

# From SYK models to charged black holes and Planckian metals

Laval University, Quebec City  
June 3, 2019

Talk online: [sachdev.physics.harvard.edu](https://sachdev.physics.harvard.edu)



# Ordinary metals and quasiparticles

- **Quasiparticles are additive excitations:**  
The low-lying excitations of the many-body system can be identified as a set  $\{n_\alpha\}$  of quasiparticles with energy  $\epsilon_\alpha$

$$E = \sum_{\alpha} n_{\alpha} \epsilon_{\alpha} + \sum_{\alpha, \beta} F_{\alpha\beta} n_{\alpha} n_{\beta} + \dots$$

In a lattice system of  $N$  sites, this parameterizes the energy of  $\sim e^{\alpha N}$  states in terms of poly( $N$ ) numbers.



# Ordinary metals and quasiparticles

- Quasiparticles eventually collide with each other. Such collisions eventually leads to thermal equilibration in a chaotic quantum state, but the equilibration takes a long time. In a Fermi liquid, this time diverges as

$$\tau_{\text{eq}} \sim \frac{\hbar E_F^3}{U^2 (k_B T)^2} \quad , \quad \text{as } T \rightarrow 0,$$

where  $U$  is the strength of interactions, and  $E_F$  is the Fermi energy.



# Ordinary metals and quasiparticles

- Similarly, a quasiparticle model implies a resistivity

$$\rho = \frac{m^*}{ne^2} \frac{1}{\tau} \sim U^2 T^2 \quad \text{with } \tau \sim \tau_{\text{eq}}$$



# Ordinary metals and quasiparticles

- These times are much longer than the ‘Planckian time’  $\hbar/(k_B T)$ , which we will find in systems without quasiparticle excitations.

$$\tau \sim \tau_{\text{eq}} \gg \frac{\hbar}{k_B T}, \quad \text{as } T \rightarrow 0.$$



A compressible many-particle state  
without quasiparticle excitations:

SYK model of fermions with random all-to-all interactions of mean-square-value  $U$ , with total fermion number  $Q$ , at temperatures  $T \ll U$

A compressible many-particle state  
without quasiparticle excitations:

SYK model of fermions with random all-to-all interactions of mean-square-value  $U$ , with total fermion number  $Q$ , at temperatures  $T \ll U$

and

Charged black holes in 3+1 dimensions of radius  $R_h$ , with total charge  $Q$ , at temperatures  $T \ll 1/R_h$

are described by a common low energy quantum theory in 0 + 1 dimensions

# Main result

The common low  $T$  path integral is  $\mathcal{Z} = \int \mathcal{D}f \mathcal{D}\phi e^{-I}$ . This can be exactly evaluated, and the action is

$$I = -s_0 + \int_0^{1/T} d\tau \left\{ \frac{K}{2} \left( \frac{\partial\phi}{\partial\tau} + i(2\pi\mathcal{E}T) \frac{\partial f}{\partial\tau} \right)^2 - \frac{\gamma}{4\pi^2} \text{Sch}[\tan(\pi T f(\tau)), \tau] \right\},$$

where  $f(\tau)$  is a monotonic reparameterization of the temporal circle with

$$f(\tau + 1/T) = f(\tau) + 1/T,$$

$\phi$  is a phase conjugate to the charge density with

$$\phi(\tau + 1/T) = \phi(\tau) + 2\pi n, \quad n \text{ integer},$$

$\text{Sch}[g[\tau], \tau]$  is the Schwarzian derivative of  $g(\tau)$ .

The couplings are related to the entropy  $S(T, \mathcal{Q})$  and the chemical potential  $\mu$  via

$$S(T \rightarrow 0, \mathcal{Q}) = s_0 + \gamma T, \quad K = \left( \frac{d\mathcal{Q}}{d\mu} \right)_{T \rightarrow 0}, \quad 2\pi\mathcal{E} = \frac{ds_0}{d\mathcal{Q}}$$

# Main result

S. Sachdev, Phys. Rev. Lett. **105**, 151602 (2010)

A. Kitaev (2015)

S. Sachdev, Phys. Rev. X **5**, 041025 (2015)

J. Maldacena and D. Stanford, Phys. Rev. D **94**, 106002 (2016)

J. Maldacena, D. Stanford, and Zhenbin Yang, PTEP 12C104 (2016)

K. Jensen, Phys. Rev. Lett. **117**, 111601 (2016)

J. Engelsoy, T.G. Mertens, and H. Verlinde, JHEP 1607 (2016) 139

R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev,  
Phys. Rev. B **95**, 155131 (2017)

A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746 P. Nayak, A. Shukla,

R.M. Soni, S.P. Trivedi, and V. Vishal, arXiv:1802.09547

P. Chaturvedi, Yingfei Gu, Wei Song, Boyang Yu, arXiv:1808.08062

U. Moitra, S. P. Trivedi, and V. Vishal, arXiv:1808.08239

S. Sachdev, arXiv:1902.04078

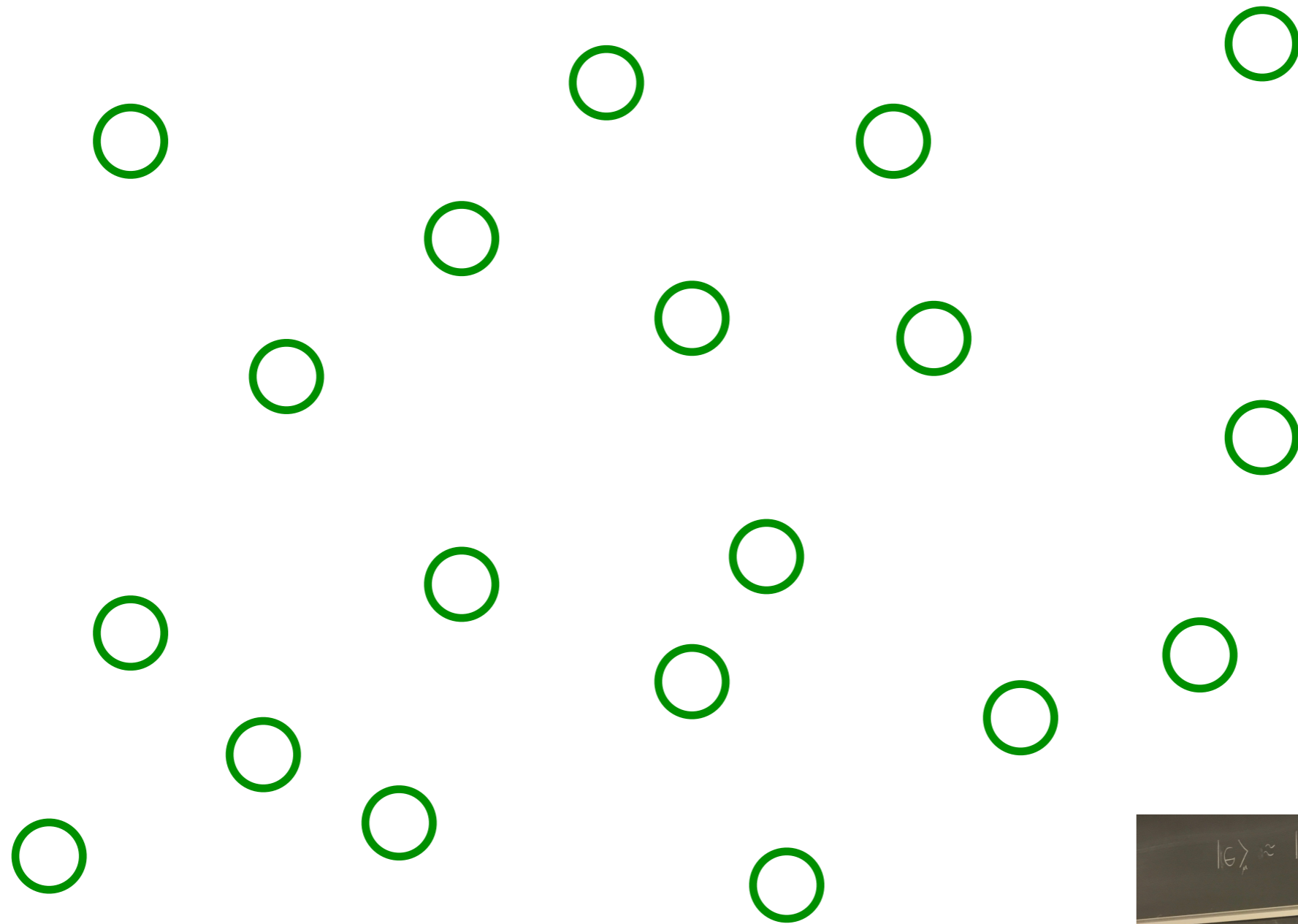
1. Quantum matter without quasiparticles:  
the complex SYK model

2. Fluctuations, and the Schwarzian

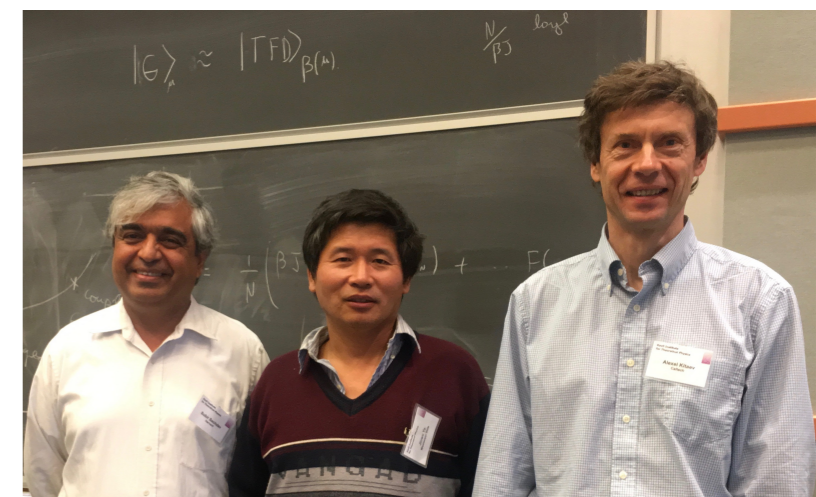
3. Einstein-Maxwell theory of charged  
black holes in AdS space

4. Planckian metals

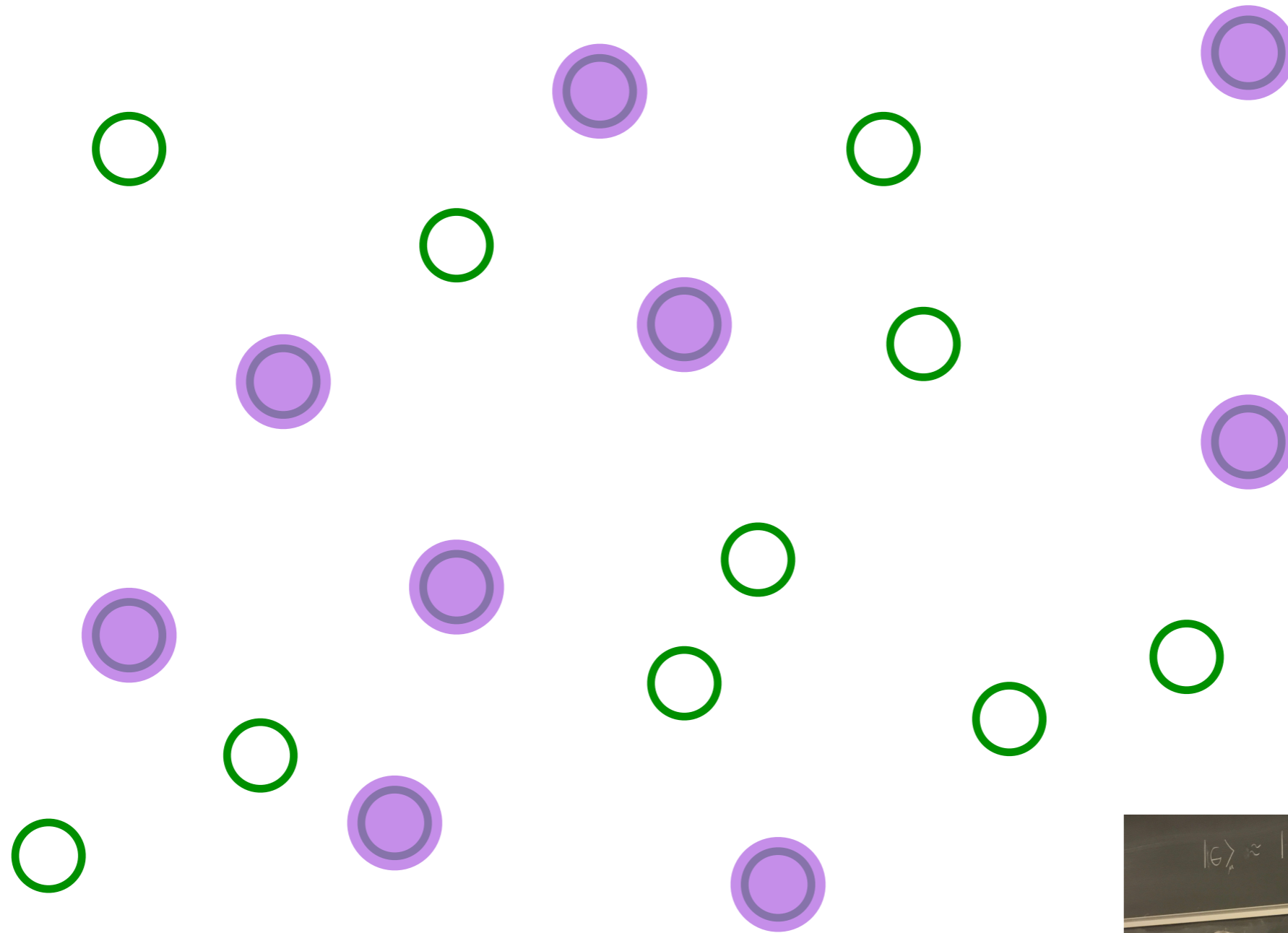
# The Sachdev-Ye-Kitaev (SYK) model



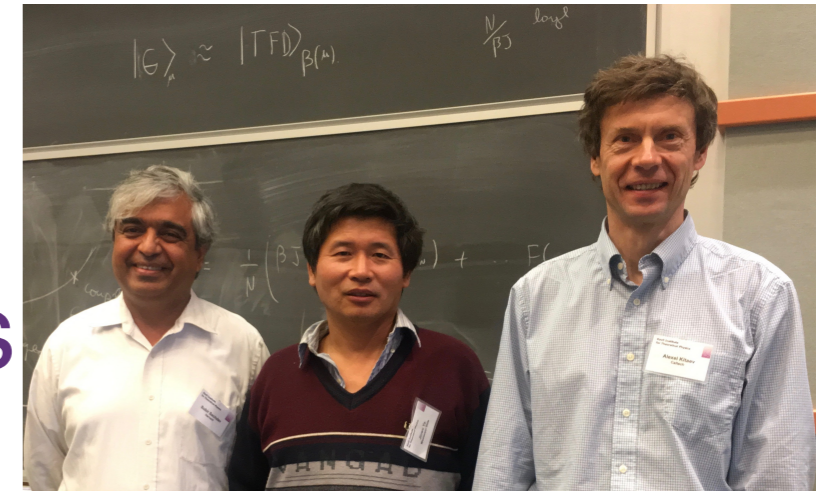
Pick a set of random positions



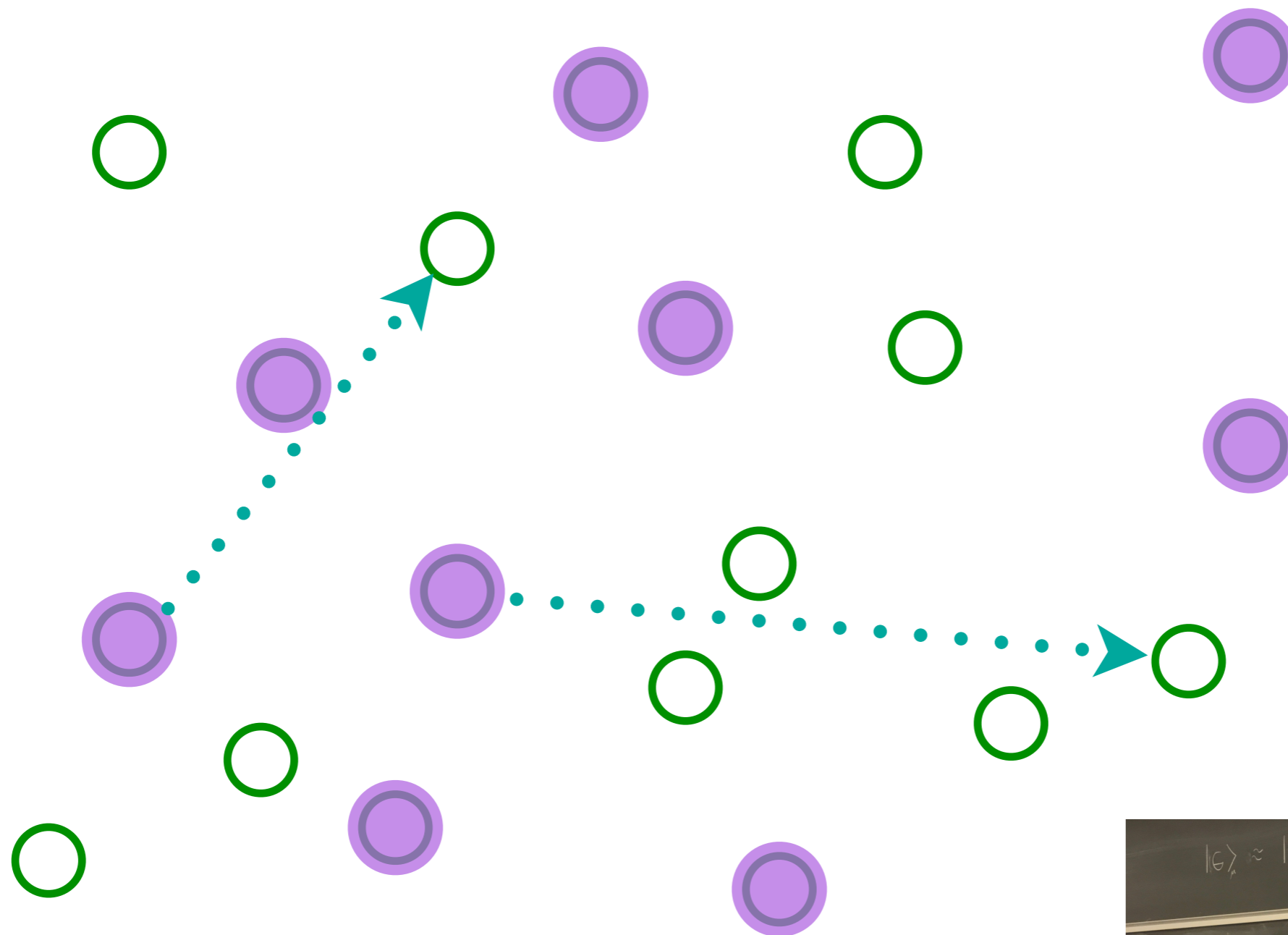
# The SYK model



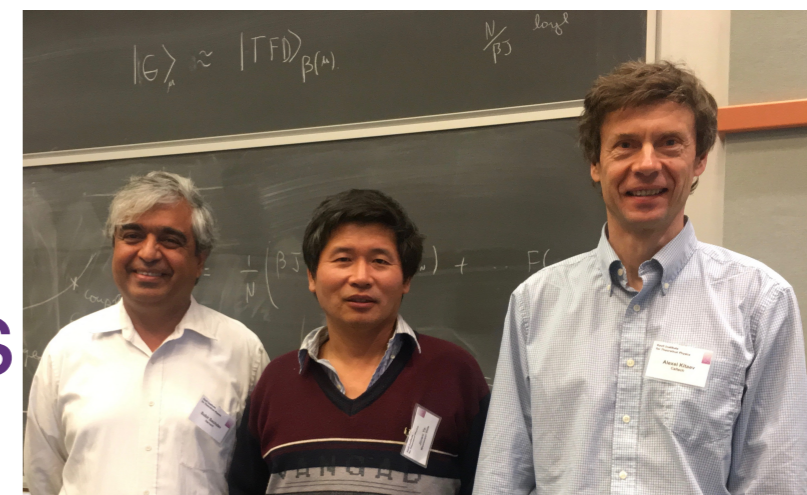
Place electrons randomly on some sites



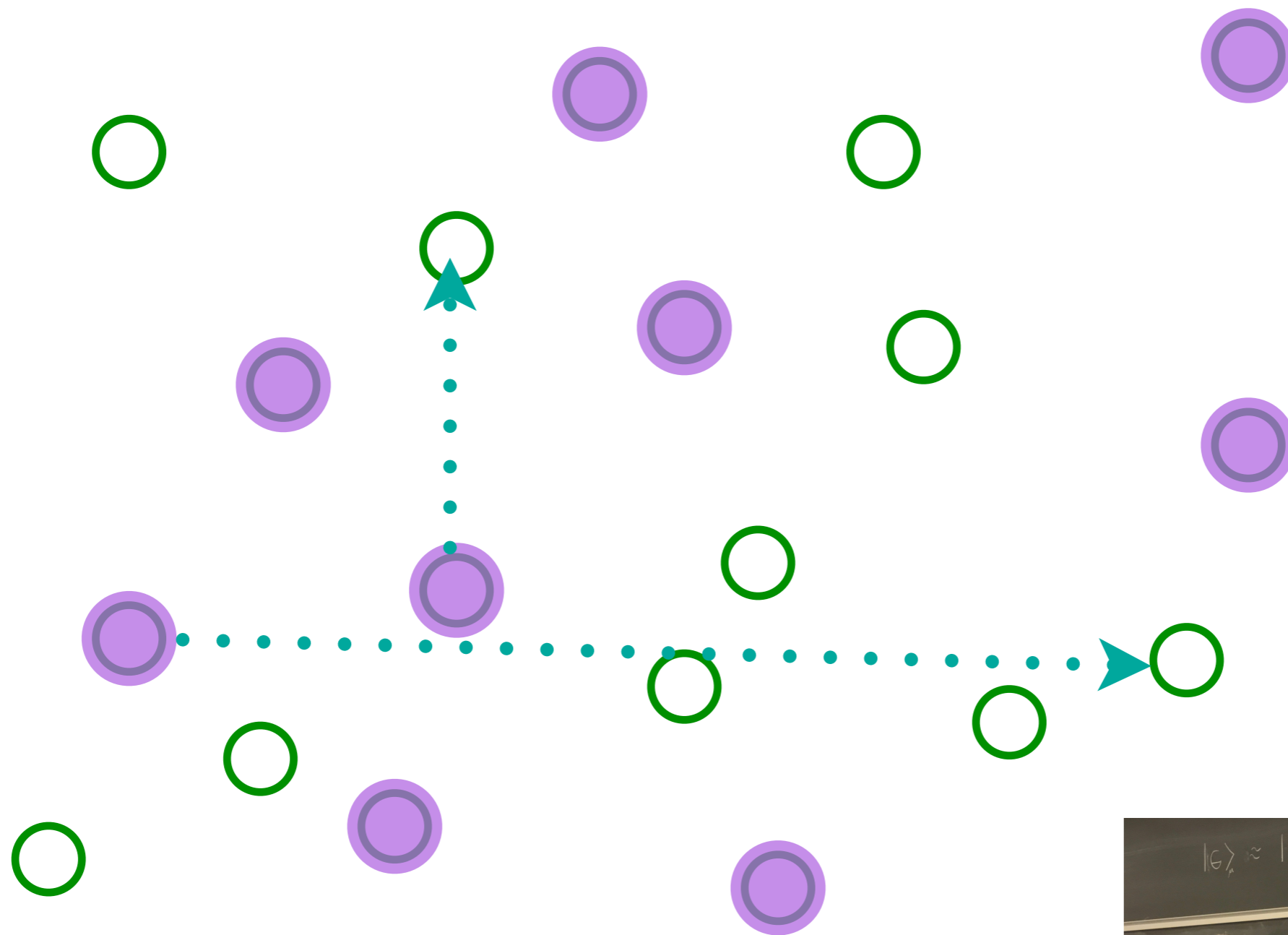
# The SYK model



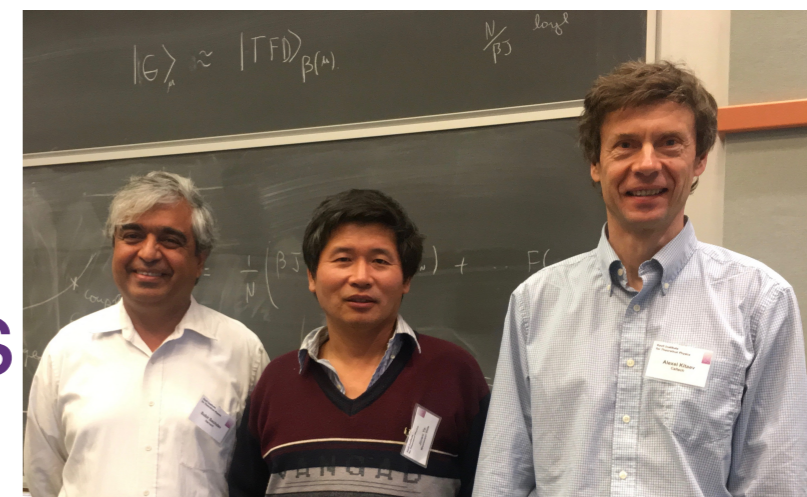
Place electrons randomly on some sites



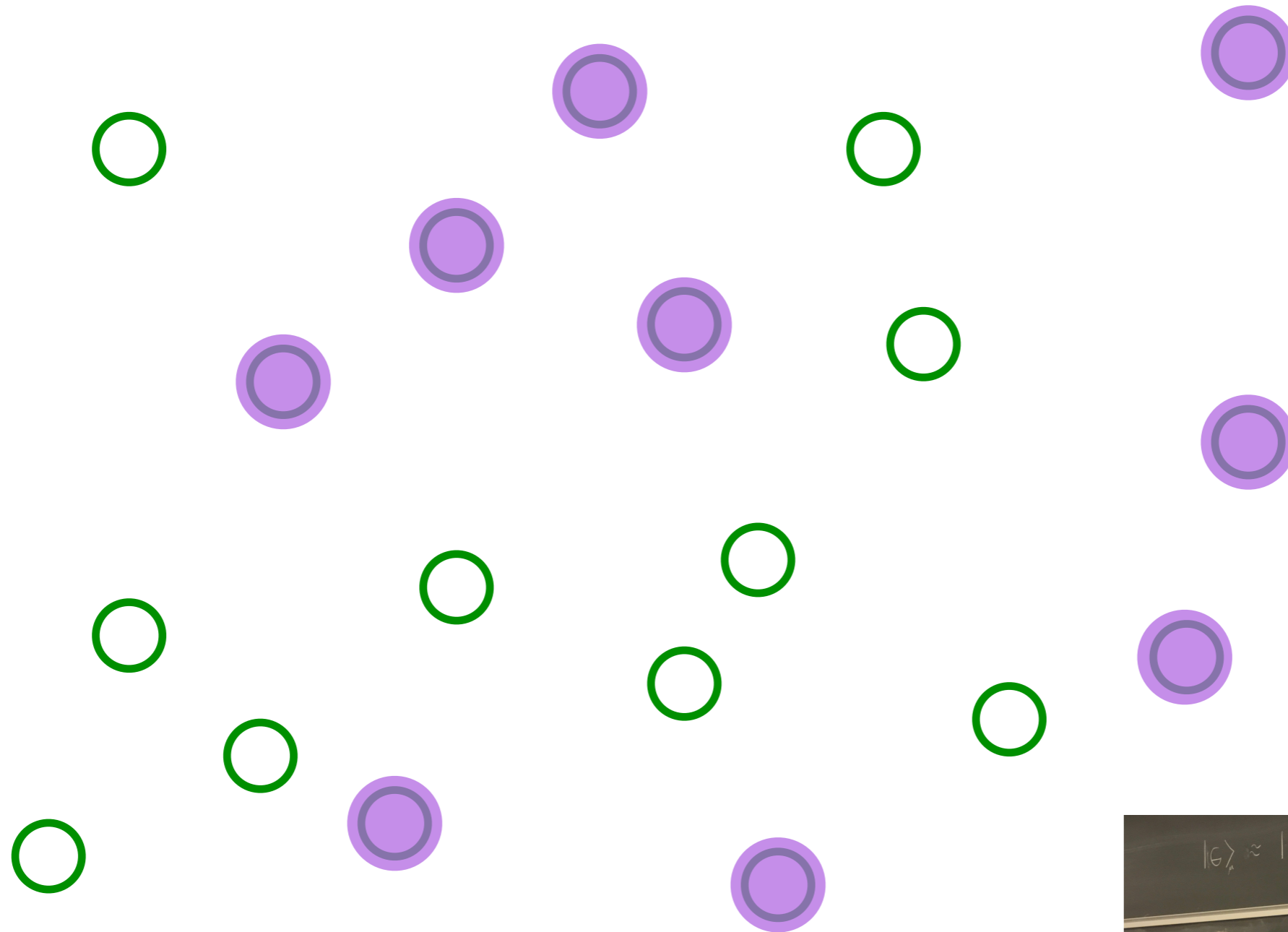
# The SYK model



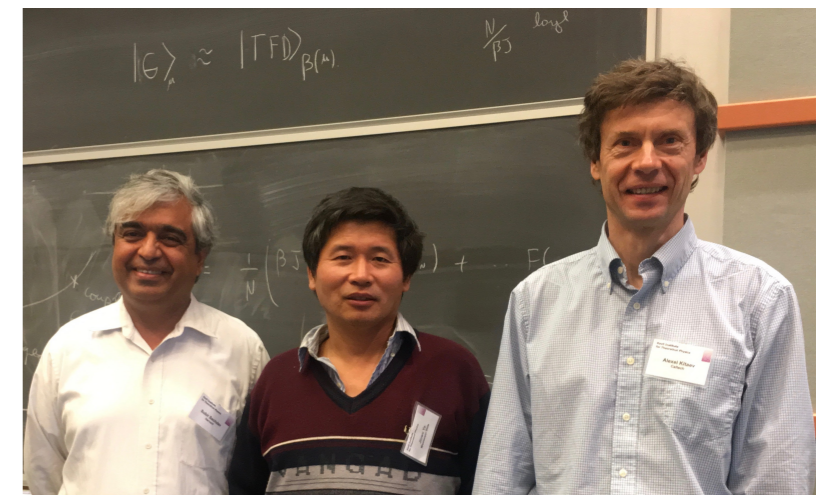
Place electrons randomly on some sites



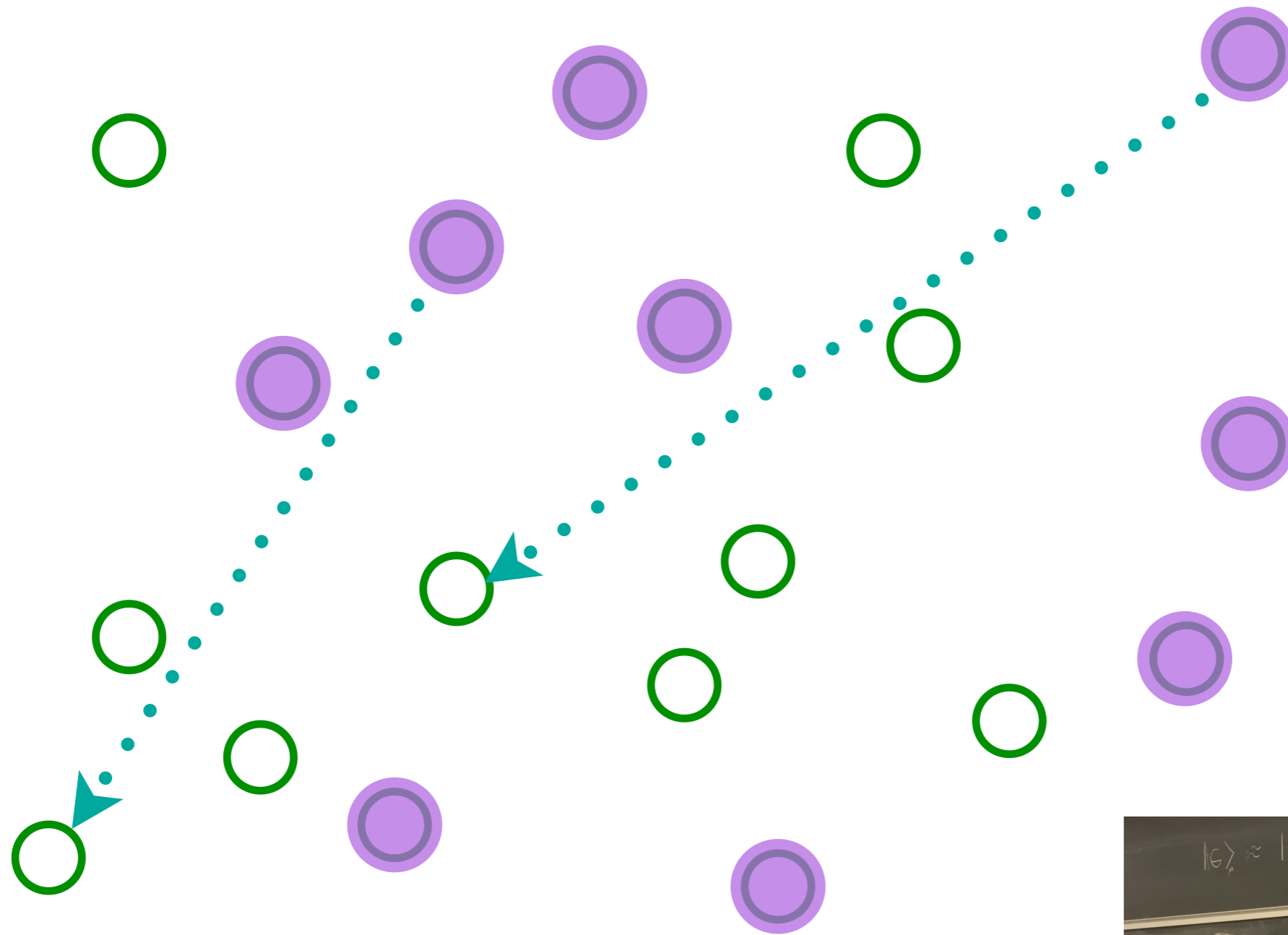
# The SYK model



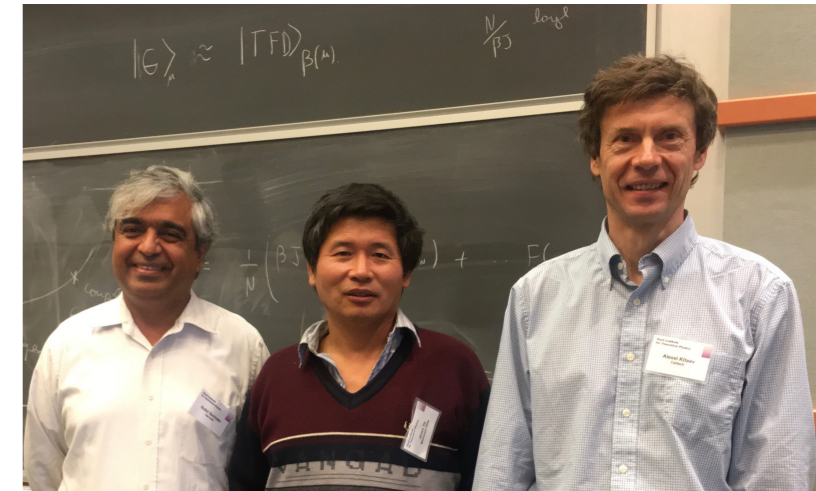
Entangle electrons pairwise randomly



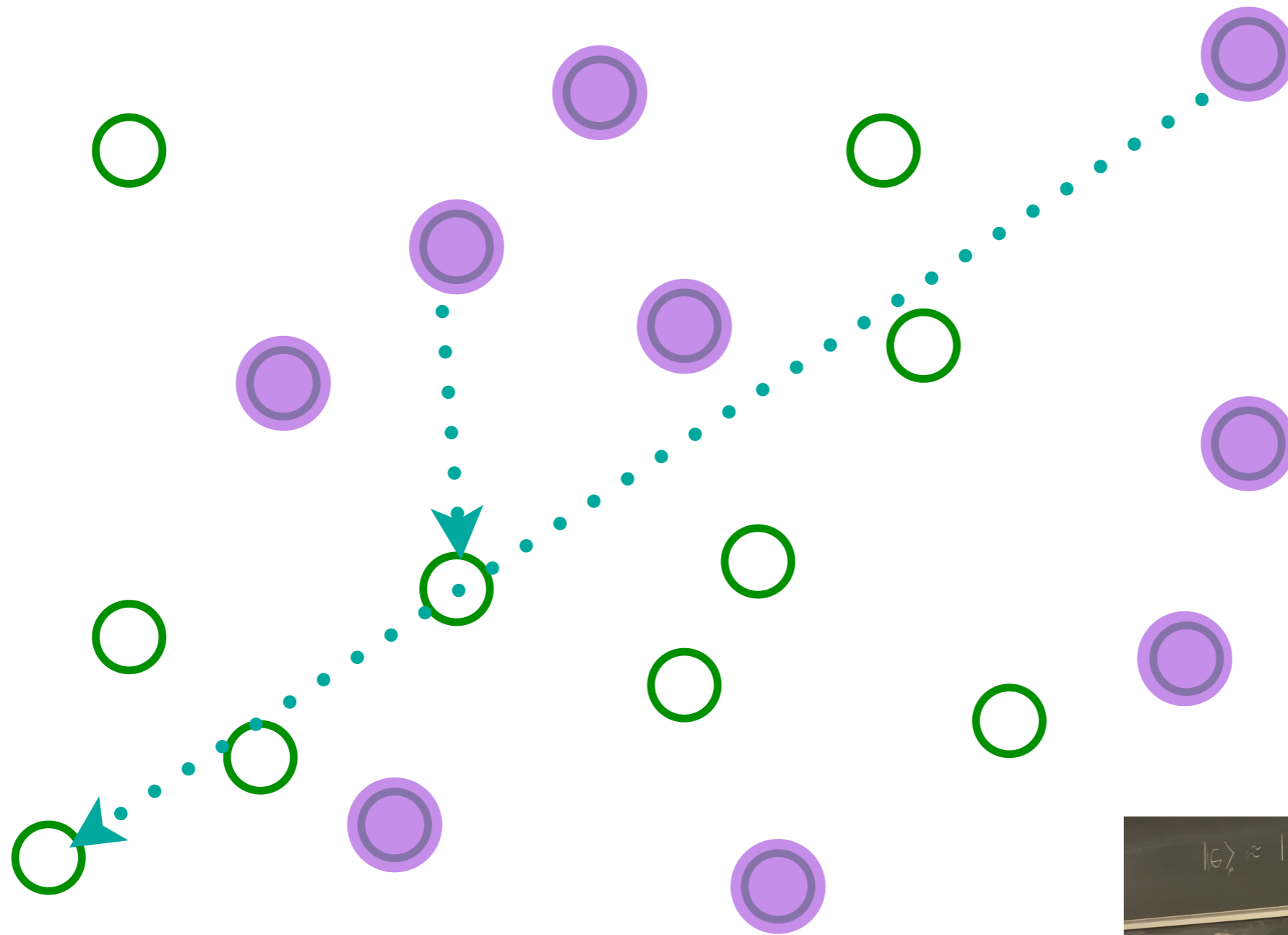
# The SYK model



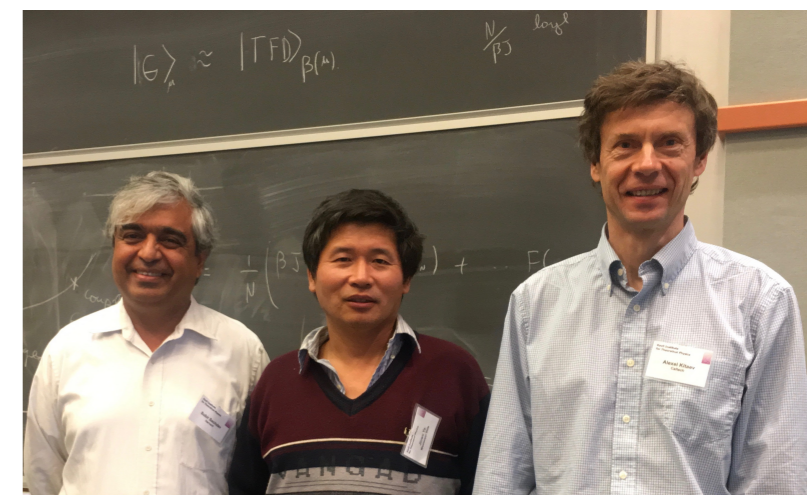
Entangle electrons pairwise randomly



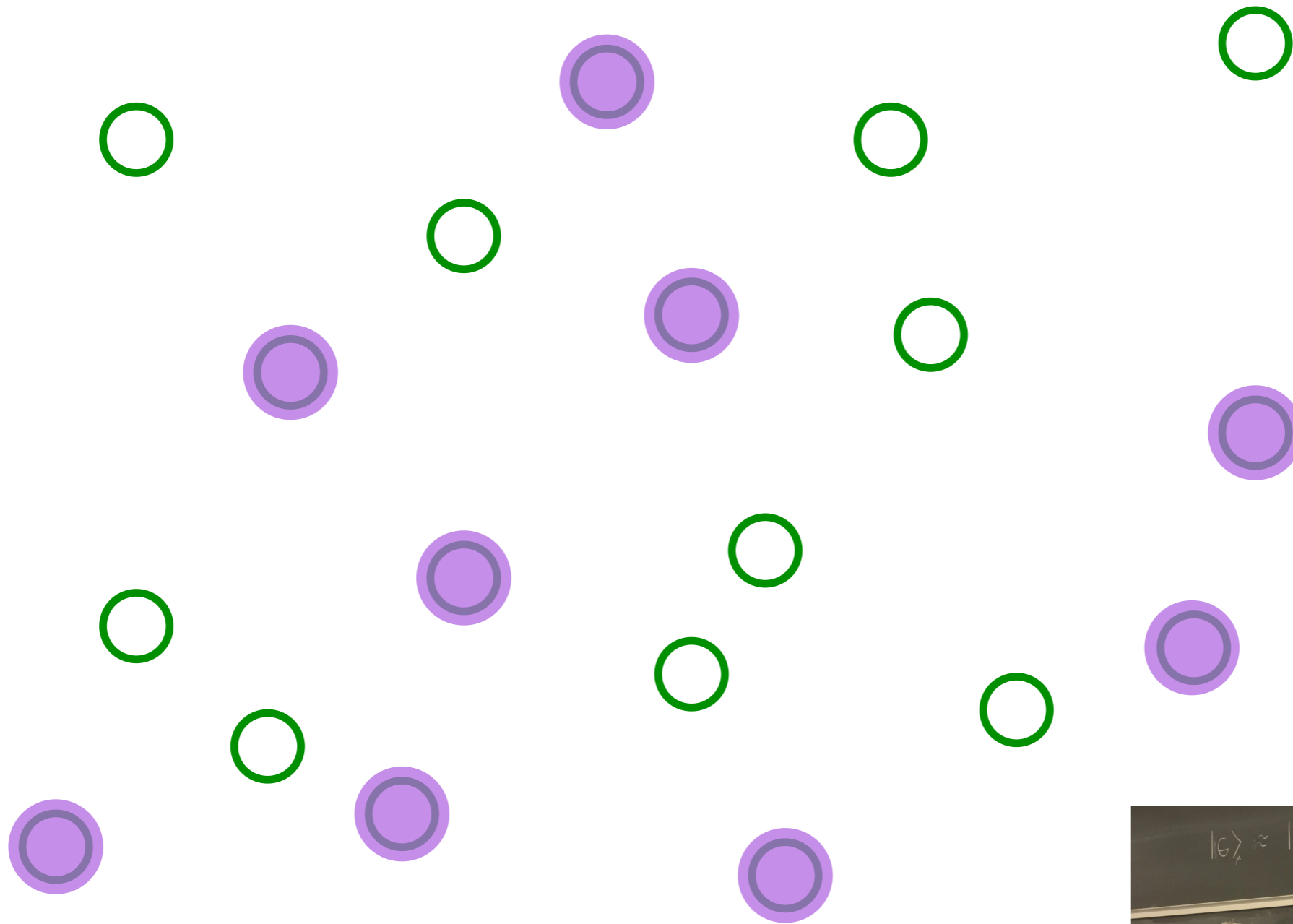
# The SYK model



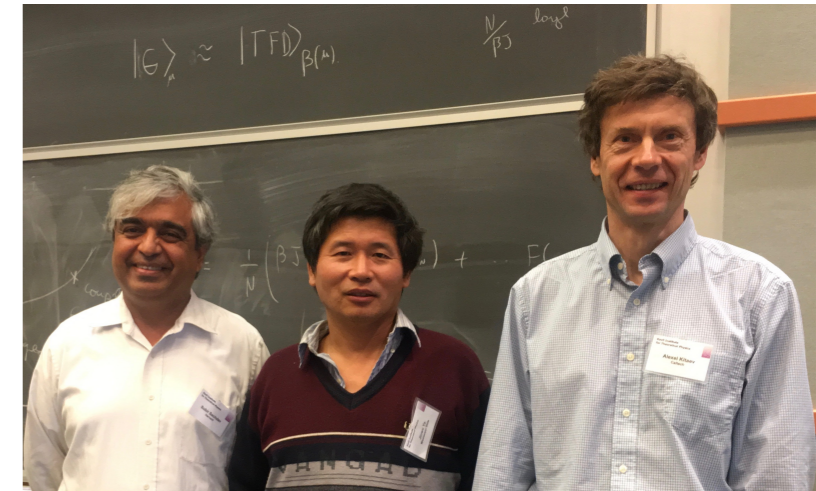
Entangle electrons pairwise randomly



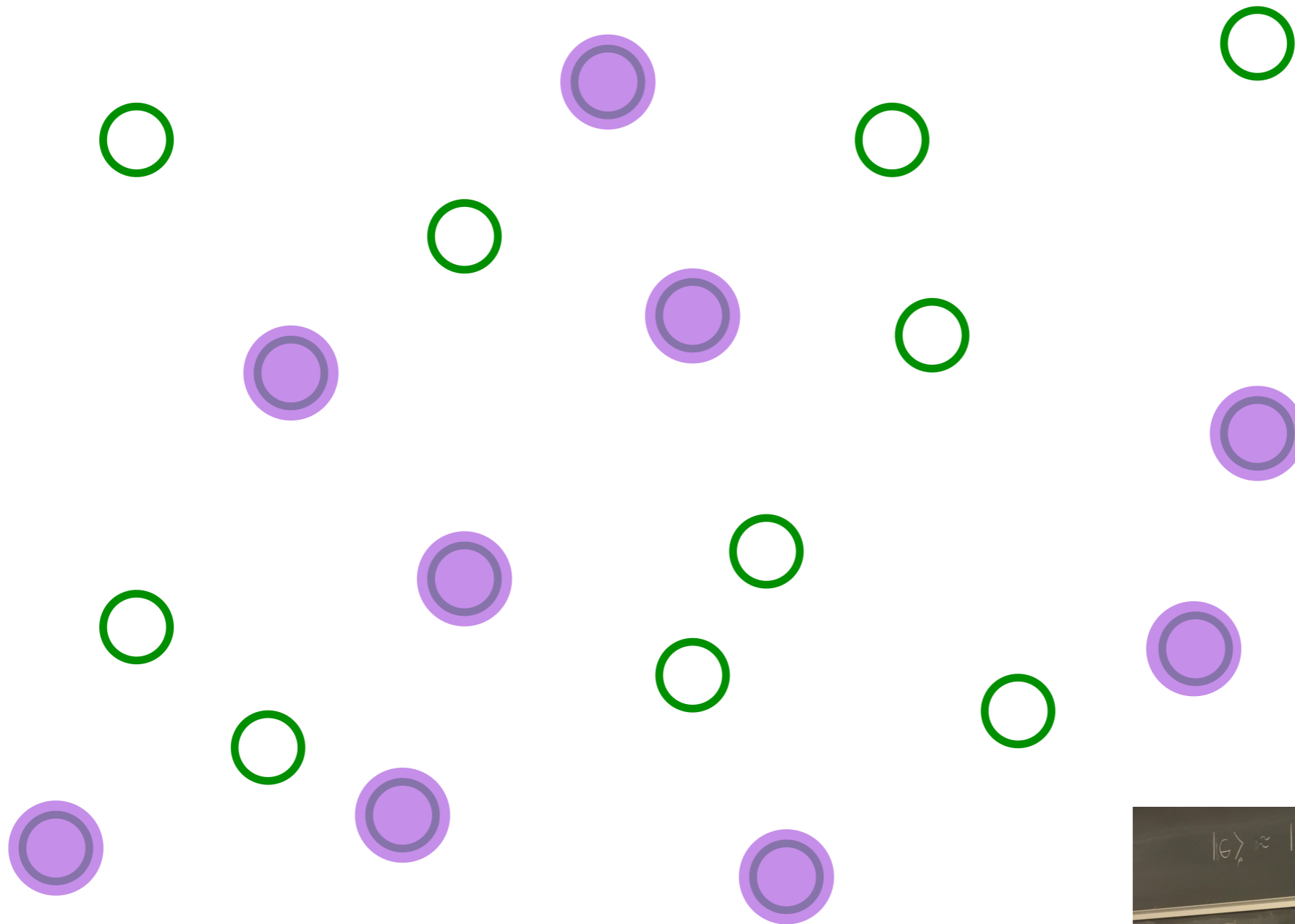
# The SYK model



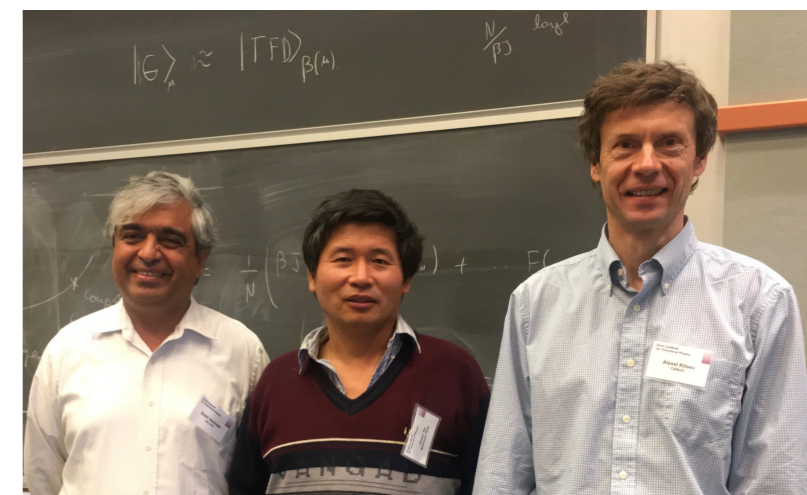
Entangle electrons pairwise randomly



# The SYK model



This describes both a strange metal  
and a black hole!



# The complex SYK model

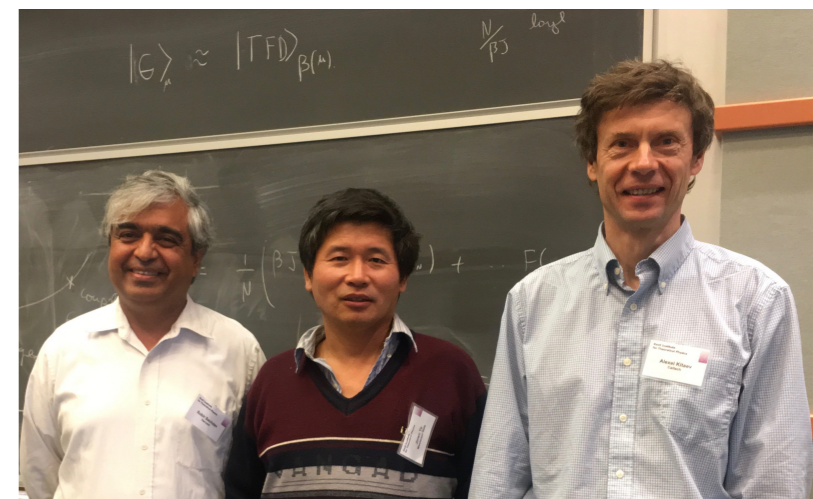
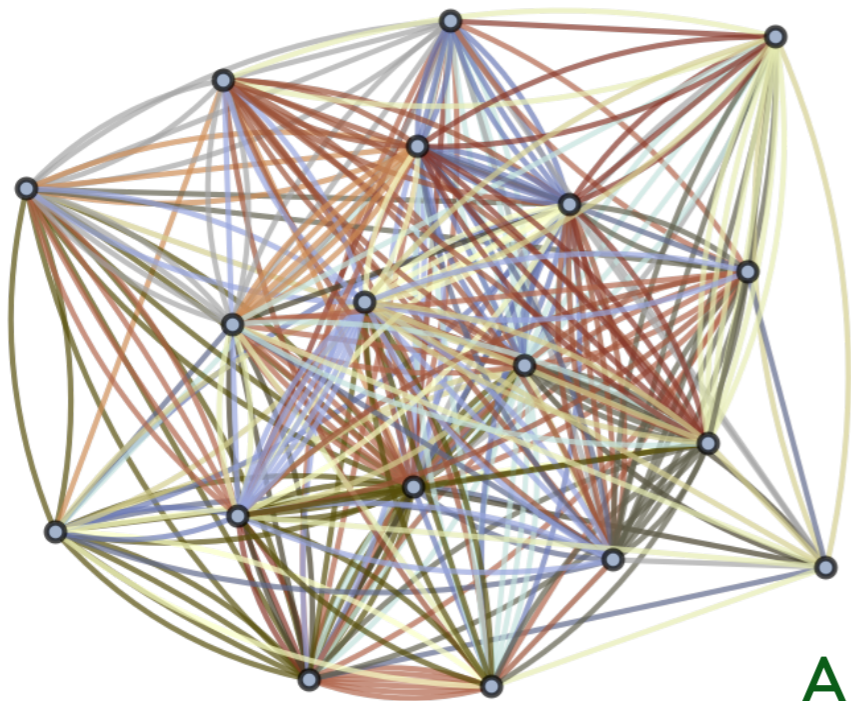
(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large  $N$  limit; T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta; \gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} + \epsilon \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$$c_{\alpha} c_{\beta} + c_{\beta} c_{\alpha} = 0 \quad , \quad c_{\alpha} c_{\beta}^{\dagger} + c_{\beta}^{\dagger} c_{\alpha} = \delta_{\alpha\beta}$$

$$Q = \frac{1}{N} \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$U_{\alpha\beta; \gamma\delta}$  are independent random variables with  $\overline{U_{\alpha\beta; \gamma\delta}} = 0$  and  $\overline{|U_{\alpha\beta; \gamma\delta}|^2} = U^2$   
 $N \rightarrow \infty$  yields critical strange metal.



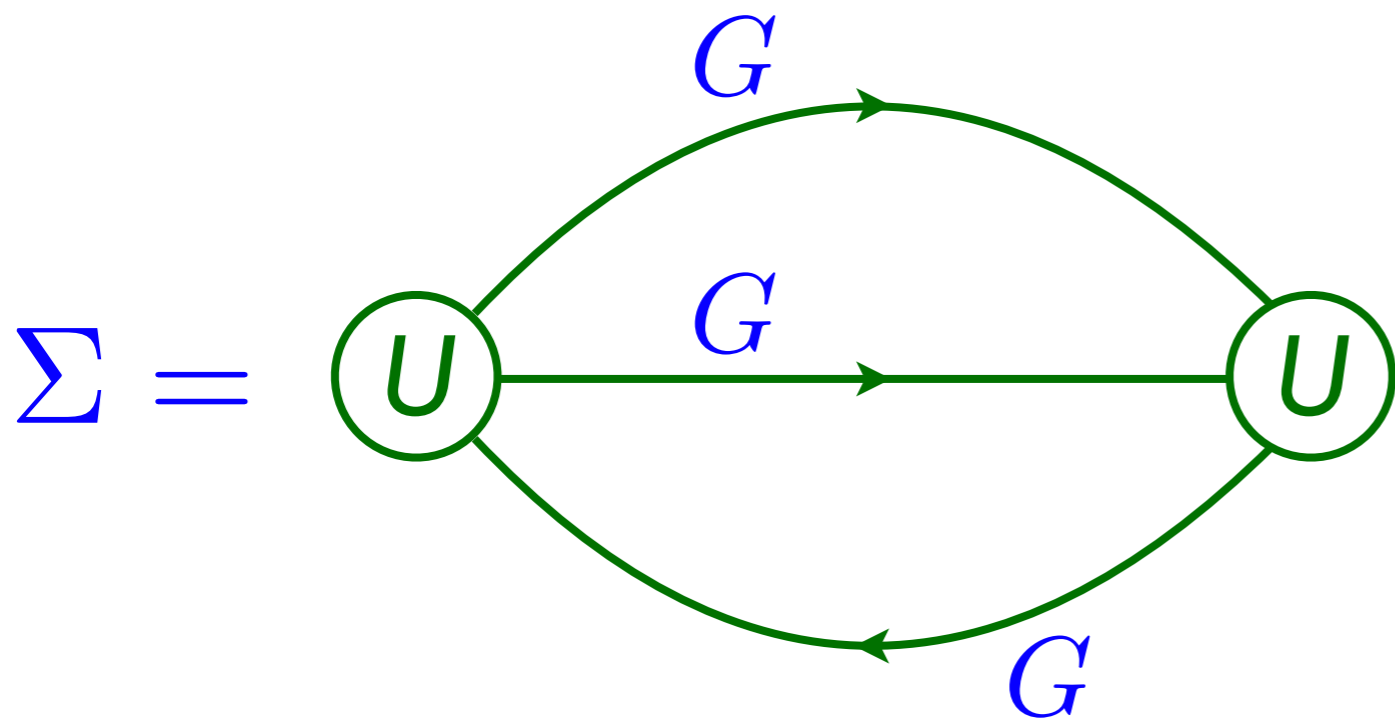
S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

# The complex SYK model

Feynman graph expansion in  $U_{\alpha\beta;\gamma\delta}$ , and graph-by-graph average, yields exact equations in the large  $N$  limit:

$$G(i\omega) = \frac{1}{i\omega - \epsilon - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = Q.$$



S. Sachdev and J. Ye,  
PRL **70**, 3339 (1993)



# The complex SYK model

The large  $N$  limit is given by the sum of “melon” Feynman graphs

For long times  $\tau > 0$

$$\langle c_\alpha(\tau) c_\alpha^\dagger(0) \rangle = e^{\pi\mathcal{E}} \frac{A}{\sqrt{\tau}}$$

$$\langle c_\alpha^\dagger(\tau) c_\alpha(0) \rangle = e^{-\pi\mathcal{E}} \frac{A}{\sqrt{\tau}}$$

The parameter  $\mathcal{E} = \mathbb{C}(\epsilon/U)$  determines the particle-hole asymmetry, and has a universal “Luttinger” relation to  $\mathcal{Q}$ .

In a Fermi liquid,

$$\langle c_\alpha(\tau) c_\alpha^\dagger(0) \rangle = \langle c_\alpha^\dagger(\tau) c_\alpha(0) \rangle = \tilde{A}/\tau$$

# The SYK model



GPS: A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)

Many-body level spacing  $\sim 2^{-N} = e^{-N \ln 2}$

Non-quasiparticle excitations with spacing  $\sim e^{-Ns_0}$

W. Fu and S. Sachdev, PRB **94**, 035135 (2016)

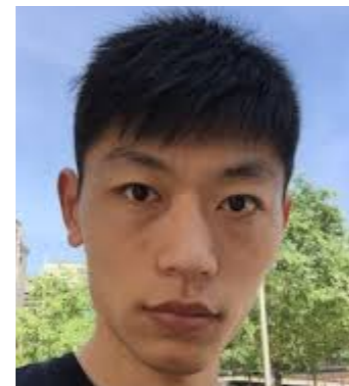
There are  $2^N$  many body levels with energy  $E$ . Shown are all values of  $E$  for a single cluster of size  $N = 12$ . The  $T \rightarrow 0$  state has an entropy  $S_{GPS} = Ns_0$ , where  $s_0 < \ln 2$  is determined by integrating

$$\frac{ds_0}{dQ} = 2\pi\mathcal{E}.$$

At  $Q = 1/2$ ,

$$s_0 = \frac{G}{\pi} + \frac{\ln(2)}{4} = 0.464848\dots$$

where  $G$  is Catalan's constant.



$\sim NU$

# The complex SYK model

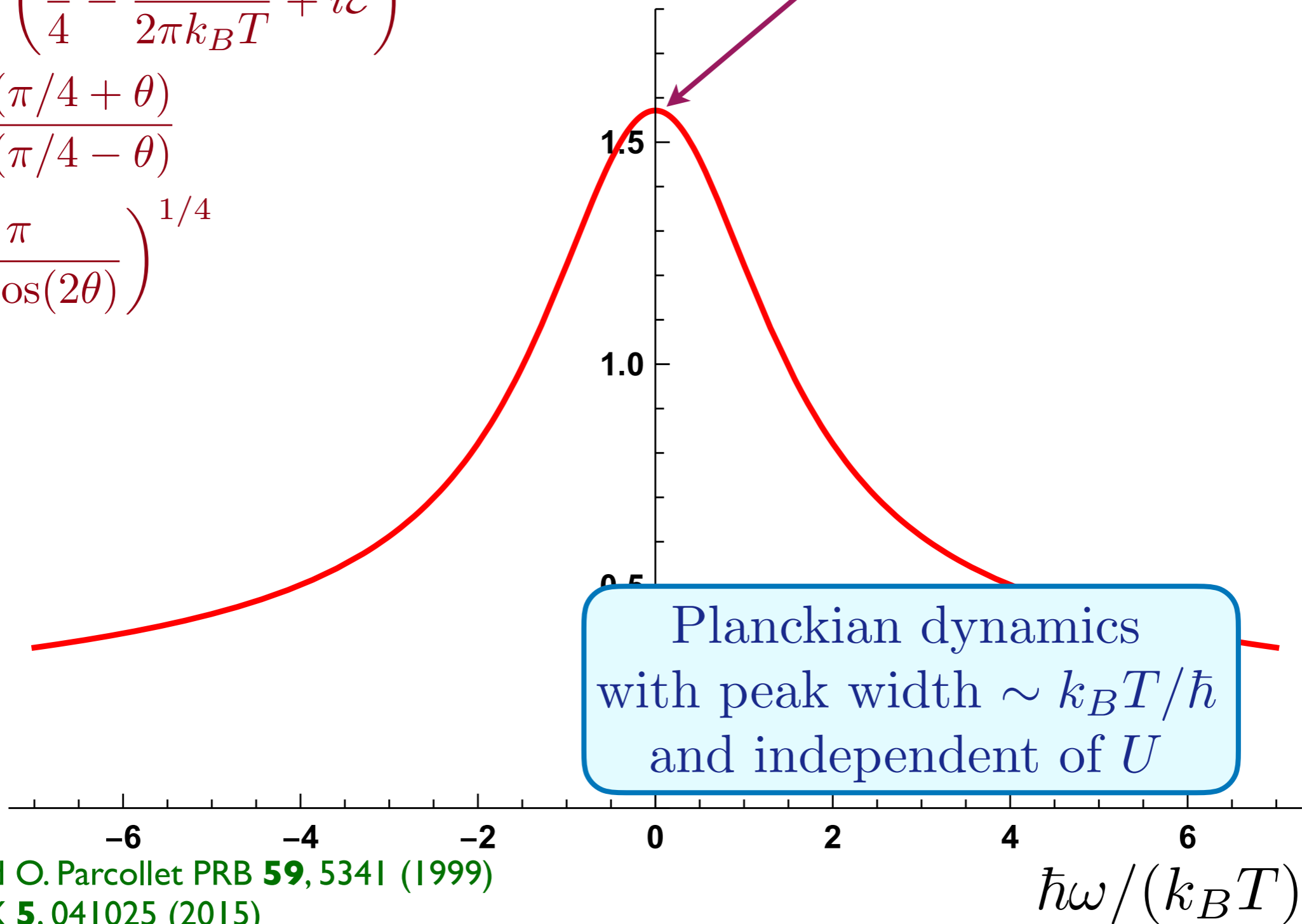
$$\mathcal{E} = \mathbb{C} \frac{\epsilon}{U}$$

$$G_{\text{SYK}}^R(\epsilon, \hbar\omega/(k_B T)) = \frac{-iC e^{-i\theta} \Gamma\left(\frac{1}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}{(2\pi T)^{1/2} \Gamma\left(\frac{3}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}$$

$$e^{2\pi\mathcal{E}} = \frac{\sin(\pi/4 + \theta)}{\sin(\pi/4 - \theta)}$$

$$C = \left(\frac{\pi}{U^2 \cos(2\theta)}\right)^{1/4}$$

$$-\text{Im}G^R(\omega) \quad \mathcal{E} = 0$$



# The complex SYK model

$$\mathcal{E} = \mathbb{C} \frac{\epsilon}{U}$$

$$G_{\text{SYK}}^R(\epsilon, \hbar\omega/(k_B T)) =$$

$$\frac{-iC e^{-i\theta} \Gamma\left(\frac{1}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}{(2\pi T)^{1/2} \Gamma\left(\frac{3}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}$$

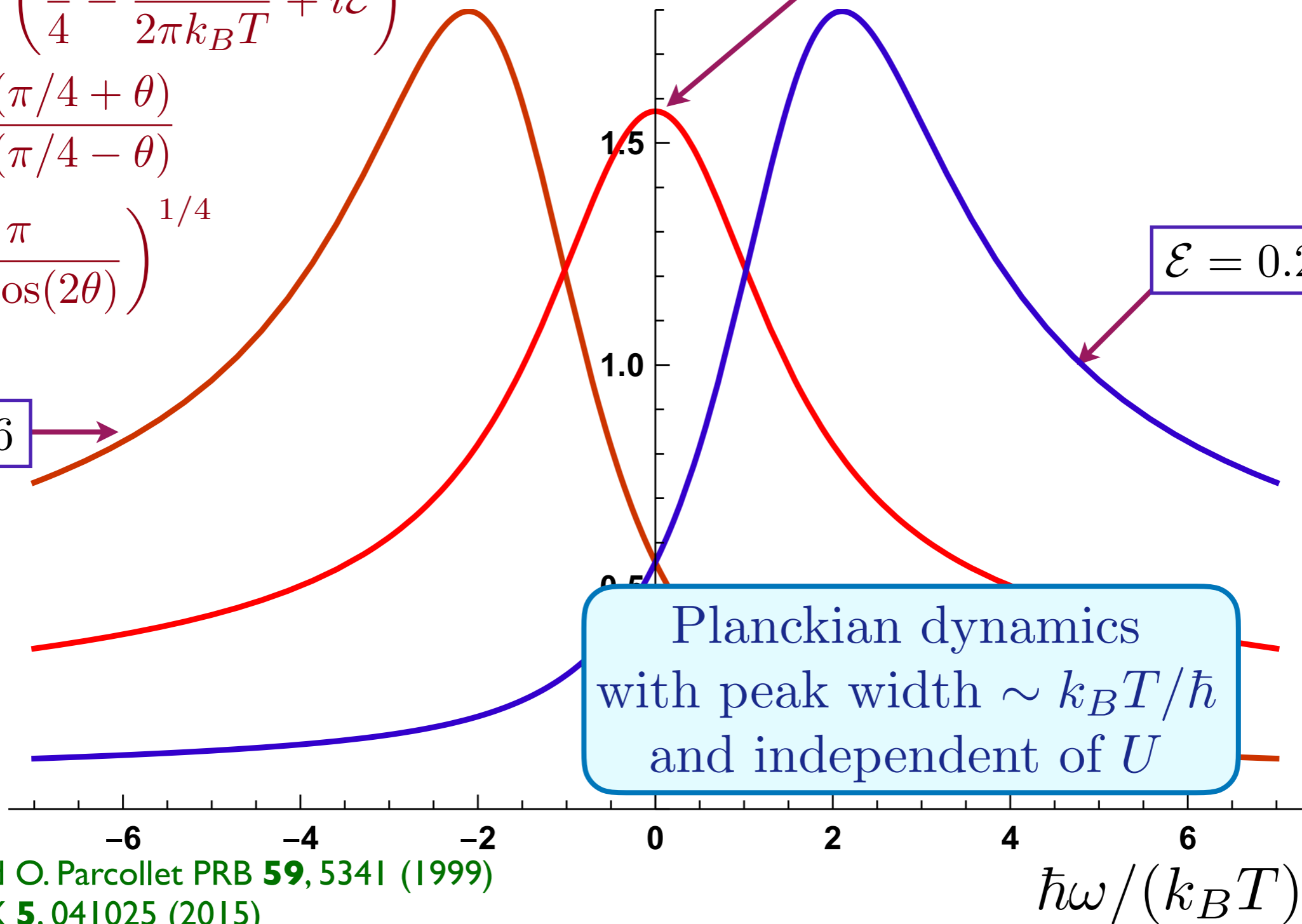
$$e^{2\pi\mathcal{E}} = \frac{\sin(\pi/4 + \theta)}{\sin(\pi/4 - \theta)}$$

$$C = \left(\frac{\pi}{U^2 \cos(2\theta)}\right)^{1/4}$$

$$-\text{Im}G^R(\omega) \quad \mathcal{E} = 0$$

$$\mathcal{E} = 0.26$$

$$\mathcal{E} = -0.26$$



Planckian dynamics  
with peak width  $\sim k_B T/\hbar$   
and independent of  $U$

1. Quantum matter without quasiparticles:  
the complex SYK model

2. Fluctuations, and the Schwarzian

3. Einstein-Maxwell theory of charged  
black holes in AdS space

4. Planckian metals

# The SYK model

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

# The SYK model

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

At frequencies  $\ll U$ , the  $i\omega + \mu$  can be dropped, and without it equations are invariant under the reparametrization and gauge transformations.

The singular part of the self-energy and the Green's function obey

$$\int_0^\beta d\tau_2 \Sigma_{\text{sing}}(\tau_1, \tau_2) G(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$

$$\Sigma_{\text{sing}}(\tau_1, \tau_2) = -U^2 G^2(\tau_1, \tau_2) G(\tau_2, \tau_1)$$

# The SYK model

$$\int_0^\beta d\tau_2 \Sigma(\tau_1, \tau_2) G(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$
$$\Sigma(\tau_1, \tau_2) = -U^2 G^2(\tau_1, \tau_2) G(\tau_2, \tau_1)$$

These equations are invariant under

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{G}(\sigma_1, \sigma_2)$$

$$\Sigma(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{\Sigma}(\sigma_1, \sigma_2)$$

where  $f(\sigma)$  and  $g(\sigma)$  are arbitrary functions.

By using  $f(\sigma) = \tan(\pi T \sigma) / (\pi T)$  we can

now obtain the  $T > 0$  solution from the  $T = 0$  solution.

# The SYK model

Let us write the large  $N$  saddle point solutions of  $S$  as

$$\begin{aligned} G_s(\tau_1 - \tau_2) &\sim (\tau_1 - \tau_2)^{-1/2} \\ \Sigma_s(\tau_1 - \tau_2) &\sim (\tau_1 - \tau_2)^{-3/2}. \end{aligned}$$

The saddle point will be invariant under a reparamaterization  $f(\tau)$  when choosing  $G(\tau_1, \tau_2) = G_s(\tau_1 - \tau_2)$  leads to a transformed  $\tilde{G}(\sigma_1, \sigma_2) = G_s(\sigma_1 - \sigma_2)$  (and similarly for  $\Sigma$ ). It turns out this is true only for the  $SL(2, \mathbb{R})$  transformations under which

$$f(\tau) = \frac{a\tau + b}{c\tau + d}, \quad ad - bc = 1.$$

So the (approximate) reparametrization symmetry is spontaneously broken down to  $SL(2, \mathbb{R})$  by the saddle point.

# Fluctuations

- The saddle-point

$$G(\tau_1 - \tau_2) = -A \frac{e^{-2\pi\mathcal{E}T(\tau_1 - \tau_2)}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left( \frac{T}{\sin(\pi T(\tau_1 - \tau_2))} \right)^{2\Delta}$$

is invariant only under  $\text{PSL}(2, \mathbb{R})$  transformations which map the thermal circle onto itself, and an associated gauge transformation

$$\frac{\tan(\pi T f(\tau))}{\pi T} = \frac{a \frac{\tan(\pi T \tau)}{\pi T} + b}{c \frac{\tan(\pi T \tau)}{\pi T} + d}, \quad ad - bc = 1,$$

$$-i\phi(\tau) = -i\phi_0 + 2\pi\mathcal{E}T(\tau - f(\tau))$$

A. Kitaev, 2015

R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, PRB **95**, 155131 (2017)

# Fluctuations

Symmetry arguments, and explicit computations, show that the effective action is

$$S_{\text{eff}}[f, \phi] = \frac{NK}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi\mathcal{E}T)\partial_\tau f)^2 - \frac{N\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

where  $f(\tau)$  is a monotonic map from  $[0, 1/T]$  to  $[0, 1/T]$ , the couplings  $K$ ,  $\gamma$ , and  $\mathcal{E}$  can be related to thermodynamic derivatives and we have used the Schwarzian:

$$\{g, \tau\} \equiv \frac{g'''}{g'} - \frac{3}{2} \left( \frac{g''}{g'} \right)^2.$$

Specifically, an argument constraining the effective at  $T = 0$  is

$$S_{\text{eff}} \left[ f(\tau) = \frac{a\tau + b}{c\tau + d}, \phi(\tau) = 0 \right] = 0,$$

and this is origin of the Schwarzian.

J. Maldacena and D. Stanford, arXiv:1604.07818;  
R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, PRB **95**, 155131 (2017);  
A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746

1. Quantum matter without quasiparticles:  
the complex SYK model

2. Fluctuations, and the Schwarzian

3. Einstein-Maxwell theory of charged  
black holes in AdS space

4. Planckian metals

# Quantum Black holes

- Black holes have an entropy and a temperature,  $T_H$
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a Planckian time  $\sim \hbar/(k_B T_H)$ .

J. D. Bekenstein, PRD **7**, 2333 (1973)  
S.W. Hawking, Nature **248**, 30 (1974)



C. V. Vishveshwara, Nature **227**, 936 (1970)

# Quantum Black holes

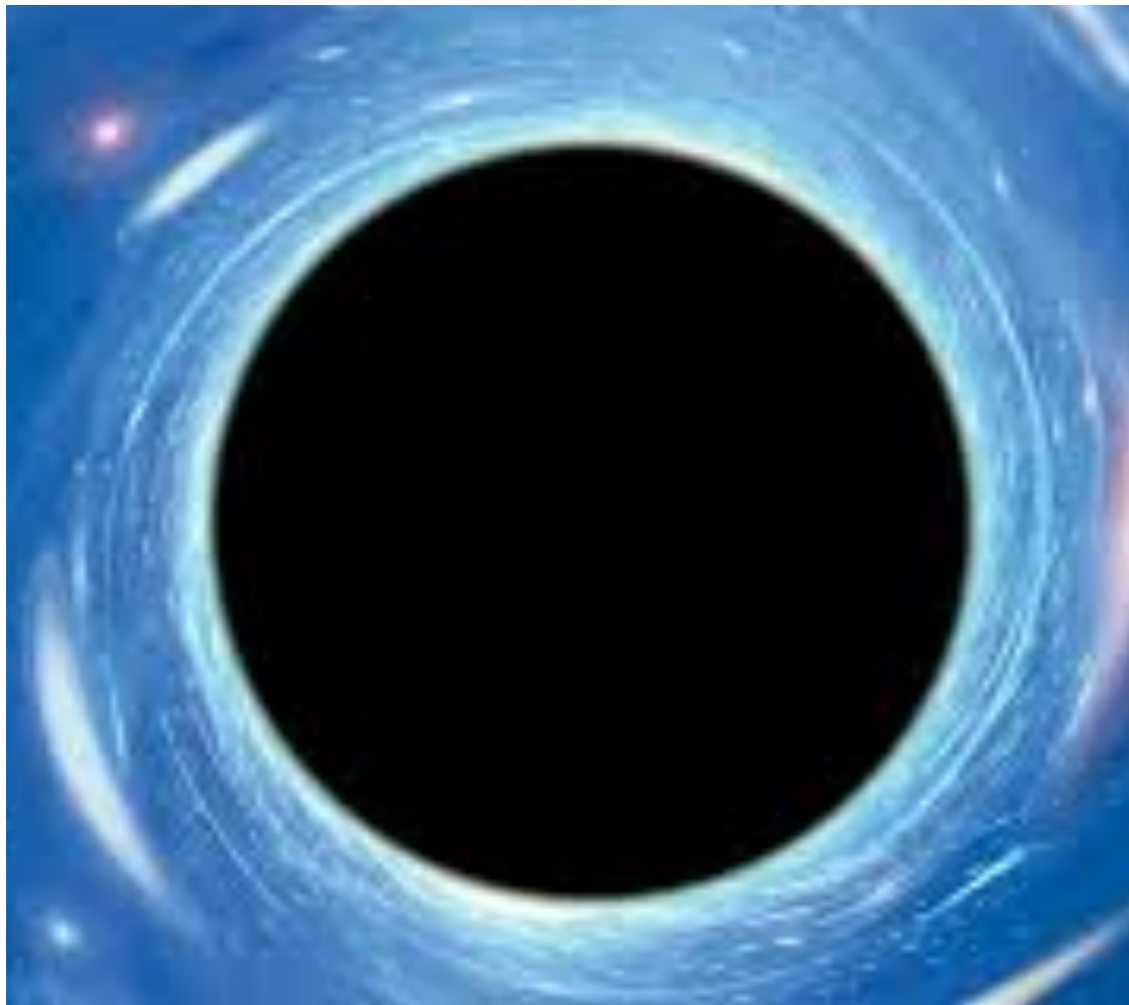
- Black holes have an entropy and a temperature,  $T_H$
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a Planckian time  $\sim \hbar/(k_B T_H)$ .

## Holography:

Quantum black holes “look like” quantum many-particle systems without quasiparticle excitations, residing “on” the surface of the black hole

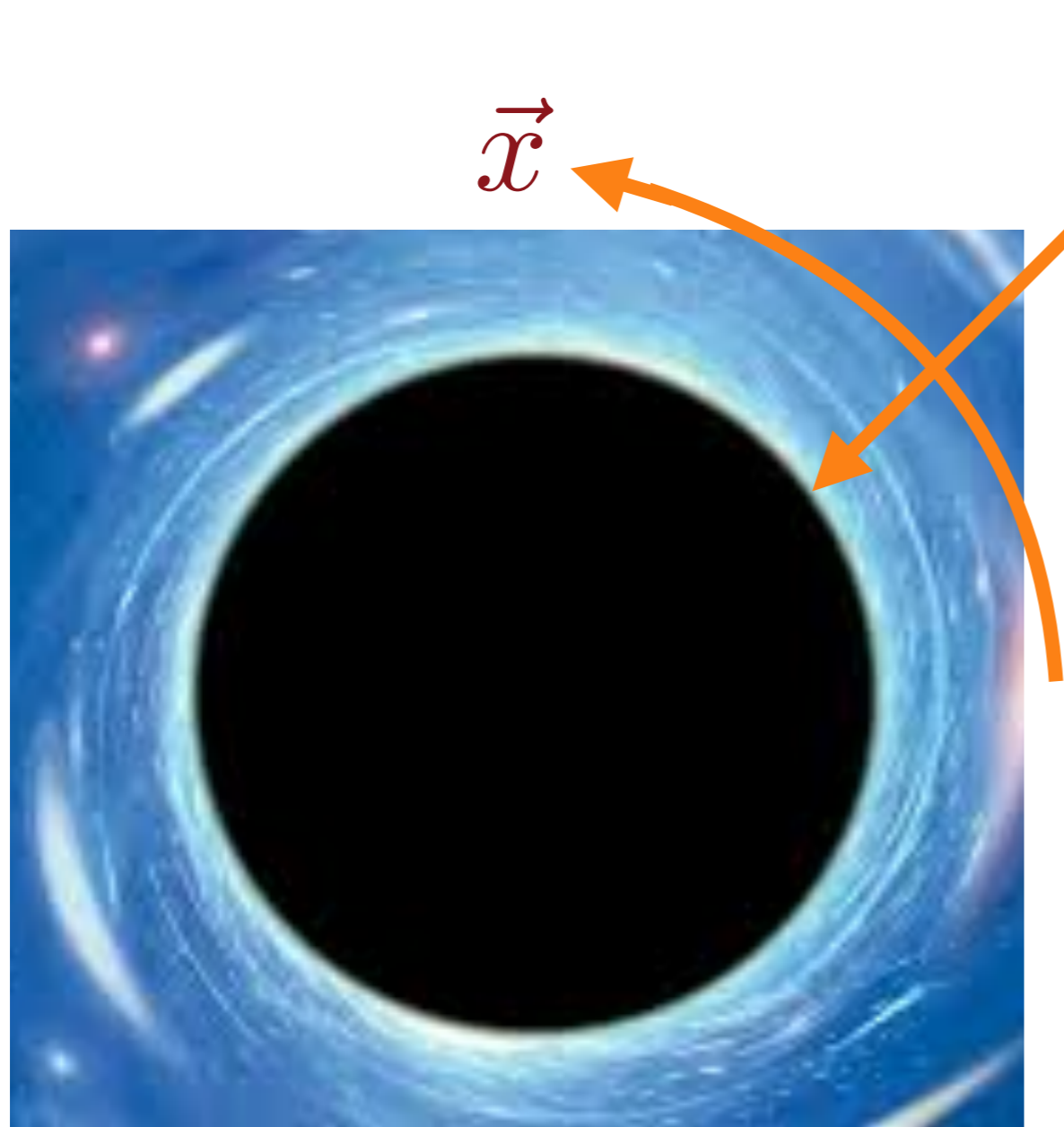


Work with a theory of Maxwell's electromagnetism and Einstein's general relativity. Include a negative cosmological constant, and examine black hole solutions with a net charge





Work with a theory of Maxwell's electromagnetism and Einstein's general relativity. Include a negative cosmological constant, and examine black hole solutions with a net charge

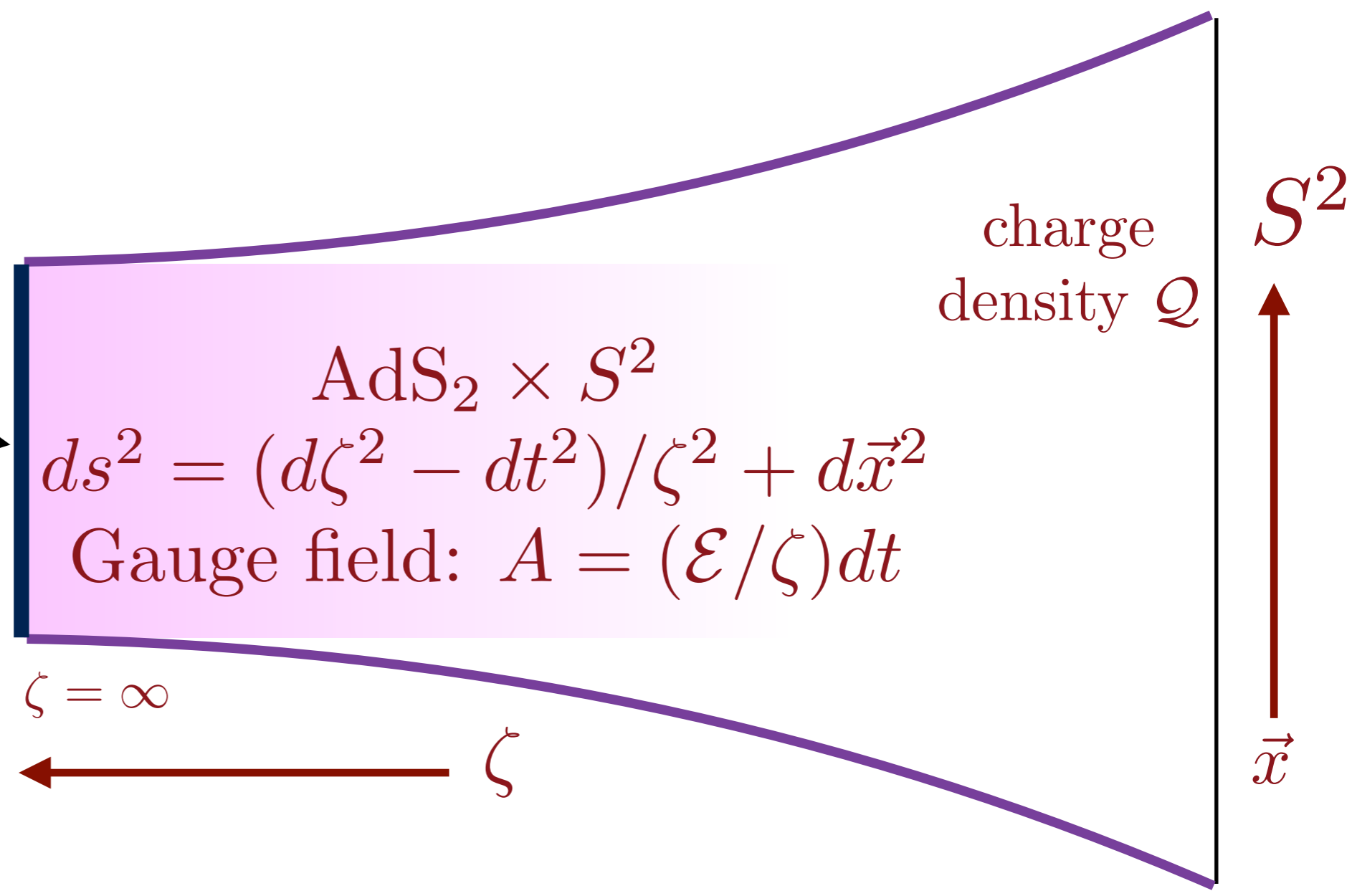


Zooming into the near-horizon region of a charged black hole at low temperature, yields a quantum theory in one space ( $\zeta$ ) and one time dimension

# SYK model and charged black holes



Black hole horizon

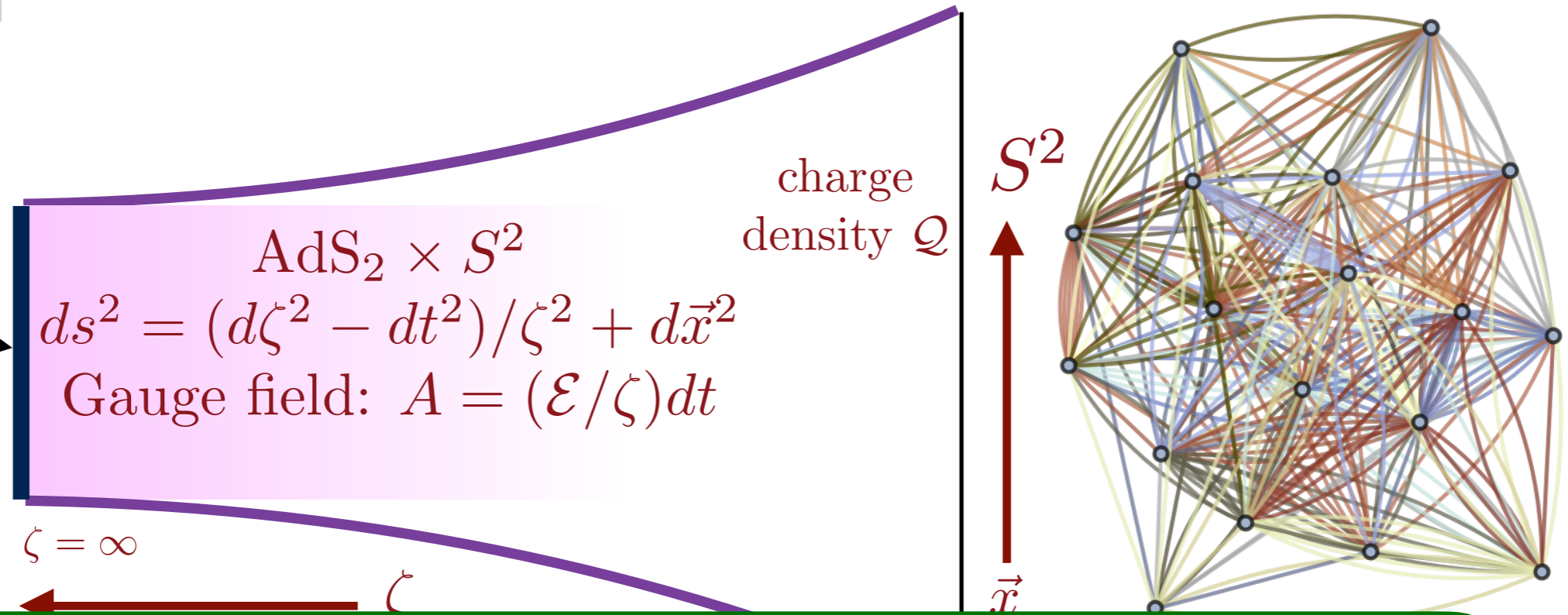


The near-horizon region of a charged black hole has the geometry of (1+1)-dimensional anti-de Sitter spacetime. By holography, this should map to a zero-dimensional quantum system: this turns out to be the SYK model

# SYK model and charged black holes



Black hole horizon



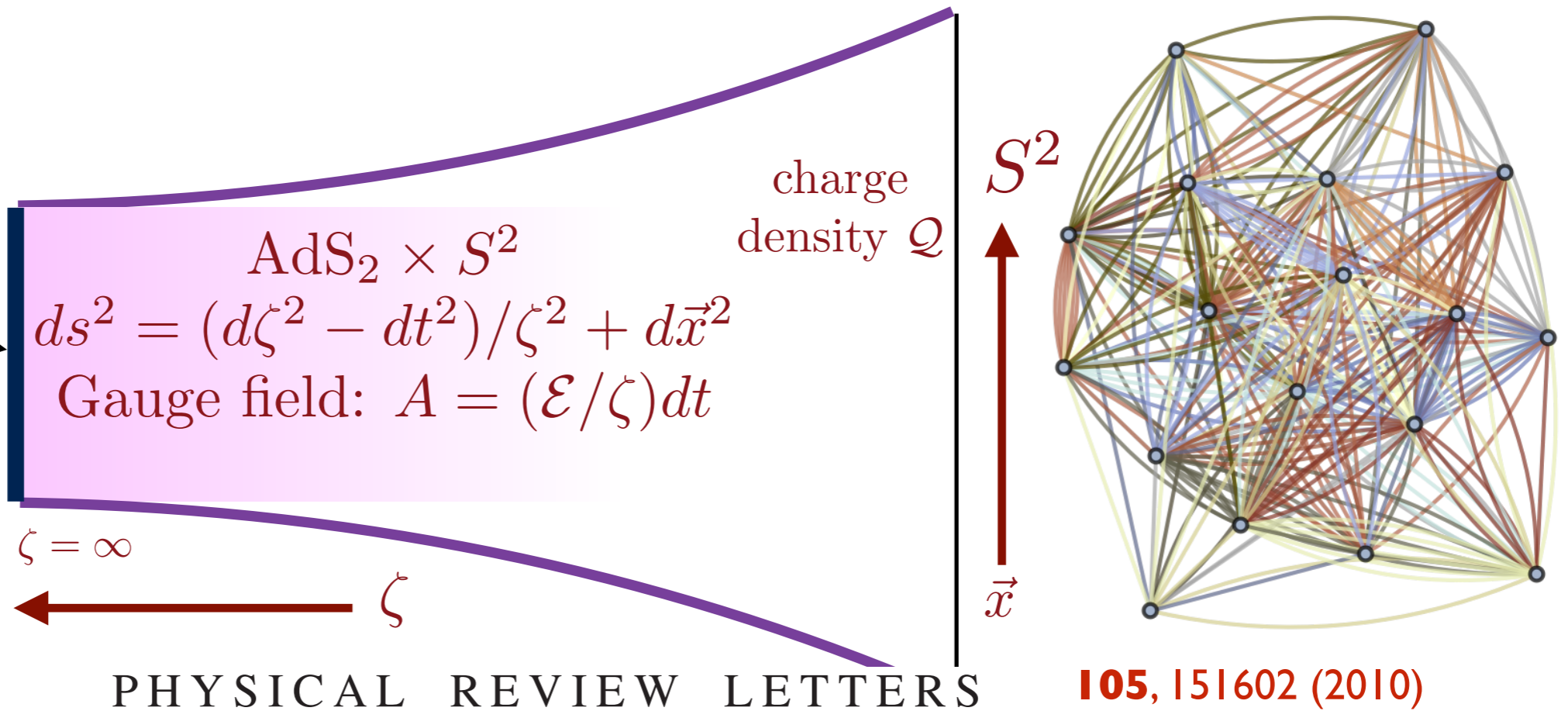
Bekenstein-Hawking entropy of  $\text{AdS}_2$  horizon at  $T = 0 \Leftrightarrow N s_0$  entropy of SYK model.

$\frac{\partial s_0}{\partial \mathcal{Q}} = 2\pi\mathcal{E}$  holds for both the black hole and the SYK model, where  $\mathcal{E}$  determines identical fermion spectral functions.

# SYK model and charged black holes



Black hole horizon



PHYSICAL REVIEW LETTERS

**105, 151602 (2010)**



## Holographic Metals and the Fractionalized Fermi Liquid

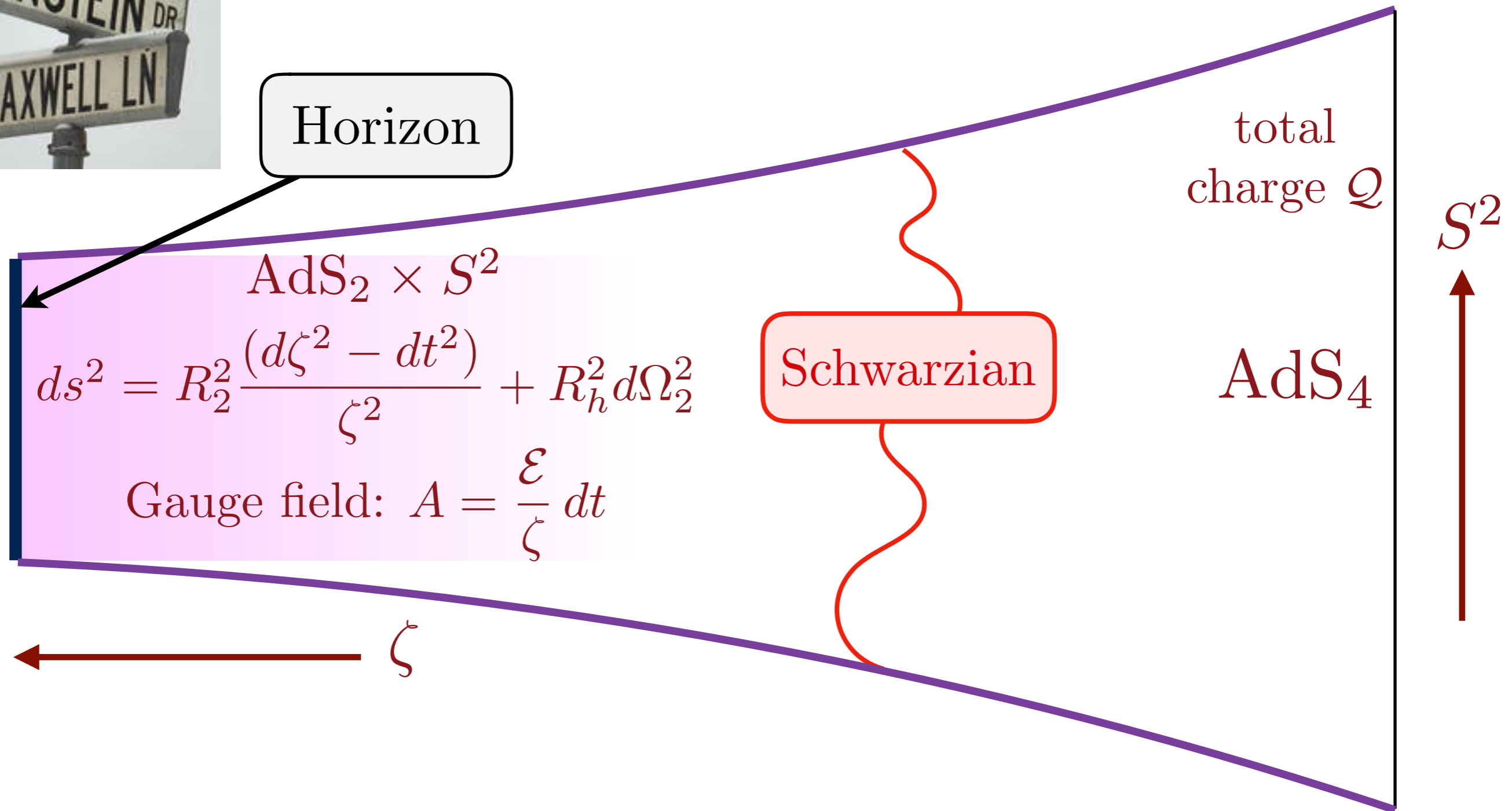
Subir Sachdev

*Department of Physics, Harvard University, Cambridge, Massachusetts 02138, USA*

*(Received 23 June 2010; published 4 October 2010)*

We show that there is a close correspondence between the physical properties of holographic metals near charged black holes in anti-de Sitter (AdS) space, and the fractionalized Fermi liquid phase of the lattice Anderson model. The latter phase has a “small” Fermi surface of conduction electrons, along with a spin liquid of local moments. This correspondence implies that certain mean-field gapless spin liquids are states of matter at nonzero density realizing the near-horizon,  $AdS_2 \times R^2$  physics of Reissner-Nordström black holes.

# SYK model and charged black holes

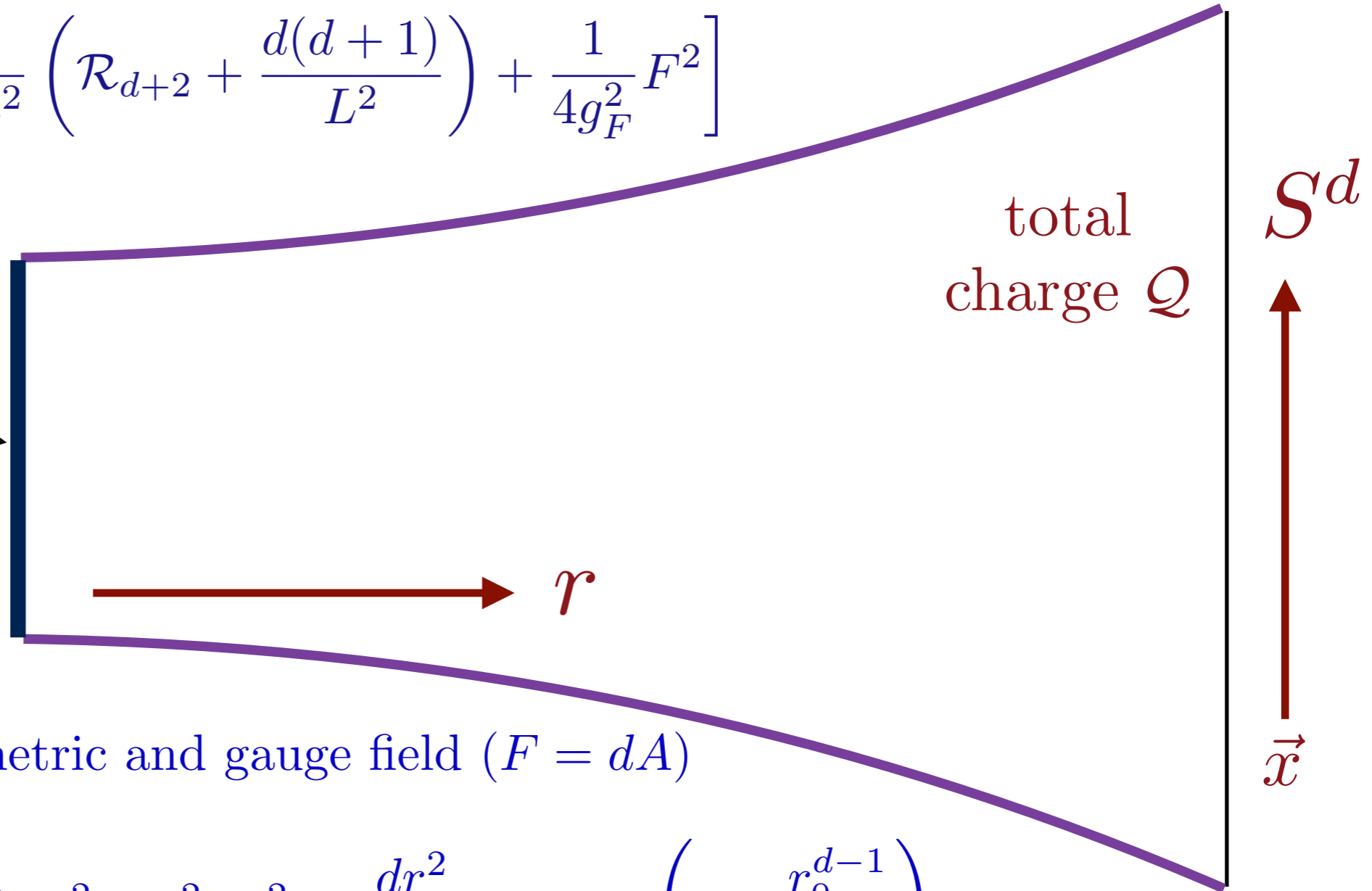
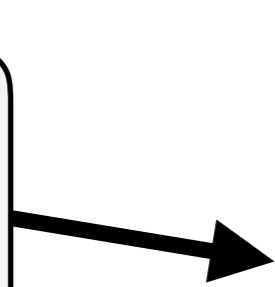


Remarkably, the correspondence between charged black holes and the SYK model also holds for the leading fluctuations at higher temperatures: both are described by a ‘Schwarzian’ theory with emergent  $SL(2, \mathbb{R})$  and  $U(1)$  gauge symmetries. For the black hole, the Schwarzian describes the fluctuations of the boundary between  $AdS_2$  and  $AdS_4$ .

# Charged black holes

$$I_{EM} = \int d^{d+2}x \sqrt{g} \left[ -\frac{1}{2\kappa^2} \left( \mathcal{R}_{d+2} + \frac{d(d+1)}{L^2} \right) + \frac{1}{4g_F^2} F^2 \right]$$

Black hole horizon of radius  $r_0$



Solutions of  $I_{EM}$  have metric and gauge field ( $F = dA$ )

$$ds^2 = V(r)d\tau^2 + r^2 d\Omega_d^2 + \frac{dr^2}{V(r)} \quad , \quad i\mu \left( 1 - \frac{r_0^{d-1}}{r^{d-1}} \right) d\tau$$

$$V(r) = 1 + \frac{r^2}{L^2} + \frac{\Theta^2}{r^{2d-2}} - \frac{M}{r^{d-1}}.$$

where  $d\Omega_d^2$  is the metric of the  $d$ -sphere. All parameters of the solution are determined in terms of the chemical potential  $\mu$ , and the Hawking temperature of horizon,  $T$ .

# Charged black holes

We write the  $(d+2)$ -dimensional metric  $g$  of  $I_{EM}$  in terms of a two-dimensional metric  $h$  and a scalar field  $\Phi$ :

$$ds^2 = \frac{ds_2^2}{\Phi^{d-1}} + \Phi^2 d\Omega_d^2.$$

The Einstein-Maxwell and Gibbons-Hawking actions reduce to an extension of Jackiw-Teitelbaum gravity ( $x \equiv (\tau, \zeta)$ )

$$I_{EM} = \int d^2x \sqrt{h} \left[ -\frac{s_d}{2\kappa^2} \Phi^d \mathcal{R}_2 + U(\Phi) + \frac{Z(\Phi)}{4g_F^2} F^2 \right]$$
$$I_{GH} = -\frac{s_d}{\kappa^2} \int_{\partial} dx \sqrt{h_b} \Phi^d \mathcal{K}_1$$

The explicit forms of the potentials  $U(\Phi)$  and  $Z(\Phi)$  are,

$$U(\Phi) = -\frac{s_d}{2\kappa^2} \left( \frac{d(d-1)}{\Phi} + \frac{d(d+1)\Phi}{L^2} \right), \quad Z(\Phi) = s_d \Phi^{2d-1}.$$

# Charged black holes

The exact saddle point of  $\Phi$  relates to  $R_h$  the horizon radius at  $T = 0$

$$\Phi(\zeta) = R_h + \frac{R_2^2}{\zeta} \quad , \quad R_h \equiv \frac{L}{g_F} \left[ \frac{(d-1)(\mu_0^2 \kappa^2 (d-1) - dg_F^2)}{d(d+1)} \right]^{1/2} \quad ,$$

while the near-horizon, low  $T \ll 1/R_h$  metric is  $\text{AdS}_2$

$$ds_2^2 = \frac{R_2^2 R_h^{d-1}}{\zeta^2} \left[ (1 - 4\pi^2 T^2 \zeta^2) d\tau^2 + \frac{d\zeta^2}{1 - 4\pi^2 T^2 \zeta^2} \right] \quad ,$$

where

$$R_2 = \frac{LR_h}{\sqrt{d(d+1)R_h^2 + (d-1)^2 L^2}}$$

The field coupling to  $\mathcal{R}_2$  is  $\Phi^d$

$$[\Phi(\zeta)]^d = R_h^d + \frac{\Phi_1}{\zeta} + \dots \quad , \quad \Phi_1 = dR_h^{d-1} R_2^2 \quad ,$$

# Charged black holes

The field coupling to  $\mathcal{R}_2$  is  $\Phi^d$

$$[\Phi(\zeta)]^d = R_h^d + \frac{\Phi_1}{\zeta} + \dots \quad , \quad \Phi_1 = dR_h^{d-1}R_2^2,$$

We choose the boundary of the  $\text{AdS}_2$  region at bulk co-ordinates  $(f(\tau), \zeta(\tau))$  with the induced boundary metric fixed at  $(R_2^2 R_h^{d-1} / \zeta_b^2) d\tau^2$  by choosing

$$\zeta(\tau) = \zeta_b f'(\tau) + \zeta_b^3 \left( \frac{[f''(\tau)]^2}{2f'(\tau)} - 2\pi^2 T^2 [f'(\tau)]^3 \right) + \dots$$

Finally, we evaluate  $I_{GH}$  along this boundary curve

$$I_1[f] = -\frac{\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

where

$$\gamma = \frac{4\pi^2 s_d \Phi_1}{\kappa^2},$$

matches the linear-in- $T$  co-efficient of the specific heat of the full Reissner-Nördstorm solution in  $d + 2$  dimensions.

1. Quantum matter without quasiparticles:  
the complex SYK model

2. Fluctuations, and the Schwarzian

3. Einstein-Maxwell theory of charged  
black holes in AdS space

4. Planckian metals

Remarkable recent observation of  
'Planckian' strange metal transport in cuprates,  
pnictides, magic-angle graphene, and  
ultracold atoms: the resistivity,  $\rho$ , is

$$\rho = \frac{m^*}{ne^2} \frac{1}{\tau}$$

with a universal scattering rate

$$\frac{1}{\tau} \approx \frac{k_B T}{\hbar},$$

independent of the strength of interactions!



Remarkable recent observation of ‘Planckian’ strange metal transport in cuprates, pnictides, magic-angle graphene, and ultracold atoms: the resistivity is associated with a universal scattering time  $\approx \hbar/(k_B T)$ .

## Universal $T$ -linear resistivity and Planckian dissipation in overdoped cuprates

NATURE PHYSICS | VOL 15 | FEBRUARY 2019 | 142-147

A. Legros<sup>1,2</sup>, S. Benhabib<sup>3</sup>, W. Tabis<sup>3,4</sup>, F. Laliberté<sup>1</sup>, M. Dion<sup>1</sup>, M. Lizaire<sup>1</sup>, B. Vignolle<sup>3</sup>, D. Vignolles<sup>3</sup>, H. Raffy<sup>5</sup>, Z. Z. Li<sup>5</sup>, P. Auban-Senzier<sup>5</sup>, N. Doiron-Leyraud<sup>1</sup>, P. Fournier<sup>1,6</sup>, D. Colson<sup>2</sup>, L. Taillefer<sup>1,6\*</sup> and C. Proust<sup>3,6\*</sup>

arXiv:1902.01034

## Planckian dissipation and scale invariance in a quantum-critical disordered pnictide

Yasuyuki Nakajima,<sup>1,2</sup> Tristin Metz,<sup>2</sup> Christopher Eckberg,<sup>2</sup> Kevin Kirshenbaum,<sup>2</sup> Alex Hughes,<sup>2</sup> Renxiong Wang,<sup>2</sup> Limin Wang,<sup>2</sup> Shanta R. Saha,<sup>2</sup> I-Lin Liu,<sup>2,3,4</sup> Nicholas P. Butch,<sup>2,4</sup> Zhonghao Liu,<sup>5,6</sup> Sergey V. Borisenko,<sup>5</sup> Peter Y. Zavalij,<sup>7</sup> and Johnpierre Paglione<sup>2,8</sup>

## Strange metal in magic-angle graphene with near Planckian dissipation

Yuan Cao,<sup>1,\*</sup> Debanjan Chowdhury,<sup>1,\*</sup> Daniel Rodan-Legrain,<sup>1</sup> Oriol Rubies-Bigordà,<sup>1</sup> Kenji Watanabe,<sup>2</sup> Takashi Taniguchi,<sup>2</sup> T. Senthil,<sup>1,†</sup> and Pablo Jarillo-Herrero<sup>1,†</sup>

arXiv:1901.03710

## Bad metallic transport in a cold atom Fermi-Hubbard system

*Science* **363**, 379–382 (2019)

Peter T. Brown<sup>1</sup>, Debayan Mitra<sup>1</sup>, Elmer Guardado-Sanchez<sup>1</sup>, Reza Nourafkan<sup>2</sup>, Alexis Reymbaut<sup>2</sup>, Charles-David Hébert<sup>2</sup>, Simon Bergeron<sup>2</sup>, A.-M. S. Tremblay<sup>2,3</sup>, Jure Kokalj<sup>4,5</sup>, David A. Huse<sup>1</sup>, Peter Schauf<sup>1\*</sup>, Waseem S. Bakr<sup>1†</sup>

Material		$n$ ( $10^{27} \text{ m}^{-3}$ )	$m^*$ ( $m_0$ )	$A_1 / d$ ( $\Omega / \text{K}$ )	$h / (2e^2 T_F)$ ( $\Omega / \text{K}$ )	$\alpha$
Bi2212	$p = 0.23$	6.8	$8.4 \pm 1.6$	$8.0 \pm 0.9$	$7.4 \pm 1.4$	$1.1 \pm 0.3$
Bi2201	$p \sim 0.4$	3.5	$7 \pm 1.5$	$8 \pm 2$	$8 \pm 2$	$1.0 \pm 0.4$
LSCO	$p = 0.26$	7.8	$9.8 \pm 1.7$	$8.2 \pm 1.0$	$8.9 \pm 1.8$	$0.9 \pm 0.3$
Nd-LSCO	$p = 0.24$	7.9	$12 \pm 4$	$7.4 \pm 0.8$	$10.6 \pm 3.7$	$0.7 \pm 0.4$
PCCO	$x = 0.17$	8.8	$2.4 \pm 0.1$	$1.7 \pm 0.3$	$2.1 \pm 0.1$	$0.8 \pm 0.2$
LCCO	$x = 0.15$	9.0	$3.0 \pm 0.3$	$3.0 \pm 0.45$	$2.6 \pm 0.3$	$1.2 \pm 0.3$
TMTSF	$P = 11 \text{ kbar}$	1.4	$1.15 \pm 0.2$	$2.8 \pm 0.3$	$2.8 \pm 0.4$	$1.0 \pm 0.3$

**Slope of  $T$ -linear resistivity vs Planckian limit in seven materials.**

$$\frac{1}{\tau} = \alpha \frac{k_B T}{\hbar}$$

A. Legros, S. Benhabib, W. Tabis, F. Laliberté, M. Dion, M. Lizaire, B. Vignolle, D. Vignolles, H. Raffy, Z. Z. Li, P. Auban-Senzier, N. Doiron-Leyraud, P. Fournier, D. Colson, L. Taillefer, and C. Proust, *Nature Physics* **15**, 142 (2019)

# The complex SYK model

$$H = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} + \epsilon \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$U_{\alpha\beta;\gamma\delta}$  are independent random variables

with  $\overline{U_{\alpha\beta;\gamma\delta}} = 0$  and  $\overline{|U_{\alpha\beta;\gamma\delta}|^2} = U^2$

# Generalized SYK models

$$H = \frac{1}{(2N)^{3/2}} \sum_{k_a} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta}(k_a) c_{k_1\alpha}^\dagger c_{k_2\beta}^\dagger c_{k_3\gamma} c_{k_4\delta} \\ + \sum_{k\alpha} \epsilon_k c_{k\alpha}^\dagger c_{k\alpha}$$

$U_{\alpha\beta;\gamma\delta}(k_a)$  is a random function of  $\alpha\beta\gamma\delta$  (as before)  
 $\epsilon_k$  has a range of values of width  $W$ .

The large  $N$  limit is still given by the sum of “melon” diagrams.

Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

Pengfei Zhang, PRB **96**, 205138 (2017)

Debanjan Chowdhury, Yochai Werman, Erez Berg, T. Senthil, PRX **8**, 031024 (2018)

Aavishkar A. Patel, John McGreevy, Daniel P. Arovas, Subir Sachdev, PRX **8**, 021049 (2018)

See also Antoine Georges and Olivier Parcollet PRB **59**, 5341 (1999)

# Generalized SYK models

$$H = \frac{1}{(2N)^{3/2}} \sum_{k_a} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta}(k_a) c_{k_1\alpha}^\dagger c_{k_2\beta}^\dagger c_{k_3\gamma} c_{k_4\delta} \\ + \sum_{k\alpha} \epsilon_k c_{k\alpha}^\dagger c_{k\alpha}$$

$U_{\alpha\beta;\gamma\delta}(k_a)$  is a random function of  $\alpha\beta\gamma\delta$  (as before)  
 $\epsilon_k$  has a range of values of width  $W$ .

The large  $N$  limit is still given by the sum of “melon” diagrams.

For many generic models in this class,  $\hbar\omega/(k_B T)$  scaling of SYK holds for  $W^2/U \ll k_B T \ll U$ , and Fermi liquid theory is recovered for  $k_B T \ll W^2/U$ .

Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

Pengfei Zhang, PRB **96**, 205138 (2017)

Debanjan Chowdhury, Yochai Werman, Erez Berg, T. Senthil, PRX **8**, 031024 (2018)

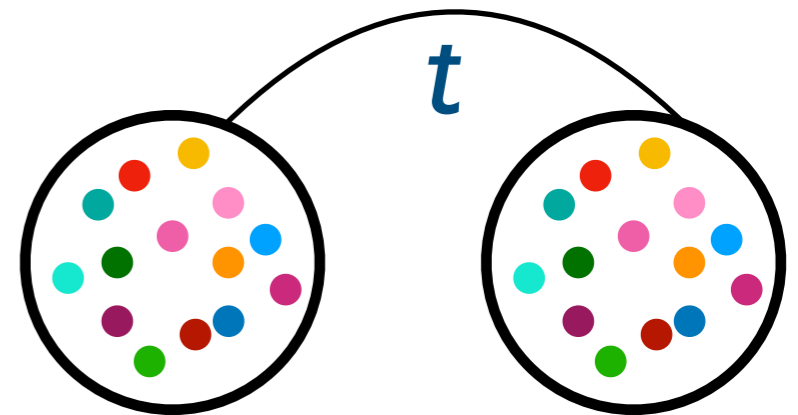
Aavishkar A. Patel, John McGreevy, Daniel P. Arovas, Subir Sachdev, PRX **8**, 021049 (2018)

See also Antoine Georges and Olivier Parcollet PRB **59**, 5341 (1999)

$$H = \frac{1}{(2N)^{3/2}} \sum_i \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta; \gamma\delta} c_{i\alpha}^\dagger c_{i\beta}^\dagger c_{i\gamma} c_{i\delta} - t \sum_{\langle ij \rangle} \sum_{\alpha} c_{i\alpha}^\dagger c_{j\alpha}$$

Equivalent to an  
 “eternal traversable wormhole”  
 between two black holes with  
 AdS<sub>2</sub> horizons

$U$



J. Maldacena and Xiao-Liang Qi, arXiv:1804.00491

# A lattice SYK model

$$H = \frac{1}{(2N)^{3/2}} \sum_i \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta; \gamma\delta} c_{i\alpha}^\dagger c_{i\beta}^\dagger c_{i\gamma} c_{i\delta} - t \sum_{\langle ij \rangle} \sum_{\alpha} c_{i\alpha}^\dagger c_{j\alpha}$$

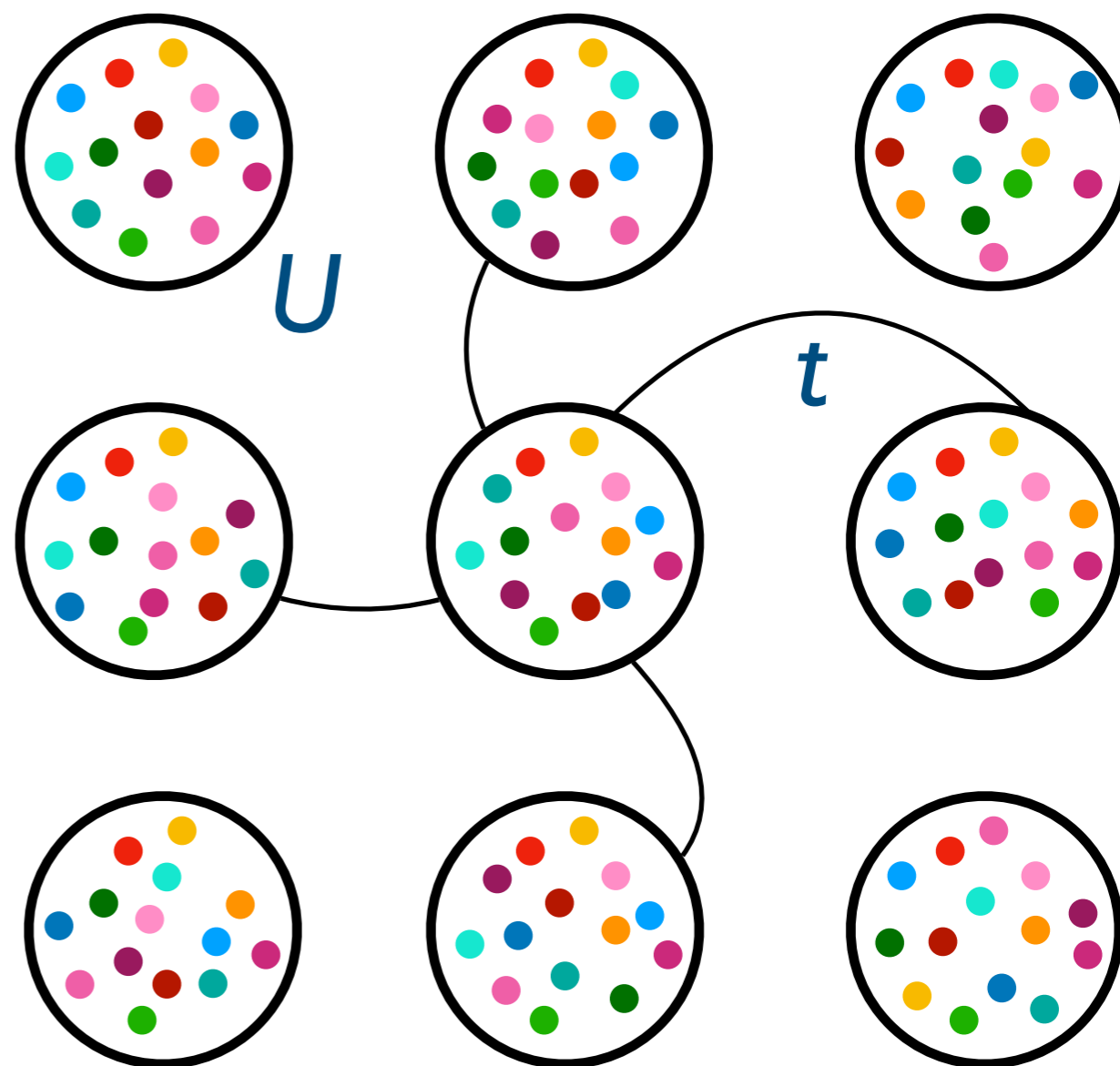
Choose  $U$  on-site,  
and uncorrelated between sites;  
yields 'incoherent metal'  
with no Fermi surface  
for  $t^2/U \ll k_B T \ll U$  with

$$G(\mathbf{k}, \omega) = G_{\text{SYK}}(\epsilon, \hbar\omega / (k_B T))$$

independent of  $\mathbf{k}$ .

There is linear-in- $T$  resistivity  
but only with  $\rho > h/e^2$ .

$$\rho \sim \frac{h}{e^2} \frac{k_B T}{t^2/U}$$



Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

Pengfei Zhang, PRB **96**, 205138 (2017)

Debanjan Chowdhury, Yochai Werman, Erez Berg, T. Senthil, PRX **8**, 031024 (2018)

Aavishkar A. Patel, John McGreevy, Daniel P. Arovas, Subir Sachdev, PRX **8**, 021049 (2018)

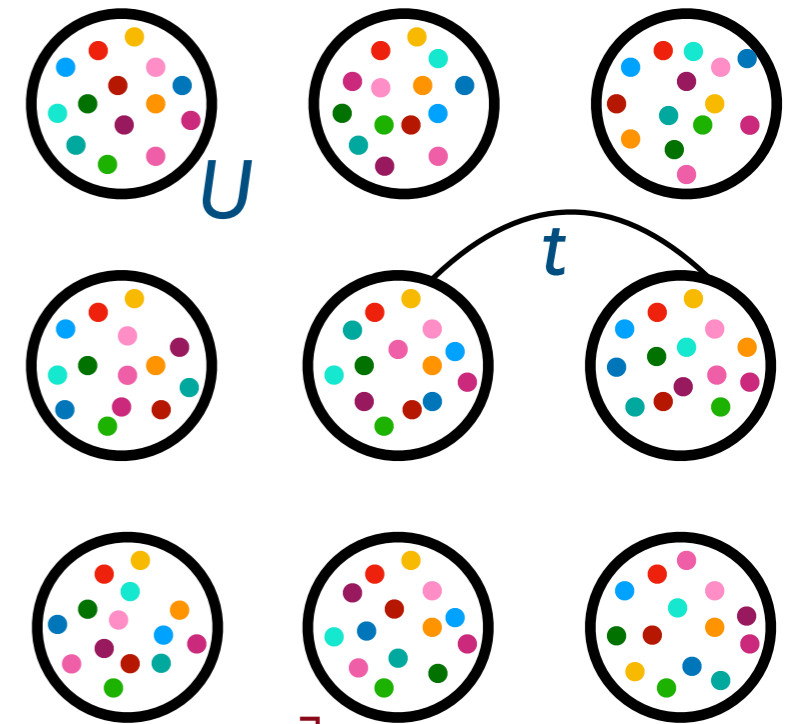
See also Antoine Georges and Olivier Parcollet PRB **59**, 5341 (1999)

# A lattice SYK model

$$H = \frac{1}{(2N)^{3/2}} \sum_{k_a} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta}(k_a) c_{k_1\alpha}^\dagger c_{k_2\beta}^\dagger c_{k_3\gamma} c_{k_4\delta} + \sum_{k\alpha} \epsilon_k c_{k\alpha}^\dagger c_{k\alpha}$$

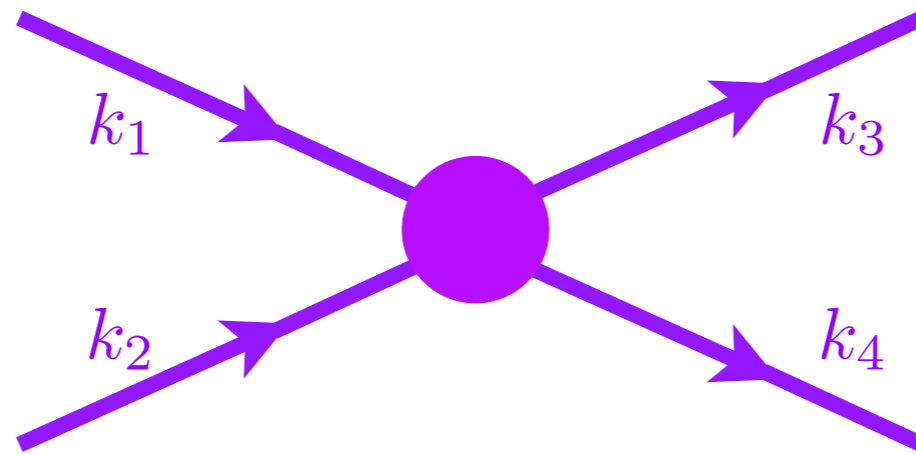
$U_{\alpha\beta;\gamma\delta}(k_a)$  is a random function of  $\alpha\beta\gamma\delta$   
 $\epsilon_k$  has a bandwidth  $W$ .

Rewriting of lattice model of incoherent and bad metal in momentum space



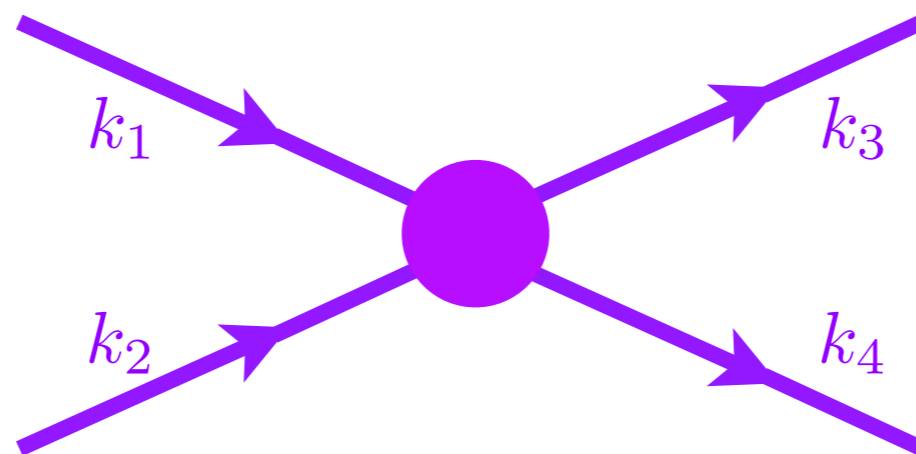
$$\overline{U(k_1, k_2, k_3, k_4) U^*(k_5, k_6, k_7, k_8)} = U^2 \left[ \delta(k_1 + k_2 - k_3 - k_4 - k_5 - k_6 + k_7 + k_8) \right]$$

# Resonant SYK model



Interactions with  $\epsilon_{k_1} + \epsilon_{k_2} \neq \epsilon_{k_3} + \epsilon_{k_4}$  are non-resonant: we “integrate these out” in a RG procedure, and assume that their main effect is a renormalization of the quasiparticle dispersion  $\epsilon_k$ , which we have already accounted for.

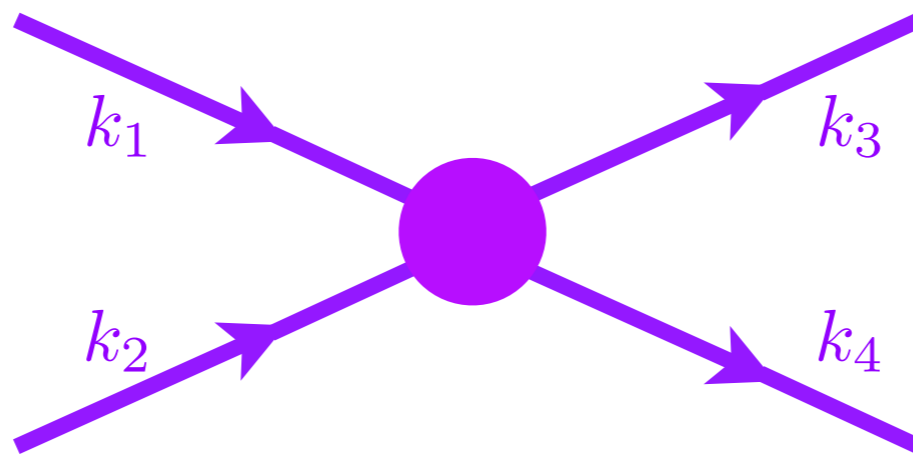
# Resonant SYK model



Interactions with  $\epsilon_{k_1} + \epsilon_{k_2} \neq \epsilon_{k_3} + \epsilon_{k_4}$  are non-resonant: we “integrate these out” in a RG procedure, and assume that their main effect is a renormalization of the quasiparticle dispersion  $\epsilon_k$ , which we have already accounted for.

Keep only the interactions resonant in the bare quasiparticle energy with  $\epsilon_{k_1} + \epsilon_{k_2} = \epsilon_{k_3} + \epsilon_{k_4}$  and account for them with a self-consistent SYK-like analysis.

# Resonant SYK model



Interactions with  $\epsilon_{k_1} + \epsilon_{k_2} \neq \epsilon_{k_3} + \epsilon_{k_4}$  are non-resonant: we “integrate these out” in a RG procedure, and assume that their main effect is a renormalization of the quasiparticle dispersion  $\epsilon_k$ , which we have already accounted for.

Keep only the interactions resonant in the bare quasiparticle energy with  $\epsilon_{k_1} + \epsilon_{k_2} = \epsilon_{k_3} + \epsilon_{k_4}$  and account for them with a self-consistent SYK-like analysis.

This is precisely the effective Hamiltonian method, when low energy states are separated from high energy states by a gap; we are assuming it can also apply in a gapless system.

# Resonant SYK model

$$H = \frac{1}{(2N)^{3/2}} \sum_{k_a} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta}(k_a) c_{k_1\alpha}^\dagger c_{k_2\beta}^\dagger c_{k_3\gamma} c_{k_4\delta} \\ + \sum_{k\alpha} \epsilon_k c_{k\alpha}^\dagger c_{k\alpha}$$

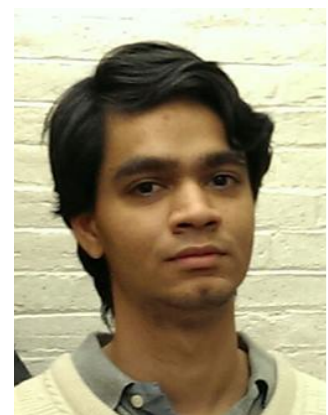
$U_{\alpha\beta;\gamma\delta}(k_a)$  is a random function of  $\alpha\beta\gamma\delta$  (as before)

The random  $k_i$  dependence of  $U$  allows only interactions resonant in the bare quasiparticle energies

with  $\epsilon_{k_1} + \epsilon_{k_2} = \epsilon_{k_3} + \epsilon_{k_4}$ .

$$\overline{U(k_1, k_2, k_3, k_4) U^*(k_5, k_6, k_7, k_8)} = \\ U^2 \left[ \delta(k_1 + k_2 - k_3 - k_4 - k_5 - k_6 + k_7 + k_8) \right] \\ \times \left[ \delta(\epsilon_{k_1} + \epsilon_{k_2} - \epsilon_{k_3} - \epsilon_{k_4}) + \delta(\epsilon_{k_5} + \epsilon_{k_6} - \epsilon_{k_7} - \epsilon_{k_8}) \right]$$

This implies off-site interactions with correlations which decay with a power-law in space.

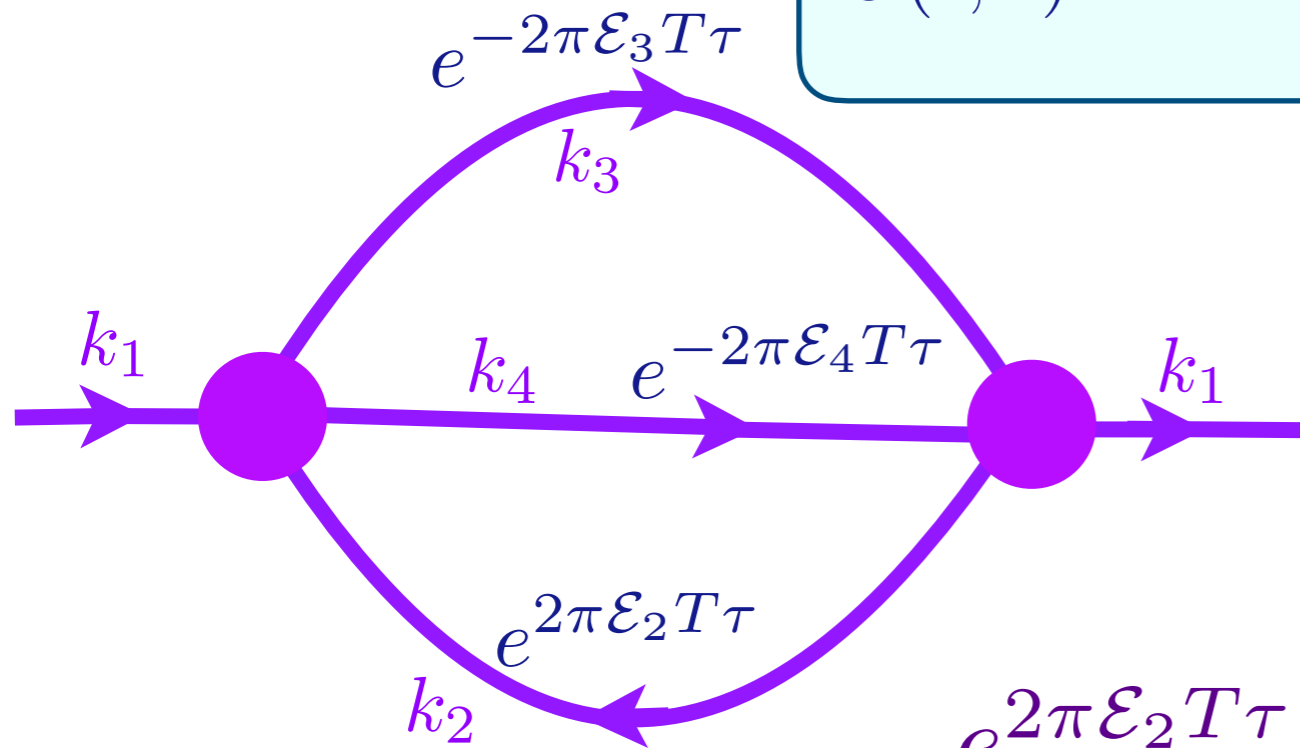


Aavishkar Patel

# Resonant SYK model

Conformal Green's function at  $T > 0$  must have the form

$$G(\epsilon, \tau) \sim e^{-2\pi\mathcal{E}T\tau} \left( \frac{T}{\sin(\pi T\tau)} \right)^{1/2}, \quad 0 < \tau < 1/T.$$



$$e^{2\pi\mathcal{E}_2 T\tau} e^{-2\pi\mathcal{E}_3 T\tau} e^{-2\pi\mathcal{E}_4 T\tau} = e^{-2\pi\mathcal{E}_1 T\tau}$$

if

$$\mathcal{E}_a = \mathbb{C}\epsilon_a/U$$

and

$$\epsilon_1 + \epsilon_2 = \epsilon_3 + \epsilon_4$$

SYK behavior in a Planckian metal as  $T \rightarrow 0$  with a remnant Fermi surface:  
 $G(k, \omega) = G_{\text{SYK}}(\epsilon_k, \hbar\omega/(k_B T))$ ,  
 with  $\mathcal{E}_k = \mathbb{C}\epsilon_k/U$

# Incoherent metal

For long times  $\tau > 0$

$$\left\langle c_k(\tau) c_k^\dagger(0) \right\rangle = e^{\pi\mathcal{E}} \frac{A}{\sqrt{\tau}}$$
$$\left\langle c_k^\dagger(\tau) c_k(0) \right\rangle = e^{-\pi\mathcal{E}} \frac{A}{\sqrt{\tau}}$$

The parameter  $\mathcal{E}$  is independent of  $k$ ,  
and determined by the total density

# Planckian metal with remnant Fermi surface

For long times  $\tau > 0$

$$\left\langle c_k(\tau) c_k^\dagger(0) \right\rangle = e^{\pi \mathbb{C} \epsilon_k / U} \frac{A}{\sqrt{\tau}}$$

$$\left\langle c_k^\dagger(\tau) c_k(0) \right\rangle = e^{-\pi \mathbb{C} \epsilon_k / U} \frac{A}{\sqrt{\tau}}$$

The particle-hole asymmetry changes as  
we cross the Fermi surface



Aavishkar Patel

# The complex SYK model

$$\mathcal{E} = \mathbb{C} \frac{\epsilon}{U}$$

$$G_{\text{SYK}}^R(\epsilon, \hbar\omega/(k_B T)) =$$

$$\frac{-iC e^{-i\theta} \Gamma\left(\frac{1}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}{(2\pi T)^{1/2} \Gamma\left(\frac{3}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}$$

$$e^{2\pi\mathcal{E}} = \frac{\sin(\pi/4 + \theta)}{\sin(\pi/4 - \theta)}$$

$$C = \left(\frac{\pi}{U^2 \cos(2\theta)}\right)^{1/4}$$

$$-\text{Im}G^R(\omega) \quad \mathcal{E} = 0$$

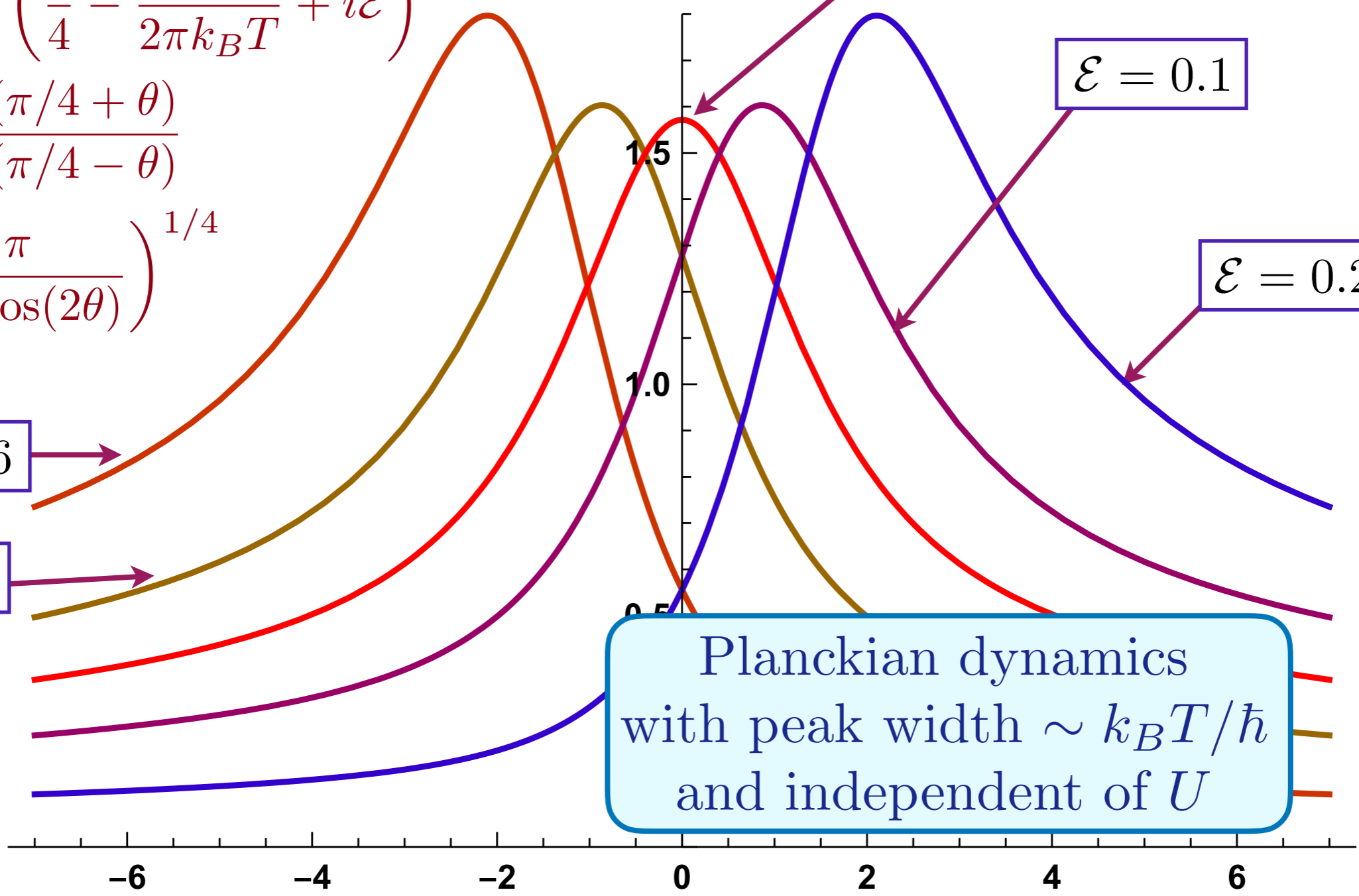
$$\mathcal{E} = 0.1$$

$$\mathcal{E} = 0.26$$

$$\mathcal{E} = -0.26$$

$$\mathcal{E} = -0.1$$

Planckian dynamics  
with peak width  $\sim k_B T/\hbar$   
and independent of  $U$



# The complex SYK model

$$\mathcal{E} = \mathbb{C} \frac{\epsilon}{U}$$

$$G_{\text{SYK}}^R(\epsilon, \hbar\omega/(k_B T)) =$$

$$\frac{-iC e^{-i\theta} \Gamma\left(\frac{1}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}{(2\pi T)^{1/2} \Gamma\left(\frac{3}{4} - \frac{i\hbar\omega}{2\pi k_B T} + i\mathcal{E}\right)}$$

$$e^{2\pi\mathcal{E}} = \frac{\sin(\pi/4 + \theta)}{\sin(\pi/4 - \theta)}$$

$$C = \left(\frac{\pi}{U^2 \cos(2\theta)}\right)^{1/4}$$

$$-\text{Im}G^R(\omega) \quad \mathcal{E} = 0$$

$$\mathcal{E} = 0.1$$

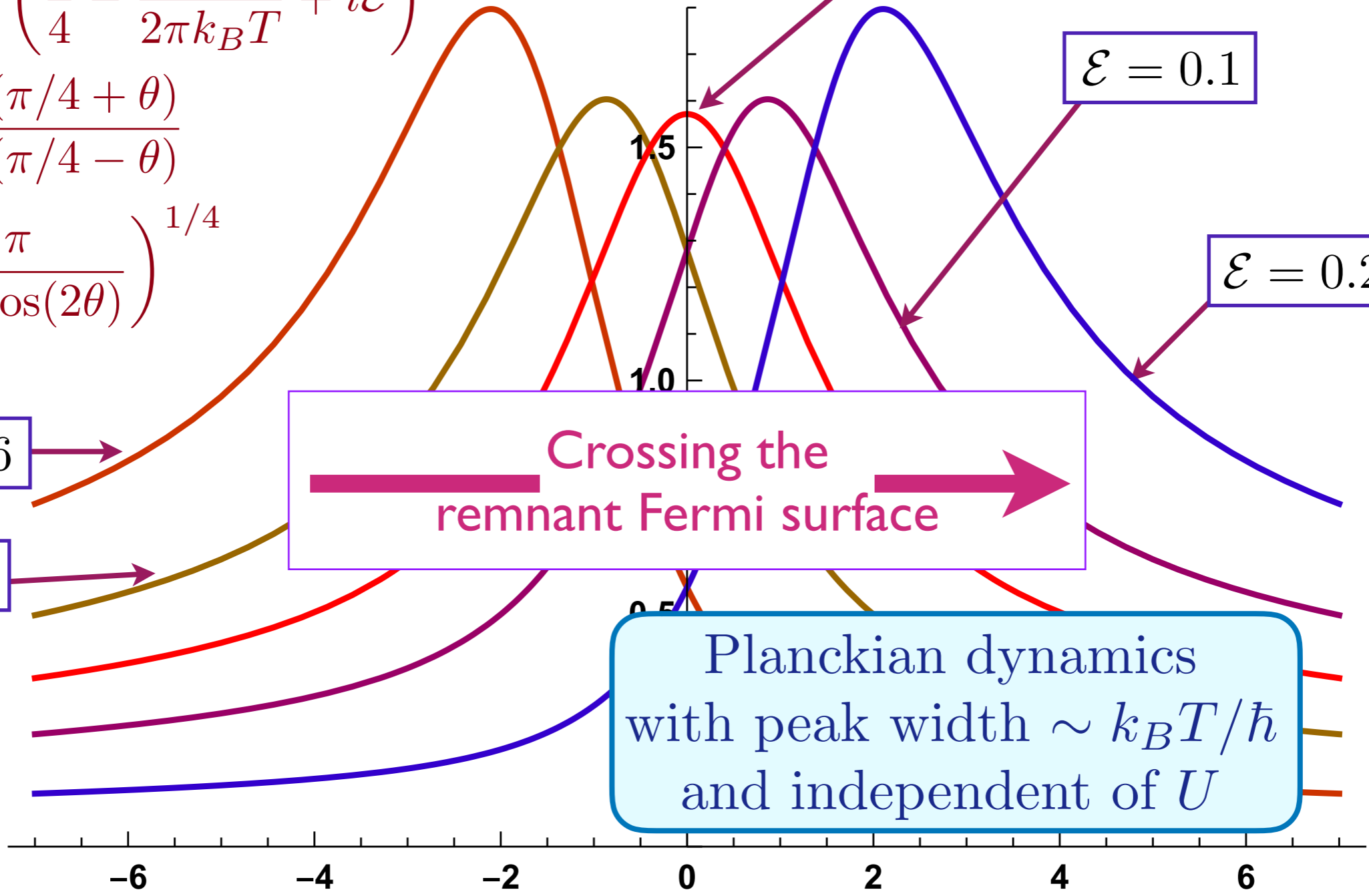
$$\mathcal{E} = 0.26$$

$$\mathcal{E} = -0.26$$

$$\mathcal{E} = -0.1$$

Crossing the remnant Fermi surface

Planckian dynamics with peak width  $\sim k_B T/\hbar$  and independent of  $U$



# Resonant SYK model

$U_{\alpha\beta;\gamma\delta}(k_a)$  is a random function of  $\alpha\beta\gamma\delta$  (as before)

The random  $k_i$  dependence of  $U$  allows only  
interactions resonant in the bare quasiparticle energies  
with  $\epsilon_{k_1} + \epsilon_{k_2} = \epsilon_{k_3} + \epsilon_{k_4}$ .

Resistivity of a [Planckian metal](#) as  $T \rightarrow 0$

From the Kubo formula, in the large  $N$  limit

$$\sigma = \frac{Ne^2 m^* v_F^2}{2T} \int_{-\infty}^{\infty} \frac{d\epsilon}{2\pi} \int_{-\infty}^{\infty} \frac{d\omega}{4\pi} \left[ \text{Im} G_{\text{SYK}}^R \left( \epsilon, \frac{\omega}{T} \right) \right]^2 \text{sech}^2 \left( \frac{\omega}{2T} \right)$$

# Resonant SYK model

$U_{\alpha\beta;\gamma\delta}(k_a)$  is a random function of  $\alpha\beta\gamma\delta$  (as before)

The random  $k_i$  dependence of  $U$  allows only interactions resonant in the bare quasiparticle energies  
with  $\epsilon_{k_1} + \epsilon_{k_2} = \epsilon_{k_3} + \epsilon_{k_4}$ .

Resistivity of a Planckian metal as  $T \rightarrow 0$

From the Kubo formula, in the large  $N$  limit

$$\sigma = \frac{Ne^2 m^* v_F^2}{2T} \int_{-\infty}^{\infty} \frac{d\epsilon}{2\pi} \int_{-\infty}^{\infty} \frac{d\omega}{4\pi} \left[ \text{Im} G_{\text{SYK}}^R \left( \epsilon, \frac{\omega}{T} \right) \right]^2 \text{sech}^2 \left( \frac{\omega}{2T} \right)$$

$$\rho = \frac{m^*}{ne^2} 2.71\mathbb{C} \frac{k_B T}{\hbar}, \quad \text{using } \mathcal{E} = \mathbb{C}\epsilon/U,$$

where

$$m^* = \frac{d V_{FS}}{\oint_{FS} |\mathbf{v}_F|},$$

where  $d$  is spatial dimensionality and  $V_{FS}$  is the volume enclosed by the Fermi surface. For a circular Fermi surface, this is the usual  $m^*$ .

# Resonant SYK model

Resistivity of a Planckian metal as  $T \rightarrow 0$

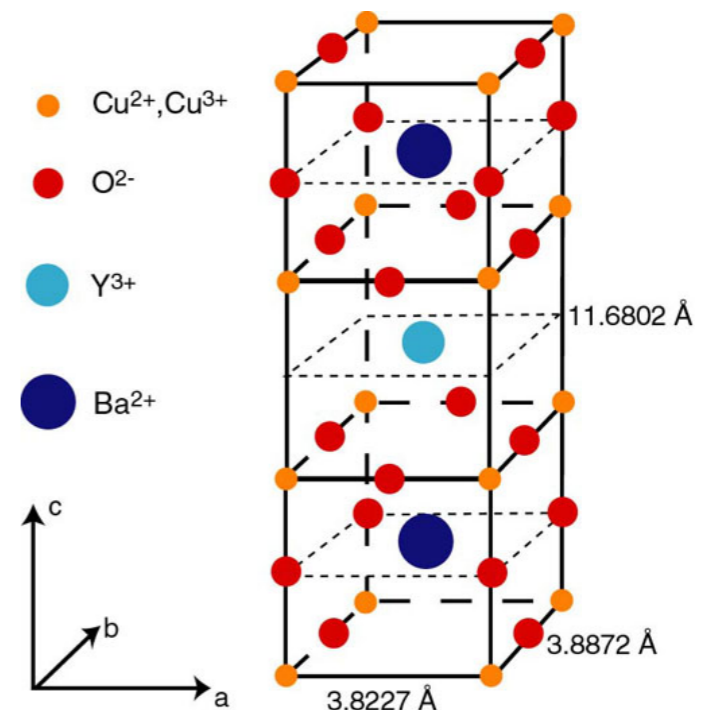
$$\rho = \frac{m^*}{ne^2} 2.71\mathbb{C} \frac{k_B T}{\hbar}$$

Note that all explicit dependence on  $U$  has cancelled out!

The number  $\mathbb{C}$  is defined by  $\mathcal{E}_k = \mathbb{C} \epsilon_k / U$  as  $|\epsilon_k| \rightarrow 0$ . This is determined by UV physics, and is very weakly dependent upon the ratio of the energy width of the interactions,  $W_U$ , to  $U$ .



Aavishkar Patel



# Resonant SYK model

Take the independent momentum shell limit,  $W_U/U \rightarrow 0$ ,

$$\overline{U(k_1, k_2, k_3, k_4)U^*(k_5, k_6, k_7, k_8)} =$$

$$U^2 \left[ \delta(k_1 + k_2 - k_3 - k_4 - k_5 - k_6 + k_7 + k_8) \right]$$

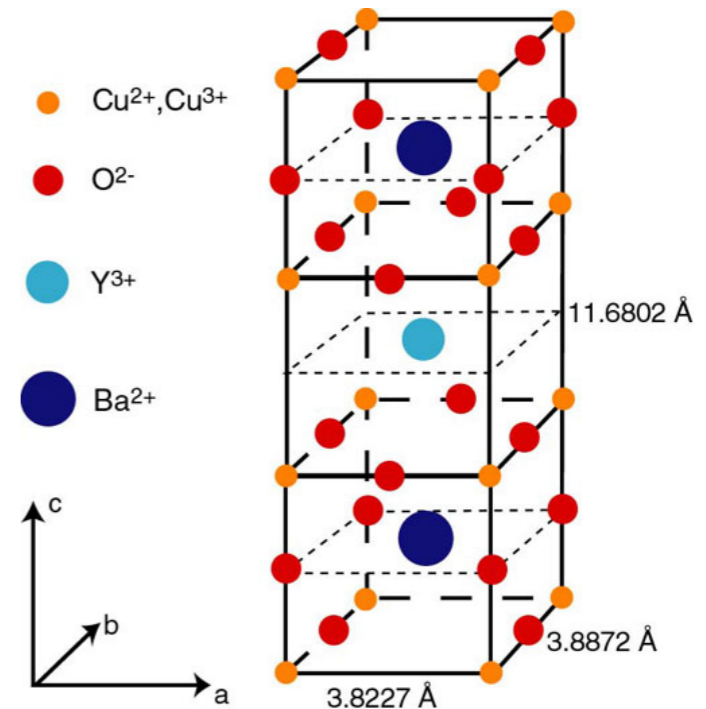
$$\times \left[ \delta(\epsilon_{k_1} - \epsilon_{k_2})\delta(\epsilon_{k_2} - \epsilon_{k_3})\delta(\epsilon_{k_3} - \epsilon_{k_4}) + \delta(\epsilon_{k_5} - \epsilon_{k_6})\delta(\epsilon_{k_6} - \epsilon_{k_7})\delta(\epsilon_{k_7} - \epsilon_{k_8}) \right]$$

$\mathbb{C} = 0.41$  as in a single SYK model,  
and we obtain a Planckian metal with

$$\rho = \frac{m^*}{ne^2} 1.11 \frac{k_B T}{\hbar}$$



Aavishkar Patel



# Charged black holes and complex SYK

- A Schwarzian theory of a time reparameterization mode, with  $SL(2, \mathbb{R})$  symmetry, describes the quantum dynamics of
  - the SYK models
  - black holes with near-extremal  $AdS_2$  horizons
- The  $T \rightarrow 0$  black hole entropy is *not* realized by a ground state degeneracy. This is in contrast to all supersymmetric and string-theoretic computations of black hole entropy.
- Resonant SYK models realize Planckian metals with remnant large Fermi surface at  $\epsilon_k = 0$ , and an effective mass  $m^*$  defined by the dispersion of  $\epsilon_k$ , and a resistivity  $\rho \sim (m^*/(ne^2))k_B T/\hbar$ .