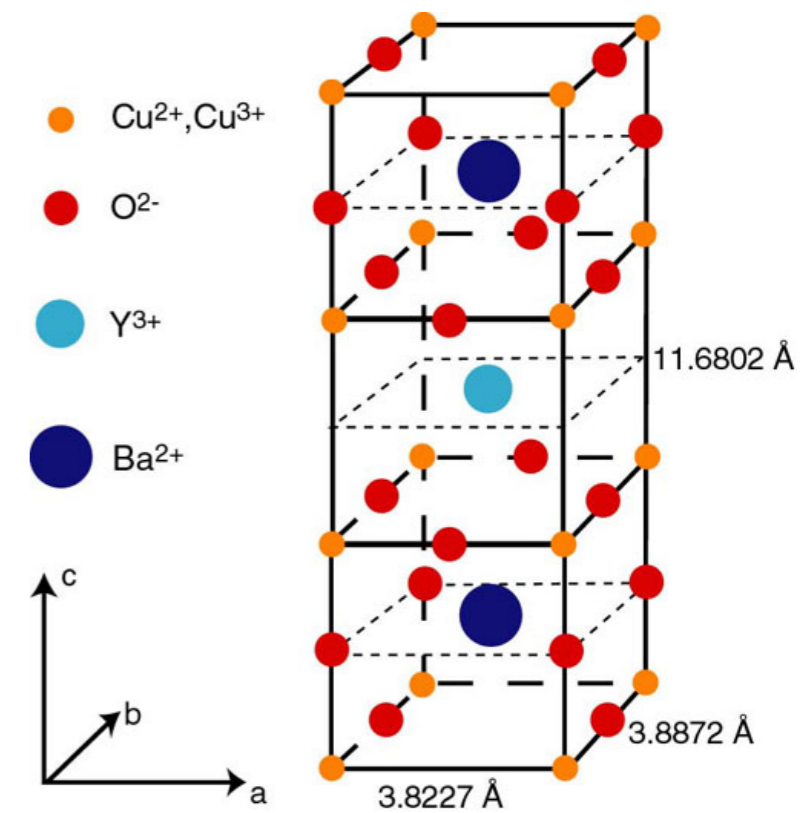
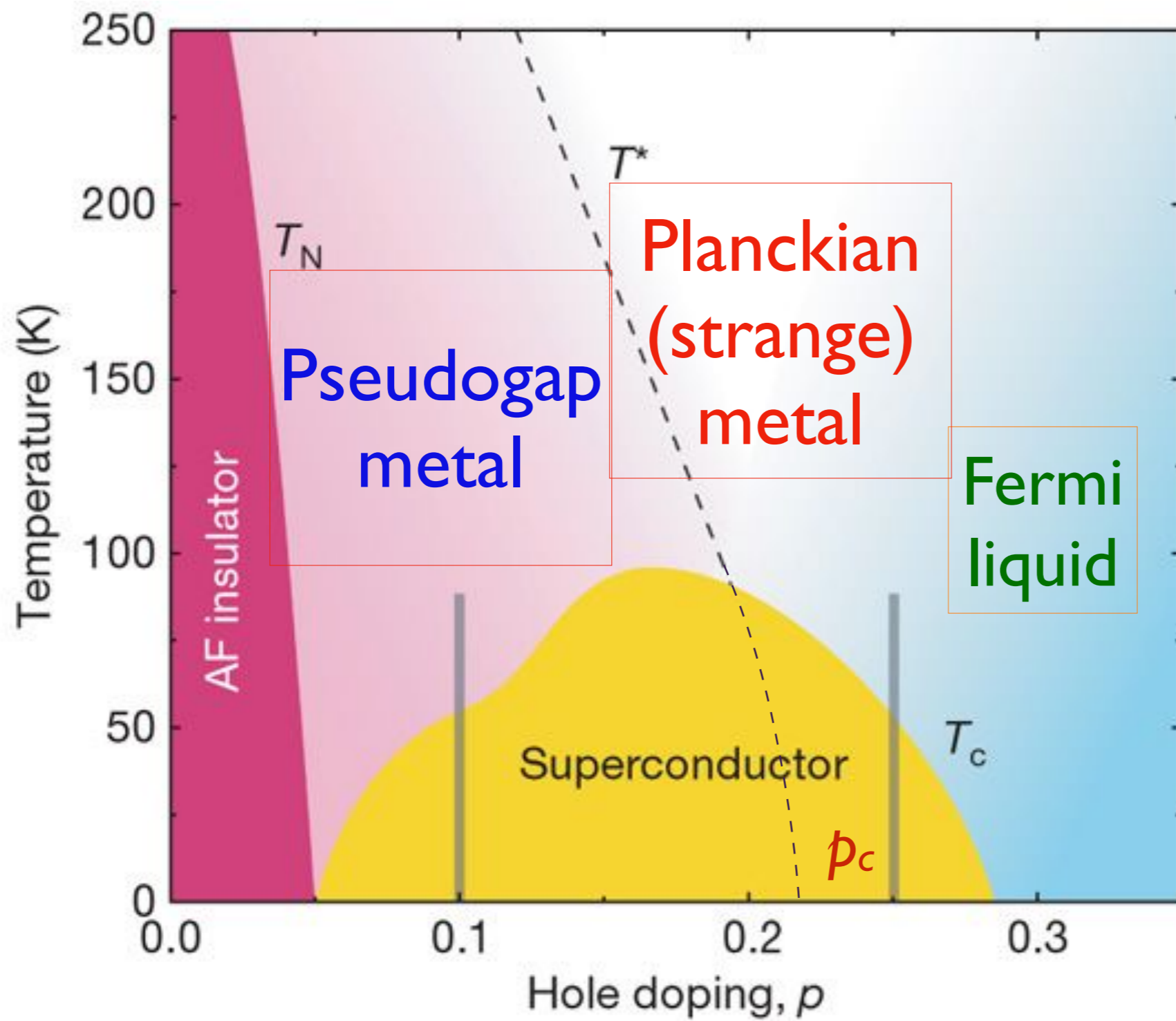
No quasiparticle decomposition of many-body states

$$s_0 = 0.464848 \dots$$

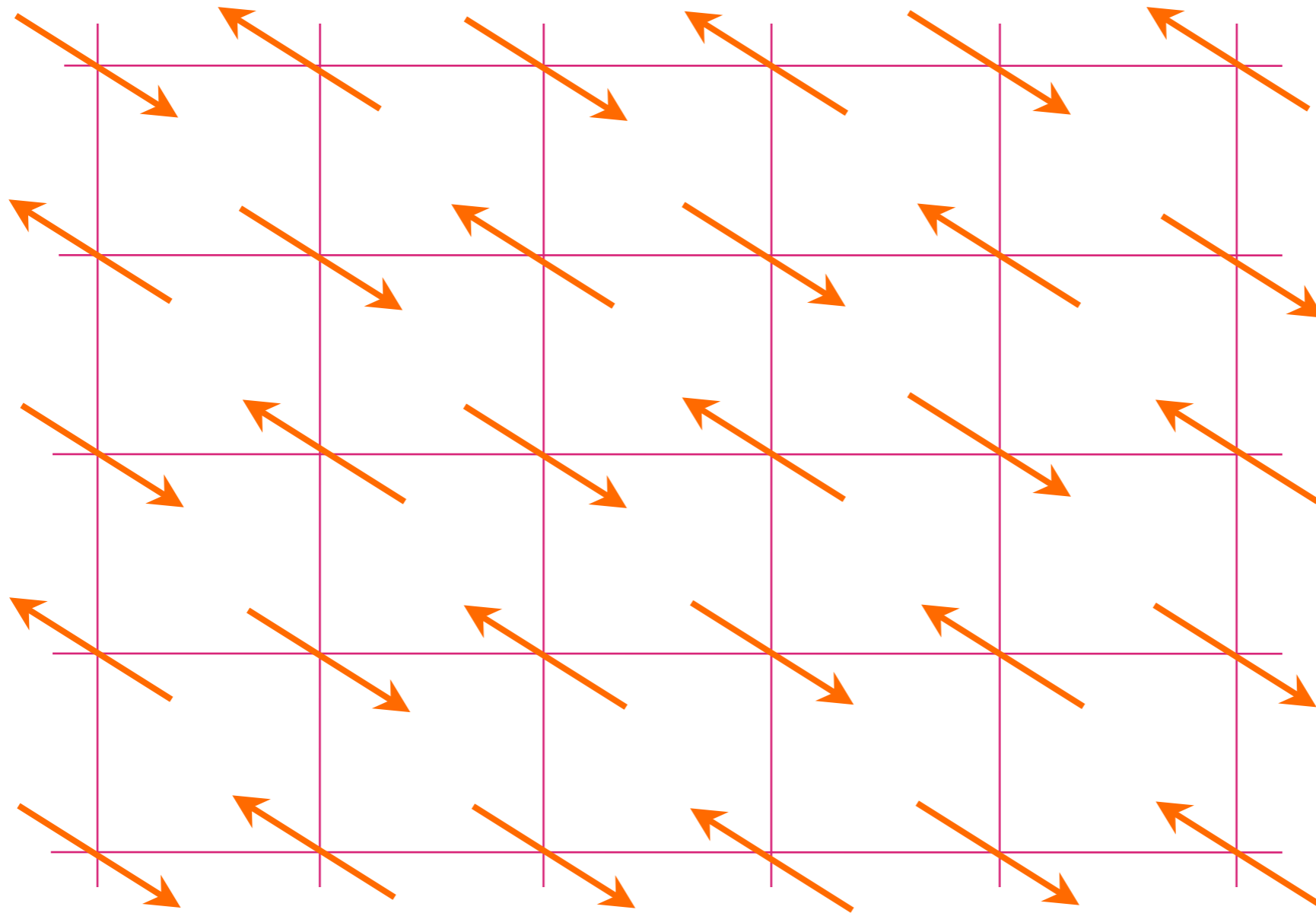
A. Georges, O. Parcollet, and S. Sachdev,
PRB **63**, 134406 (2001)

Complex SYK model

1. Introduction to the cuprates
2. Random matrix and SYK models
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4. Back to the cuprates

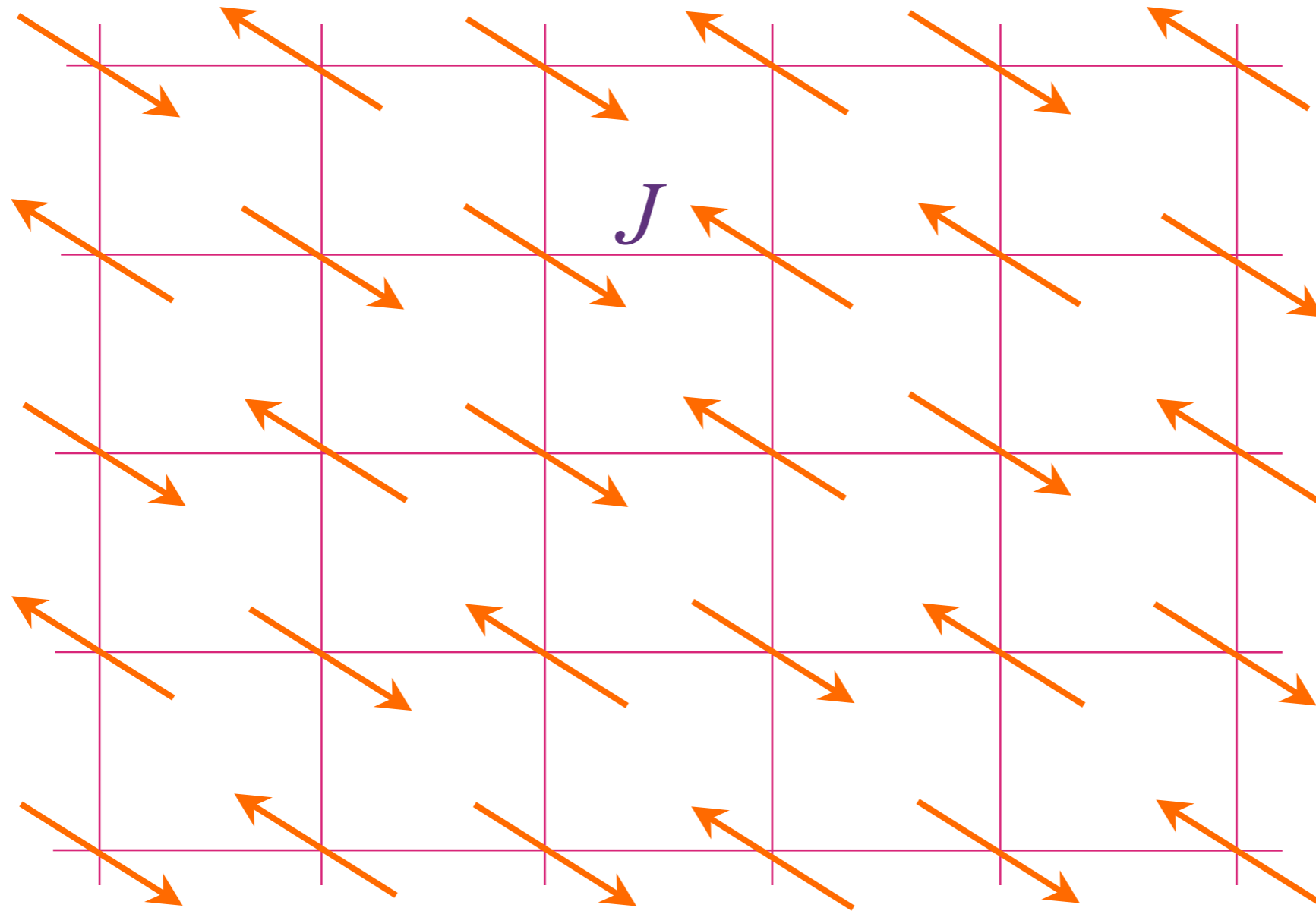


Insulating antiferromagnet



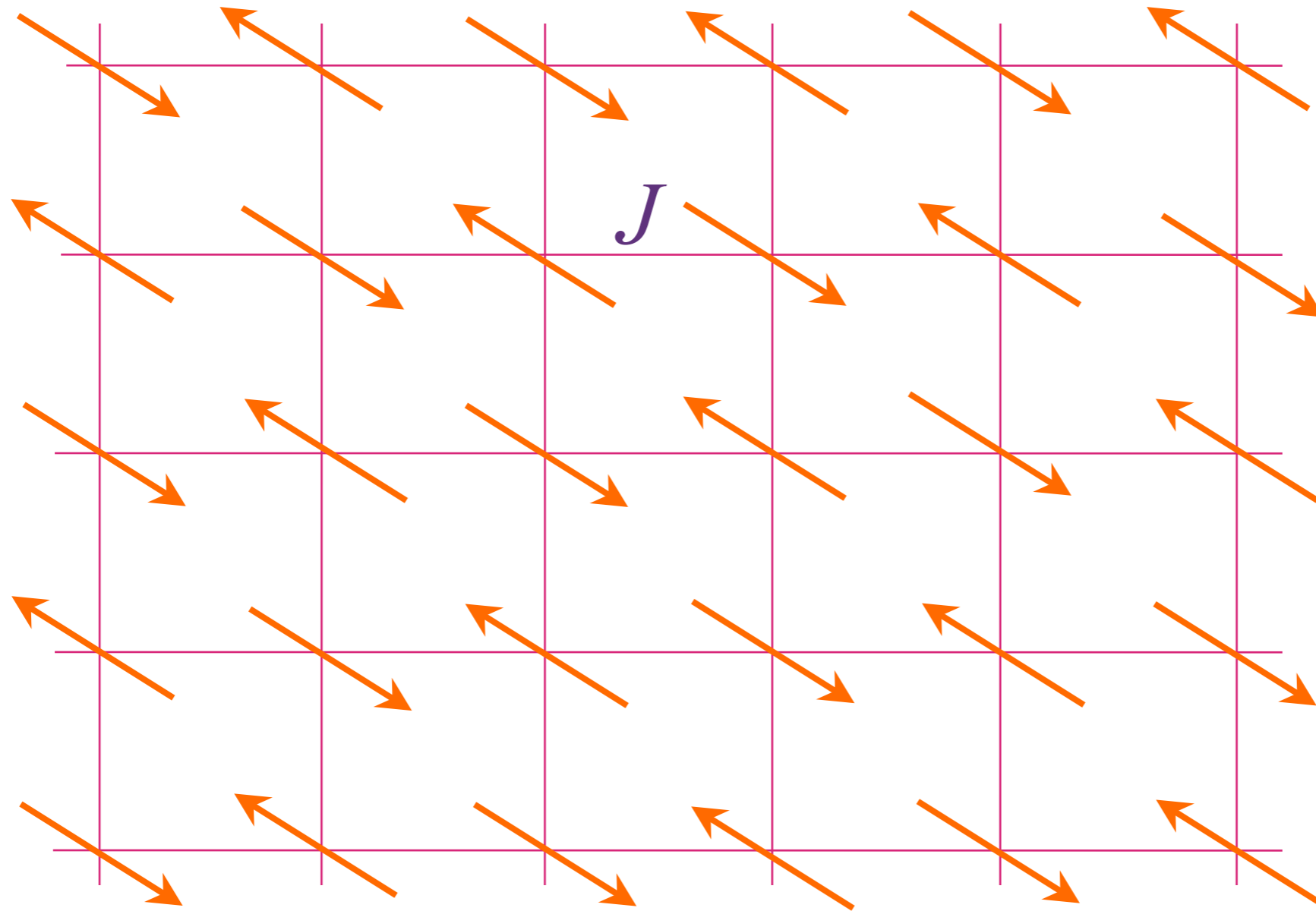
$$p=0$$

Insulating antiferromagnet



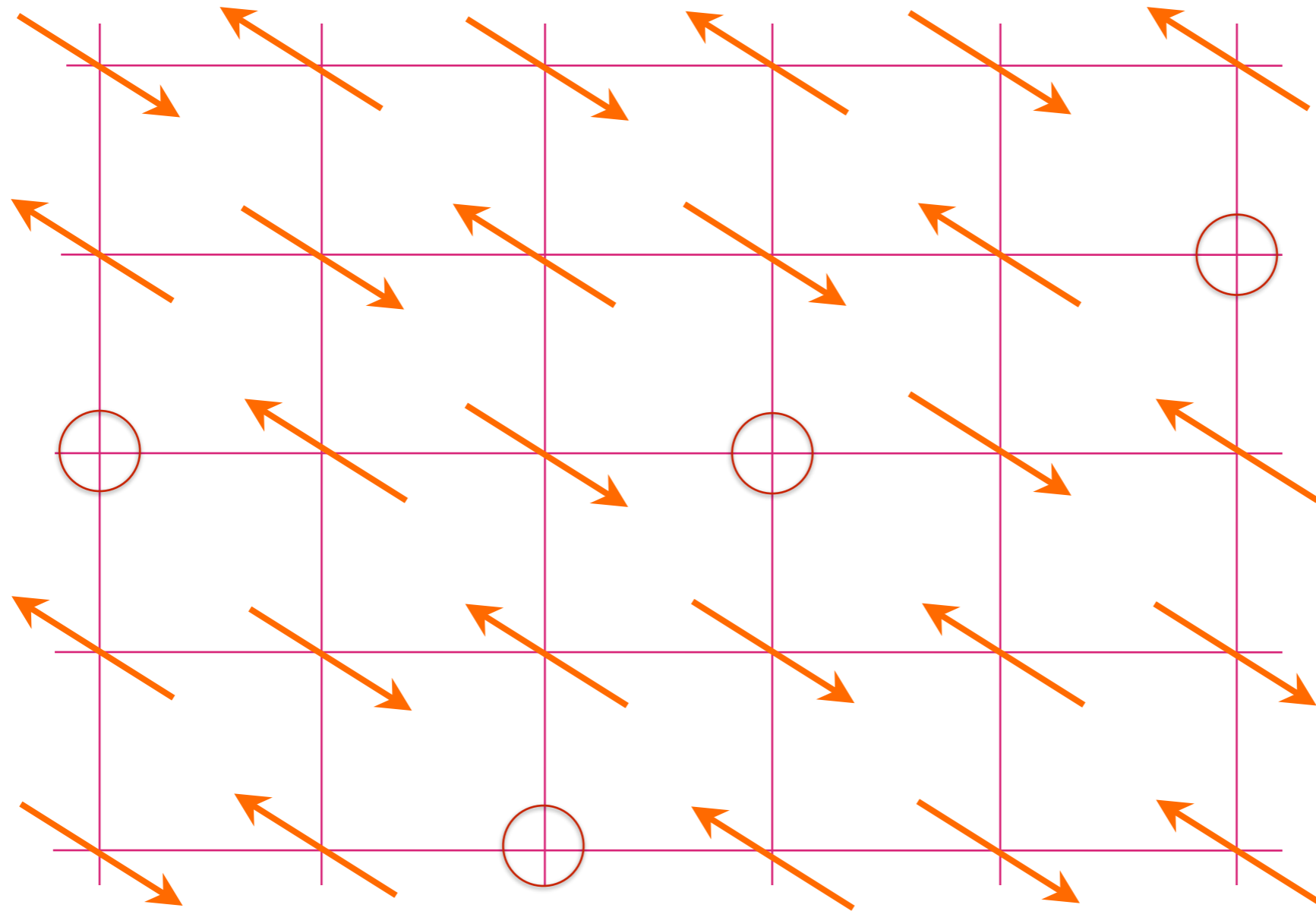
$$p=0$$

Insulating antiferromagnet

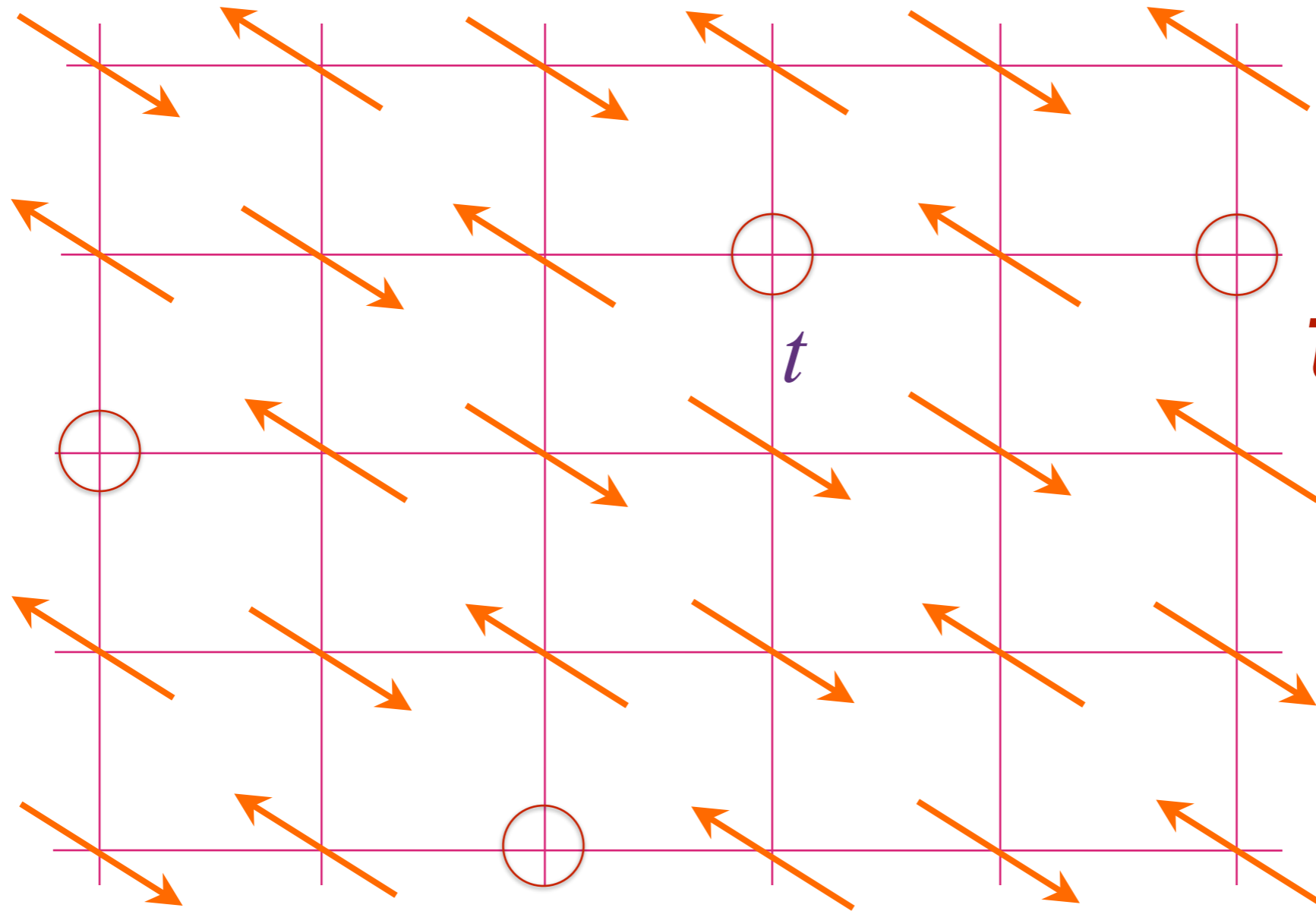


$$p=0$$

Antiferromagnet doped with hole density p

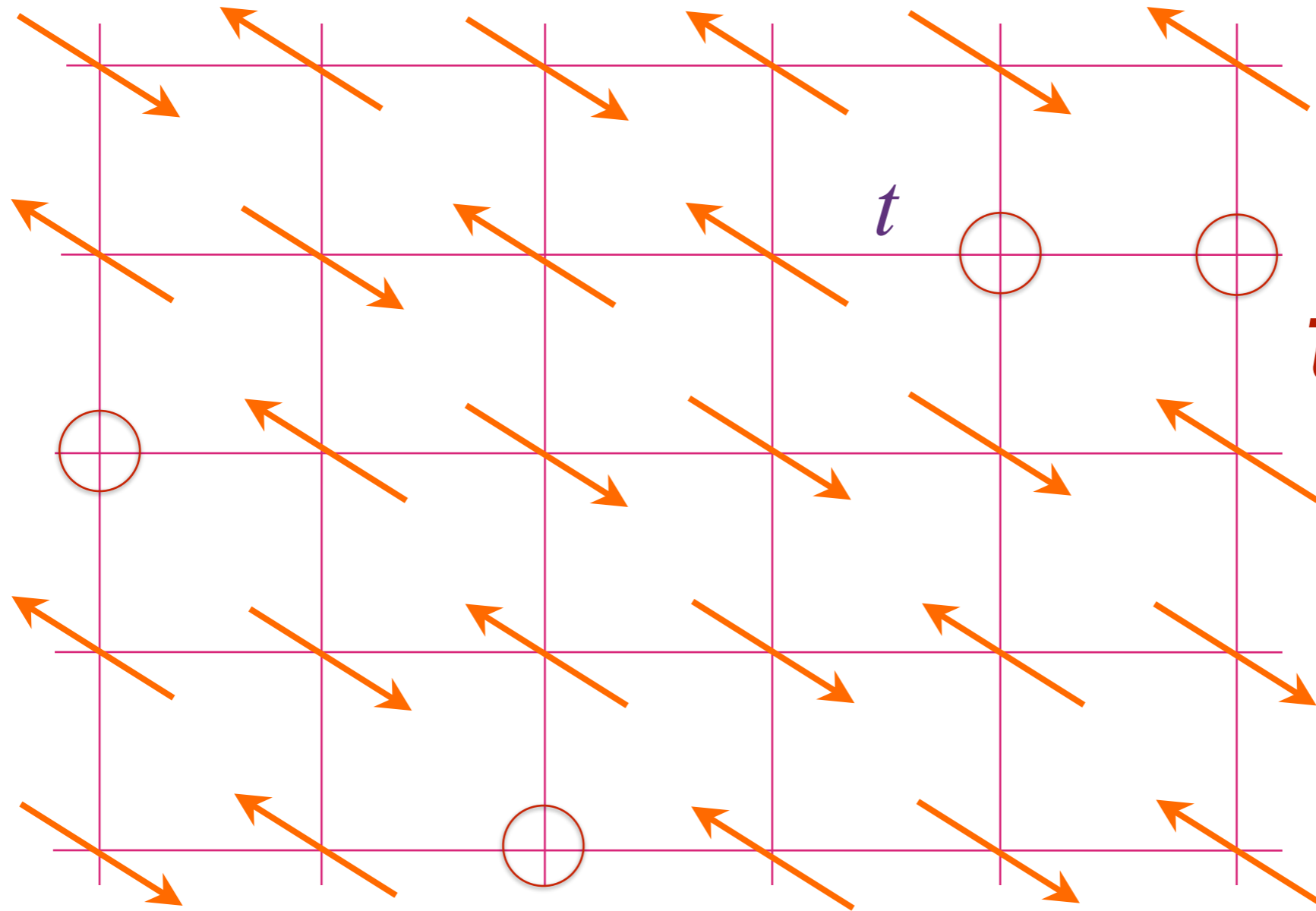


Antiferromagnet doped with hole density p



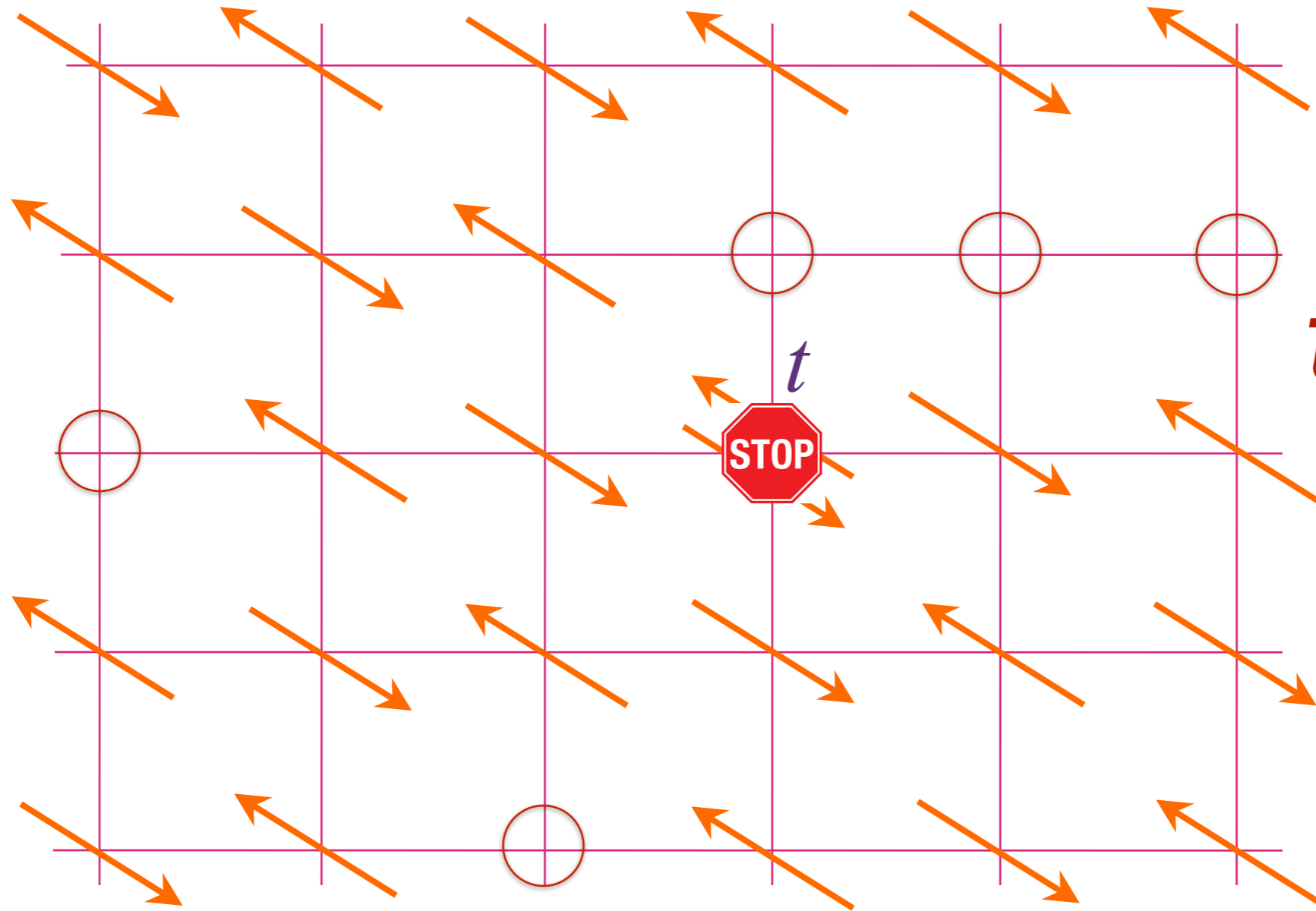
t - J model

Antiferromagnet doped with hole density p



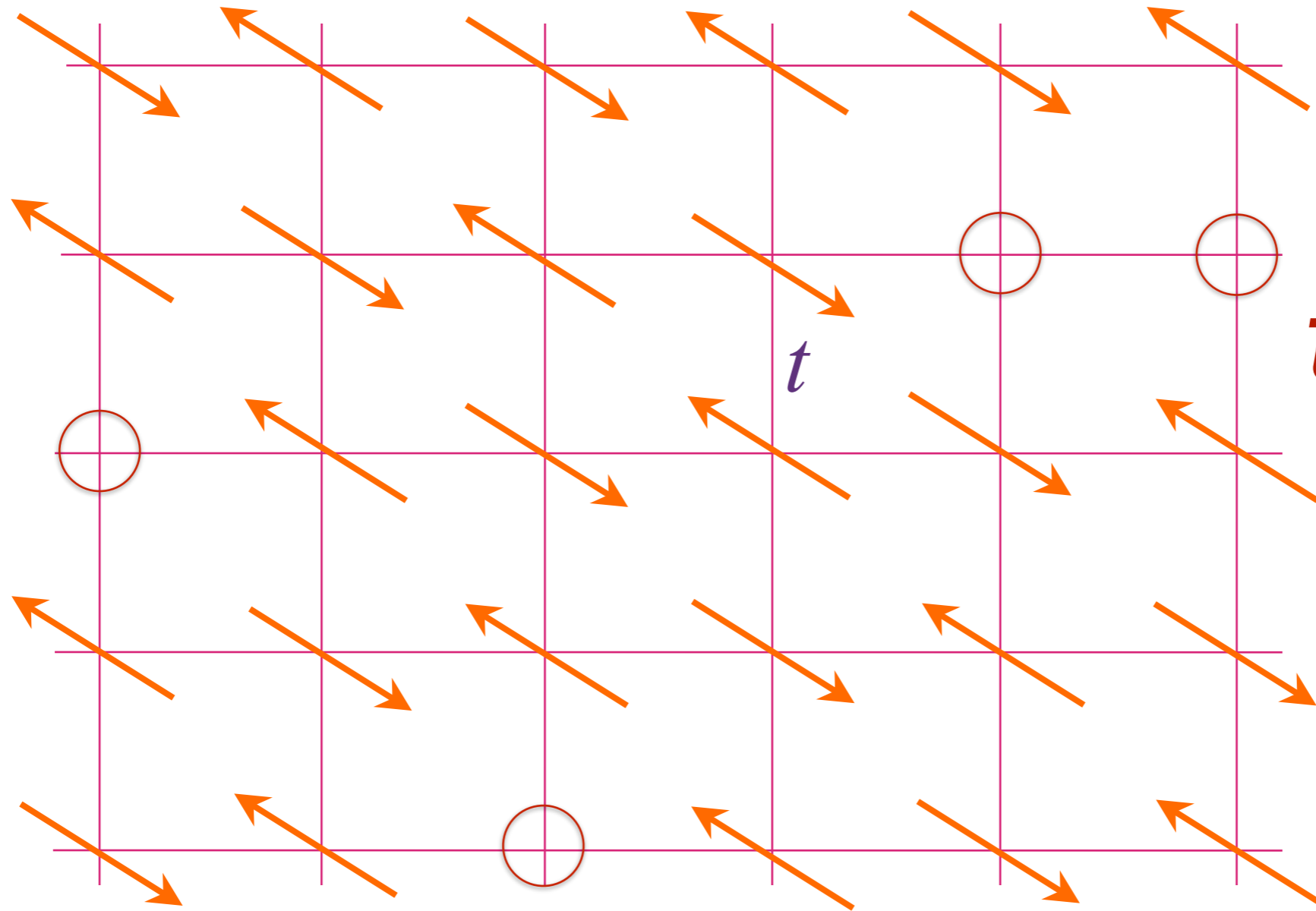
t - J model

Antiferromagnet doped with hole density p



t - J model

Antiferromagnet doped with hole density p



t - J model

Antiferromagnet doped with hole density p

$$H = -t \sum_{\langle ij \rangle} \mathcal{P}_d c_{i\alpha}^\dagger c_{j\alpha} \mathcal{P}_d + J \sum_{\langle ij \rangle} \vec{S}_i \cdot \vec{S}_j$$

$$\vec{S}_i = \frac{1}{2} c_{i\alpha}^\dagger \vec{\sigma} c_{i\alpha}$$

\mathcal{P}_d projects out doubly-occupied sites.

t - J model

Random t - J model doped with hole density p

$$H = -\frac{1}{\sqrt{N}} \sum_{i,j=1}^N t_{ij} \mathcal{P}_d c_{i\alpha}^\dagger c_{j\alpha} \mathcal{P}_d + \frac{1}{\sqrt{N}} \sum_{i<j=1}^N J_{ij} \vec{S}_i \cdot \vec{S}_j$$

$$\vec{S}_i = \frac{1}{2} c_{i\alpha}^\dagger \vec{\sigma} c_{i\alpha}$$

\mathcal{P}_d projects out doubly-occupied sites.

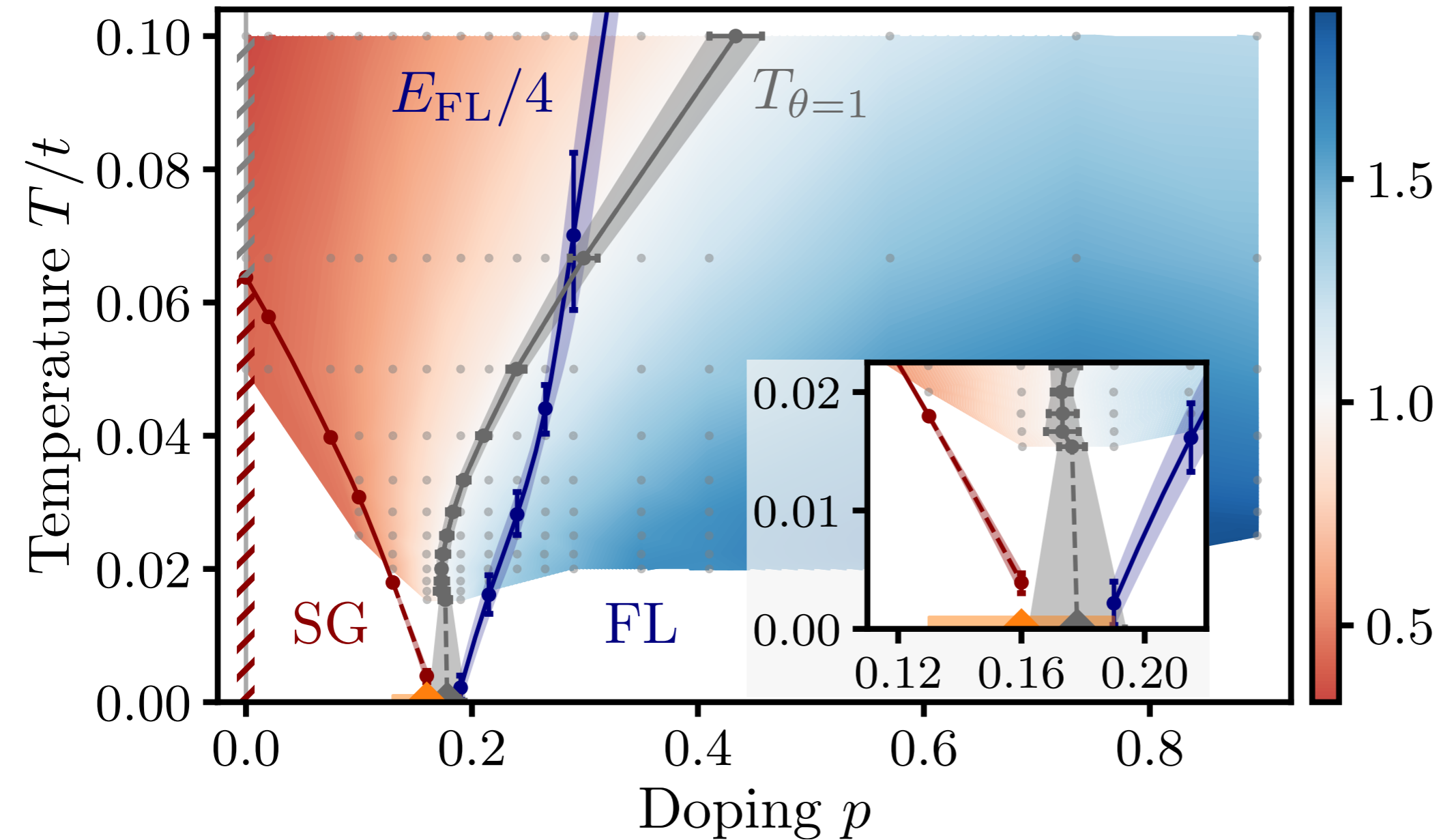
$$J_{ij} \text{ random, } \overline{J_{ij}} = 0, \overline{J_{ij}^2} = J^2$$

$$t_{ij} \text{ random, } \overline{t_{ij}} = 0, \overline{t_{ij}^2} = t^2$$

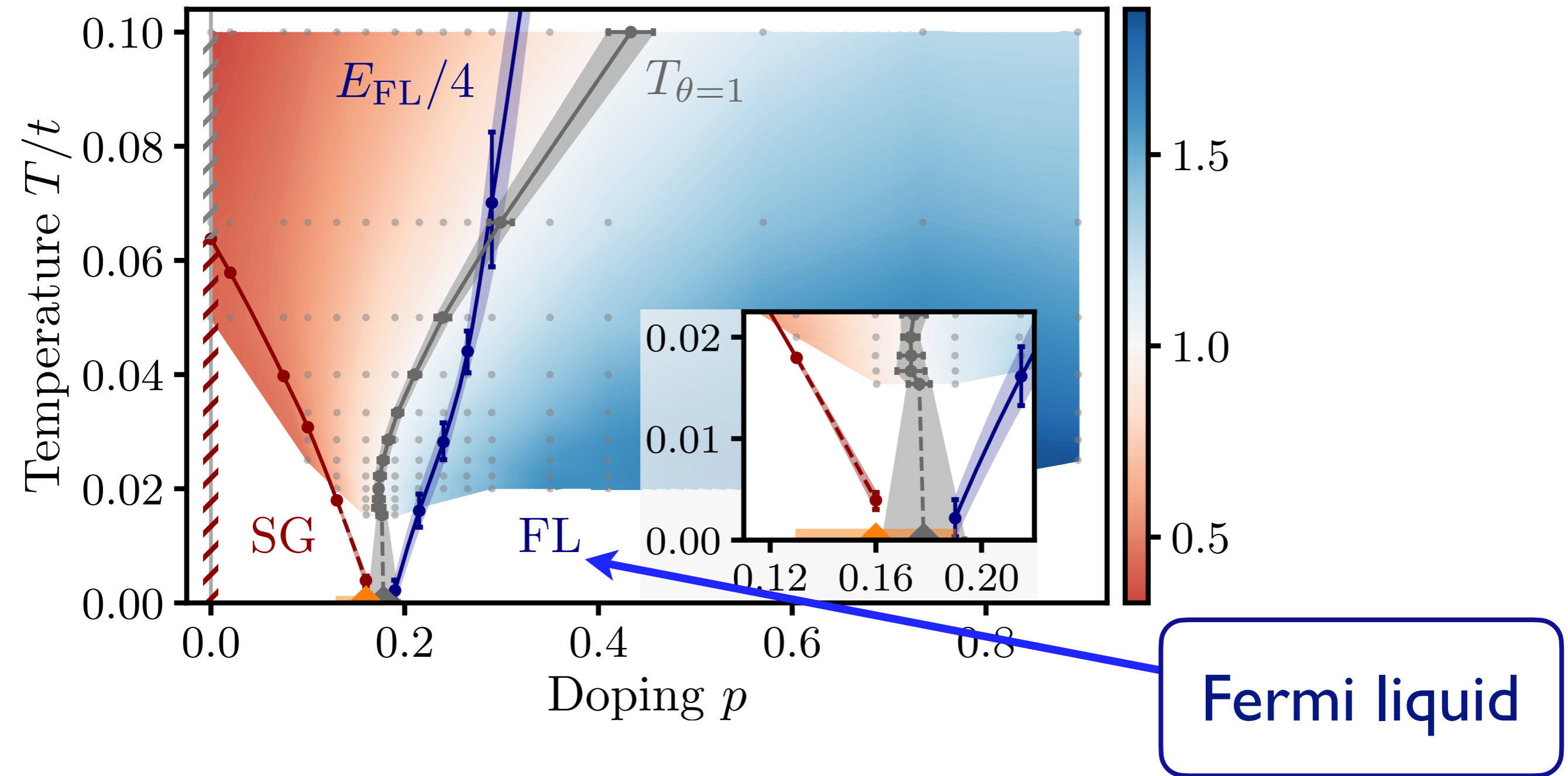
$J \Rightarrow$ two-particle interaction, as in SYK

$t \Rightarrow$ one-particle hopping, as in random matrices

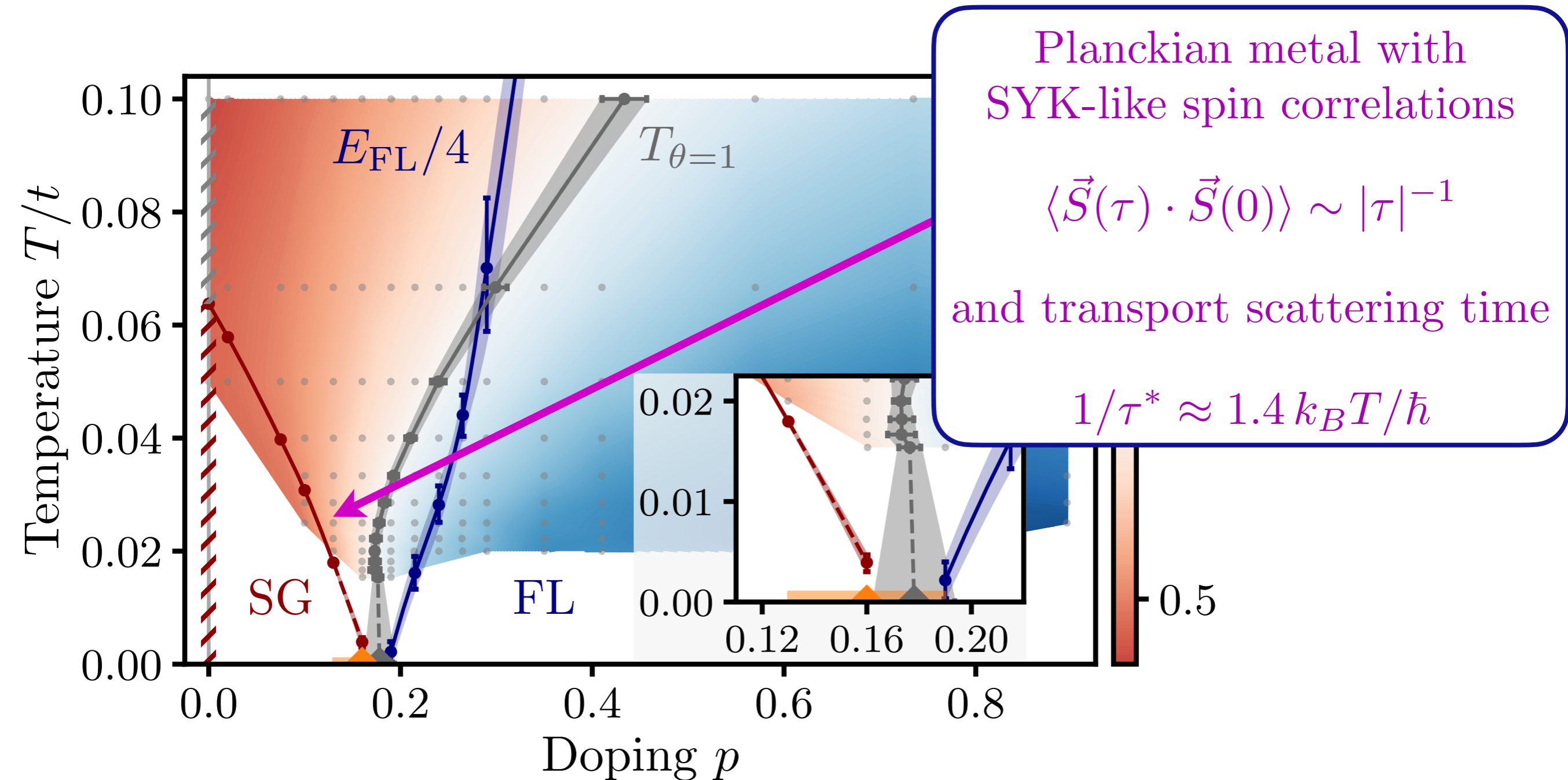
Numerical solution of t - J model on a fully-connected cluster with all-to-all and random t_{ij} and J_{ij}



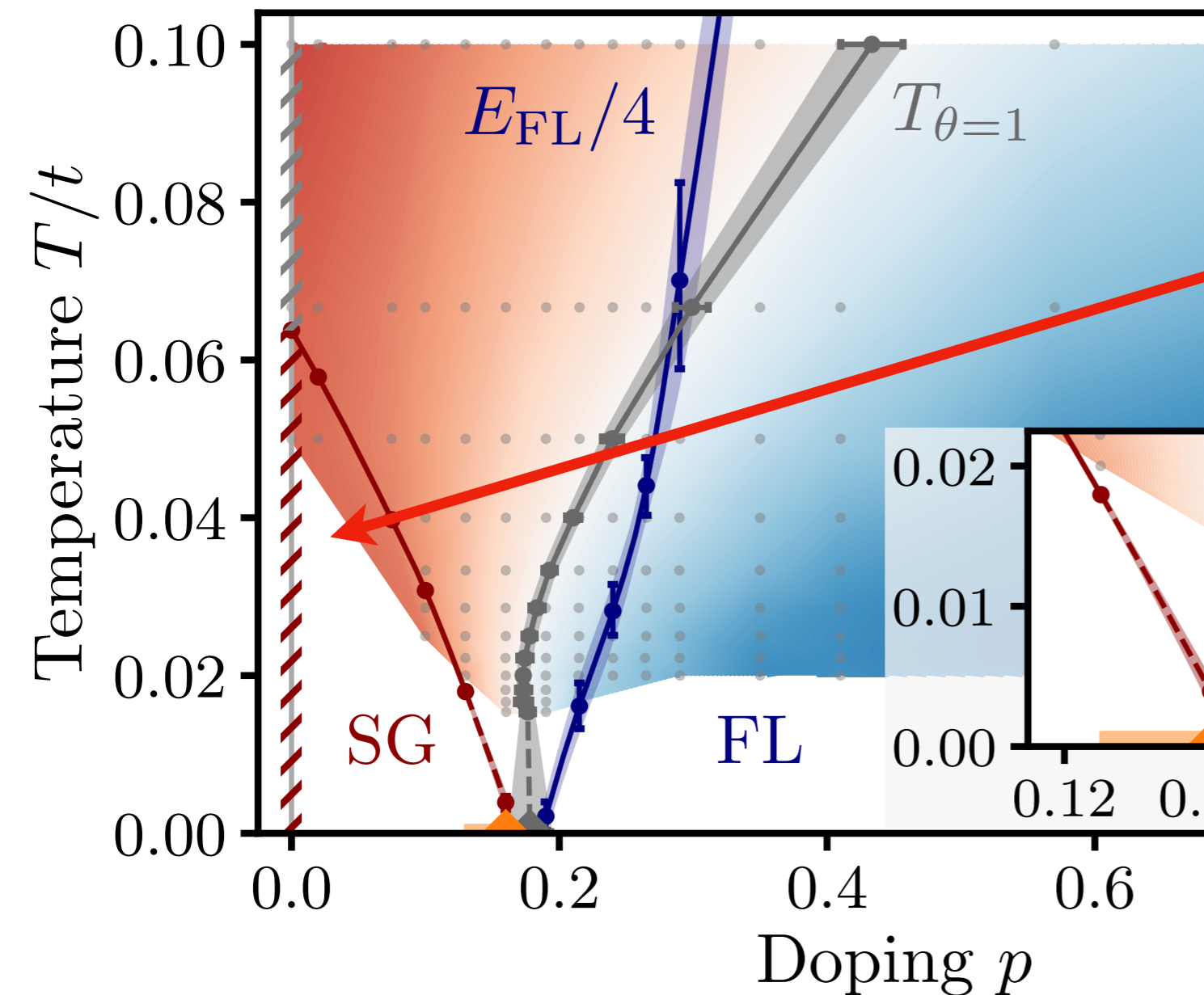
Numerical solution of t - J model on a fully-connected cluster with all-to-all and random t_{ij} and J_{ij}



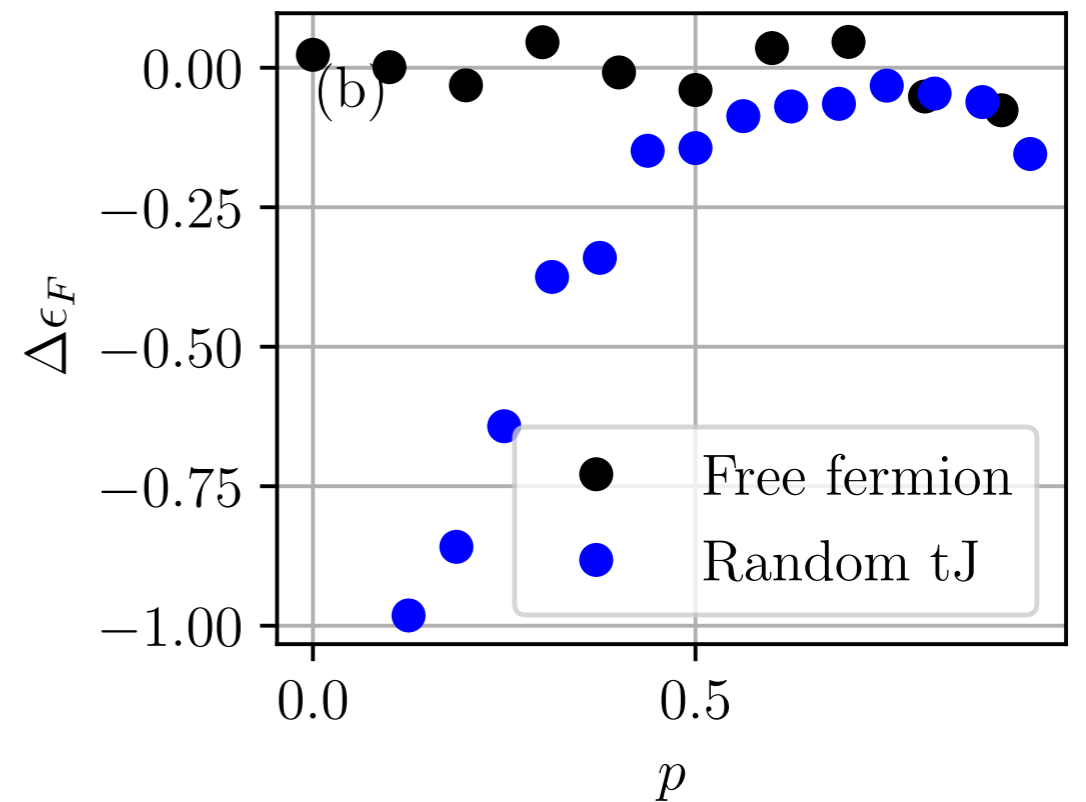
Numerical solution of t - J model on a fully-connected cluster with all-to-all and random t_{ij} and J_{ij}



Numerical solution of t - J model on a fully-connected cluster with all-to-all and random t_{ij} and J_{ij}



Pseudogap metal with spin glass order and non-Luttinger carrier density



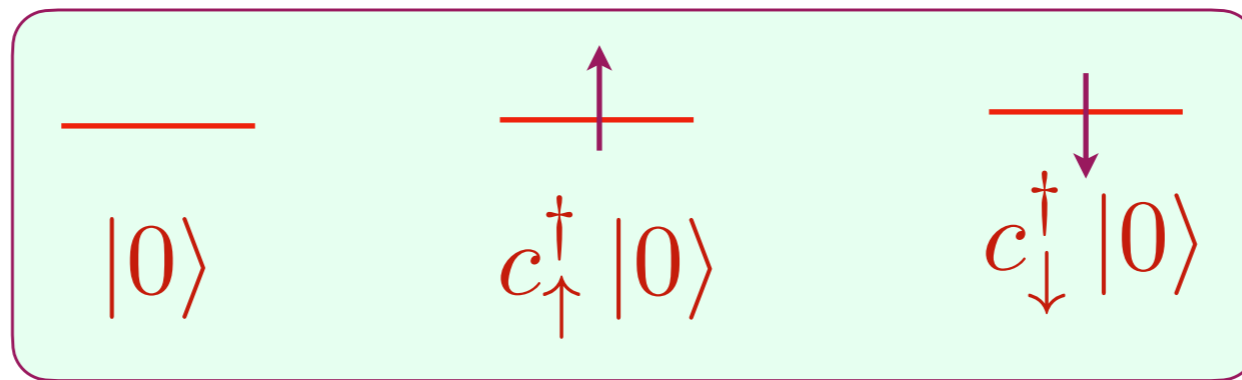
- How do we understand the appearance of SYK criticality in the random t - J model, which has a random *single* particle hopping term ?

- How do we understand the appearance of SYK criticality in the random t - J model, which has a random *single* particle hopping term ?
- What is the origin of the non-Luttinger Fermi energy in the pseudogap metal ?

- How do we understand the appearance of SYK criticality in the random t - J model, which has a random *single* particle hopping term ?
- What is the origin of the non-Luttinger Fermi energy in the pseudogap metal ?

Fractionalization!

t - J model

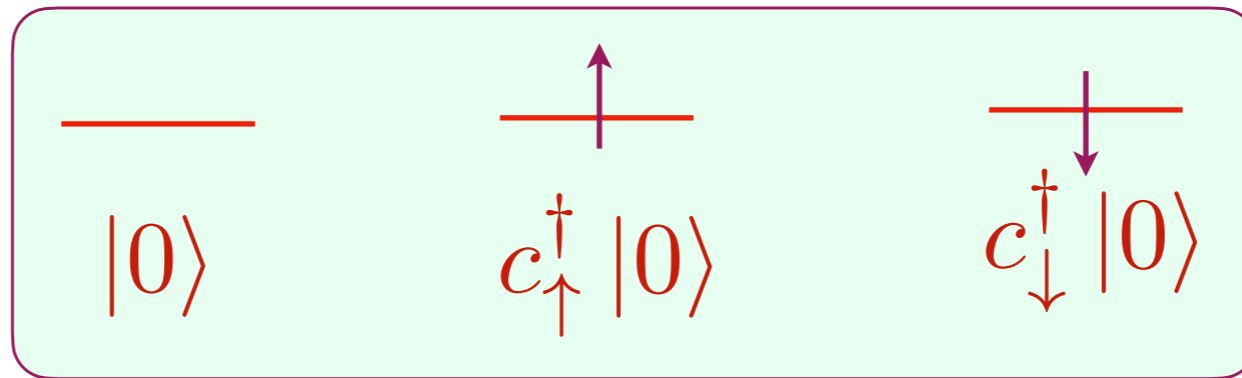


$$\vec{S}_i = \frac{1}{2} c_{i\alpha}^{\dagger} \vec{\sigma}_{\alpha\beta} c_{i\beta},$$

$$\sum_{\alpha} c_{i\alpha}^{\dagger} c_{i\alpha} \leq 1,$$

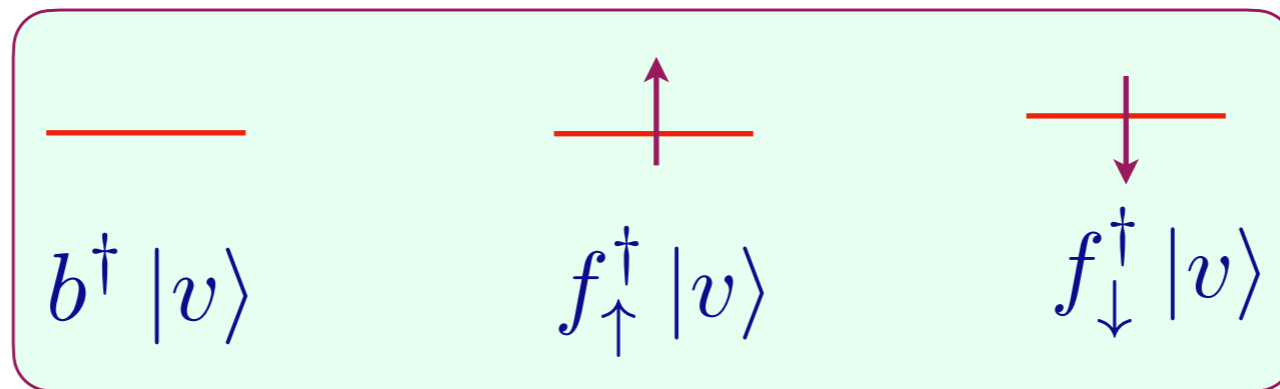
$$\frac{1}{N} \sum_{i\alpha} c_{i\alpha}^{\dagger} c_{i\alpha} = 1 - p$$

t - J model



$$\vec{S}_i = \frac{1}{2} c_{i\alpha}^{\dagger} \vec{\sigma}_{\alpha\beta} c_{i\beta}, \quad \sum_{\alpha} c_{i\alpha}^{\dagger} c_{i\alpha} \leq 1, \quad \frac{1}{N} \sum_{i\alpha} c_{i\alpha}^{\dagger} c_{i\alpha} = 1 - p$$

Fractionalization



$$c_{\alpha} = f_{\alpha} b^{\dagger} \quad ; \quad \vec{S} = \frac{1}{2} f_{\alpha}^{\dagger} \sigma_{\alpha\beta} f_{\beta}$$

$$f_{\alpha}^{\dagger} f_{\alpha} + b^{\dagger} b = 1$$

$$\text{U(1) gauge invariance:} \quad b \rightarrow b e^{i\phi}, \quad f_{\alpha} \rightarrow f_{\alpha} e^{i\phi}$$

t - J model

$$H = -\frac{1}{\sqrt{N}} \sum_{i,j=1}^N t_{ij} \mathcal{P}_d c_{i\alpha}^\dagger c_{j\alpha} \mathcal{P}_d + \frac{1}{\sqrt{N}} \sum_{i<j=1}^N J_{ij} \vec{S}_i \cdot \vec{S}_j$$

\mathcal{P}_d projects out doubly-occupied sites.

Fractionalized form:

$$H = -\frac{1}{\sqrt{N}} \sum_{i,j=1}^N t_{ij} f_{i\alpha}^\dagger f_{j\alpha} b_j^\dagger b_i + \frac{1}{\sqrt{N}} \sum_{i<j=1}^N \frac{J_{ij}}{4} f_{i\alpha}^\dagger \sigma_{\alpha\beta} f_{i\beta} \cdot f_{j\gamma}^\dagger \sigma_{\gamma\delta} f_{j\delta}$$

Both t_{ij} and J_{ij} are 4-parton terms, and this allows SYK criticality in an intermediate regime.

t - J model

Fractionalization and emergent gauge fields leading to non-Luttinger volume phases in the pseudogap regime

PHYSICAL REVIEW X **10**, 041057 (2020)

Fermi Surface Reconstruction without Symmetry Breaking

Snir Gazit 

*Racah Institute of Physics and The Fritz Haber Research Center for Molecular Dynamics,
The Hebrew University of Jerusalem, Jerusalem 91904, Israel*

Fakher F. Assaad

*Institut für Theoretische Physik und Astrophysik, Universität Würzburg,
Am Hubland, D-97074 Würzburg, Germany,
and Würzburg-Dresden Cluster of Excellence ct.qmat, Universität Würzburg,
Am Hubland, D-97074 Würzburg, Germany*

Subir Sachdev 

Department of Physics, Harvard University, Cambridge, Massachusetts 02138, USA

Summary

- SYK: a solvable model without quasiparticle excitations, exhibiting thermalization and many-body chaos in a time of order $\hbar/(k_B T)$, independent of microscopic energy scales.

Summary

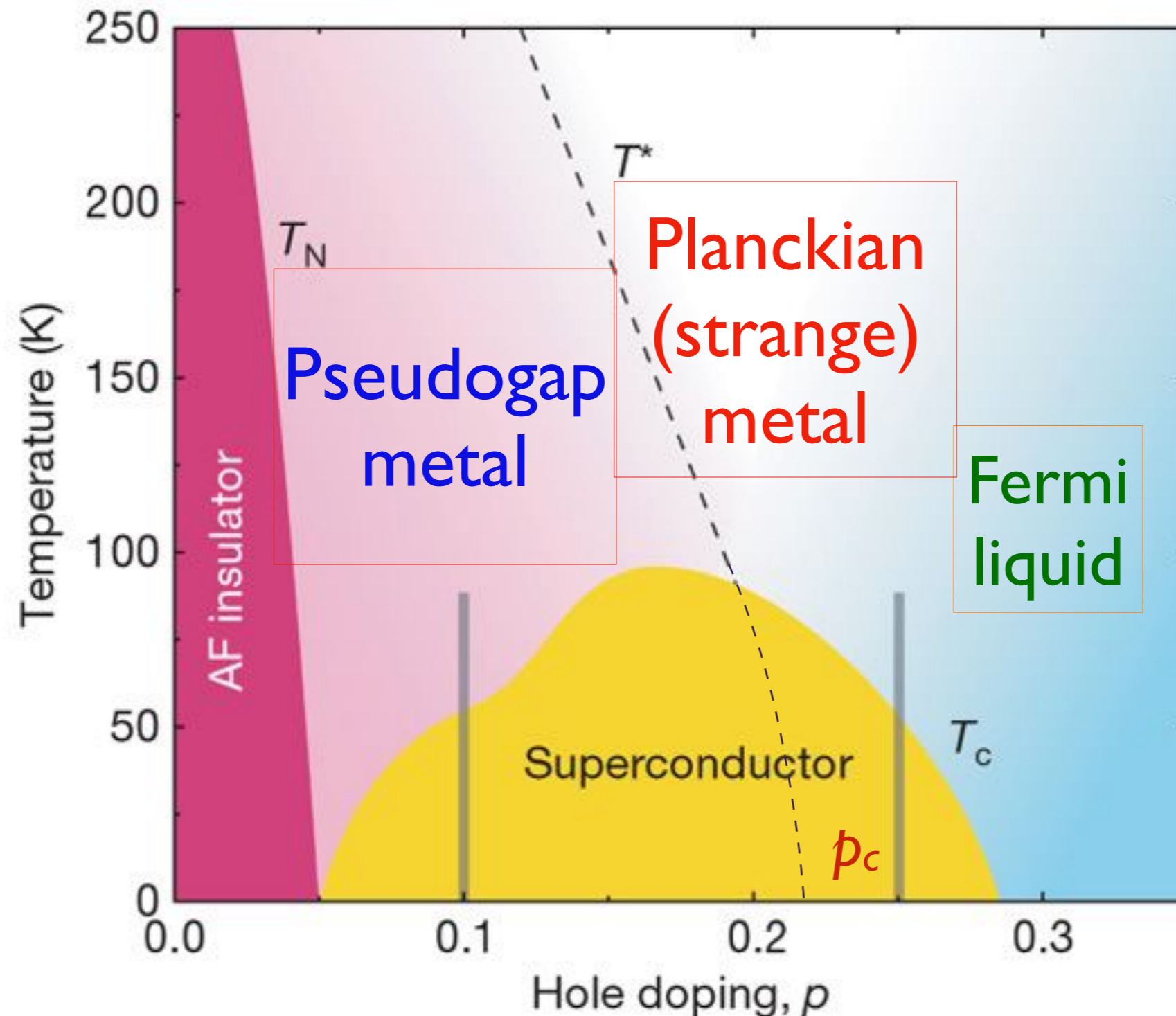
- SYK: a solvable model without quasiparticle excitations, exhibiting thermalization and many-body chaos in a time of order $\hbar/(k_B T)$, independent of microscopic energy scales.
- Low energy theory of time reparameterizations in the SYK model is identical to the theory of a near-horizon boundary graviton of charged black holes. This leads to a logarithmic correction to the Bekenstein-Hawking entropy at fixed charge Q :
$$S(T_H \rightarrow 0, Q) = A/(4G) - (3/2) \ln(1/T_H).$$

Summary

- SYK: a solvable model without quasiparticle excitations, exhibiting thermalization and many-body chaos in a time of order $\hbar/(k_B T)$, independent of microscopic energy scales.
- Low energy theory of time reparameterizations in the SYK model is identical to the theory of a near-horizon boundary graviton of charged black holes. This leads to a logarithmic correction to the Bekenstein-Hawking entropy at fixed charge Q :
$$S(T_H \rightarrow 0, Q) = A/(4G) - (3/2) \ln(1/T_H).$$
- Similar corrections to the entropy in ‘wormholes’ have recently lead to progress in understanding the flow of quantum information in evaporating black holes.

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 - Linear-in- T resistivity in the Planckian metal from the time reparameterization soft-mode.
(the same mode that yields the leading correction to the BH entropy of charged black holes, and the contribution of ‘wormholes’)