

Universal low temperature theory of SYK models of strange metals and charged black holes

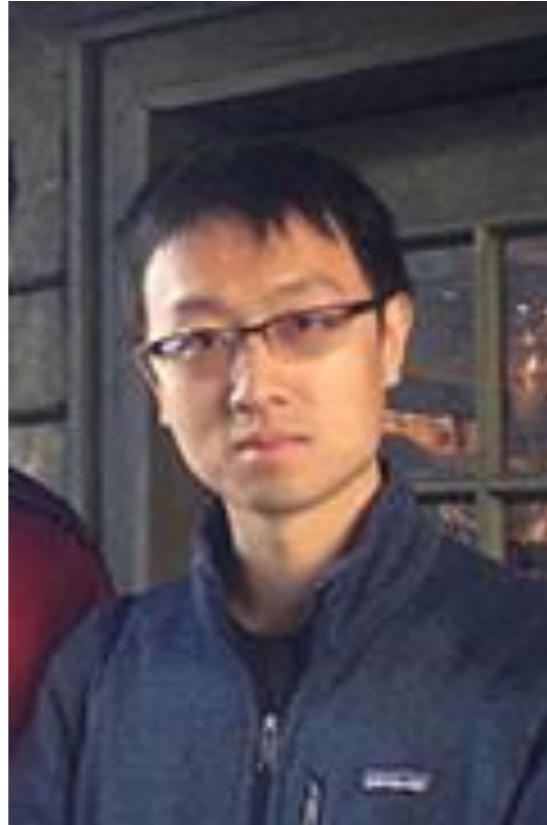
Subir Sachdev

New Directions in Theoretical Physics 3,
Higgs Centre for Theoretical Physics
University of Edinburgh
January 11, 2019

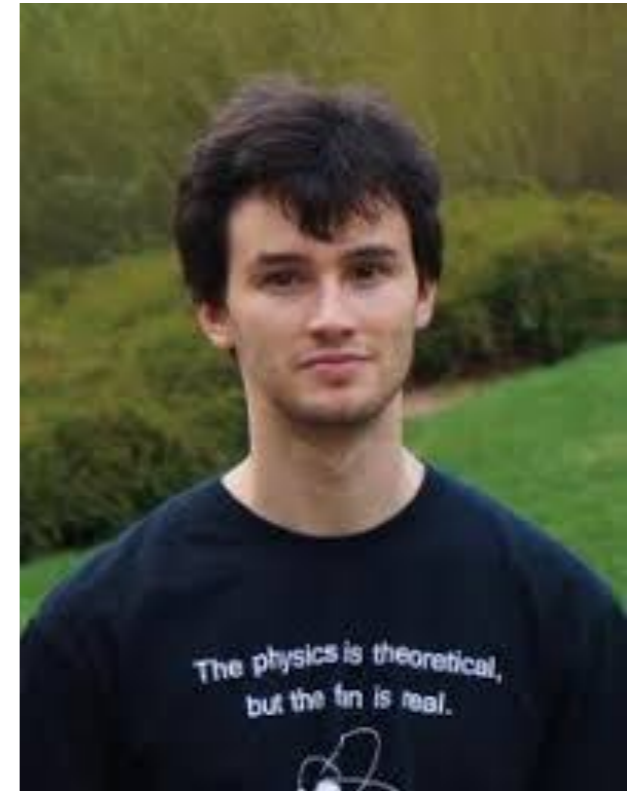




Wenbo Fu



Yingfei Gu



Grigory Tarnopolsky



Main result

SYK model of fermions with random interactions of mean-square-value J , with total fermion number Q ,
at temperatures $T \ll J$

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SYK model of fermions with random interactions of mean-square-value J , with total fermion number Q ,
at temperatures $T \ll J$

and

Charged black holes in 3+1 dimensions of radius R_h ,
with total charge Q , at temperatures $T \ll 1/R_h$

are described by a common low energy quantum
theory in 0 + 1 dimensions

Main result

The common low T path integral is $\mathcal{Z} = \int \mathcal{D}f \mathcal{D}\phi e^{-I}$. This can be exactly evaluated, and the action is

$$I = -s_0 + \int_0^{1/T} d\tau \left\{ \frac{K}{2} \left(\frac{\partial\phi}{\partial\tau} + i(2\pi\mathcal{E}T) \frac{\partial f}{\partial\tau} \right)^2 - \frac{\gamma}{4\pi^2} \text{Sch}[\tan(\pi T f(\tau)), \tau] \right\},$$

where $f(\tau)$ is a monotonic reparameterization of the temporal circle with

$$f(\tau + 1/T) = f(\tau) + 1/T,$$

ϕ is a phase conjugate to the charge density with

$$\phi(\tau + 1/T) = \phi(\tau) + 2\pi n, \quad n \text{ integer},$$

$\text{Sch}[g[\tau], \tau]$ is the Schwarzian derivative of $g(\tau)$.

The couplings are related to the entropy $S(T, Q)$ and the chemical potential μ via

$$S(T \rightarrow 0, Q) = s_0 + \gamma T, \quad K = \left(\frac{dQ}{d\mu} \right)_{T \rightarrow 0}, \quad 2\pi\mathcal{E} = \frac{ds_0}{dQ}$$

Main result

- Not the AdS/CFT correspondence, which involves only neutral black holes at $T > 0$.
- Unlike the AdS/CFT correspondence, *both* sides of the duality are solvable. This has enabled numerous recent studies of black holes quantum information.

Main result

S. Sachdev, Phys. Rev. Lett. **105**, 151602 (2010)

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J. Maldacena and D. Stanford, Phys. Rev. D **94**, 106002 (2016)

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J. Engelsoy, T.G. Mertens, and H. Verlinde, JHEP 1607 (2016) 139

R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev,
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A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746 P. Nayak, A. Shukla,

R.M. Soni, S.P. Trivedi, and V. Vishal, arXiv:1802.09547

P. Chaturvedi, Yingfei Gu, Wei Song, Boyang Yu, arXiv:1808.08062

U. Moitra, S. P. Trivedi, and V. Vishal, arXiv:1808.08239

1. Normal and strange metals

(a) with quasiparticles (normal):

random matrix model

(b) without quasiparticles (strange):

the complex SYK model

2. Einstein-Maxwell theory of charged black holes in AdS space

3. Fluctuations, and the Schwarzian

I. Normal and strange metals

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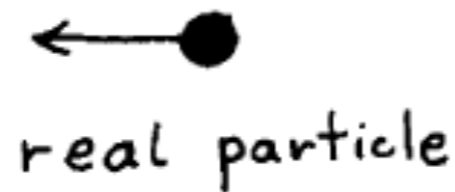
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Ordinary metals

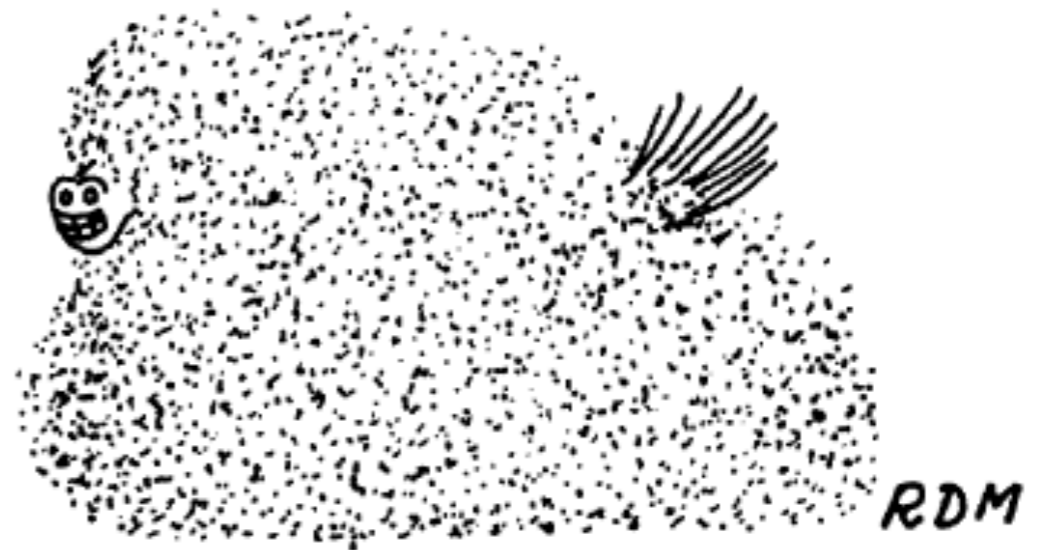


Ordinary metals are shiny, and they conduct heat and electricity efficiently. Each atom donates electrons which are delocalized throughout the entire crystal

Almost all many-electron systems are described by the quasiparticle concept: a quasiparticle is an “excited lump” in the many-electron state which responds just like an ordinary particle.



real horse



quasi horse

What are quasiparticles ?

- **Quasiparticles are additive excitations:**

The low-lying excitations of the many-body system can be identified as a set $\{n_\alpha\}$ of quasiparticles with energy ε_α

$$E = \sum_{\alpha} n_{\alpha} \varepsilon_{\alpha} + \sum_{\alpha, \beta} F_{\alpha\beta} n_{\alpha} n_{\beta} + \dots$$

In a lattice system of N sites, this parameterizes the energy of $\sim e^{\alpha N}$ states in terms of poly(N) numbers.

What are quasiparticles ?

- Quasiparticles eventually collide with each other. Such collisions eventually leads to thermal equilibration in a chaotic quantum state, but the equilibration takes a long time. In a Fermi liquid, this time diverges as

$$\tau_{\text{eq}} \sim \frac{\hbar E_F}{(k_B T)^2} \quad , \quad \text{as } T \rightarrow 0,$$

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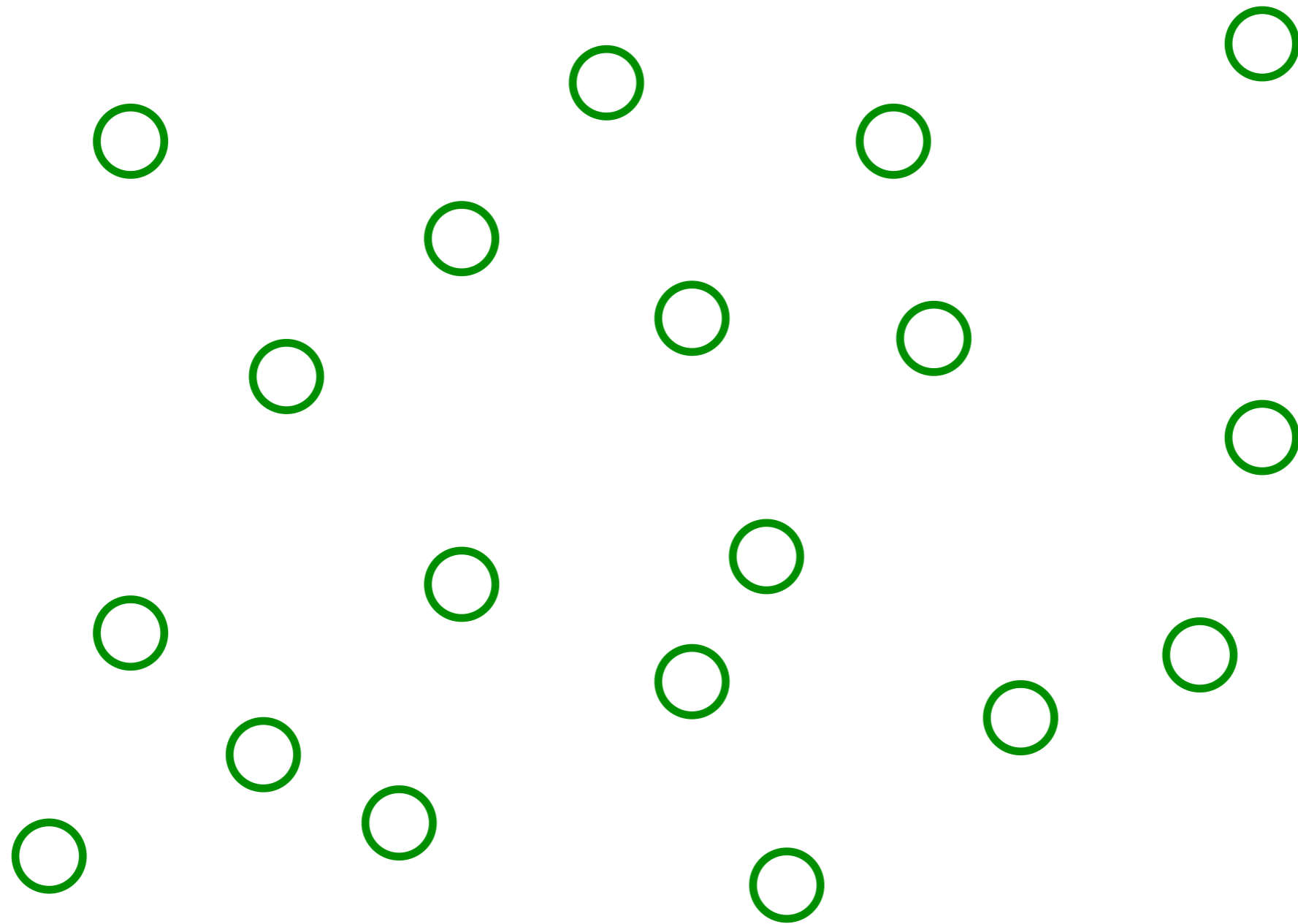
$$\tau_{\text{eq}} \sim \frac{\hbar E_F}{(k_B T)^2} \quad , \quad \text{as } T \rightarrow 0,$$

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- This time is much longer than the ‘Planckian time’ $\hbar/(k_B T)$, which we will find in systems without quasiparticle excitations.

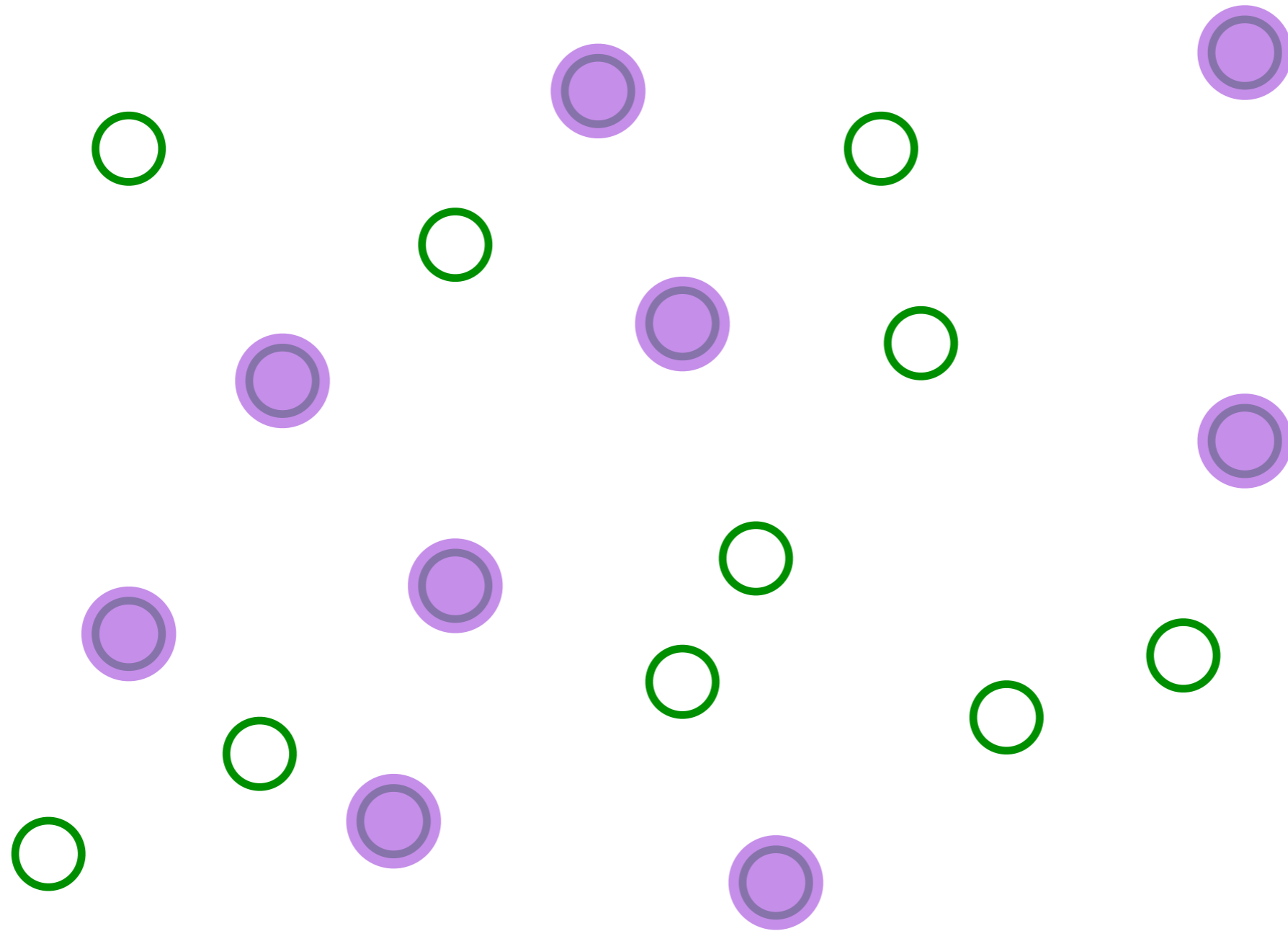
$$\tau_{\text{eq}} \gg \frac{\hbar}{k_B T} \quad , \quad \text{as } T \rightarrow 0.$$

A simple model of a metal with quasiparticles



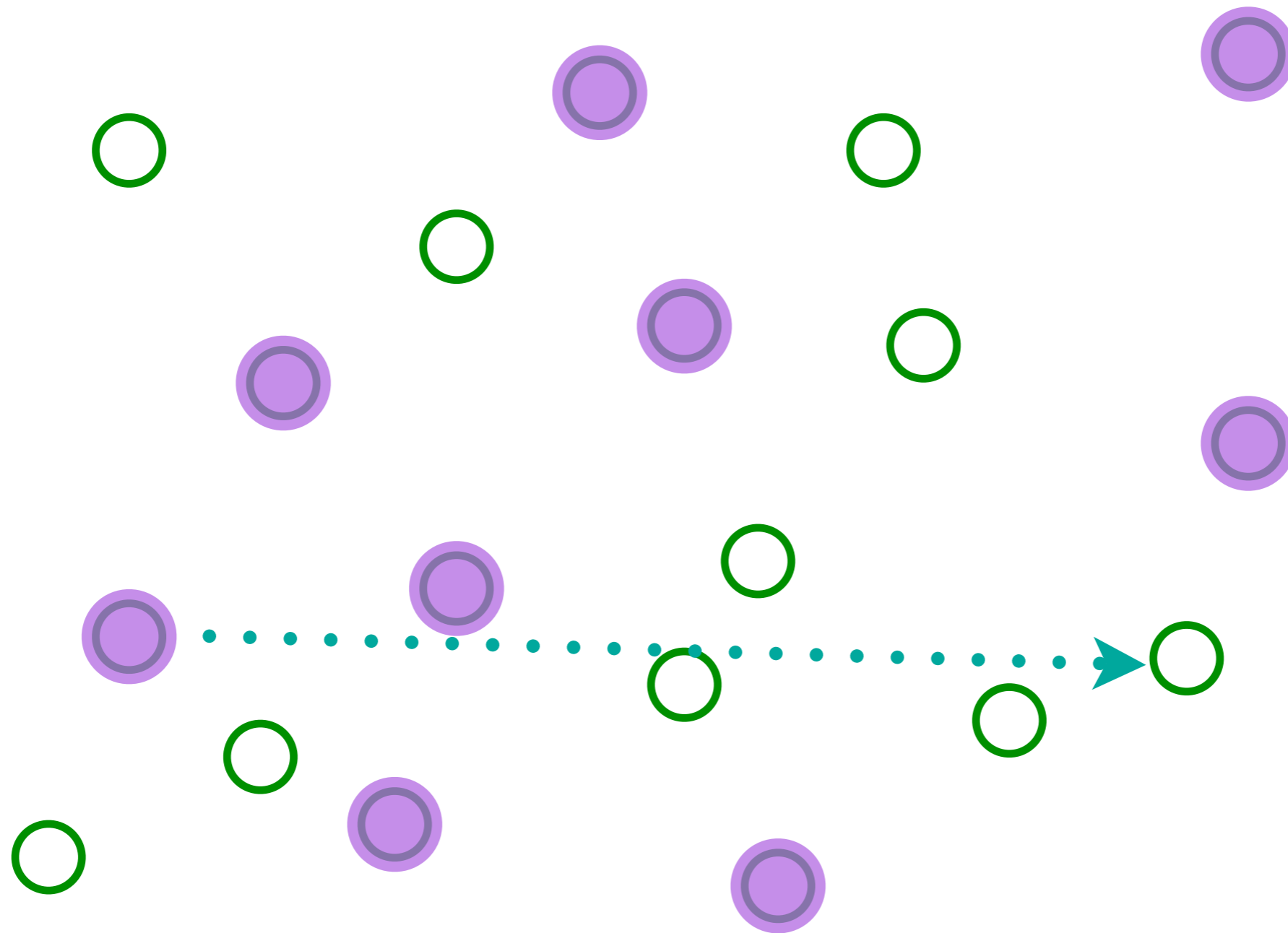
Pick a set of random positions

A simple model of a metal with quasiparticles



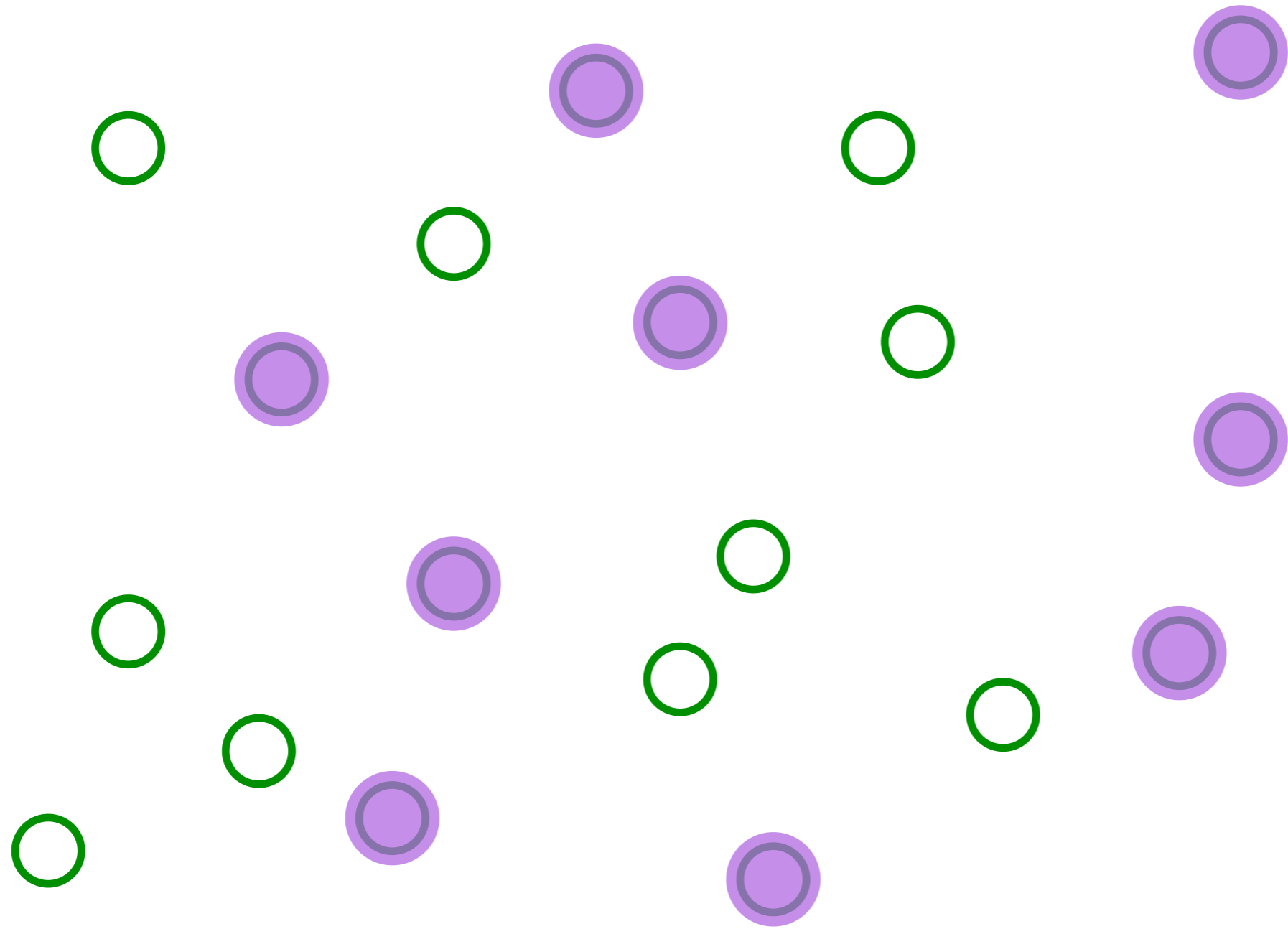
Place electrons randomly on some sites

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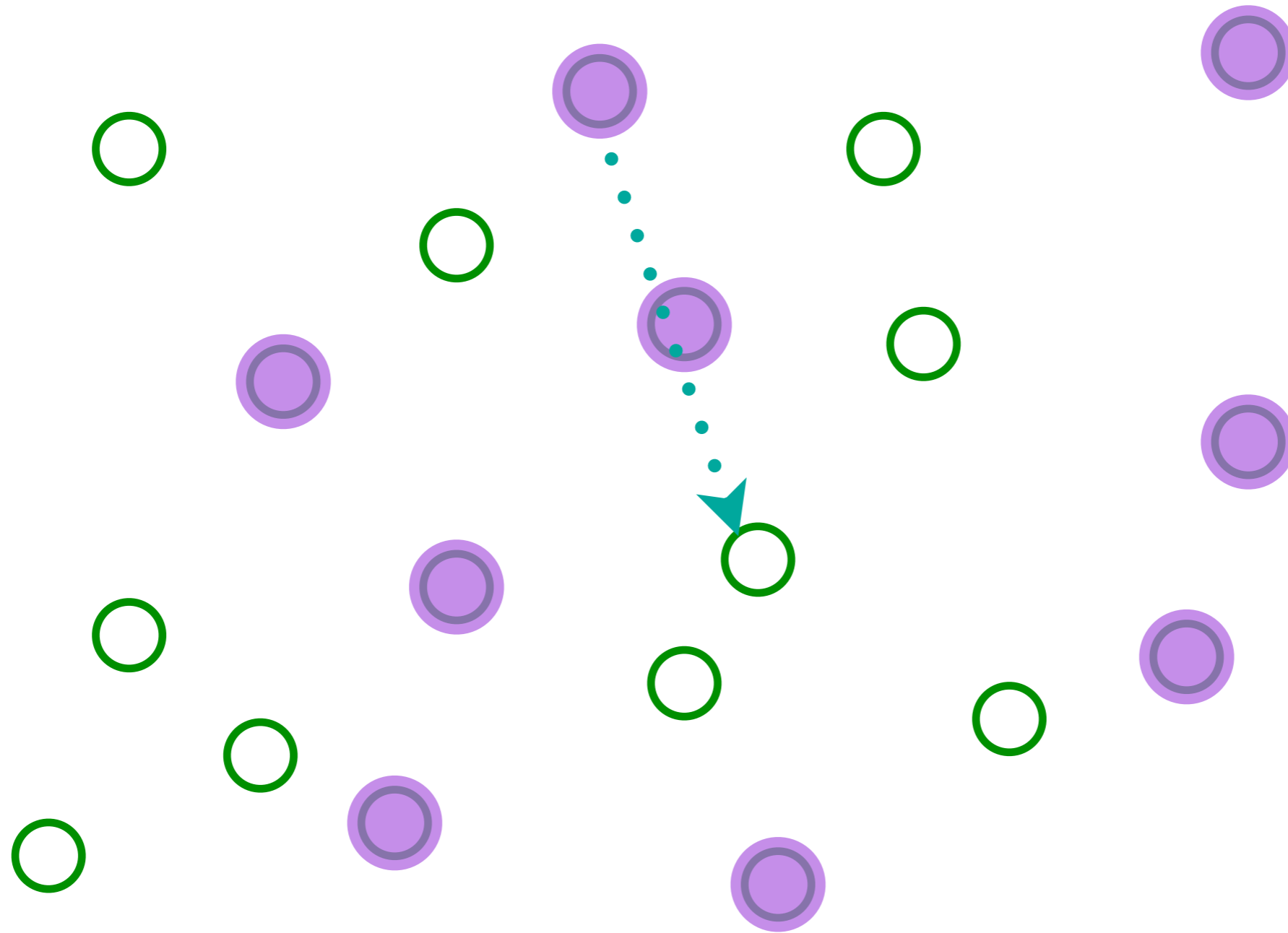
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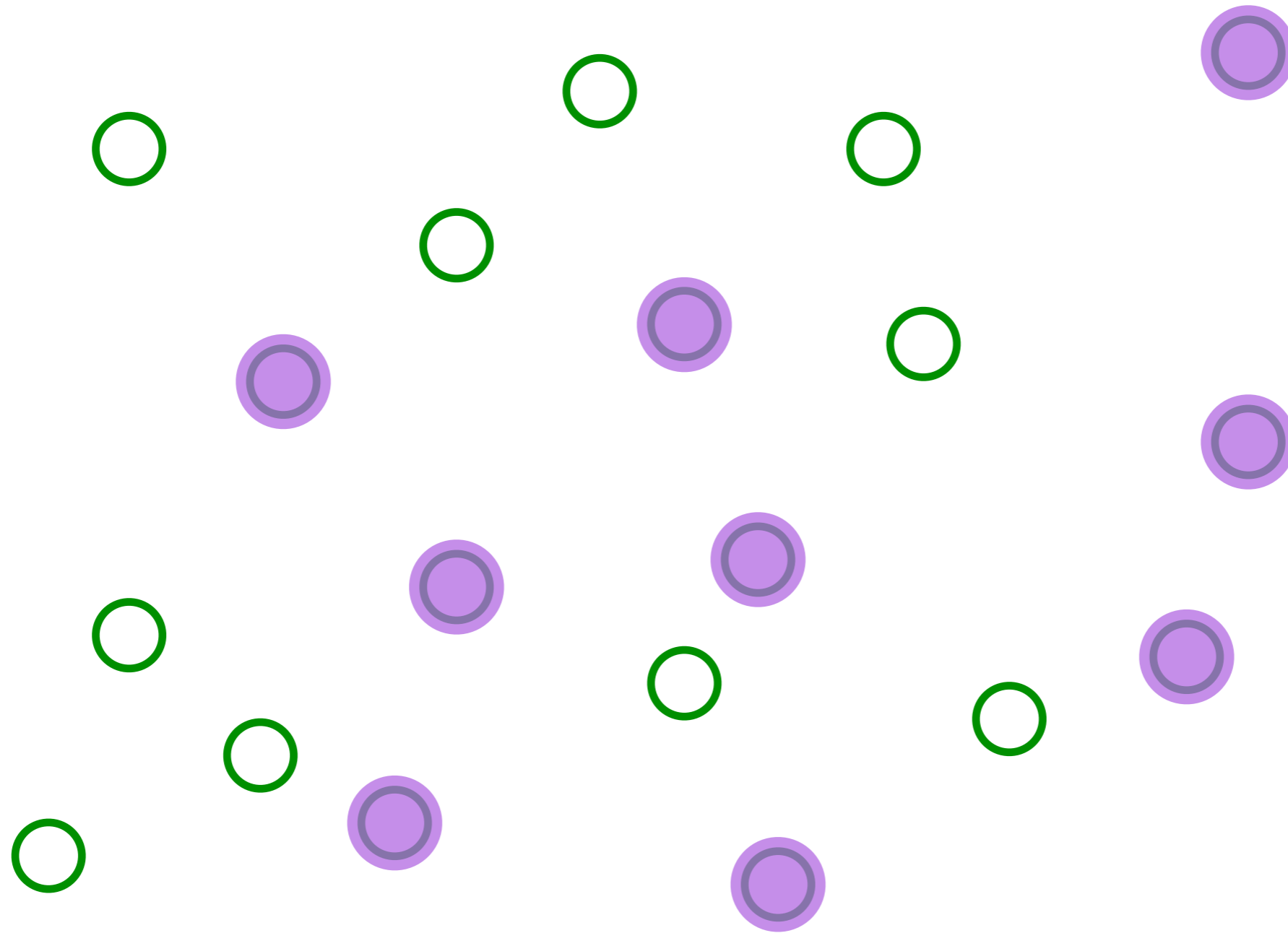
Electrons move one-by-one randomly

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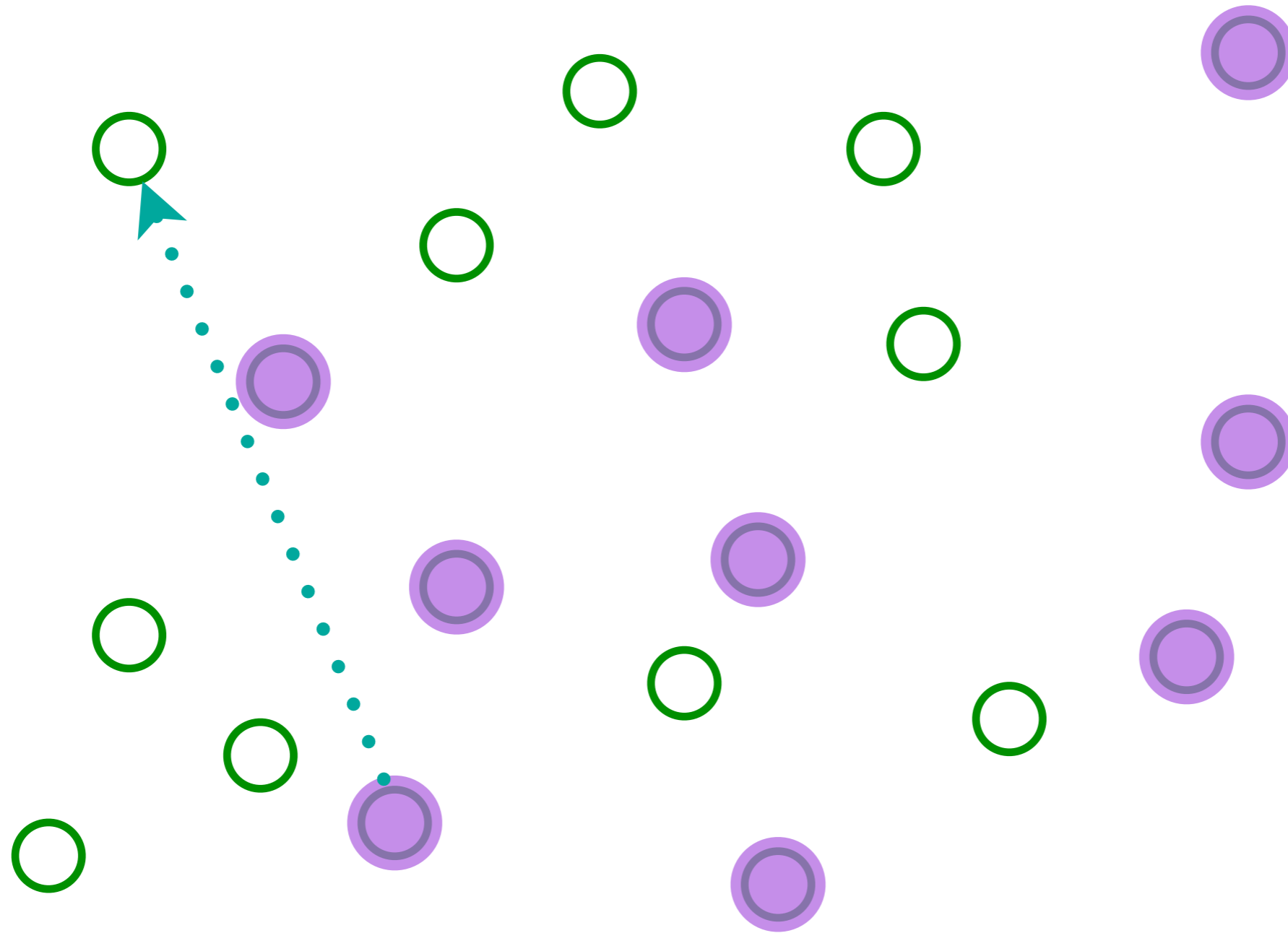
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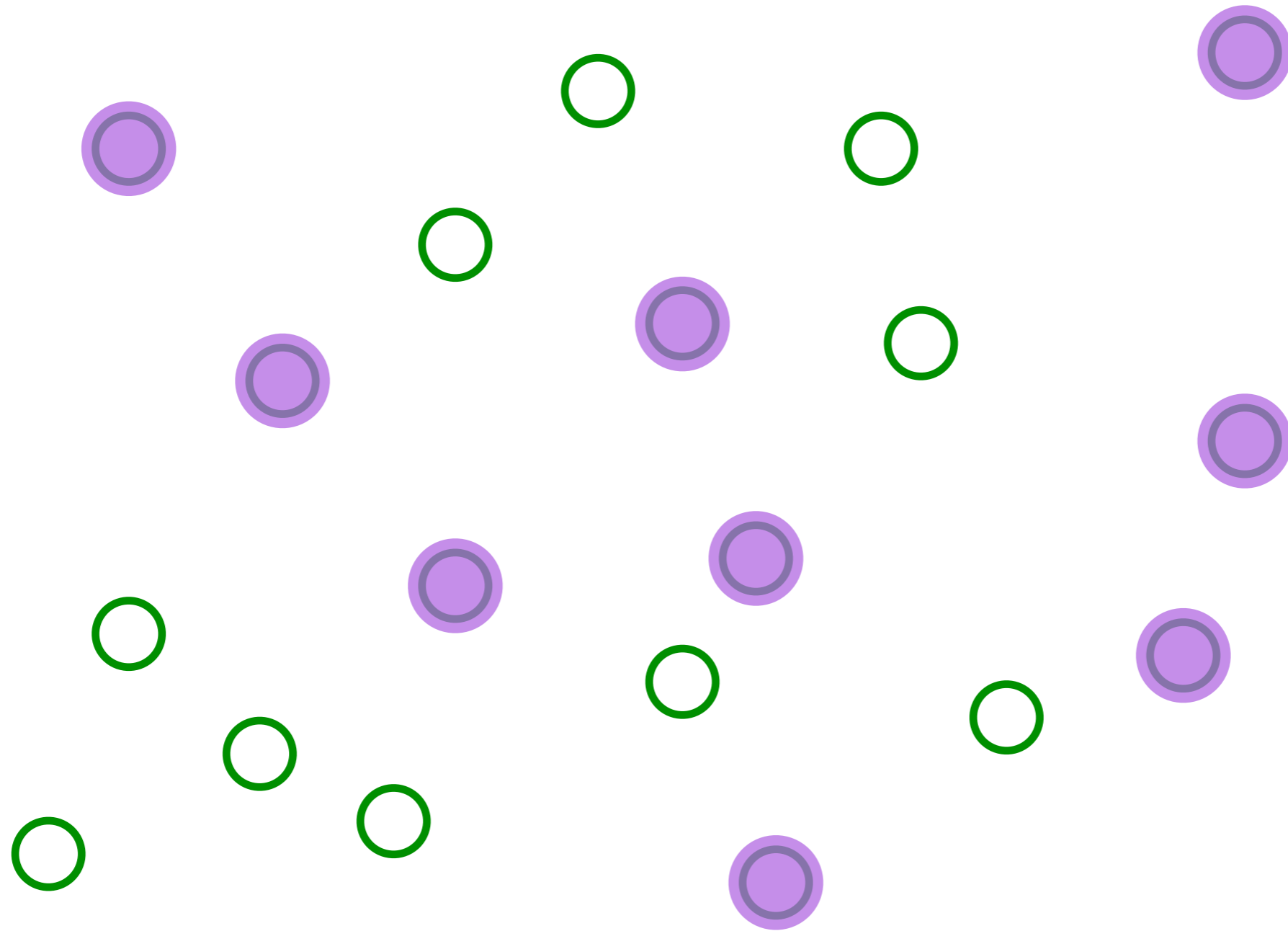
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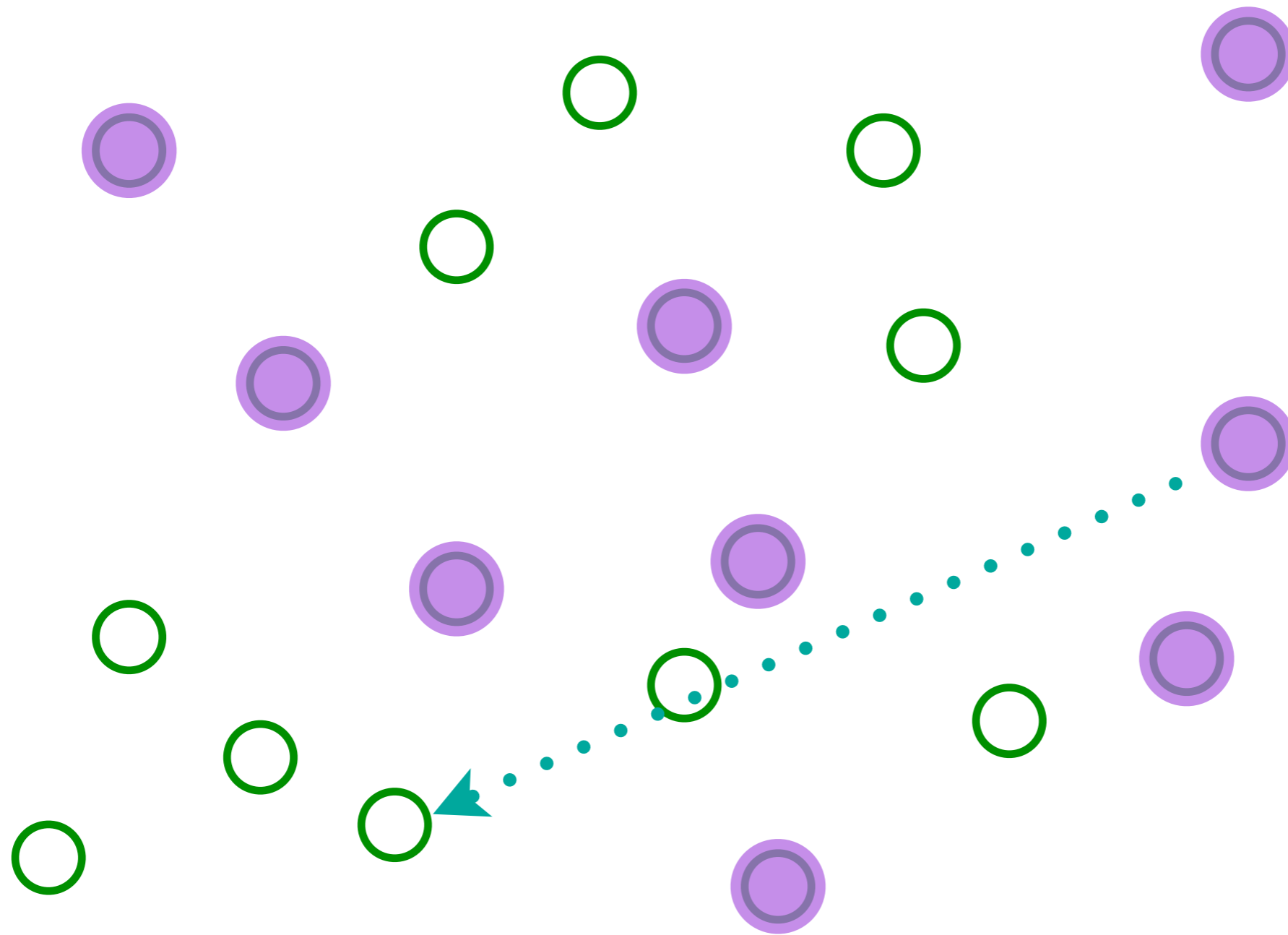
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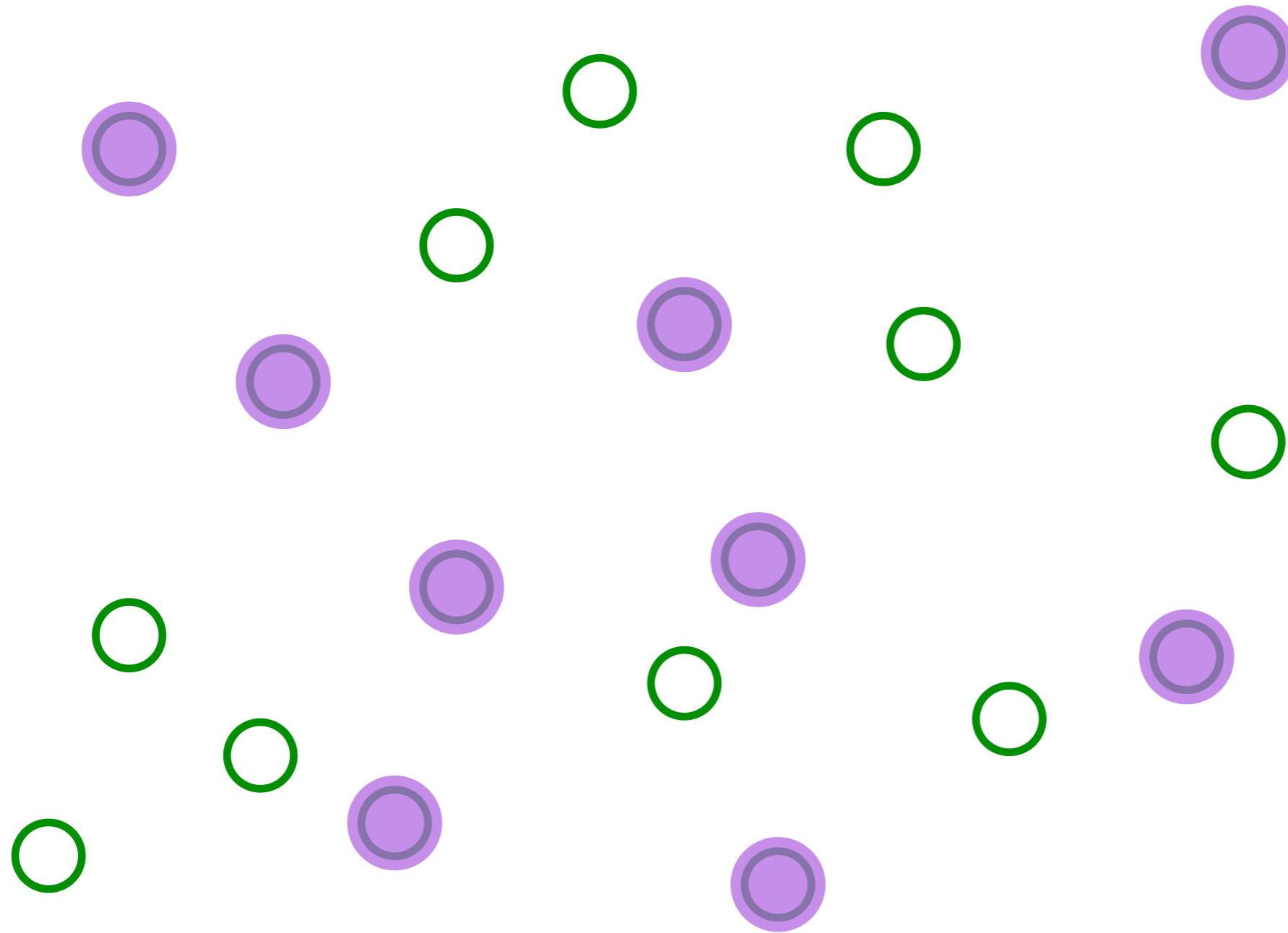
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A simple model of a metal with quasiparticles

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$\frac{1}{N} \sum_i c_i^\dagger c_i = Q$$

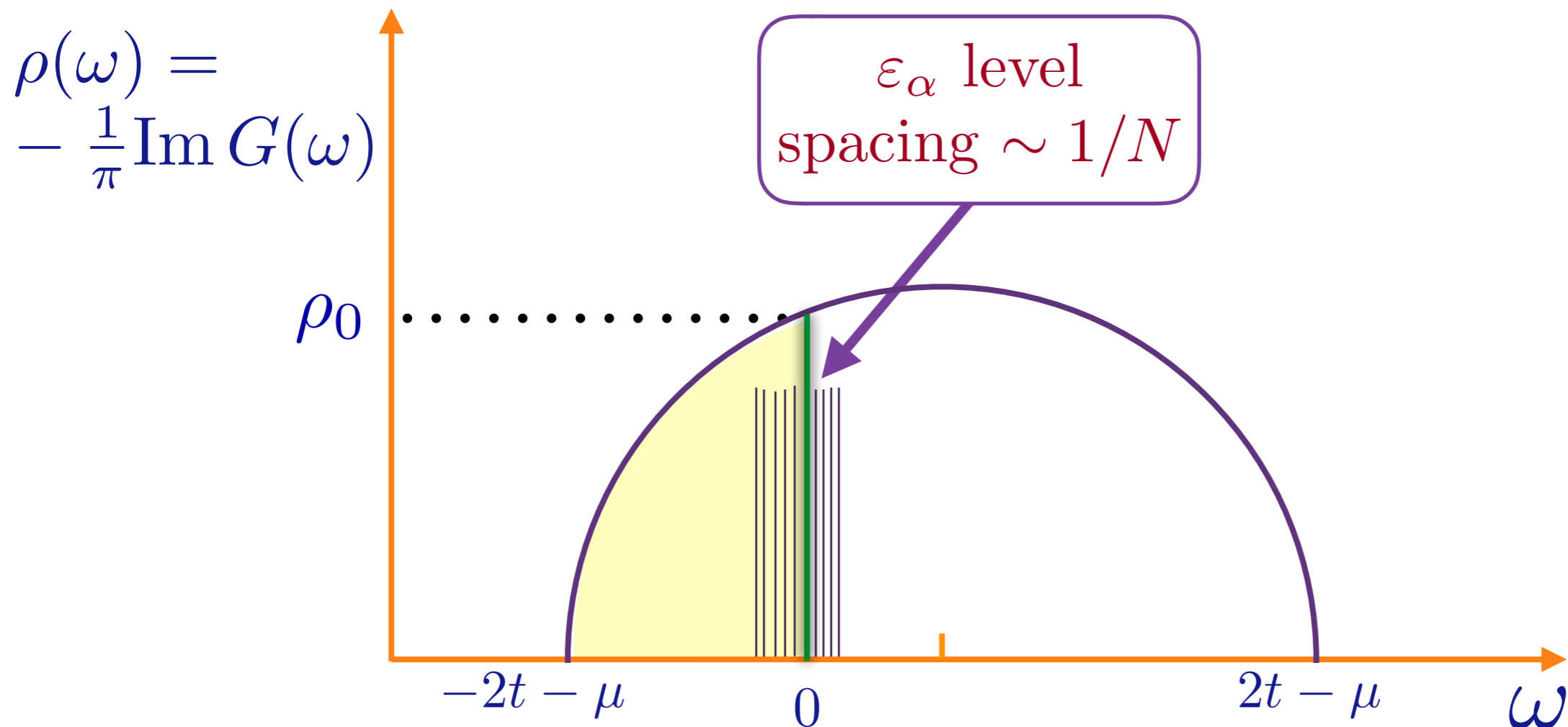
t_{ij} are independent random variables with $\overline{t_{ij}} = 0$ and $\overline{|t_{ij}|^2} = t^2$

**Fermions occupying the eigenstates of a
 $N \times N$ random matrix**

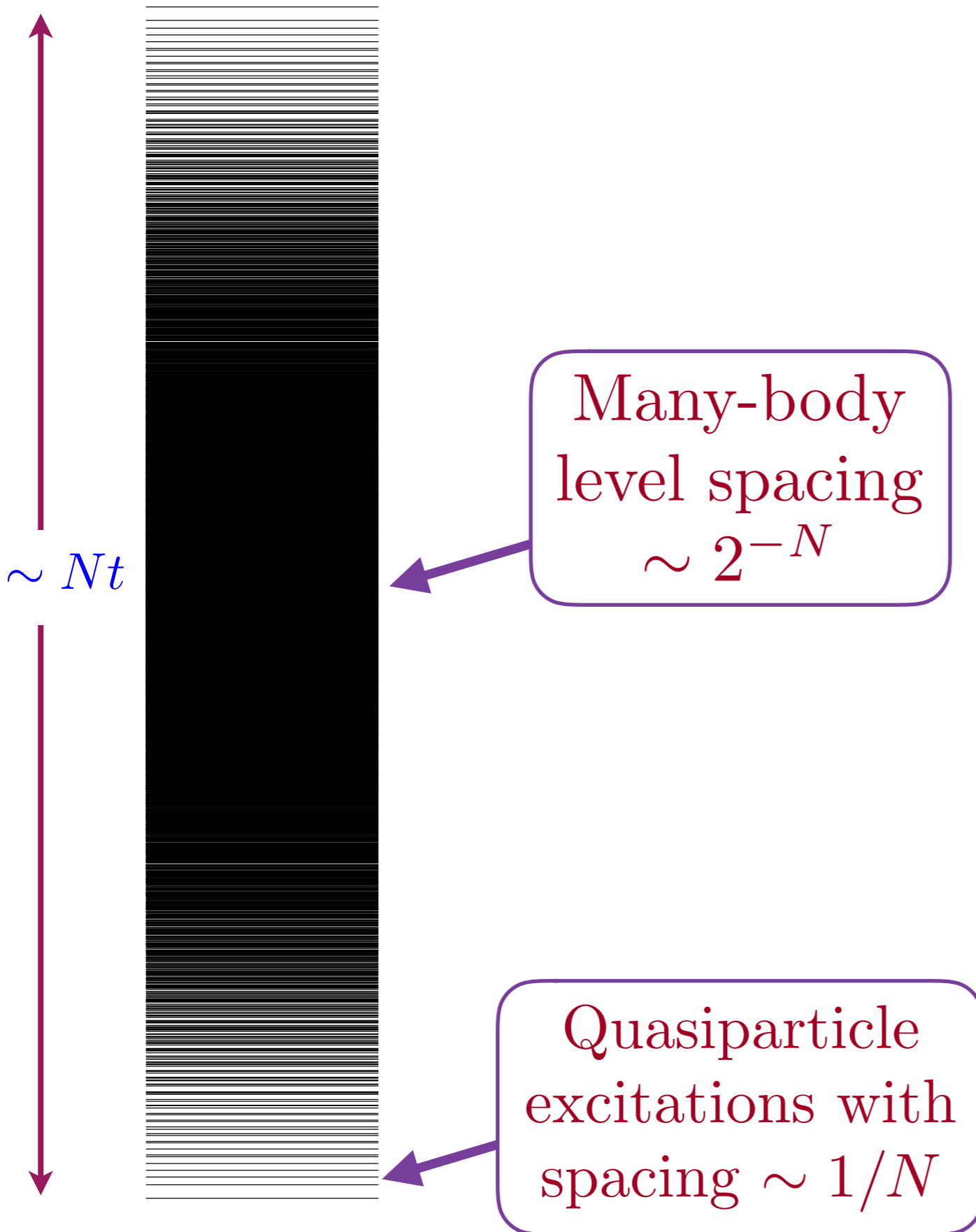
A simple model of a metal with quasiparticles

Let ε_α be the eigenvalues of the matrix t_{ij}/\sqrt{N} . The fermions will occupy the lowest NQ eigenvalues, upto the Fermi energy E_F . The single-particle density of states is

$$\rho(\omega) = (1/N) \sum_\alpha \delta(\omega - \varepsilon_\alpha), \text{ and } \rho_0 \equiv \rho(\omega = 0).$$



A simple model of a metal with quasiparticles



There are 2^N many body levels with energy

$$E = \sum_{\alpha=1}^N n_{\alpha} \varepsilon_{\alpha},$$

where $n_{\alpha} = 0, 1$. Shown are all values of E for a single cluster of size $N = 12$. The ε_{α} have a level spacing $\sim 1/N$.

A simple model of a metal with quasiparticles

Now add weak interactions

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i + \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N U_{ij;kl} c_i^\dagger c_j^\dagger c_k c_\ell$$

$U_{ij;kl}$ are independent random variables with $\overline{U_{ij;kl}} = 0$ and $|\overline{U_{ij;kl}}|^2 = U^2$. We compute the lifetime of a quasiparticle, τ_α , in an exact eigenstate $\psi_\alpha(i)$ of the free particle Hamiltonian with energy ε_α . By Fermi's Golden rule, for ε_α at the Fermi energy

$$\begin{aligned} \frac{1}{\tau_\alpha} &= \pi U^2 \rho_0^2 \int d\varepsilon_\beta d\varepsilon_\gamma d\varepsilon_\delta f(\varepsilon_\beta)(1 - f(\varepsilon_\gamma))(1 - f(\varepsilon_\delta)) \delta(\varepsilon_\alpha + \varepsilon_\beta - \varepsilon_\gamma - \varepsilon_\delta) \\ &= \frac{\pi^3 U^2 \rho_0^2}{4} T^2 \end{aligned}$$

where ρ_0 is the density of states at the Fermi energy, and $f(\varepsilon) = 1/(e^{\varepsilon/T} + 1)$ is the Fermi function.

Fermi liquid state: Two-body interactions lead to a scattering time of quasiparticle excitations from in (random) single-particle eigenstates which diverges as $\sim T^{-2}$ at the Fermi level.

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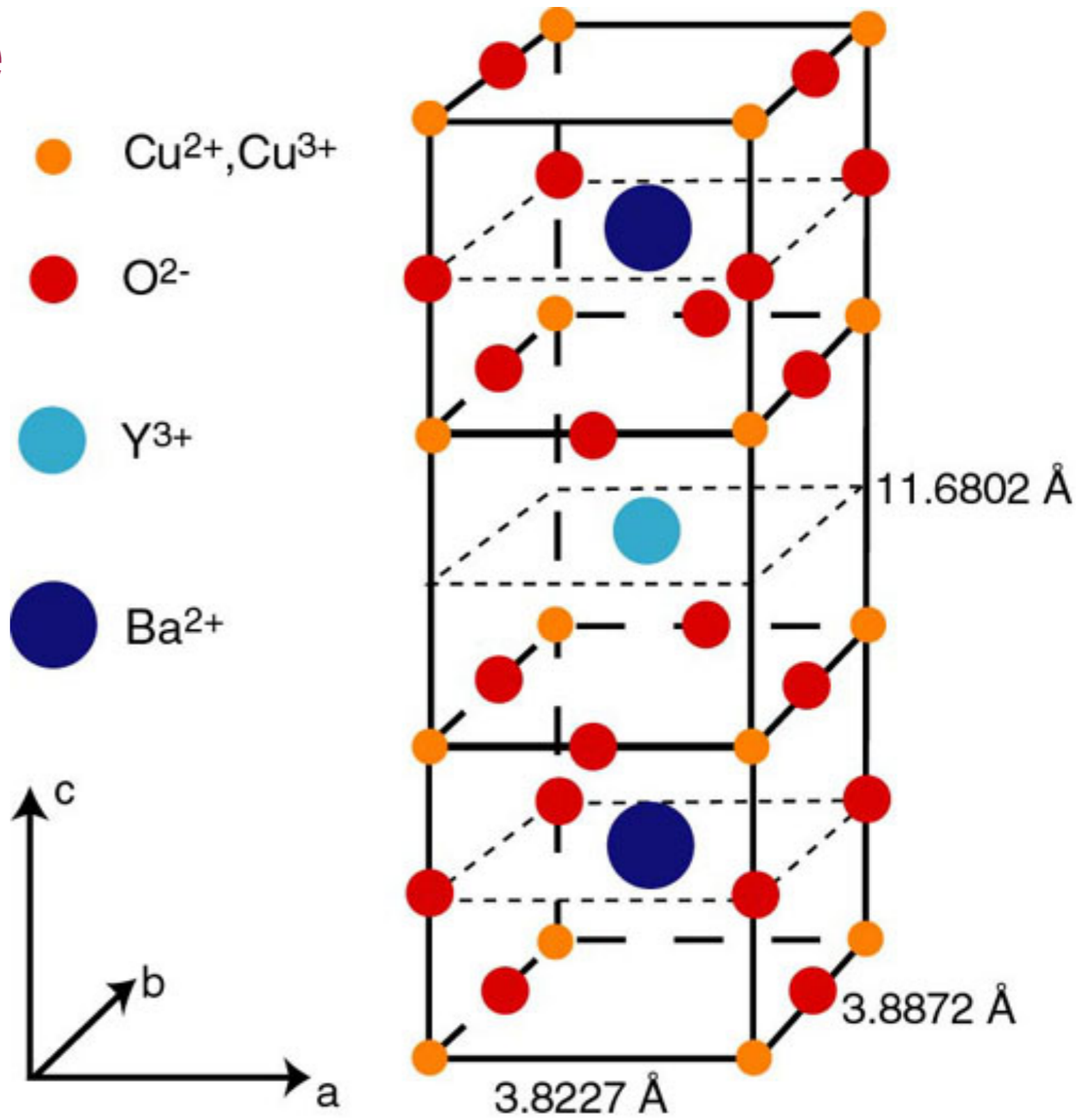
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High temperature superconductors





“Strange”,

“Bad”,



or “Incoherent”,

H. Takagi, B. Batlogg,
H. L. Kao, J. Kwo,
R. J. Cava,
J. J. Krajewski, and
W. F. Peck, Jr.,
Phys. Rev. Lett. **69**,
2975 (1992)

metal found ubiquitously at temperatures

$T > T_c$ (the superconducting critical temperature)

has a resistivity, ρ , which obeys

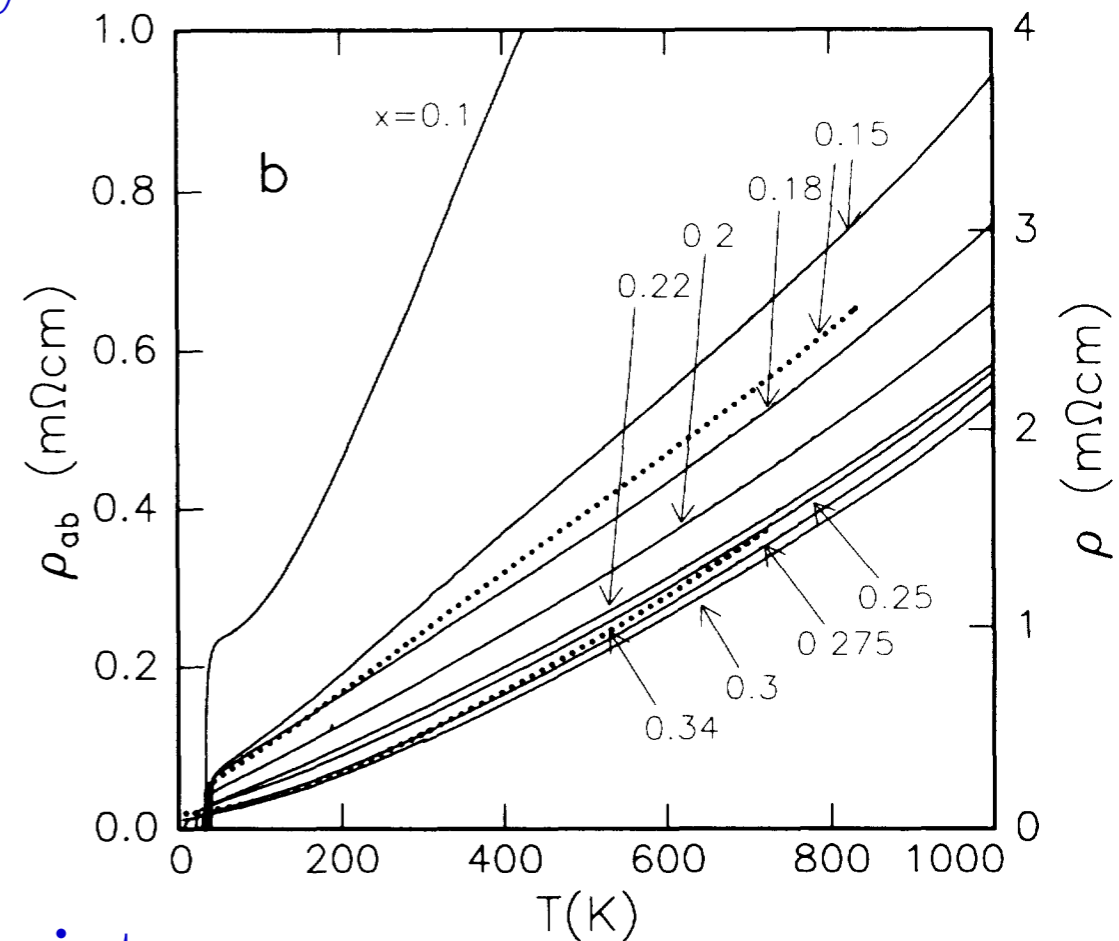
$$\rho \sim T,$$

and

in some cases $\rho \gg h/e^2$

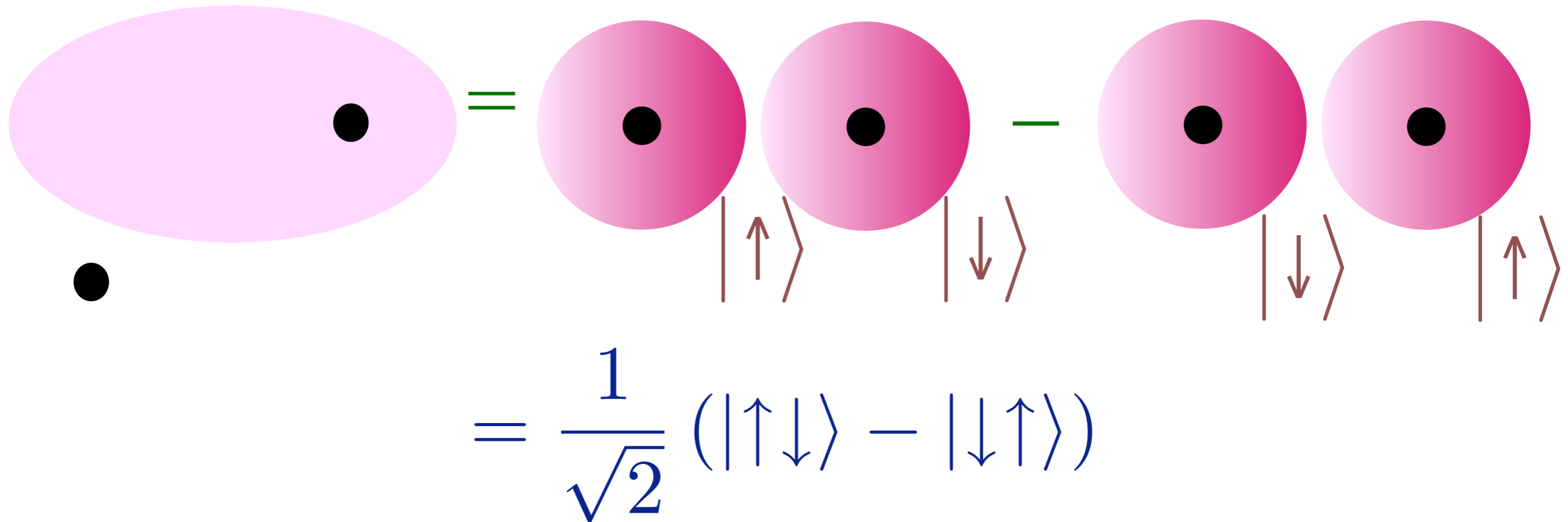
(in two dimensions),

where h/e^2 is the quantum unit of resistance.

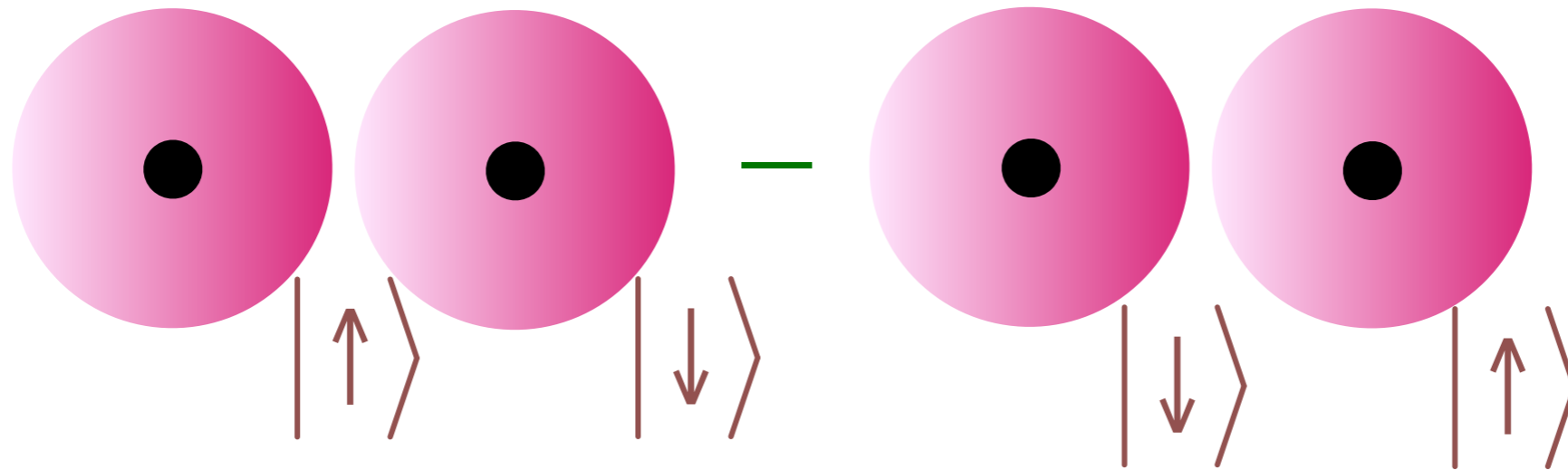


Quantum Entanglement: quantum superposition with more than one particle

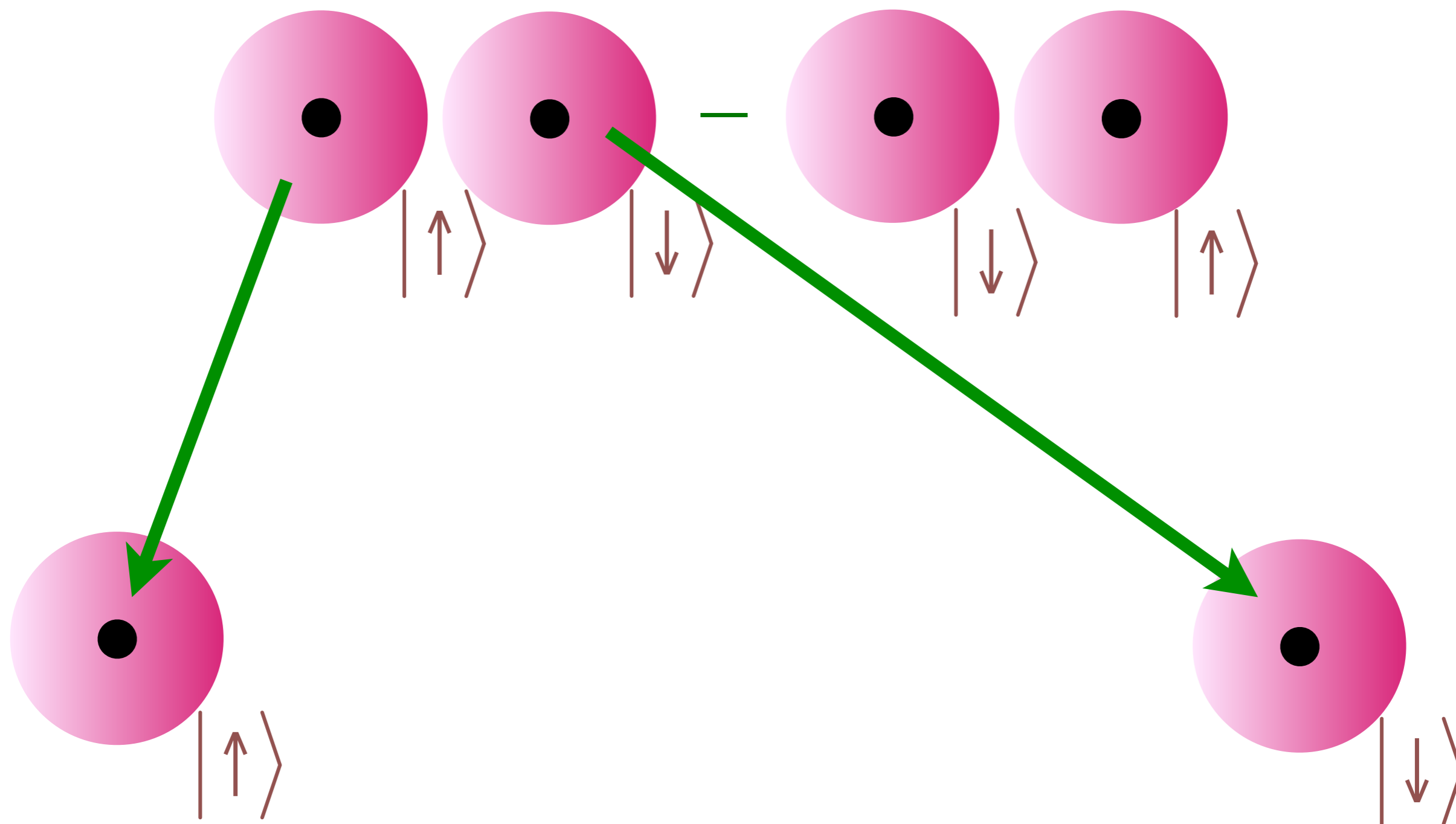
Hydrogen molecule:



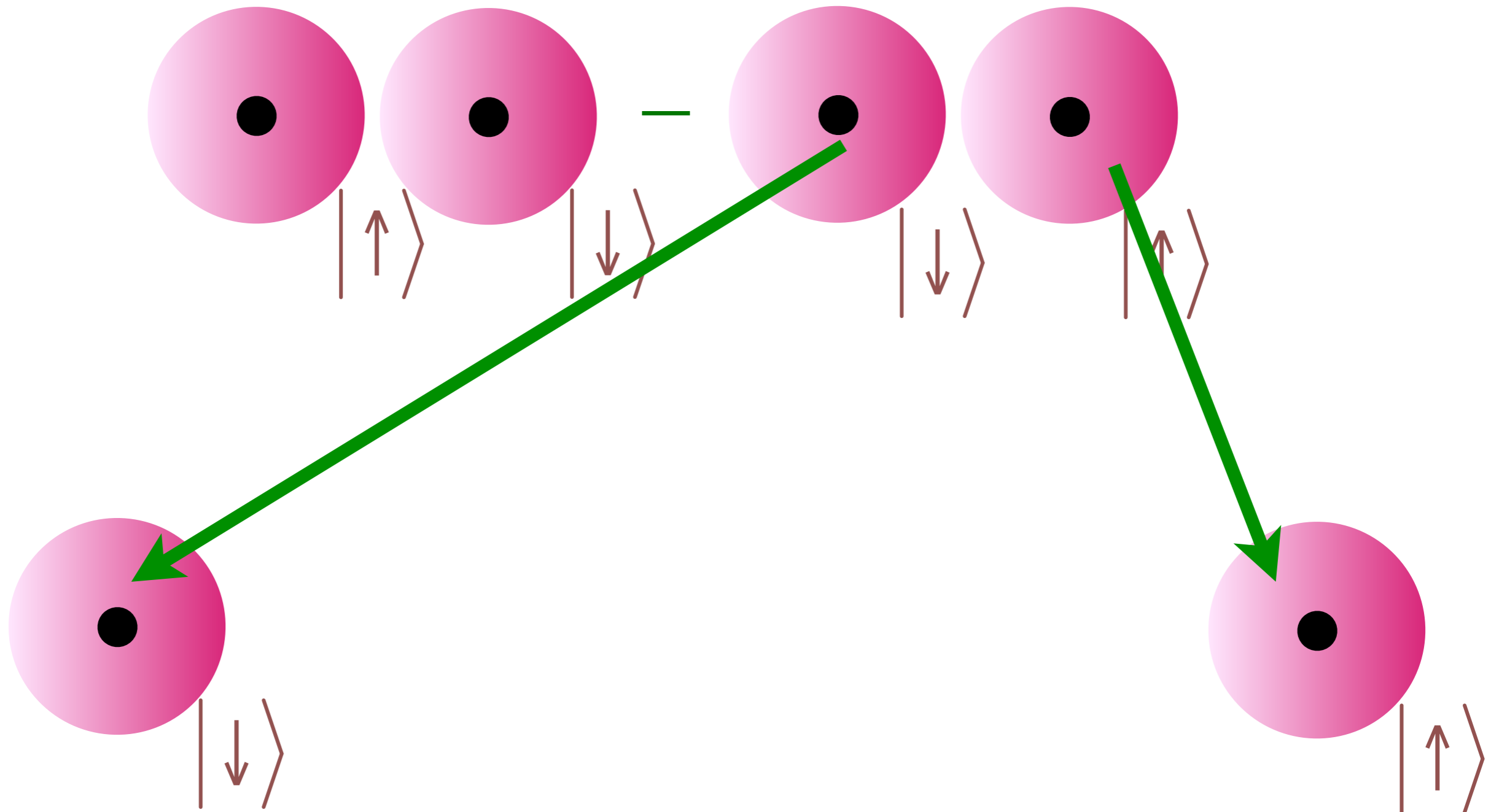
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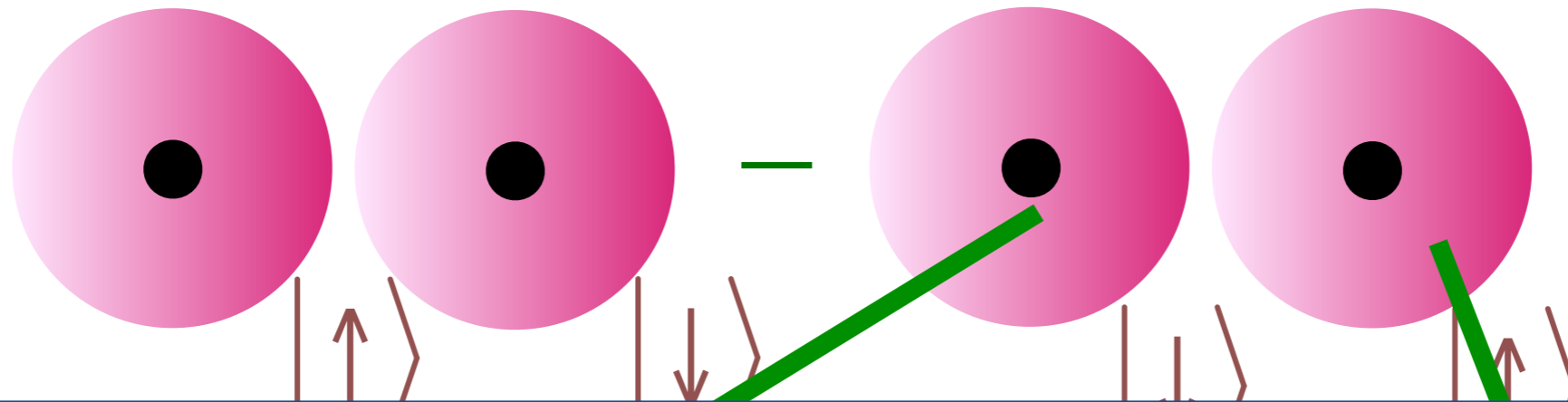
Quantum Entanglement: quantum superposition with more than one particle



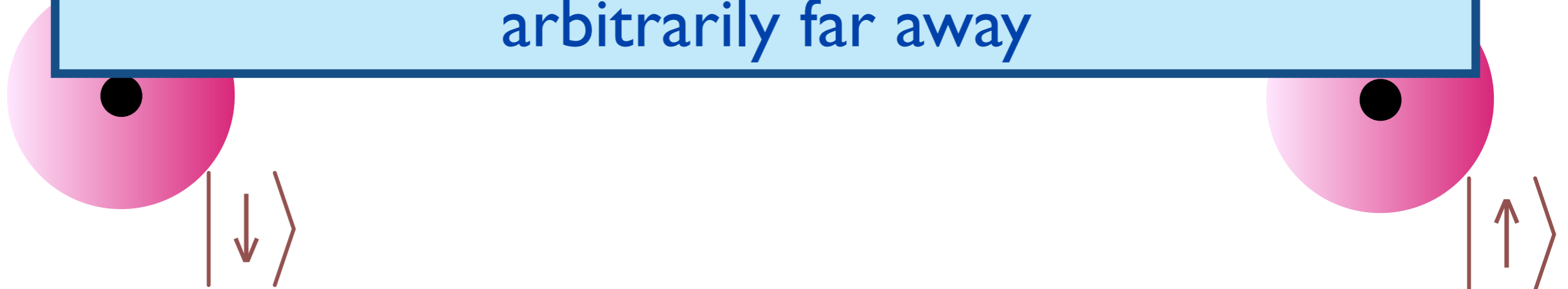
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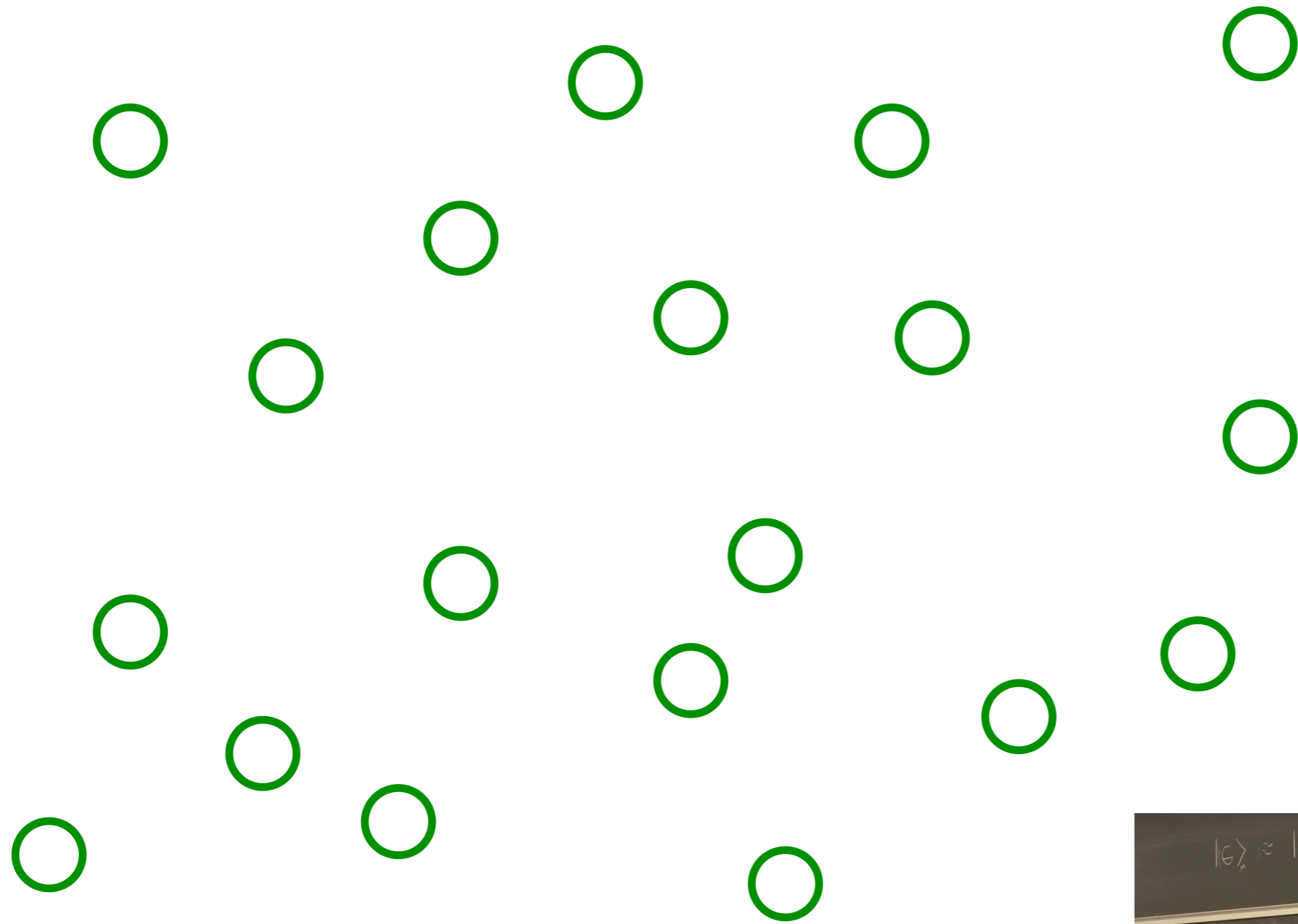
Quantum Entanglement: quantum superposition with more than one particle



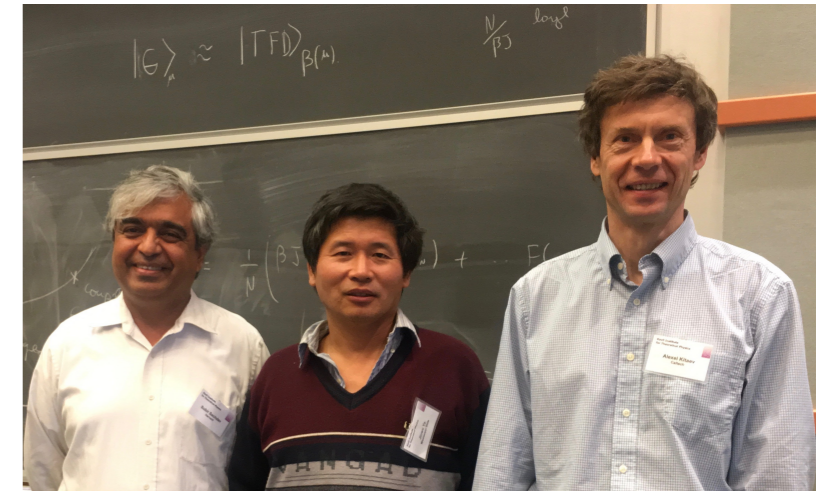
Einstein-Podolsky-Rosen “paradox” (1935):
Measurement of one particle instantaneously
determines the state of the other particle
arbitrarily far away



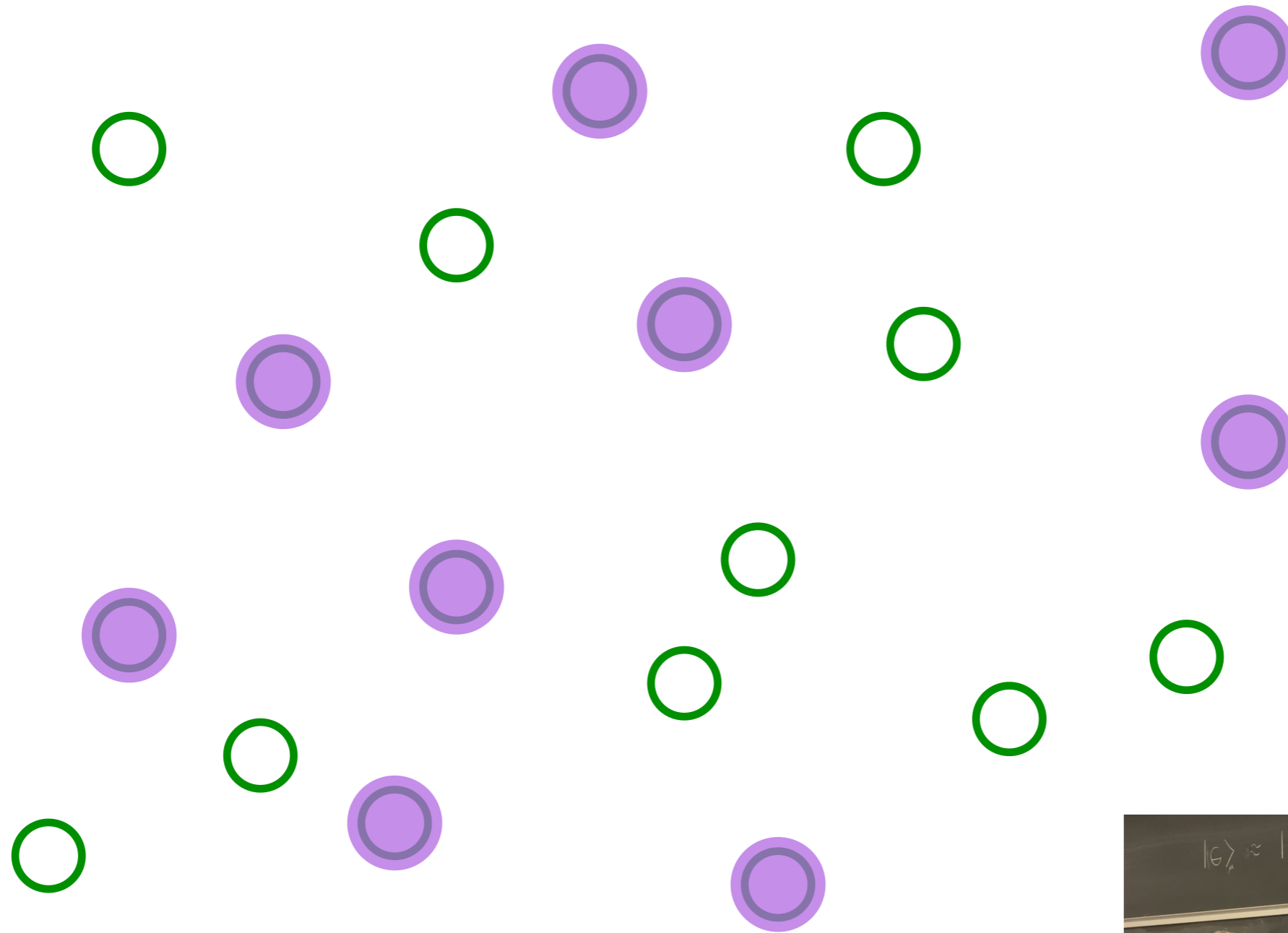
The Sachdev-Ye-Kitaev (SYK) model



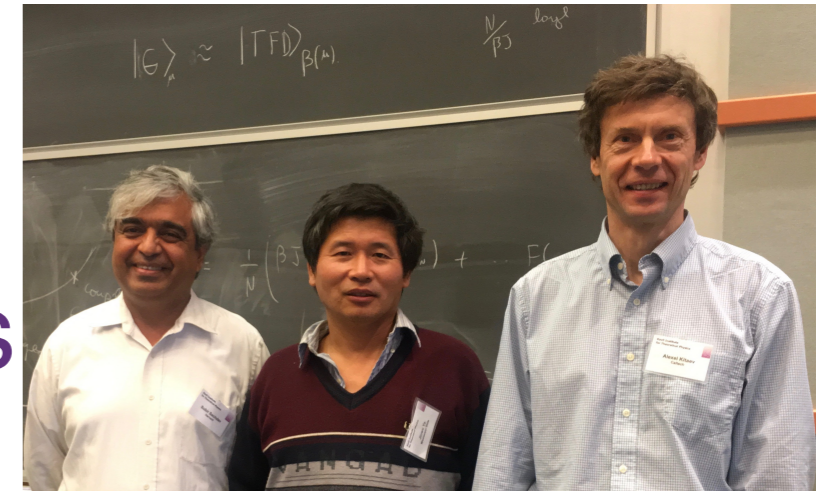
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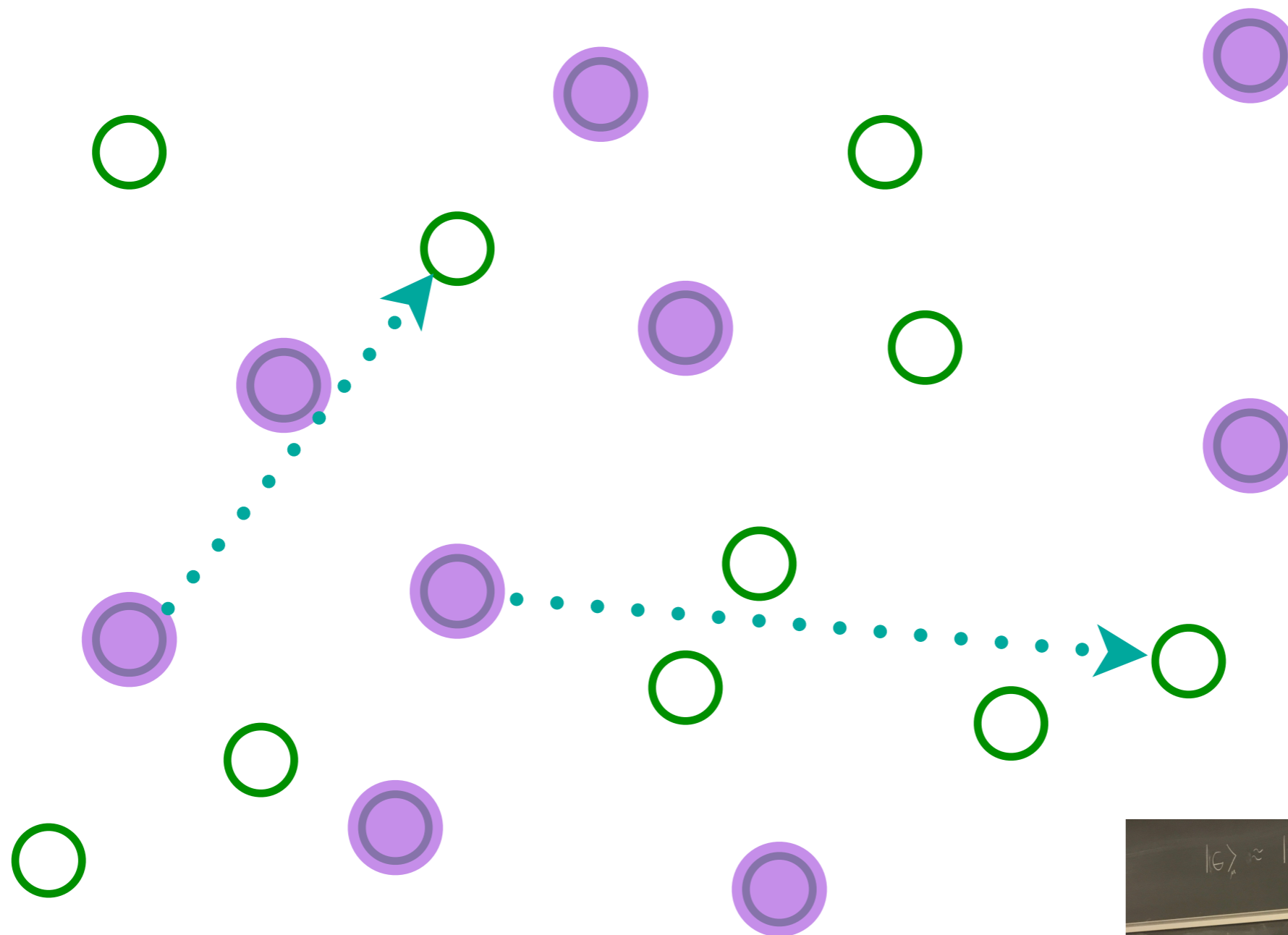
The SYK model



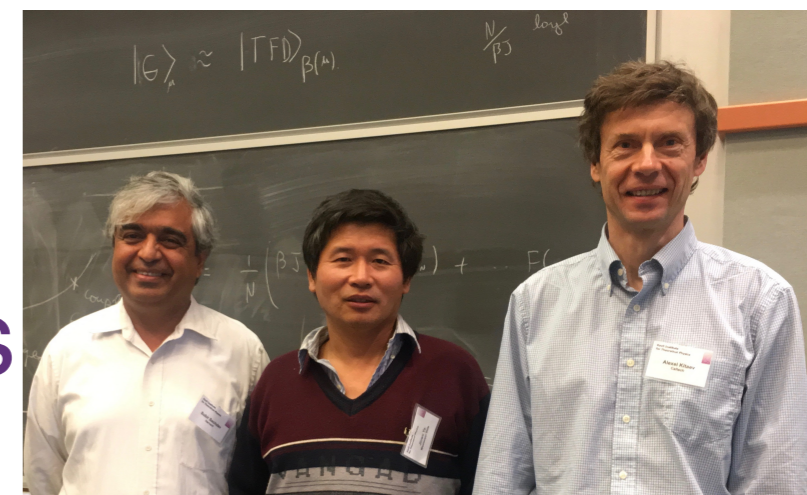
Place electrons randomly on some sites



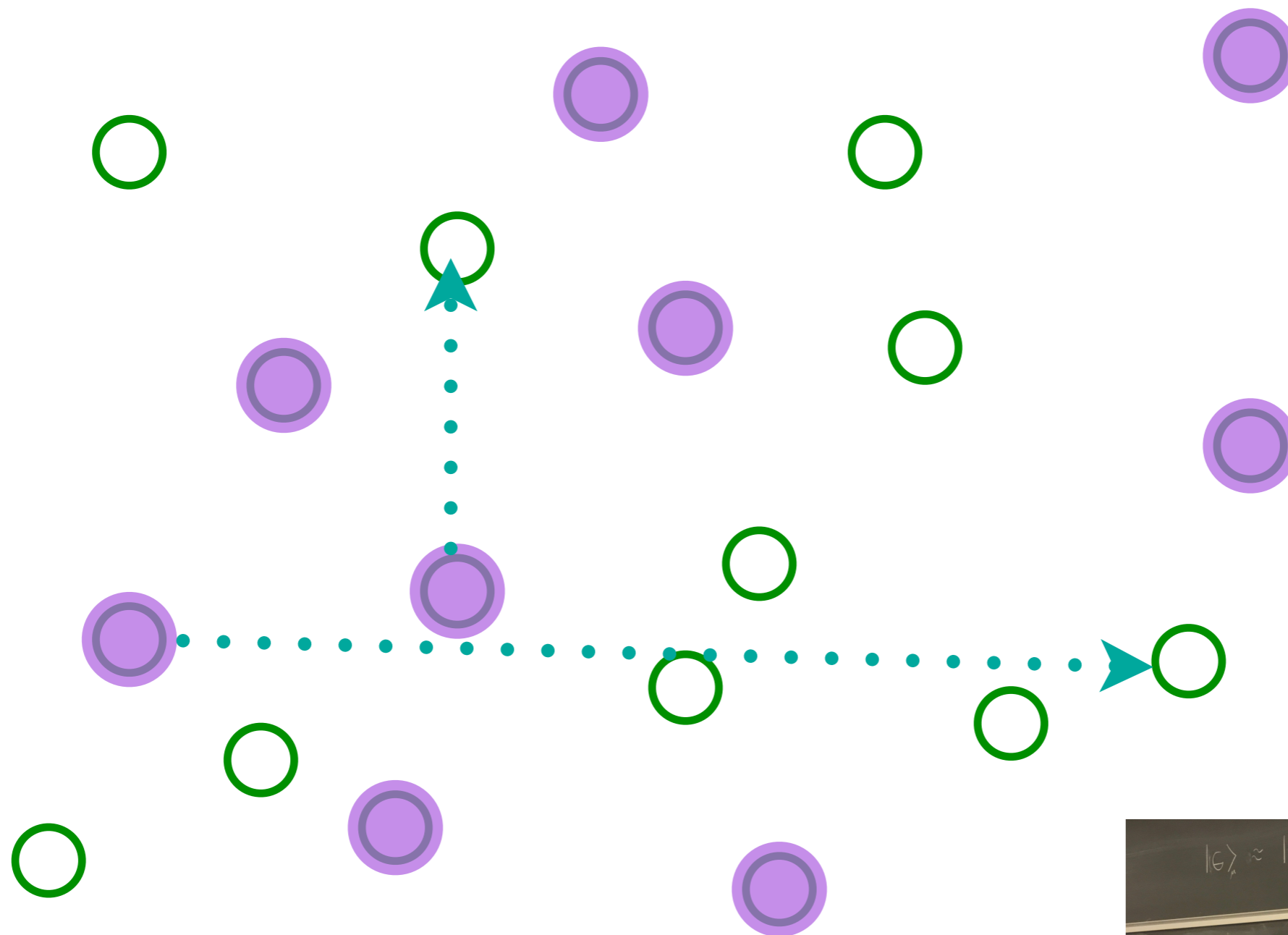
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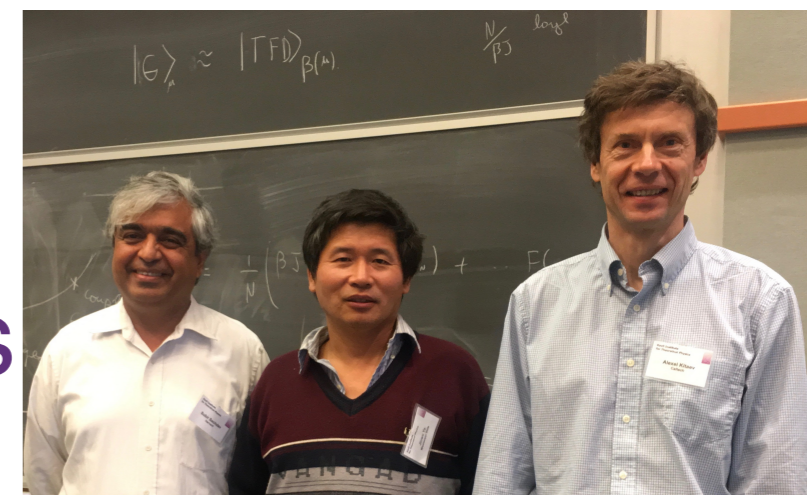
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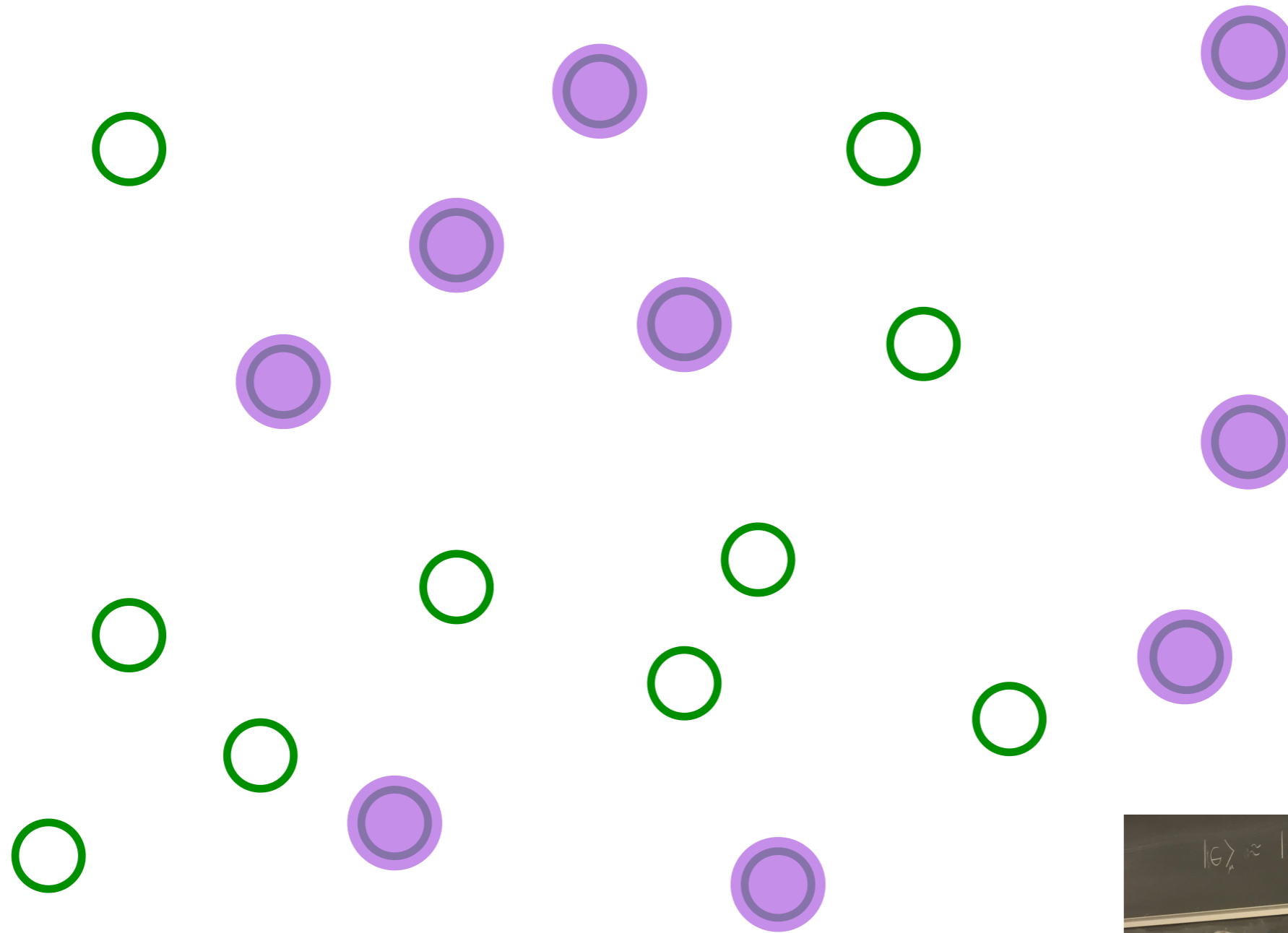
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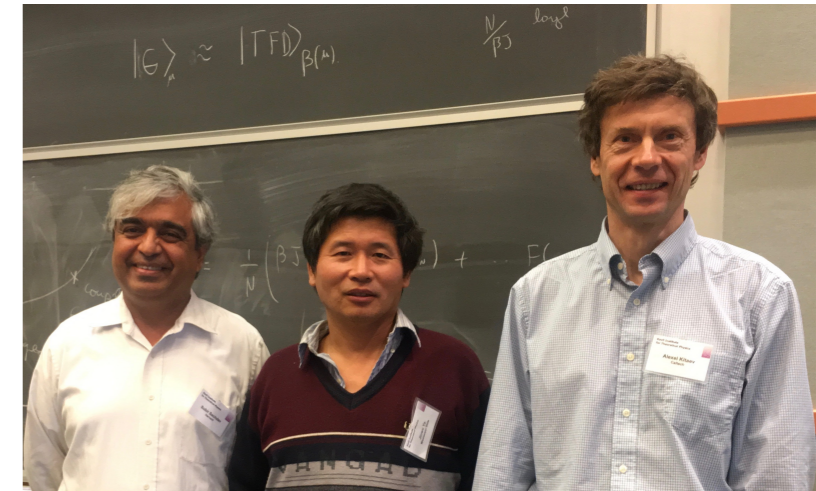
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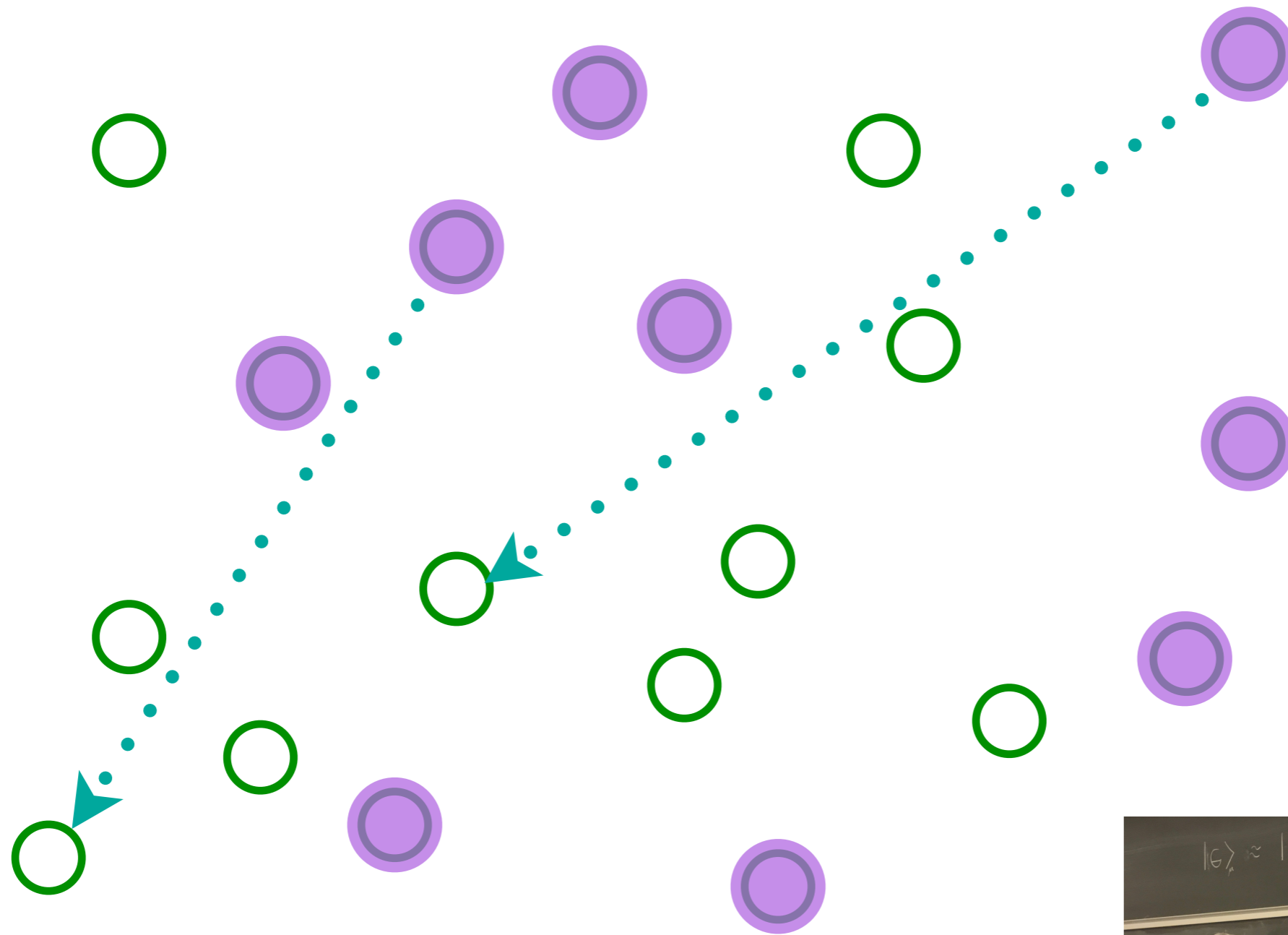
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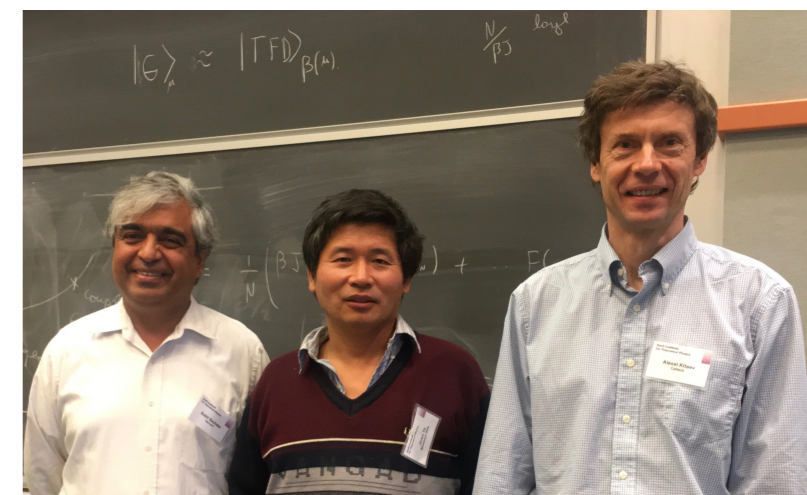
Entangle electrons pairwise randomly



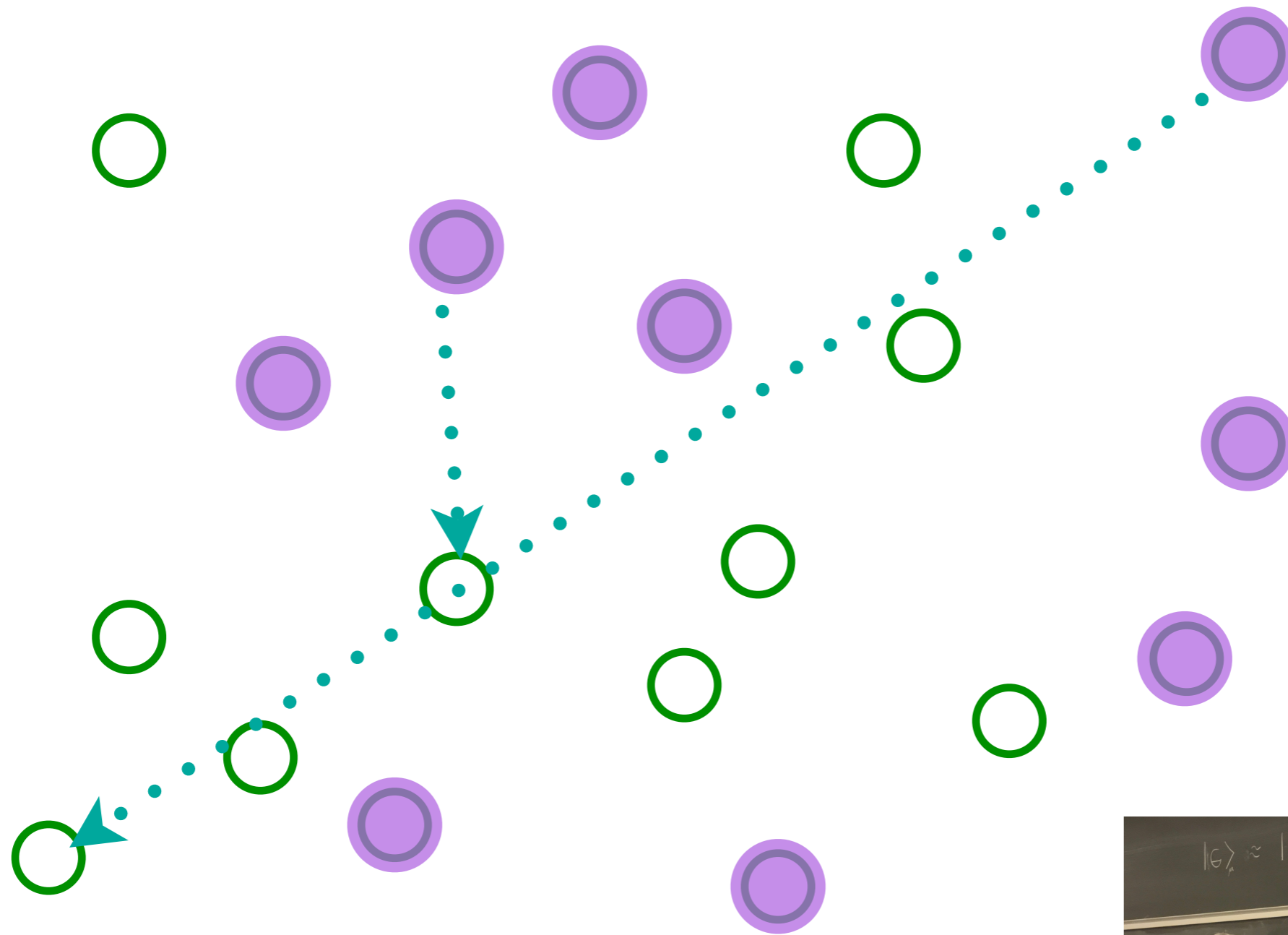
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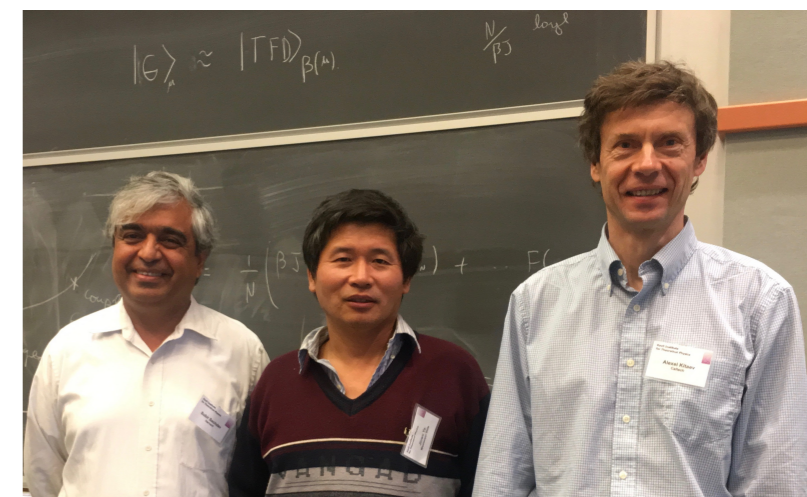
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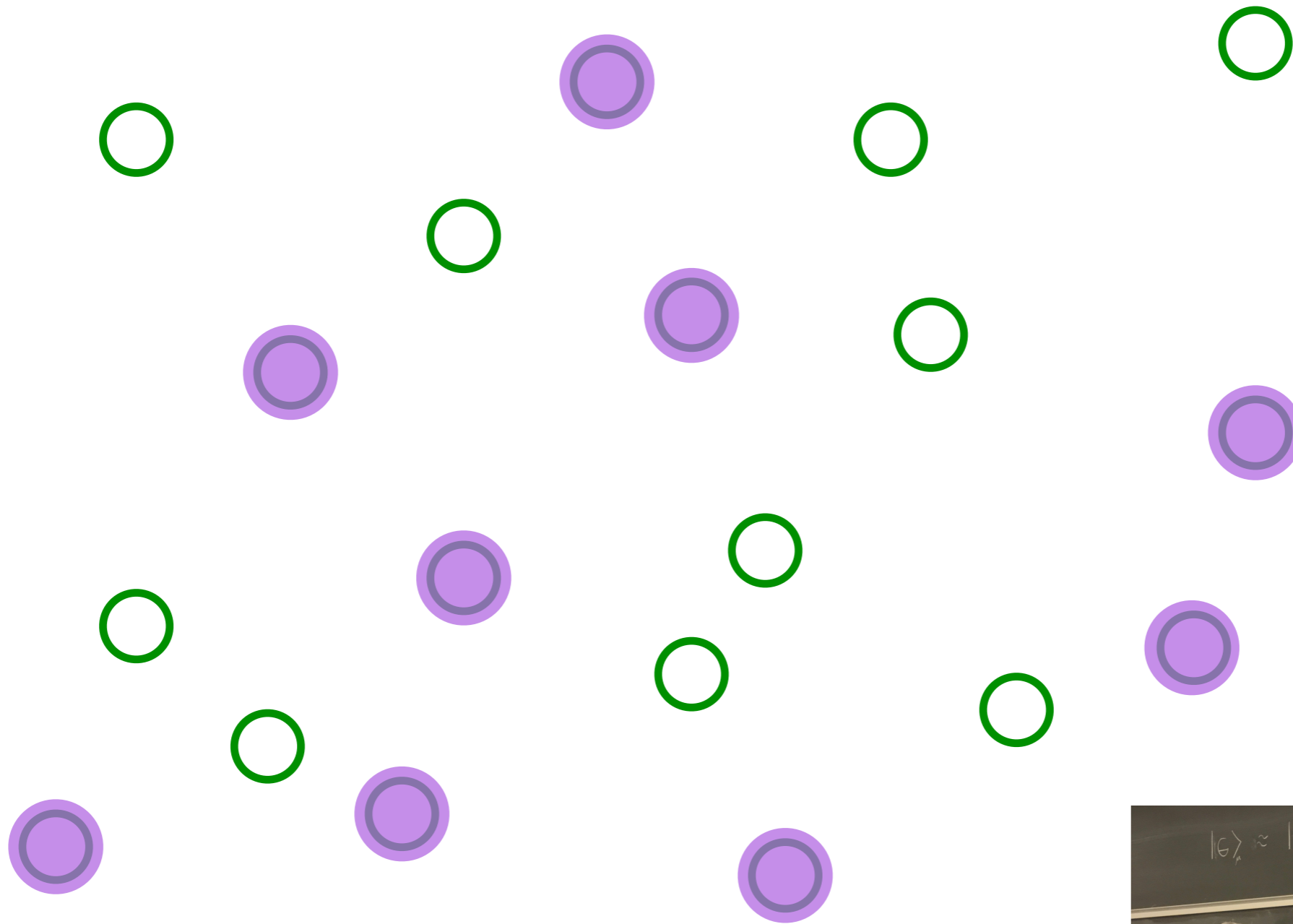
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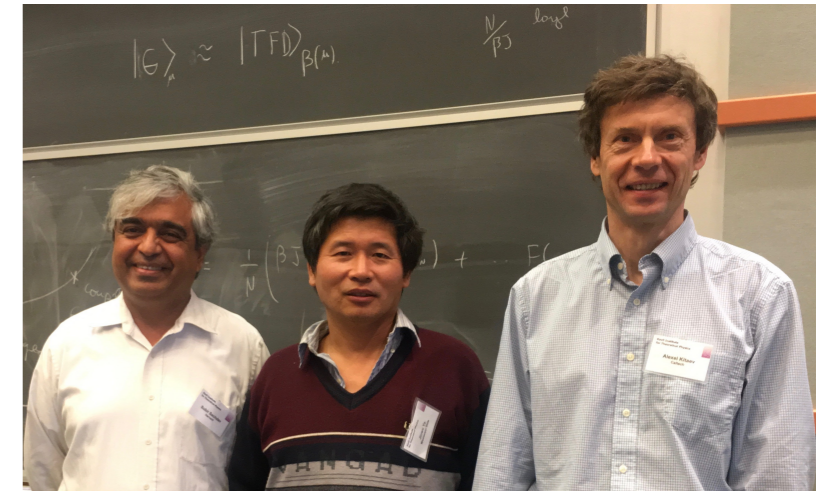
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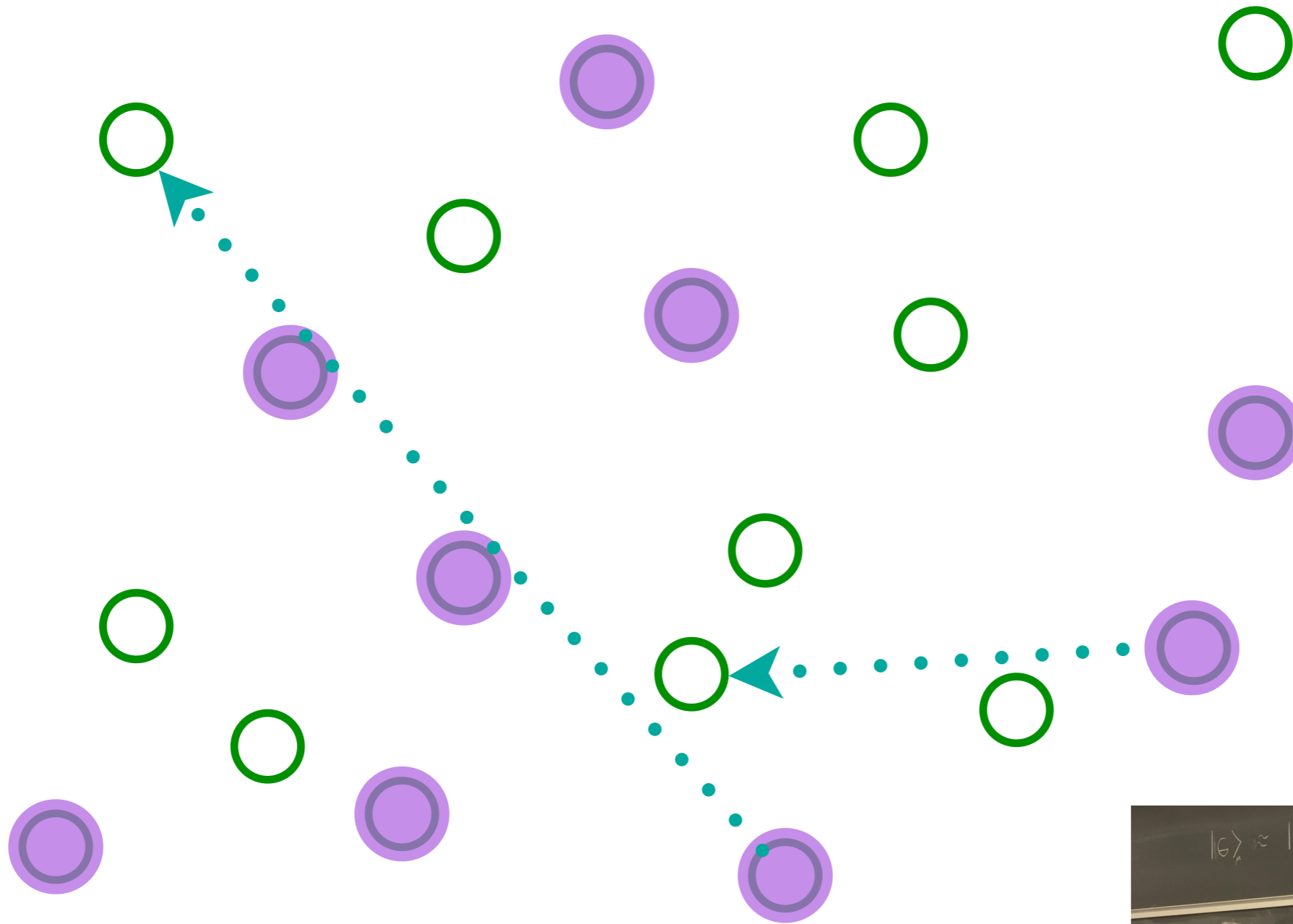
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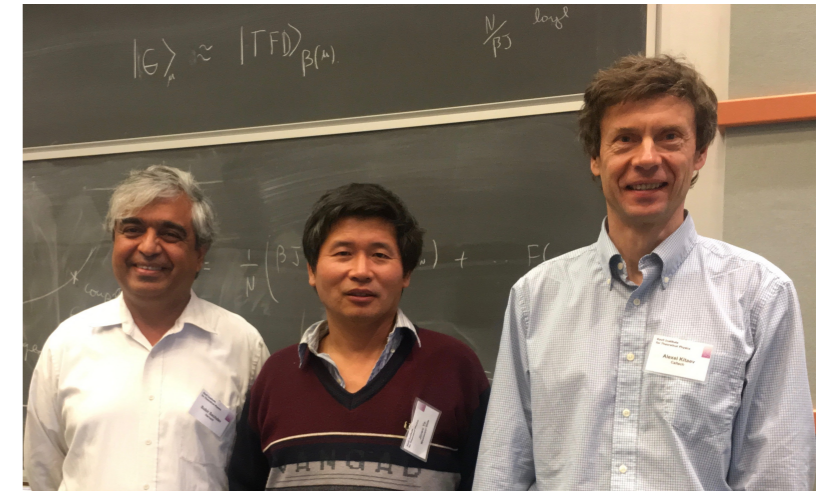
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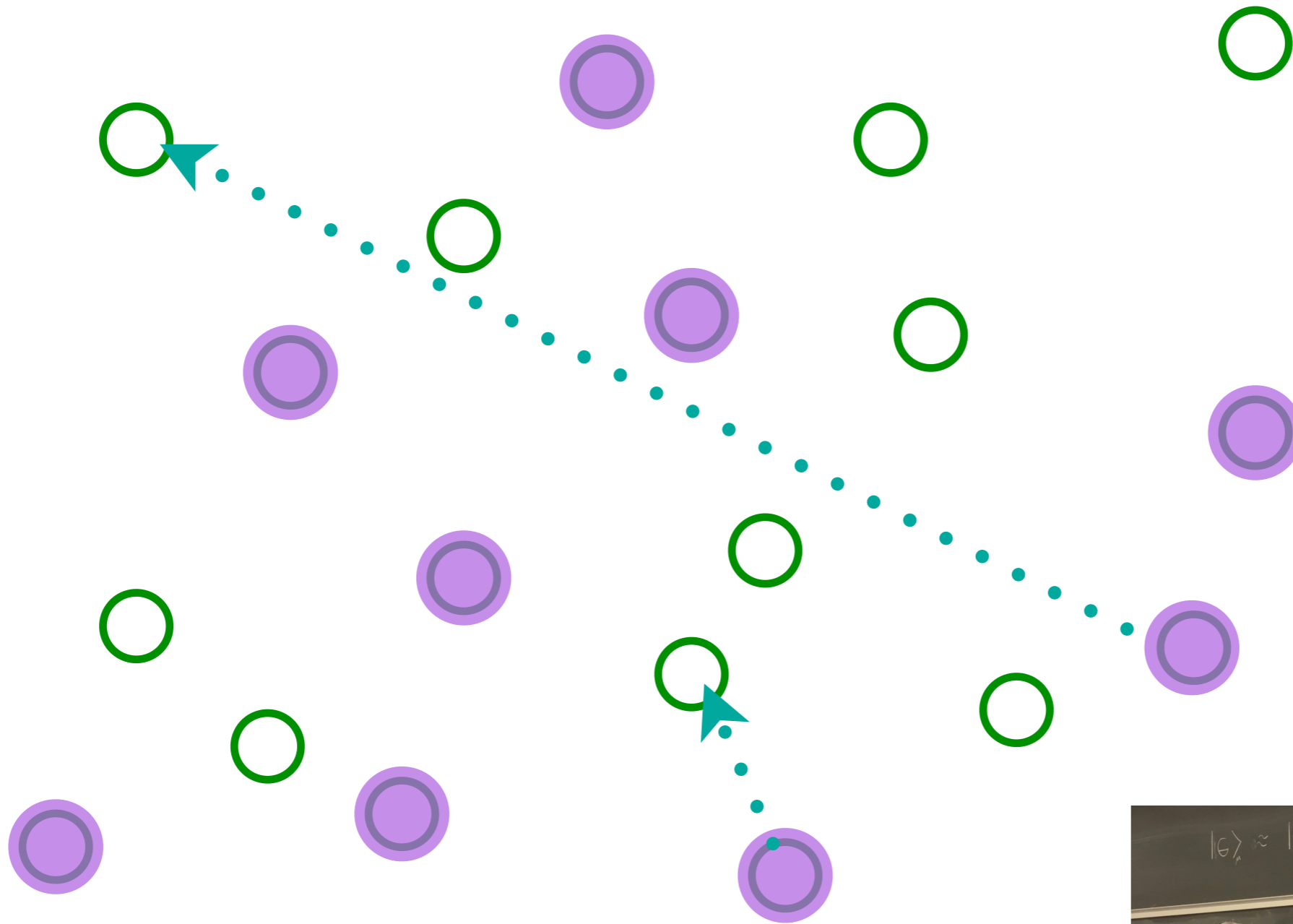
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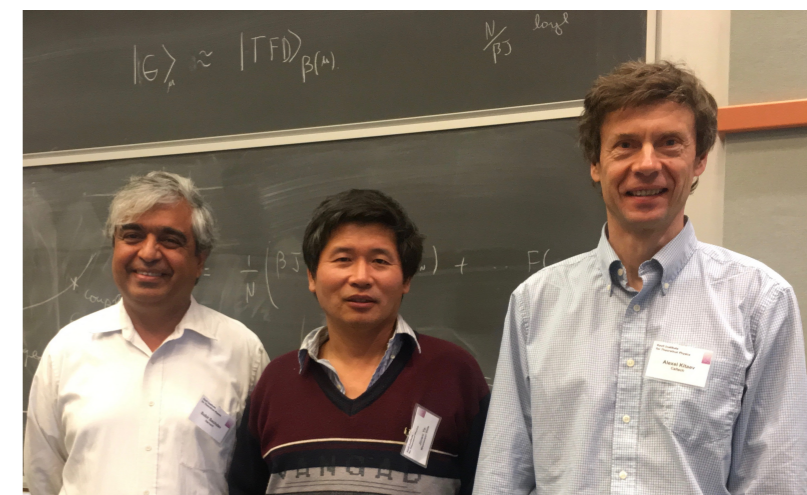
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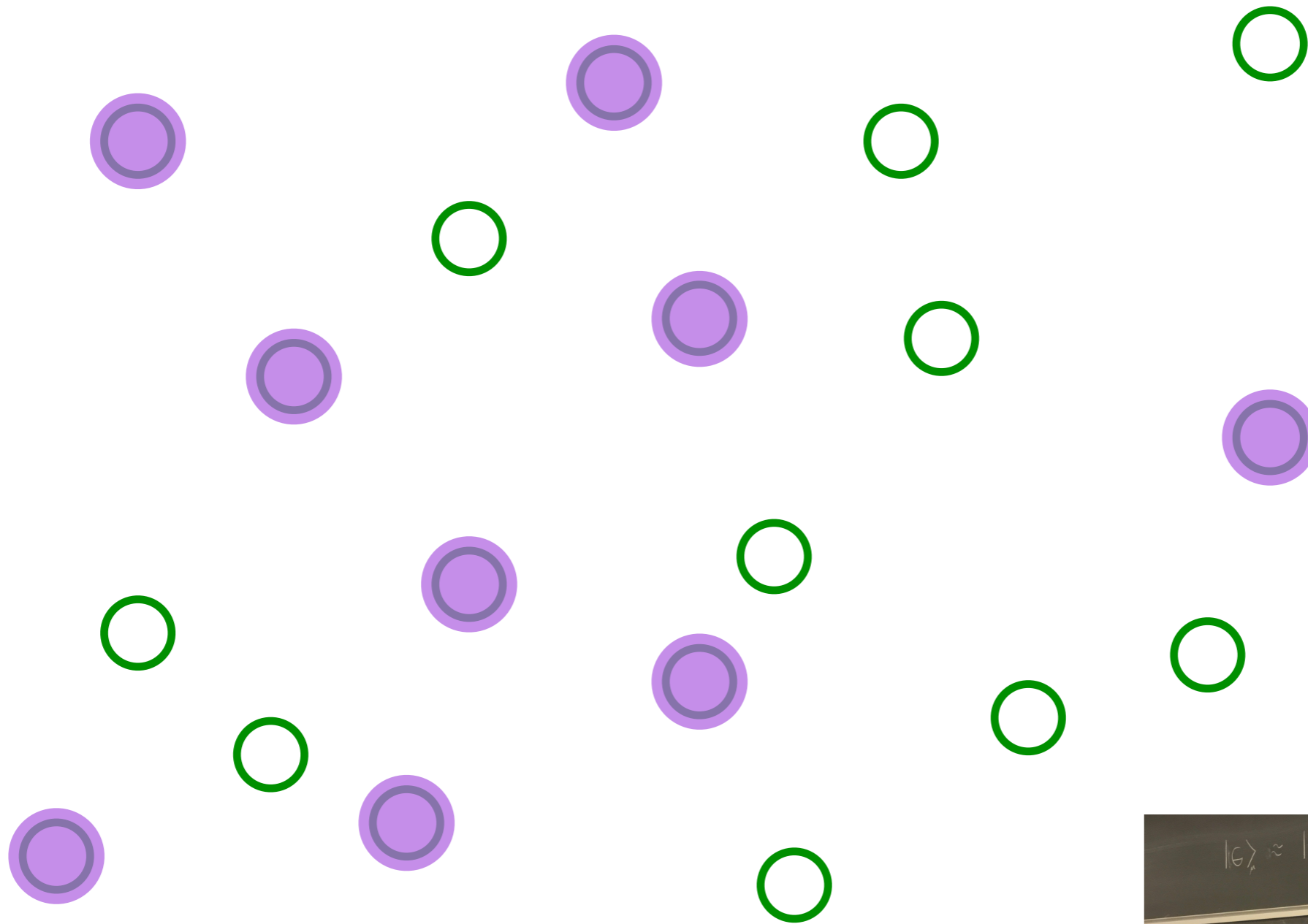
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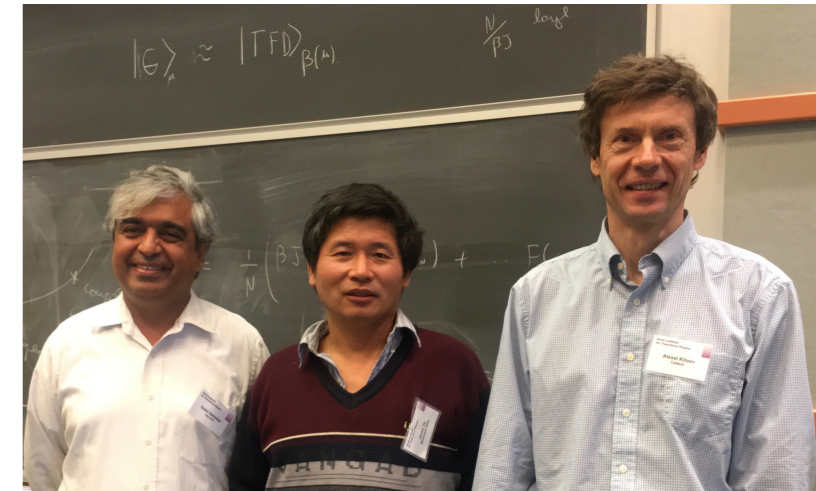
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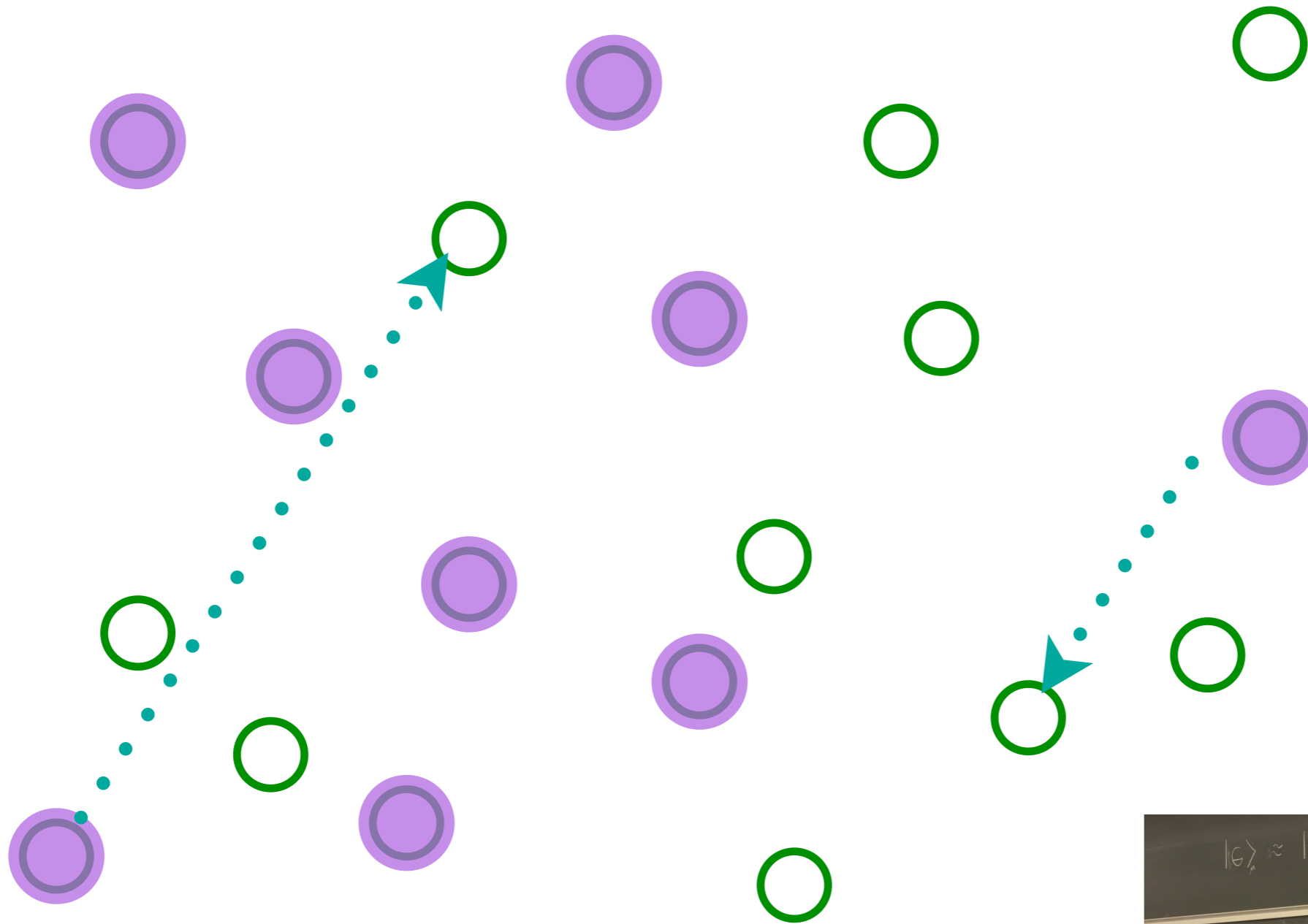
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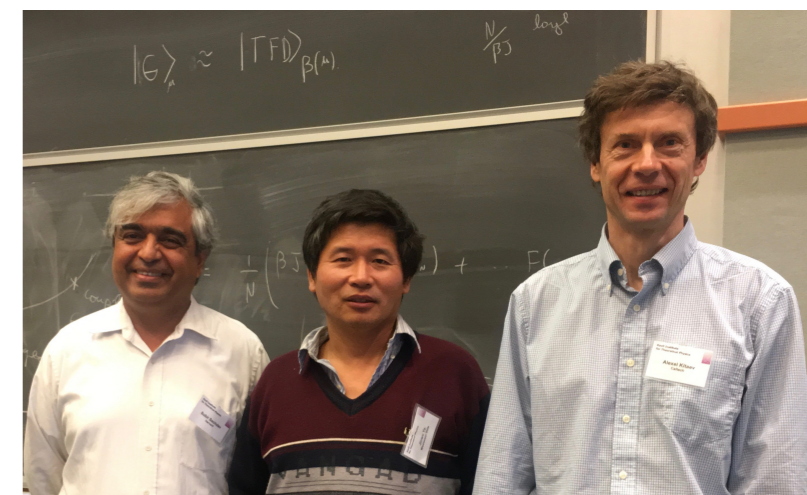
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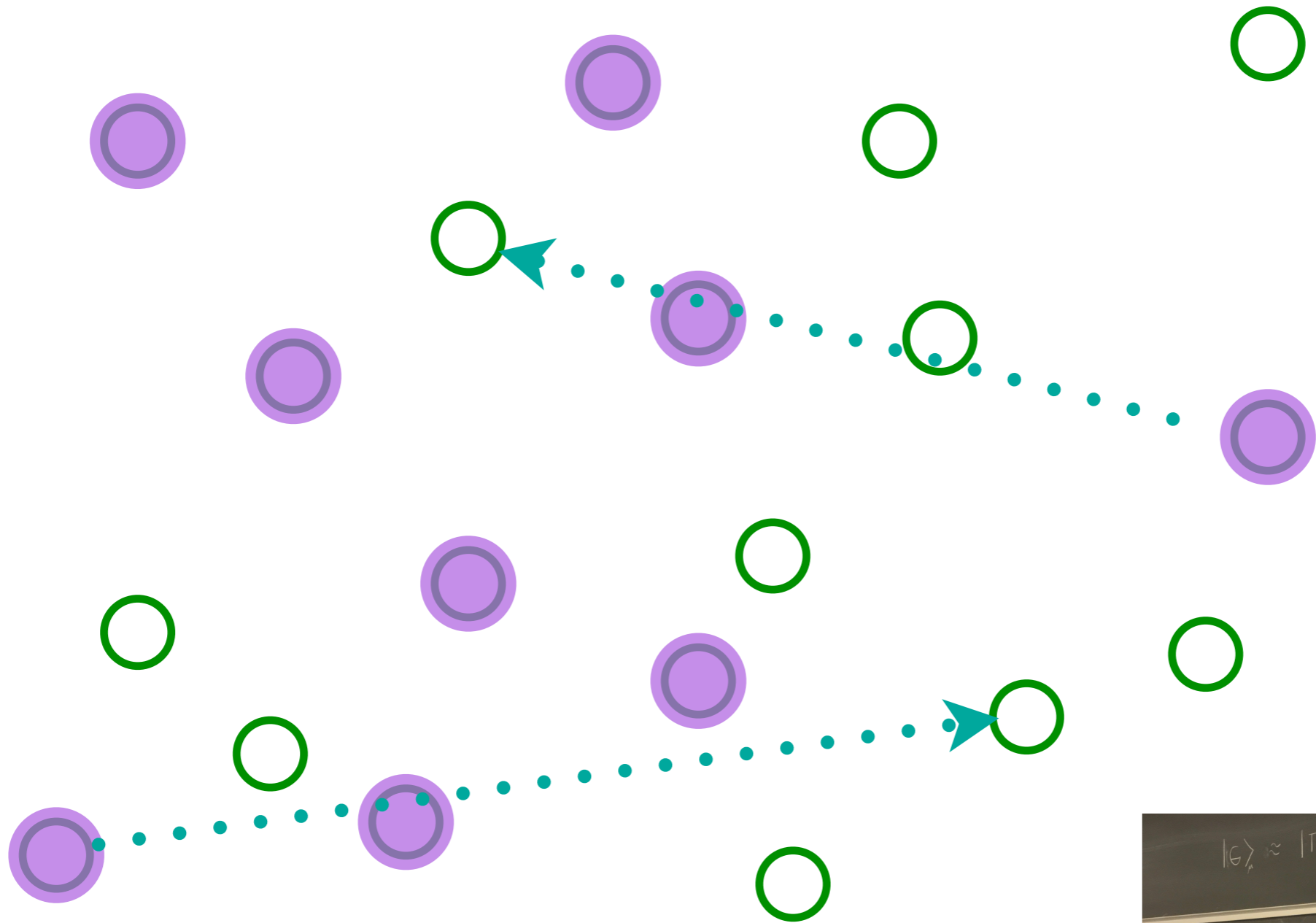
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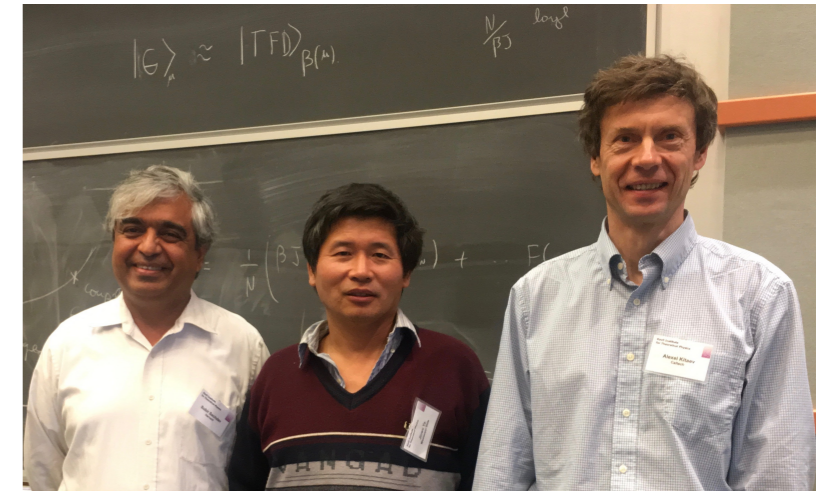
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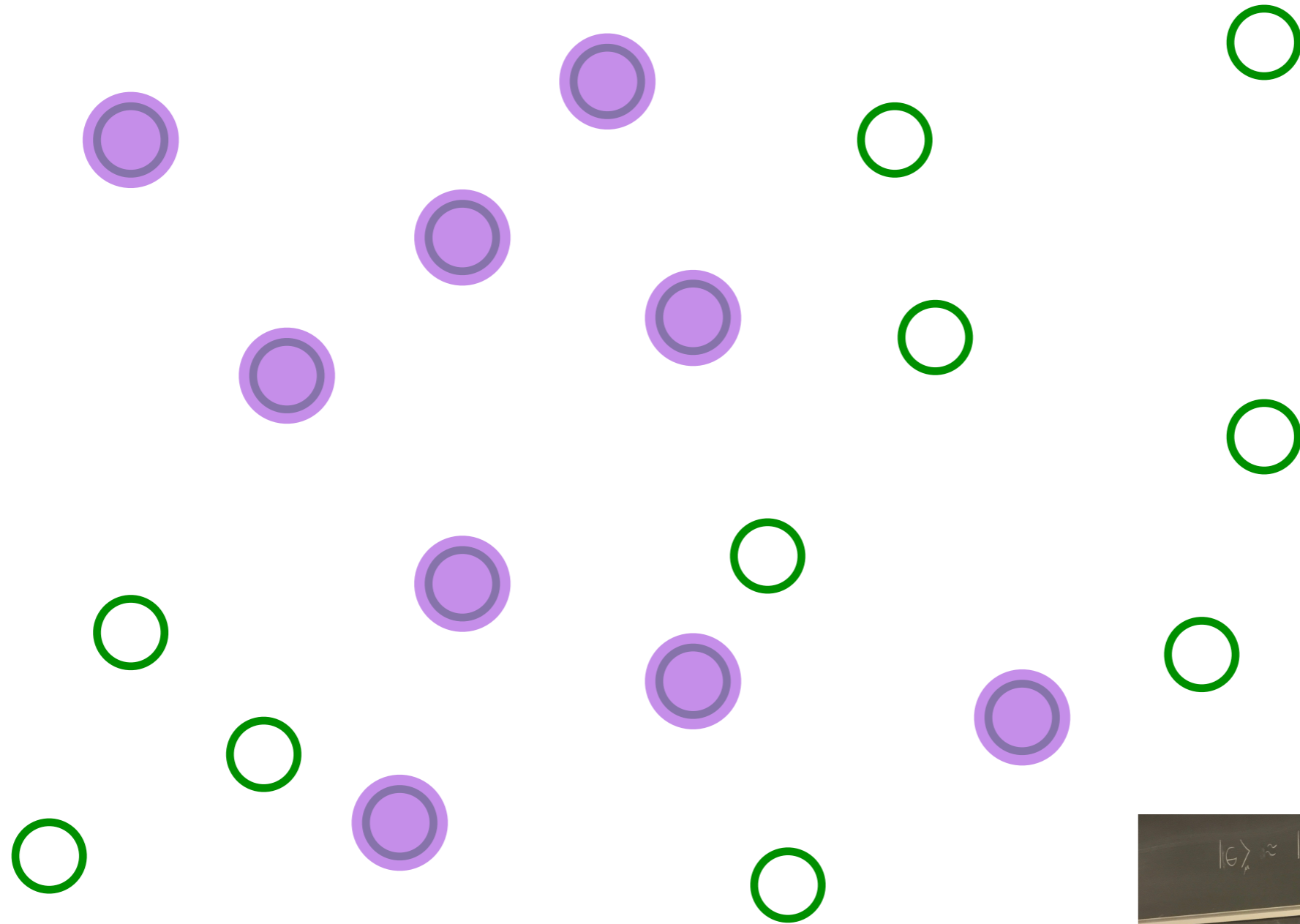
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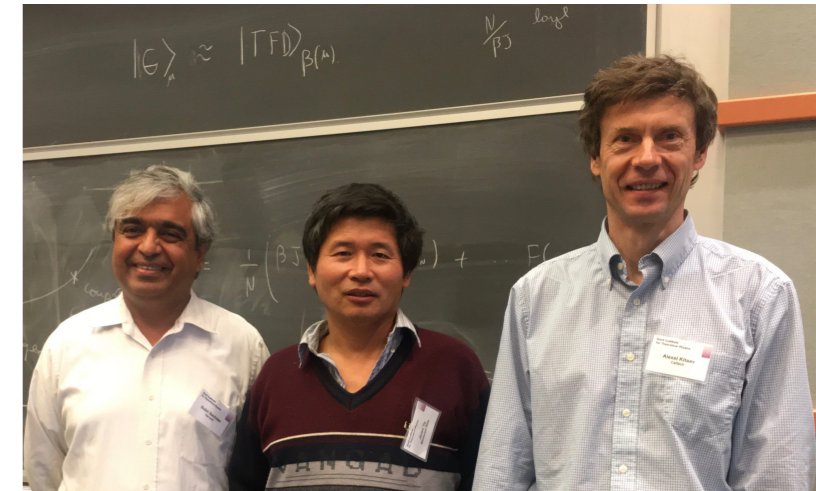
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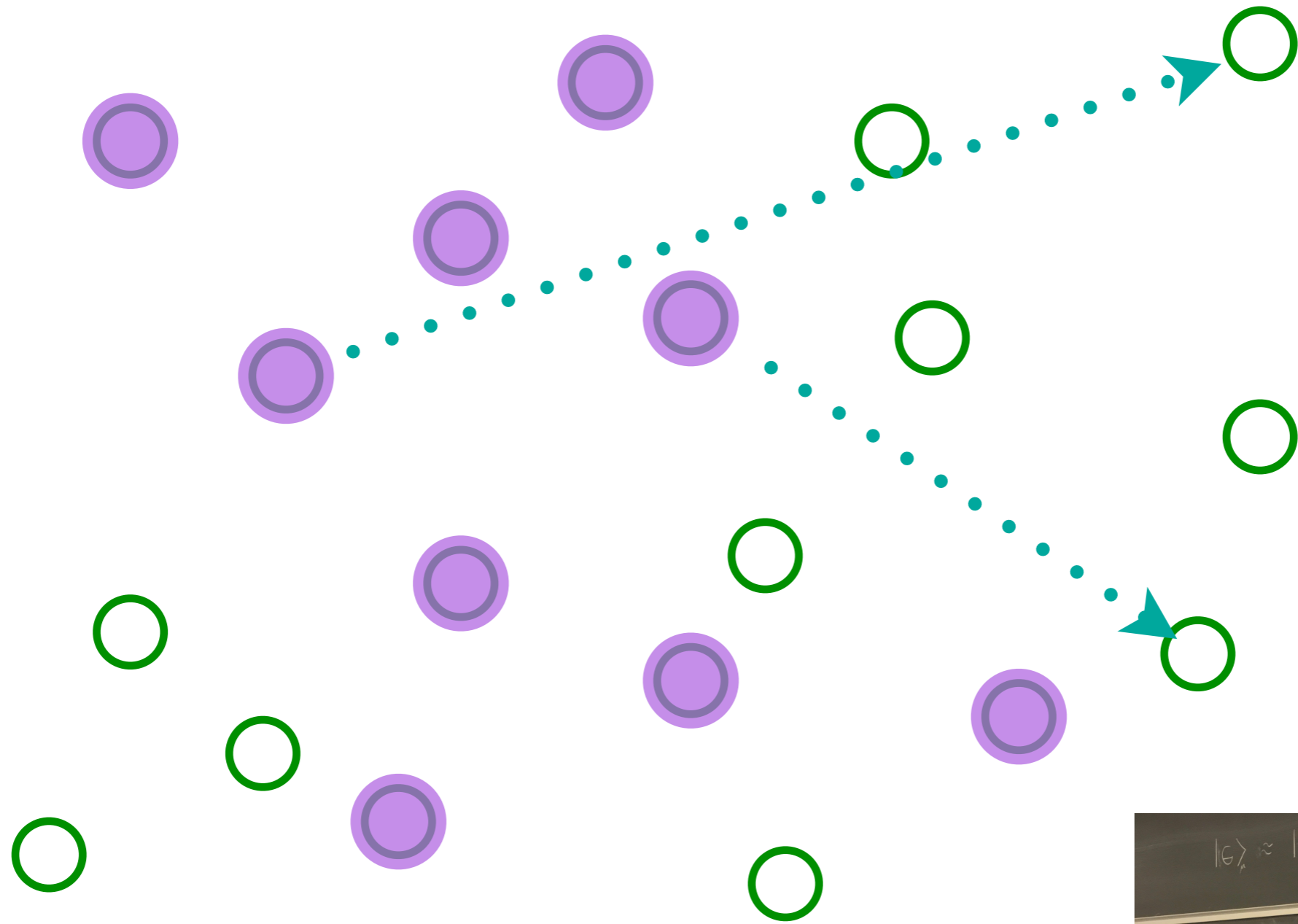
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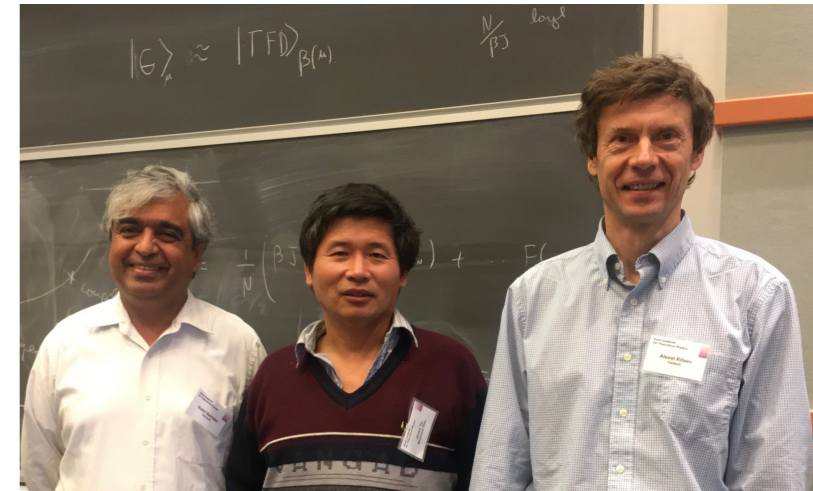
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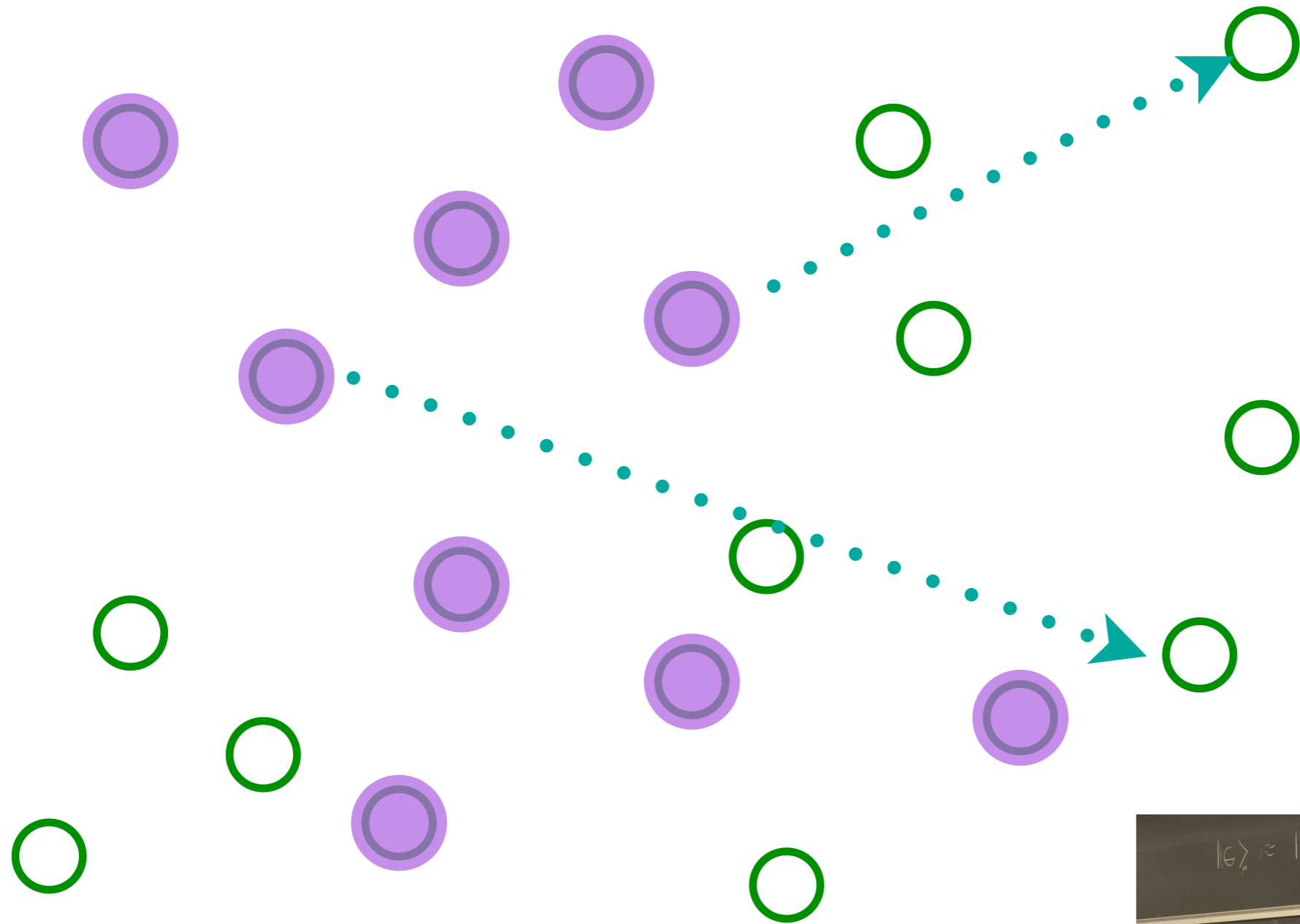
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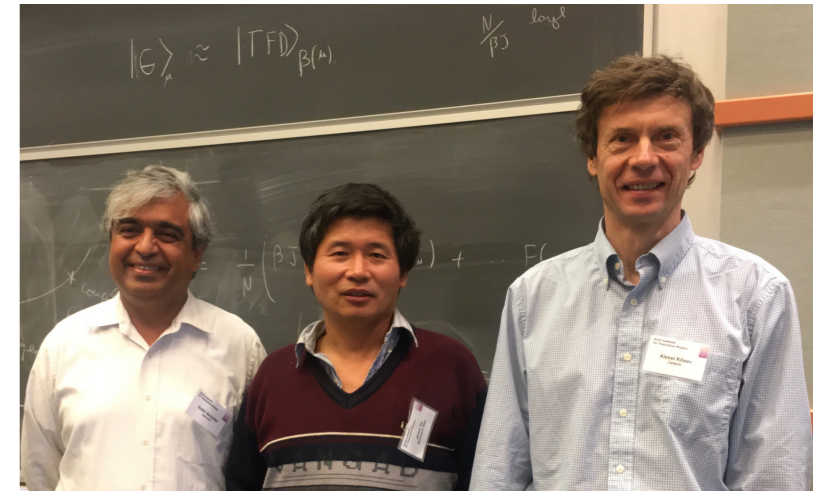
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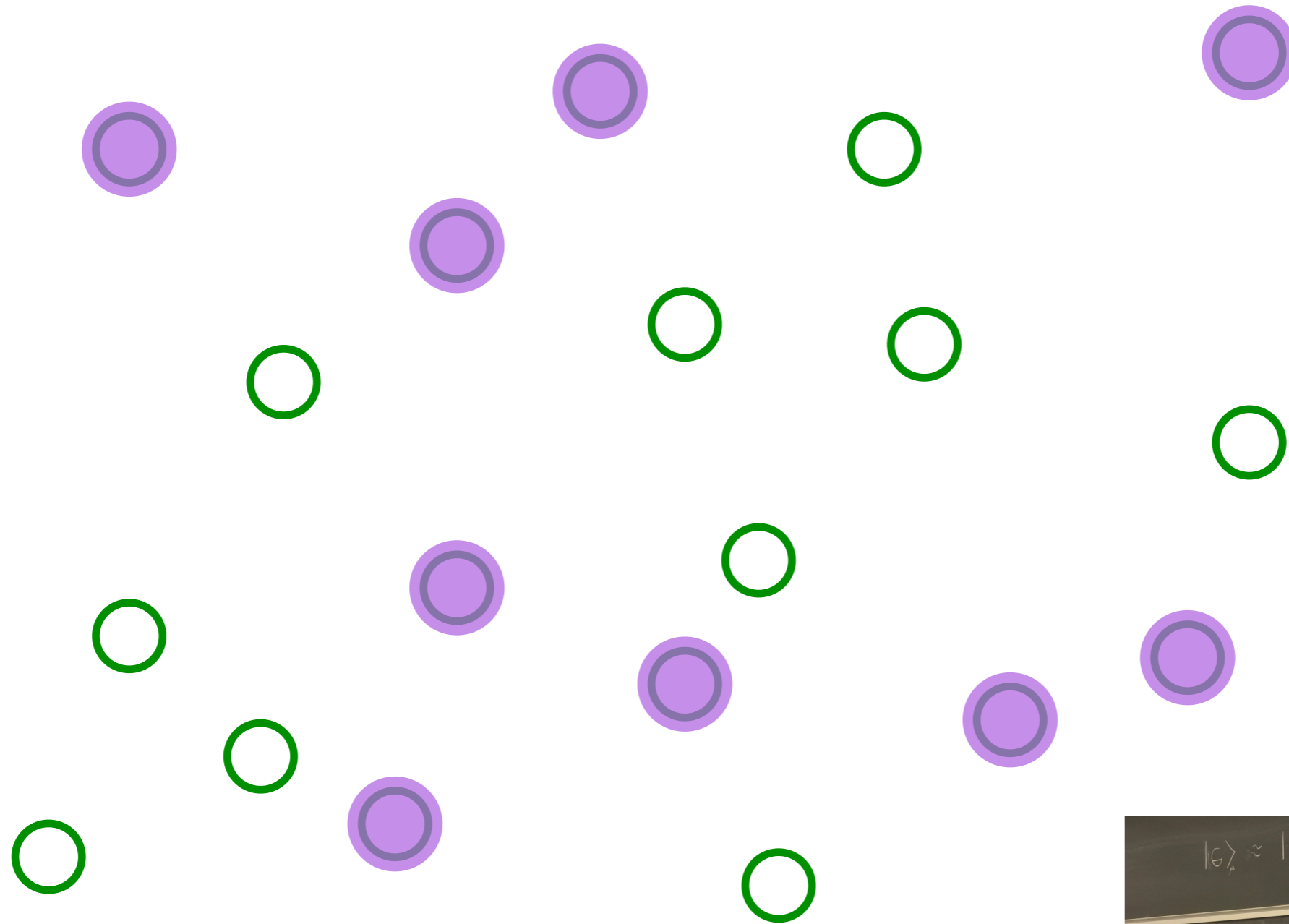
The SYK model



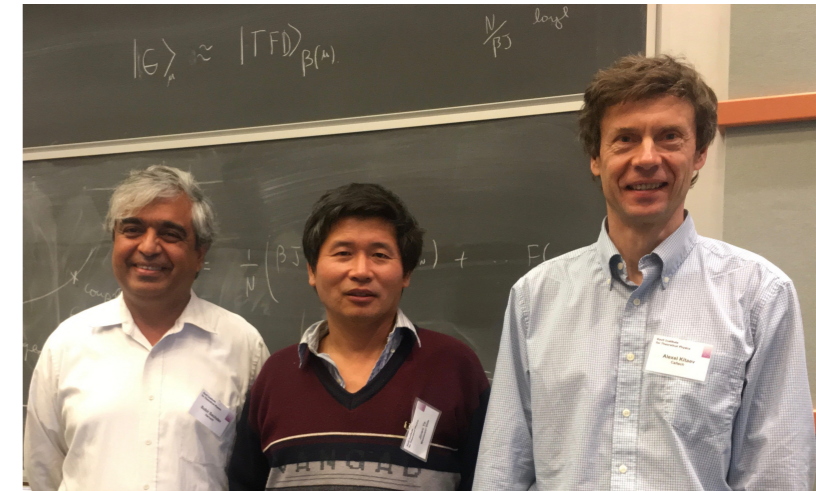
Entangle electrons pairwise randomly



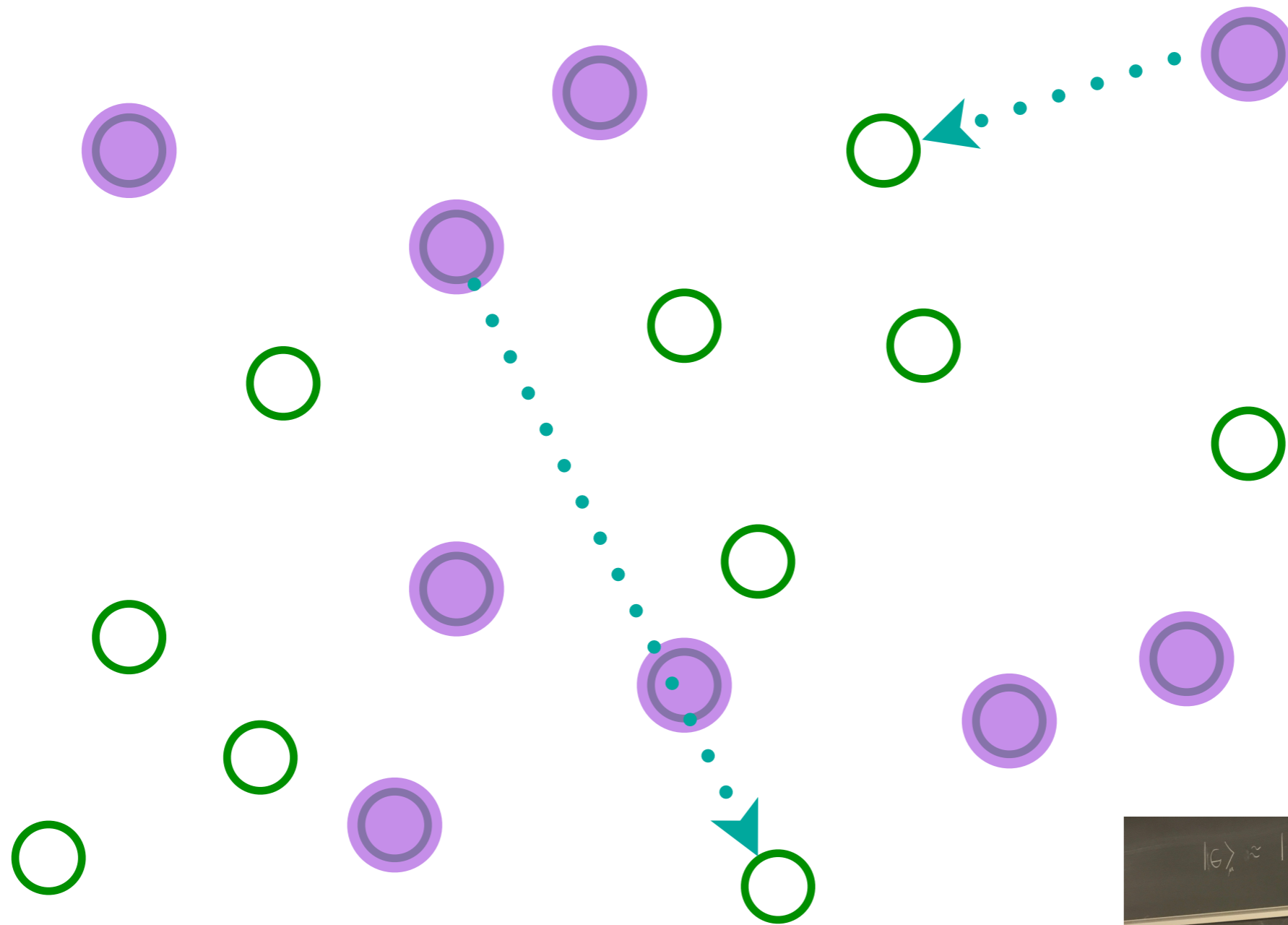
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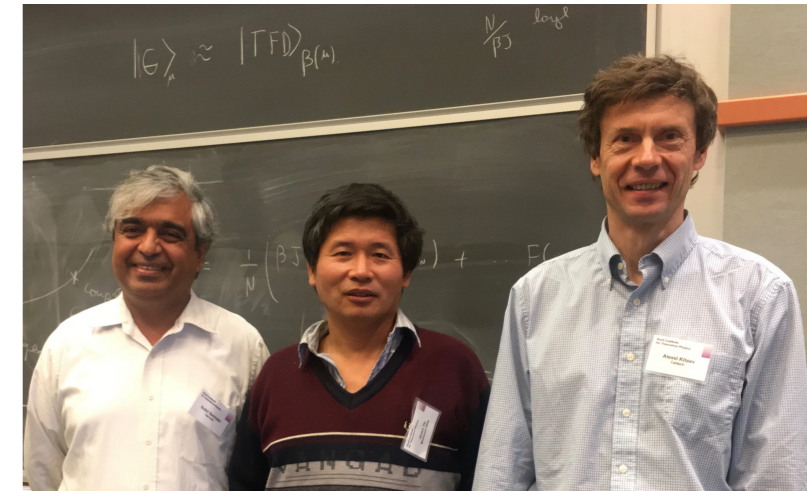
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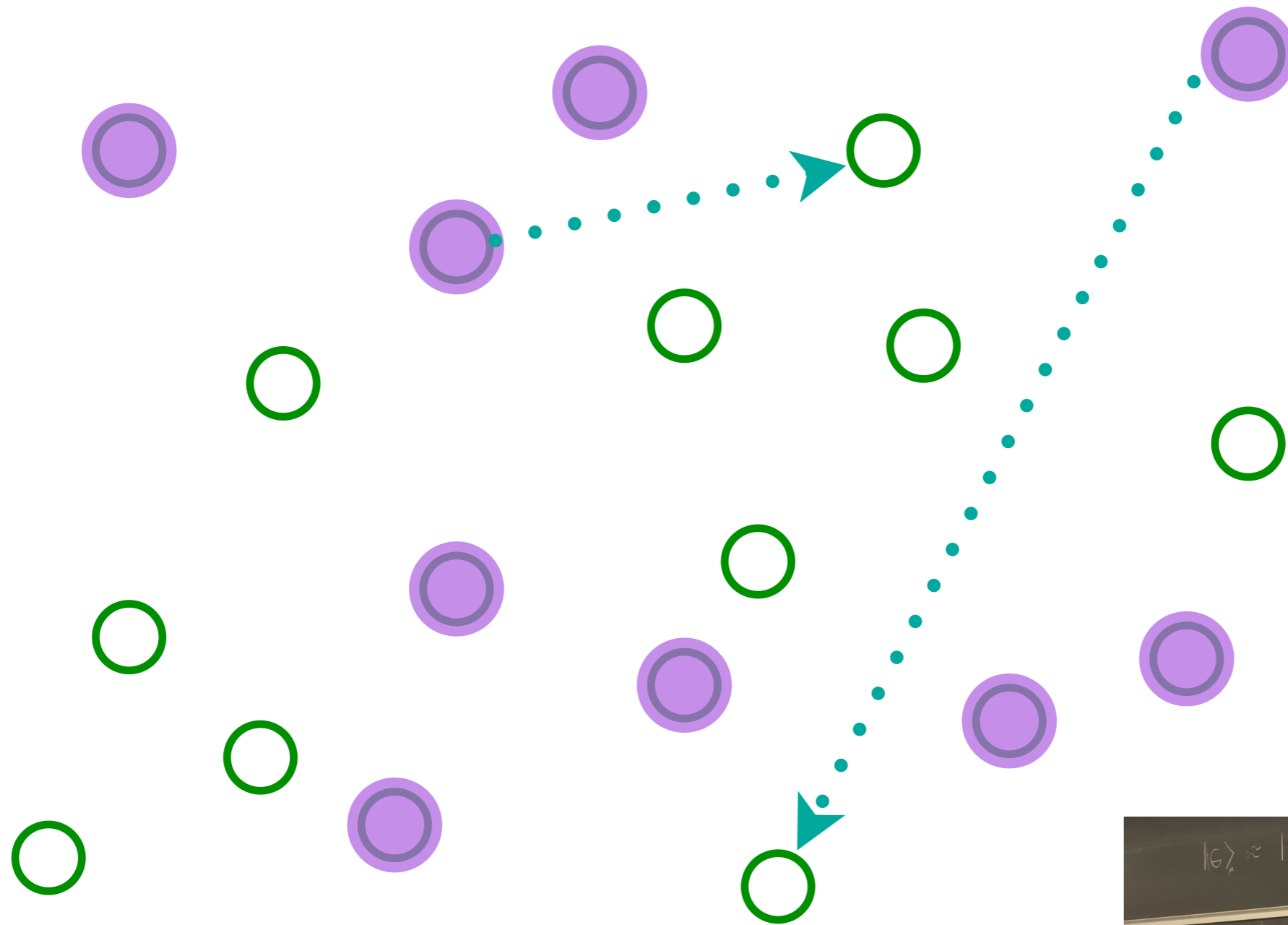
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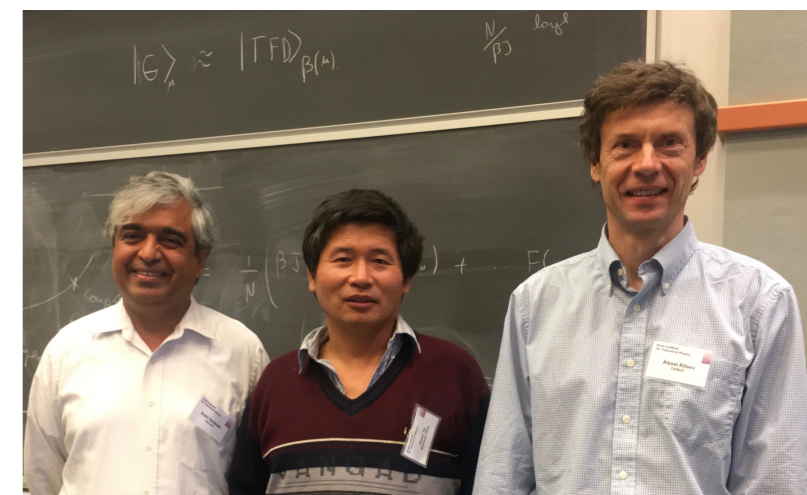
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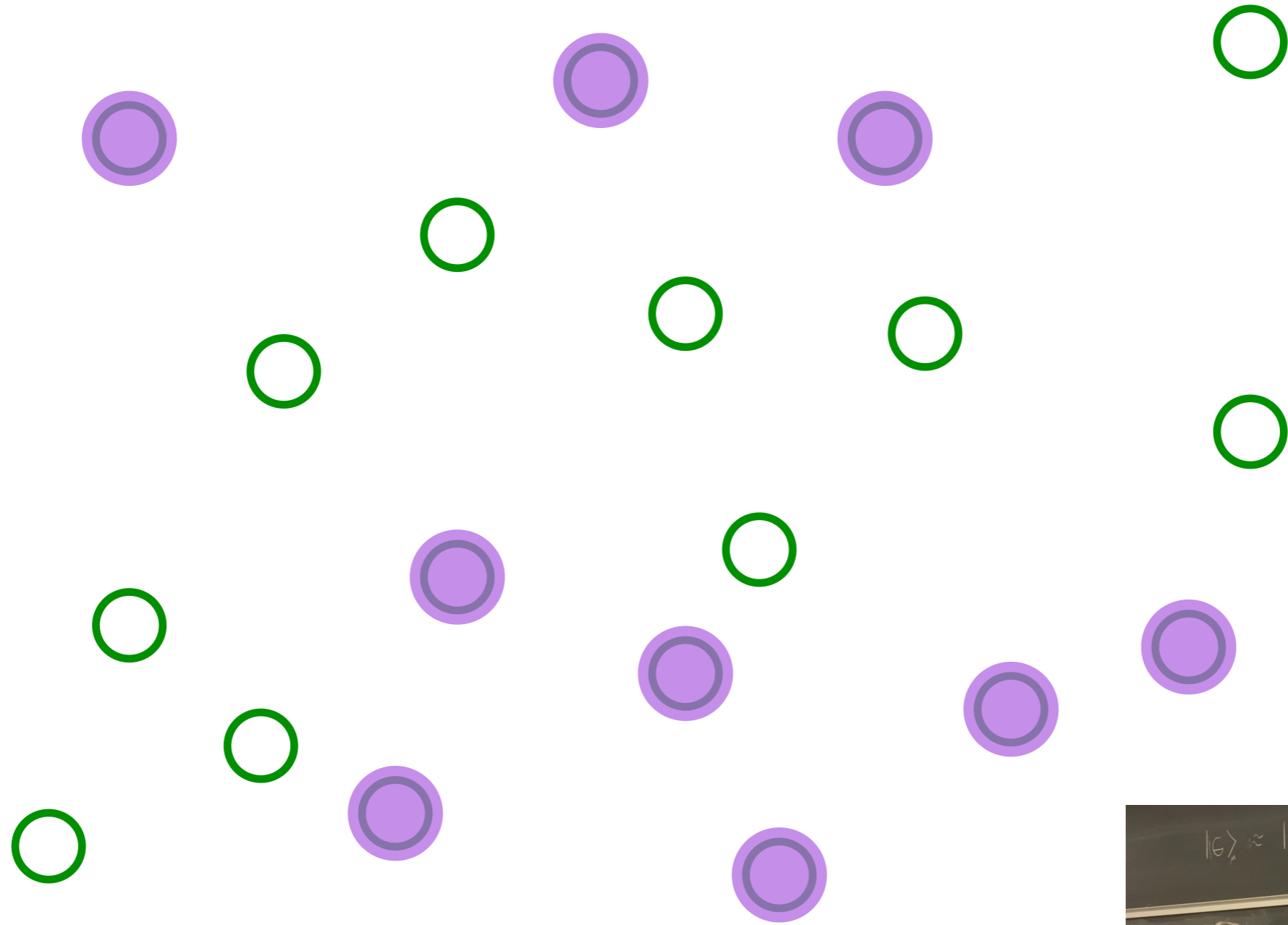
The SYK model



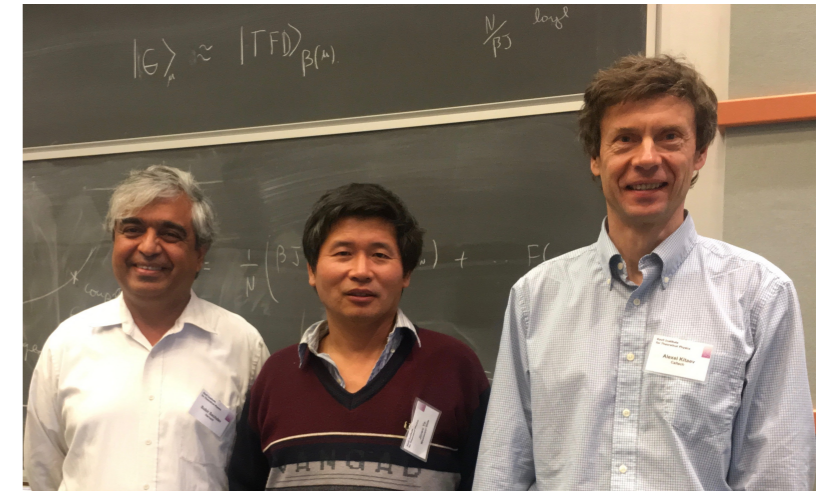
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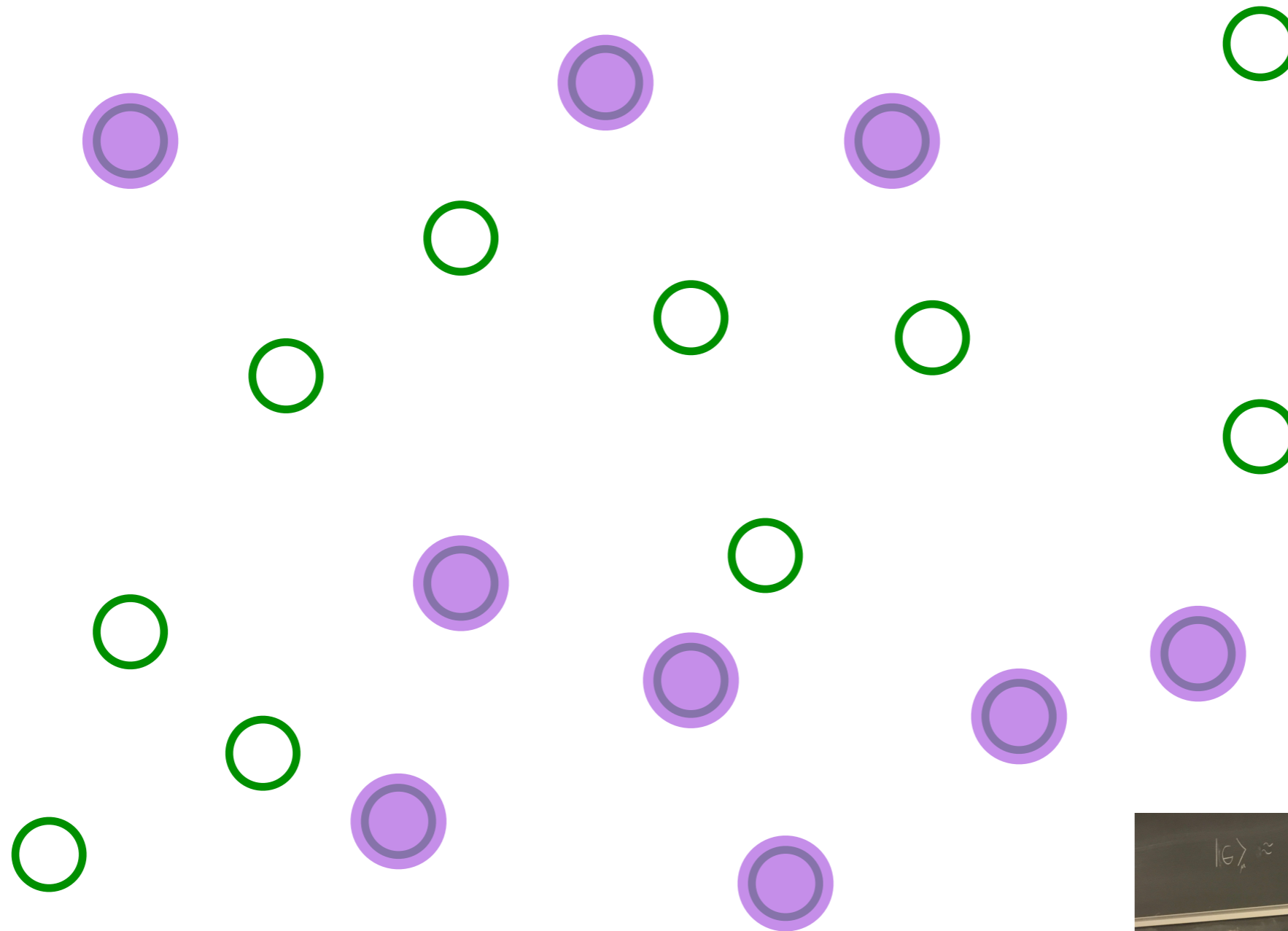
The SYK model



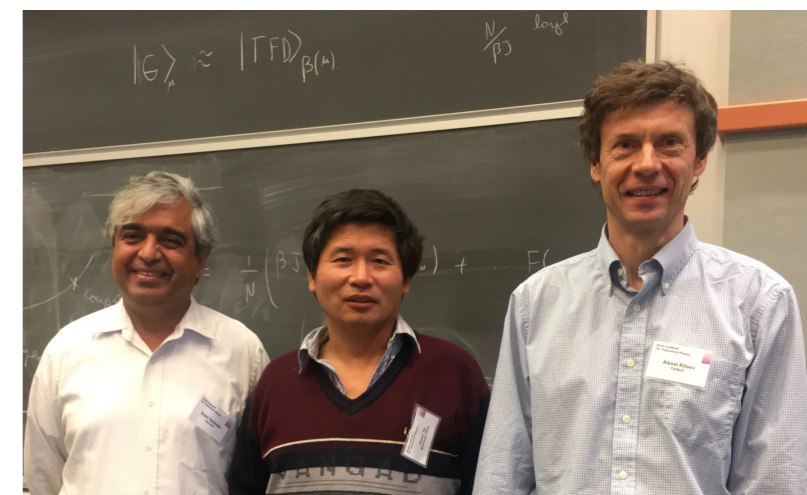
Entangle electrons pairwise randomly



The SYK model



This describes both a strange metal
and a black hole!



The SYK model

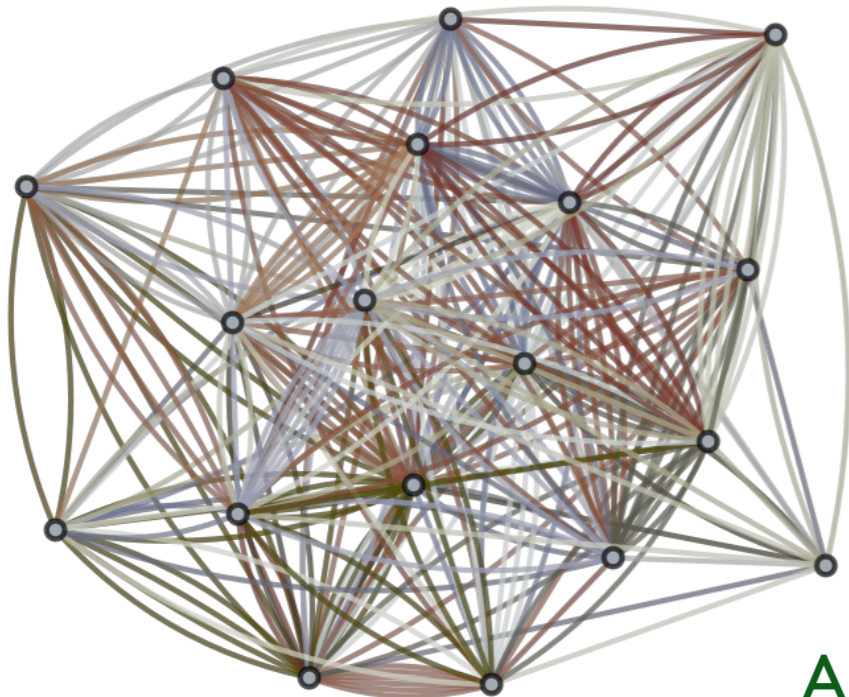
(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large N limit; T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N U_{ij;k\ell} c_i^\dagger c_j^\dagger c_k c_\ell - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$Q = \frac{1}{N} \sum_i c_i^\dagger c_i$$

$U_{ij;k\ell}$ are independent random variables with $\overline{U_{ij;k\ell}} = 0$ and $\overline{|U_{ij;k\ell}|^2} = U^2$
 $N \rightarrow \infty$ yields critical strange metal.

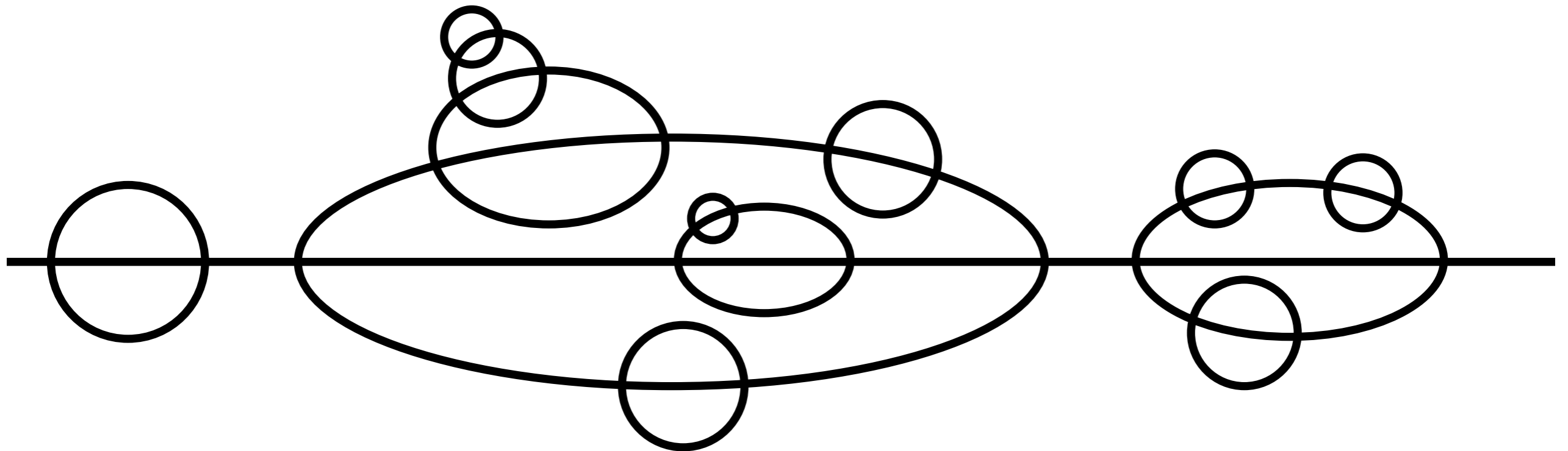


S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

The SYK model

The large N limit is given by the sum of “melon” Feynman graphs



S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

The SYK model

- Solution of the melon diagrams yields a Green's function at $T = 0$ of the form

$$G(\tau) \sim \begin{cases} -\tau^{-2\Delta} & \tau > 0 \\ e^{-2\pi\mathcal{E}}(-\tau)^{-2\Delta} & \tau < 0 \end{cases}, \quad T = 0$$

where $\Delta = 1/4$ is the fermion scaling dimension. The *particle-hole asymmetry* is determined by the parameter \mathcal{E} which universally depends only upon \mathcal{Q} .

- At $T > 0$ this has the conformal form

$$G(\tau) = -A \frac{e^{-2\pi\mathcal{E}T\tau}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left(\frac{T}{\sin(\pi T\tau)} \right)^{2\Delta}, \quad 0 < \tau < 1/T.$$

The SYK model



There are 2^N many body levels with energy E . Shown are all values of E for a single cluster of size $N = 12$. The $T \rightarrow 0$ state has an entropy $S_{GPS} = N s_0$, where $s_0 < \ln 2$ is determined by integrating

$$\frac{ds_0}{dQ} = 2\pi\mathcal{E}.$$

At $Q = 1/2$,

$$s_0 = \frac{G}{\pi} + \frac{\ln(2)}{4} = 0.464848\dots$$

where G is Catalan's constant.

GPS: A. Georges, O. Parcollet, and S. Sachdev, PRB **63**, 134406 (2001)

$\sim NU$

Many-body level spacing $\sim 2^{-N} = e^{-N \ln 2}$

Non-quasiparticle excitations with spacing $\sim e^{-N s_0}$

The SYK model

No quasiparticles



Julia Steinberg

- Rapid local thermal equilibration (of fermion correlators) in a ‘Planckian’ time

$$\tau_{\text{eq}} \sim \frac{\hbar}{k_B T} \quad , \quad \text{as } T \rightarrow 0.$$

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

A. Eberlein, V. Kasper, S. Sachdev, and
J. Steinberg, PRB **96**, 205123 (2017)

Established by solution of Schwinger-Keldysh equations for a quench.

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A. Eberlein, V. Kasper, S. Sachdev, and
J. Steinberg, PRB **96**, 205123 (2017)

Established by solution of Schwinger-Keldysh equations for a quench.

- Presence of quasiparticles should slow down thermalization, so *all* quantum systems obey

$$\tau_{\text{eq}} > C \frac{\hbar}{k_B T} \quad , \quad \text{as } T \rightarrow 0.$$

S. Sachdev, *Quantum Phase Transitions*,
Cambridge (1999)

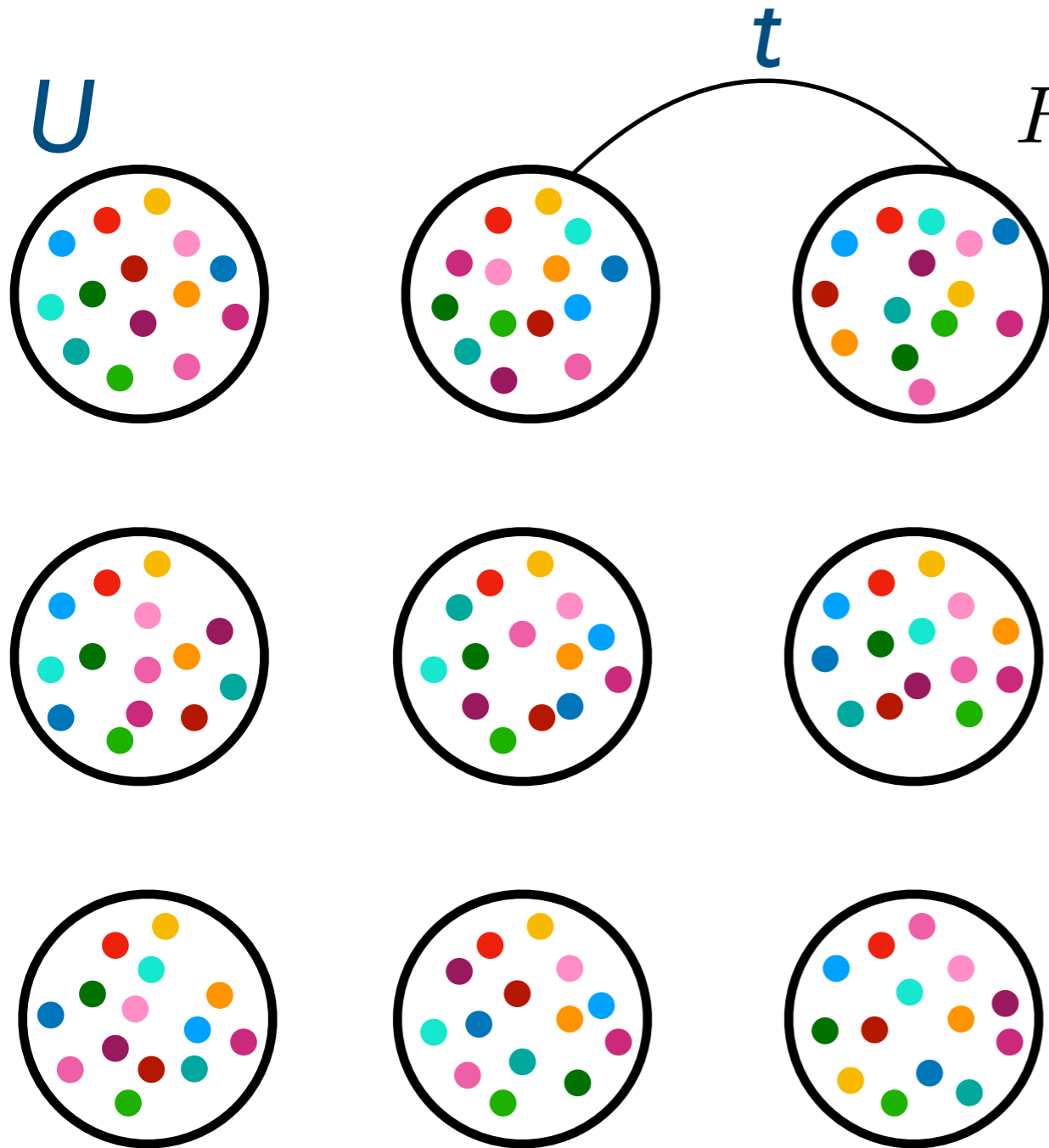
Absence of quasiparticles \Leftrightarrow Fastest possible thermalization



Coupled SYK Islands



SYK quantum islands of electrons with random or regular hopping between them.



$$H = \sum_x \sum_{i < j, k < l} U_{ijkl,x} c_{ix}^\dagger c_{jx}^\dagger c_{kx} c_{lx} + \sum_{\langle xx' \rangle} \sum_{i,j} t_{ij,xx'} c_{i,x}^\dagger c_{j,x'}$$

$$\overline{|U_{ijkl}|^2} = \frac{2U^2}{N^3}$$

$$\overline{|t_{ij,xx'}|^2} = t_0^2/N$$

Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

Pengfei Zhang, PRB **96**, 205138 (2017)

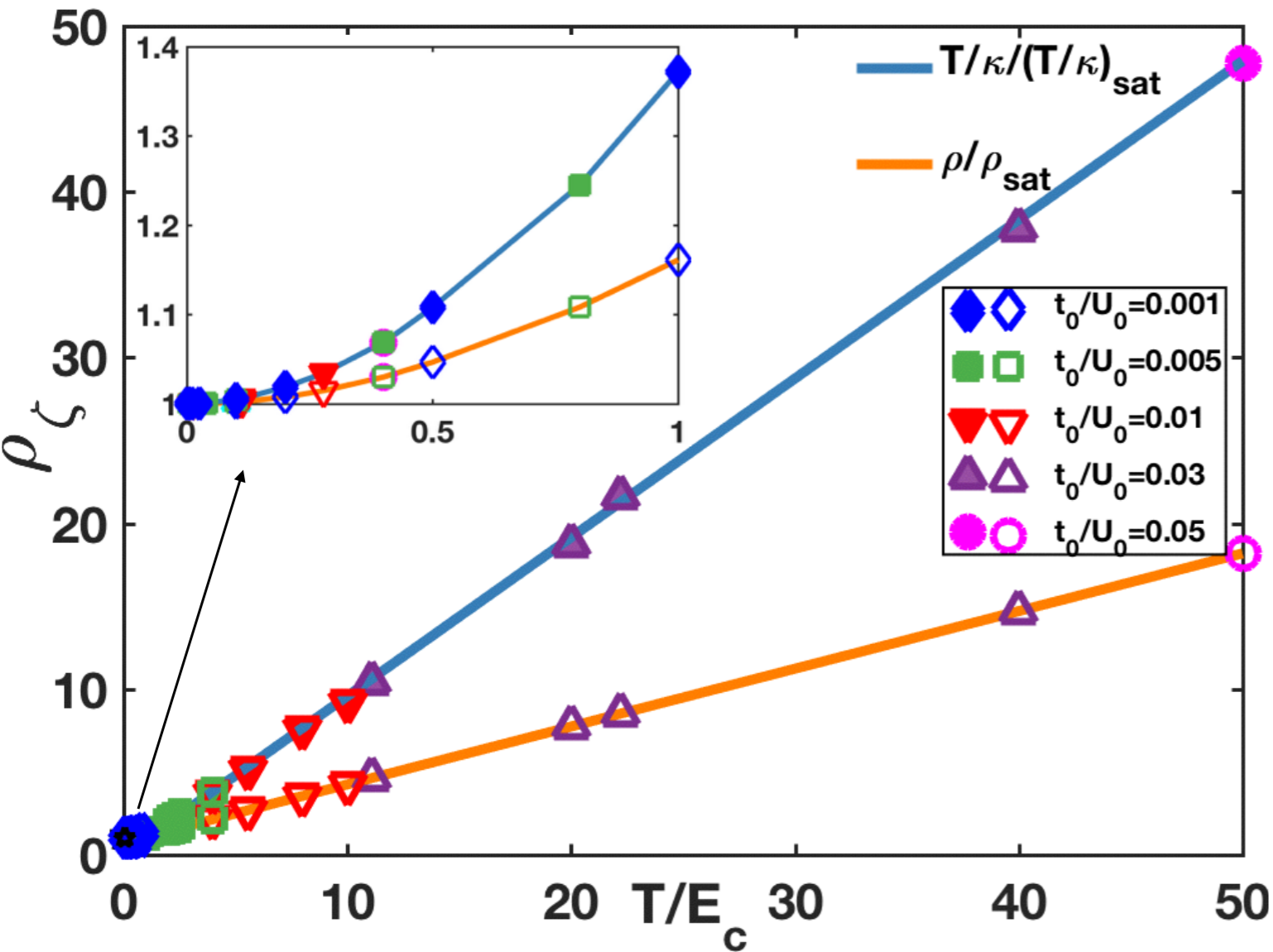
Debanjan Chowdhury, Yochai Werman, Erez Berg, T. Senthil, PRX **8**, 021049 (2018)

Aavishkar A. Patel, John McGreevy, Daniel P. Arovas, Subir Sachdev, PRX **8**, 021049 (2018)

See also A. Georges and O. Parcollet PRB **59**, 5341 (1999)

Coupled SYK Islands

Low 'coherence' scale



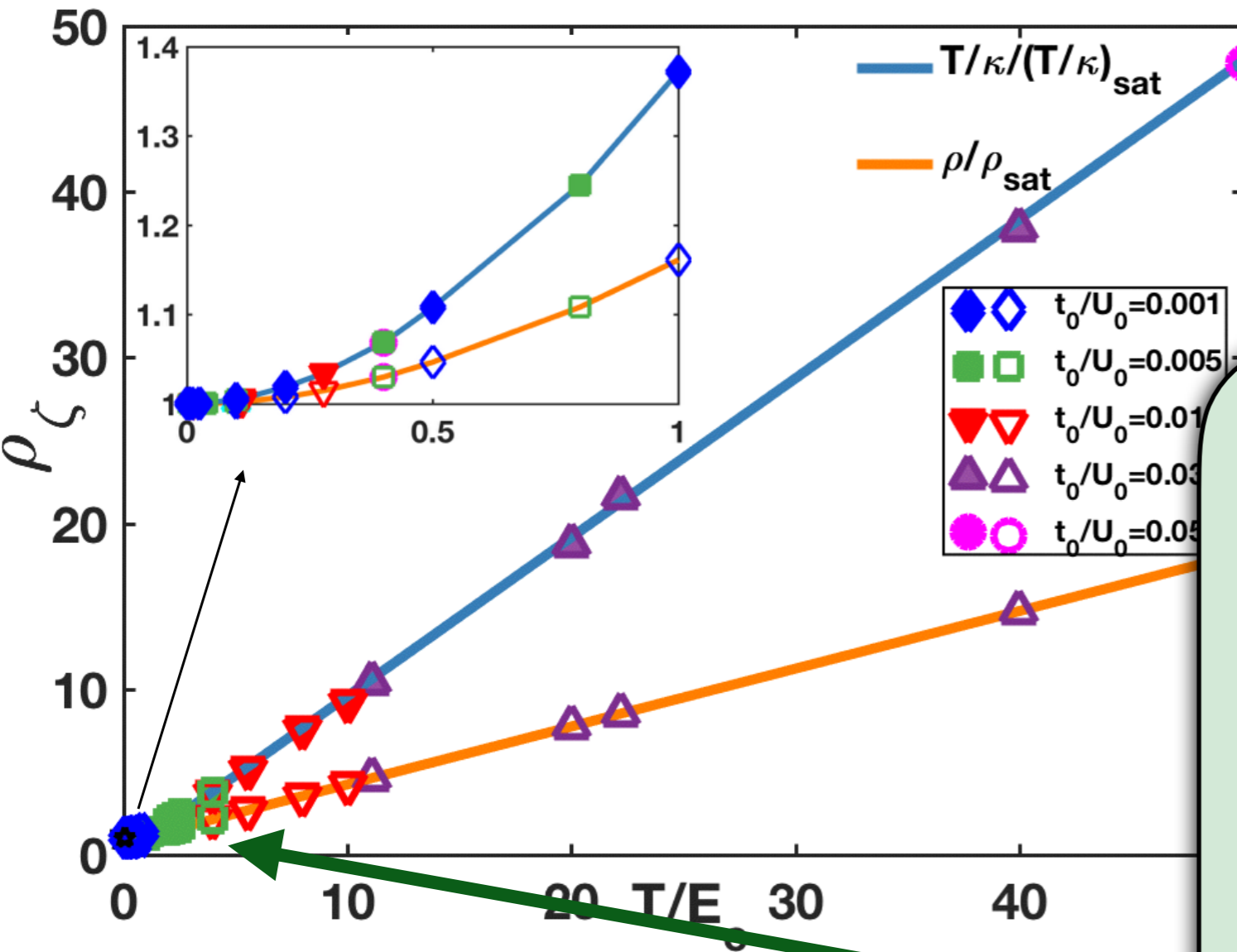
$$E_c \sim \frac{t_0^2}{U}$$

Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

See also A. Georges and O. Parcollet PRB **59**, 5341 (1999)

Coupled SYK Islands

Low 'coherence' scale



$$E_c \sim \frac{t_0^2}{U}$$

For $T < E_c$, the resistivity, ρ , and entropy density, s , are

$$\rho = \frac{h}{e^2} \left[c_1 + c_2 \left(\frac{T}{E_c} \right)^2 \right]$$

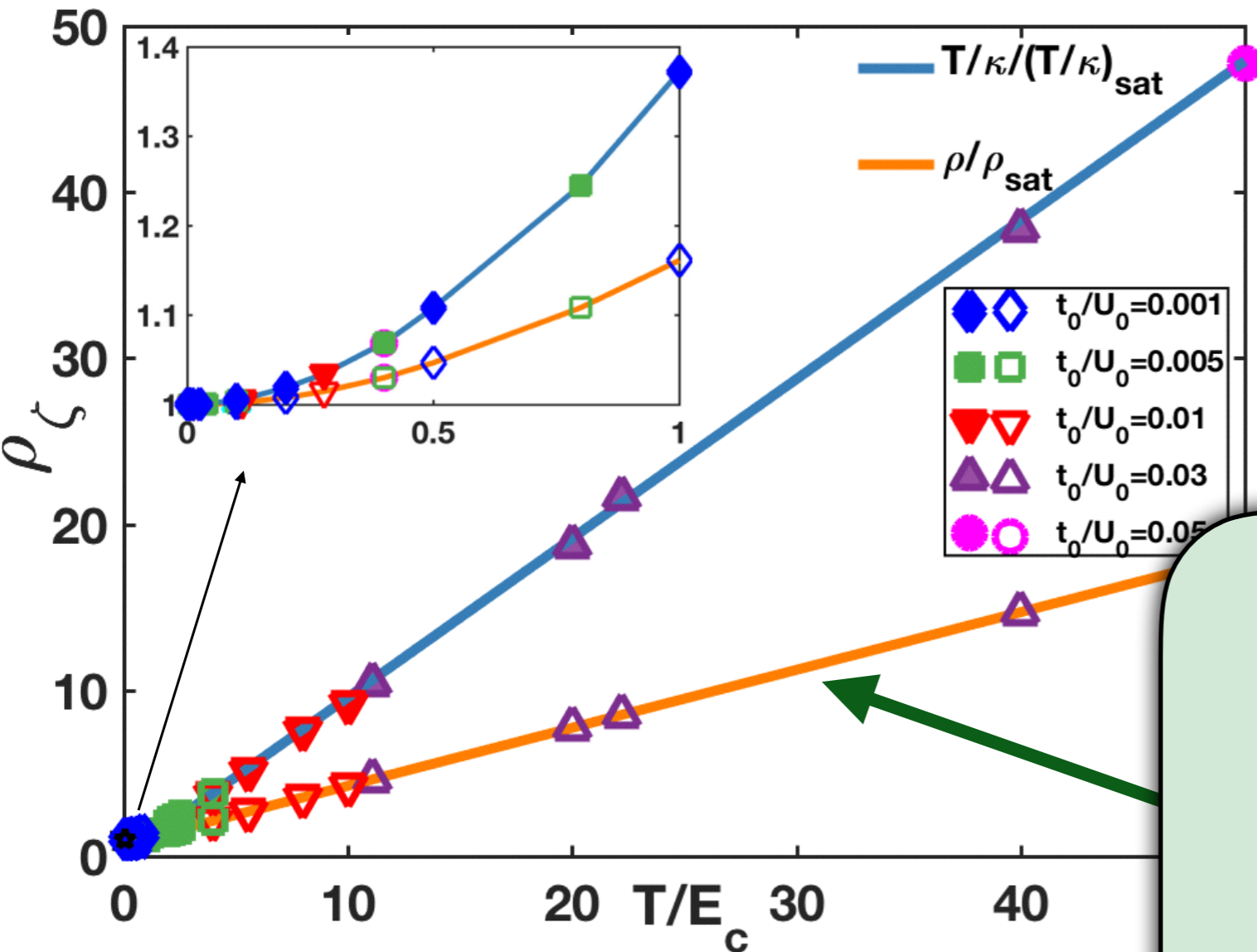
$$s \sim s_0 \left(\frac{T}{E_c} \right)$$

Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

See also A. Georges and O. Parcollet PRB **59**, 5341 (1999)

Coupled SYK Islands

Low 'coherence' scale



$$E_c \sim \frac{t_0^2}{U}$$

For $E_c < T < U$, the resistivity, ρ , and entropy density, s , are

$$\rho \sim \frac{h}{e^2} \left(\frac{T}{E_c} \right), \quad s = s_0$$

Xue-Yang Song, Chao-Ming Jian, and L. Balents, PRL **119**, 216601 (2017)

See also A. Georges and O. Parcollet PRB **59**, 5341 (1999)

I. Normal and strange metals

(a) with quasiparticles (normal):

random matrix model

(b) without quasiparticles (strange):

the complex SYK model

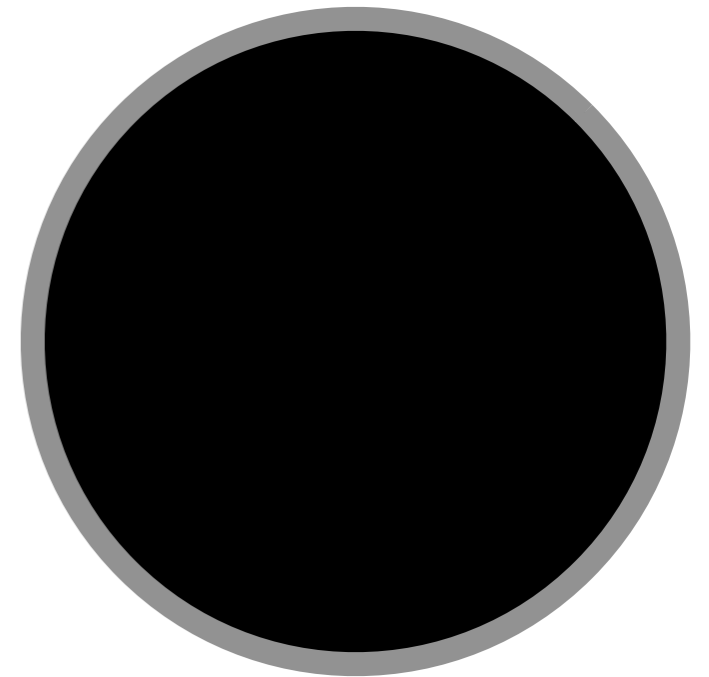
2. Einstein-Maxwell theory of charged black holes in AdS space

3. Fluctuations, and the Schwarzian

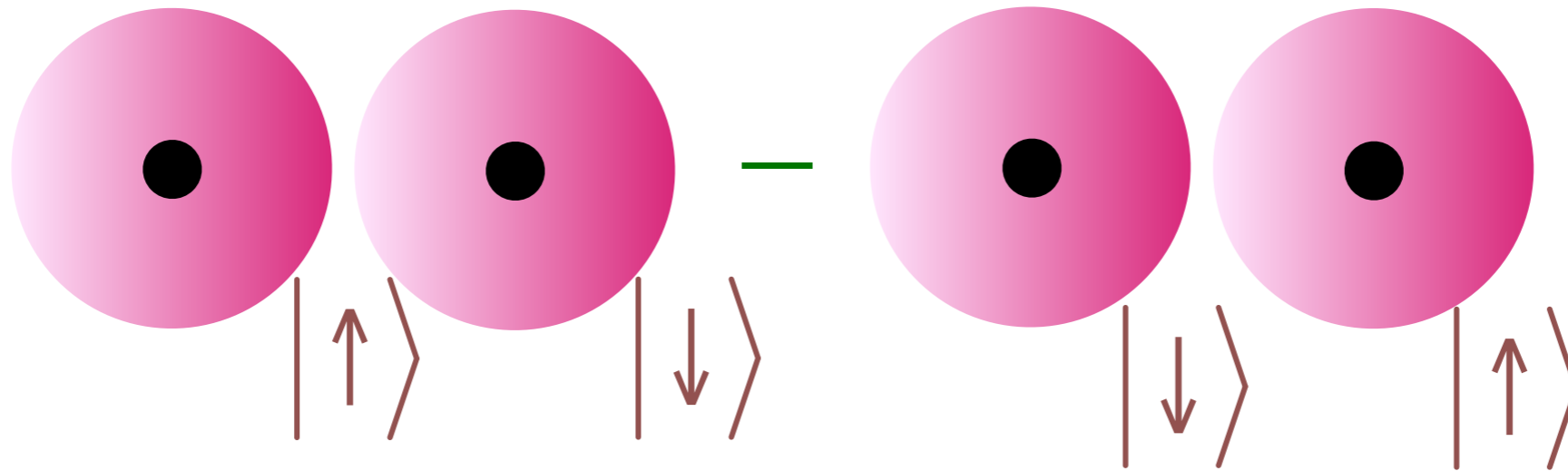
Black Holes

Objects so dense that light is gravitationally bound to them.

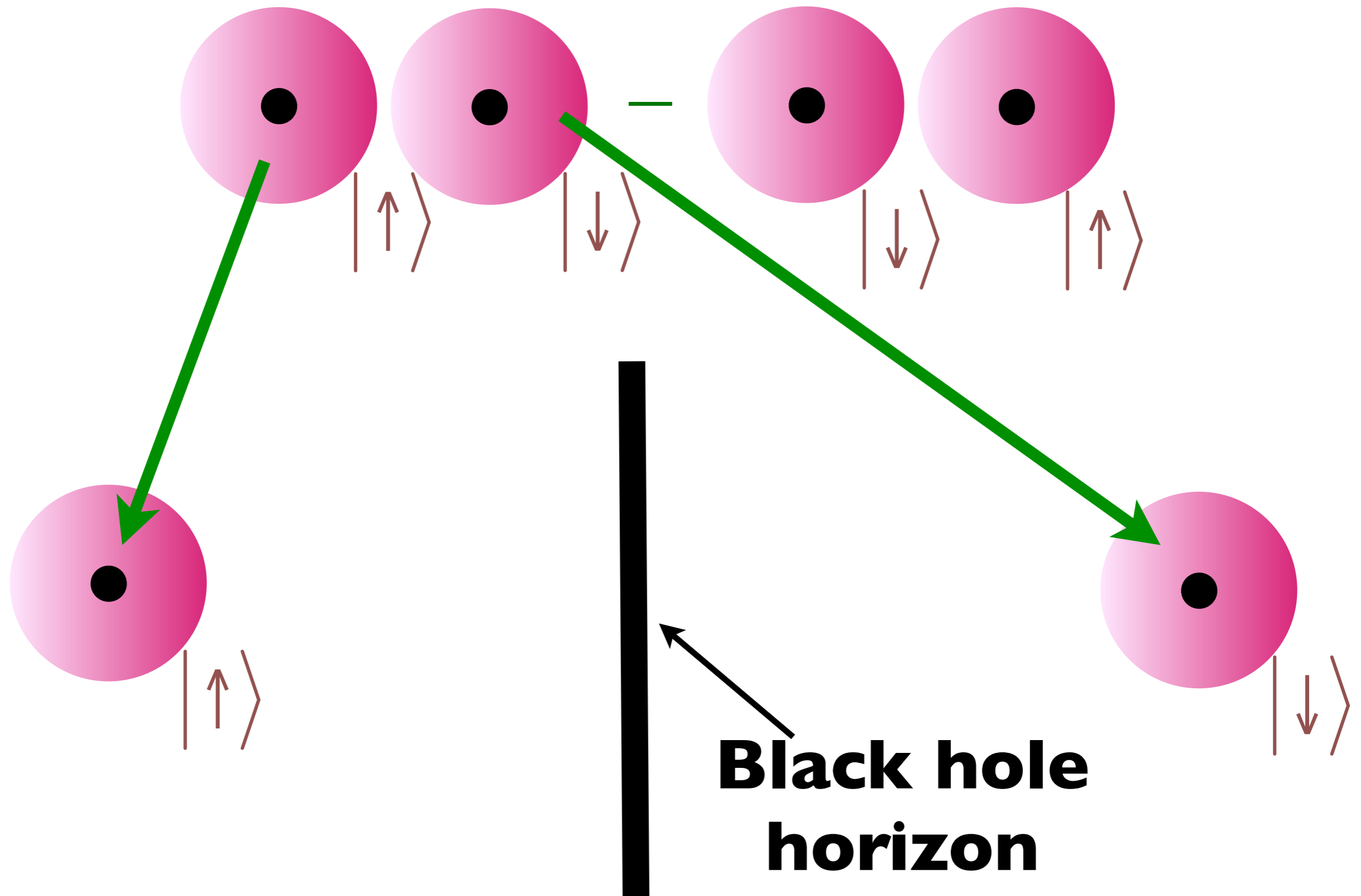
In Einstein's theory, the region inside the black hole **horizon** is disconnected from the rest of the universe.



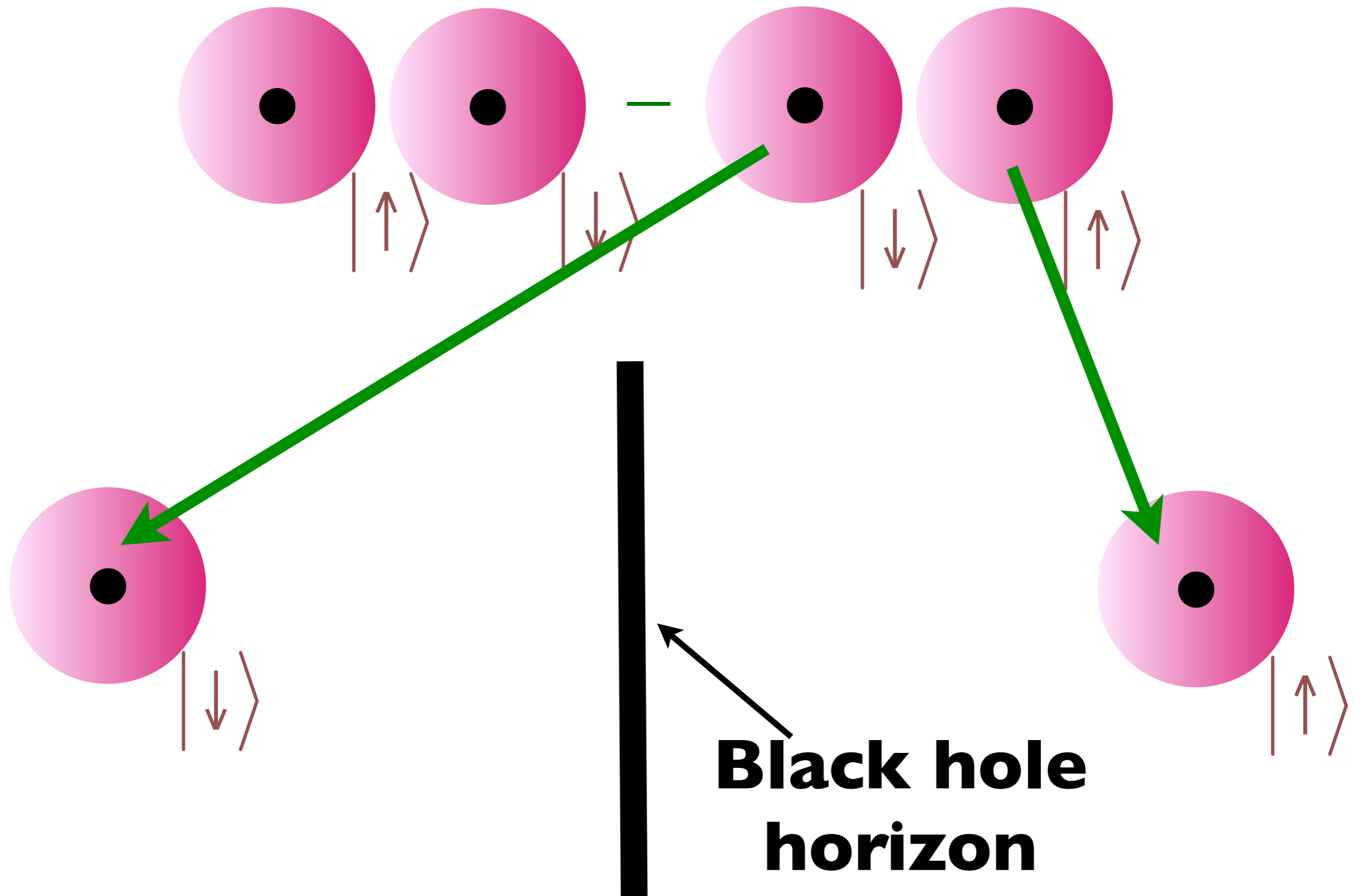
Quantum Entanglement across a black hole horizon



Quantum Entanglement across a black hole horizon

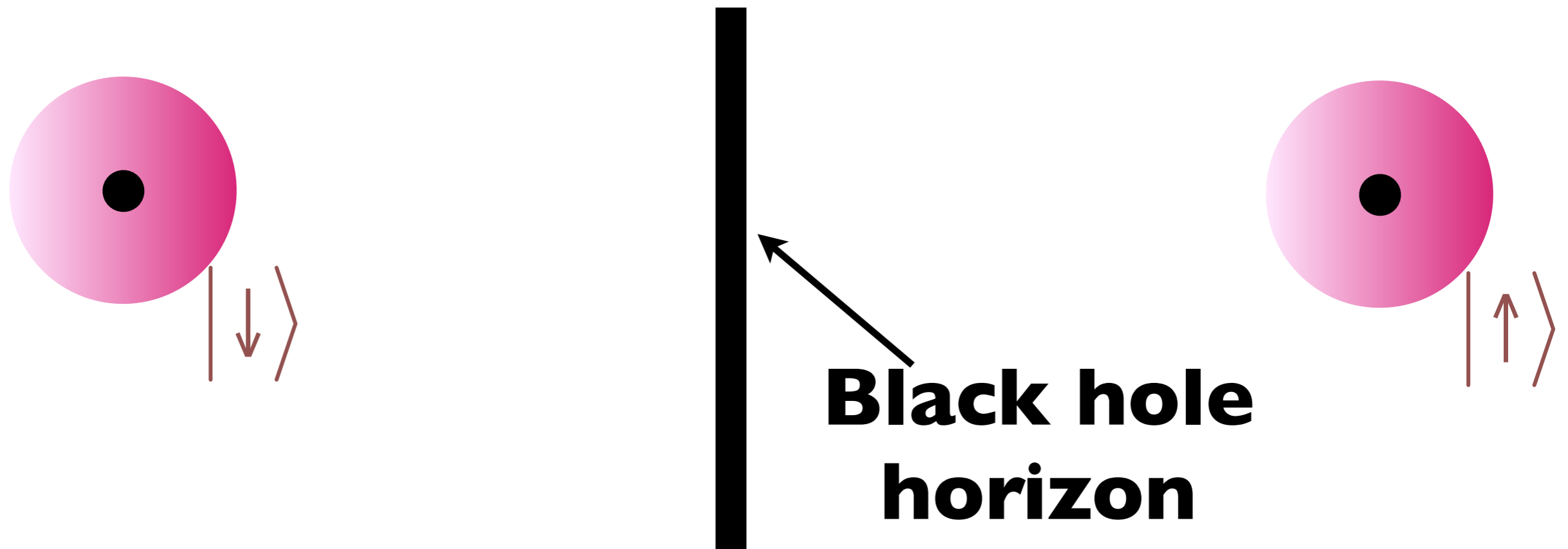


Quantum Entanglement across a black hole horizon



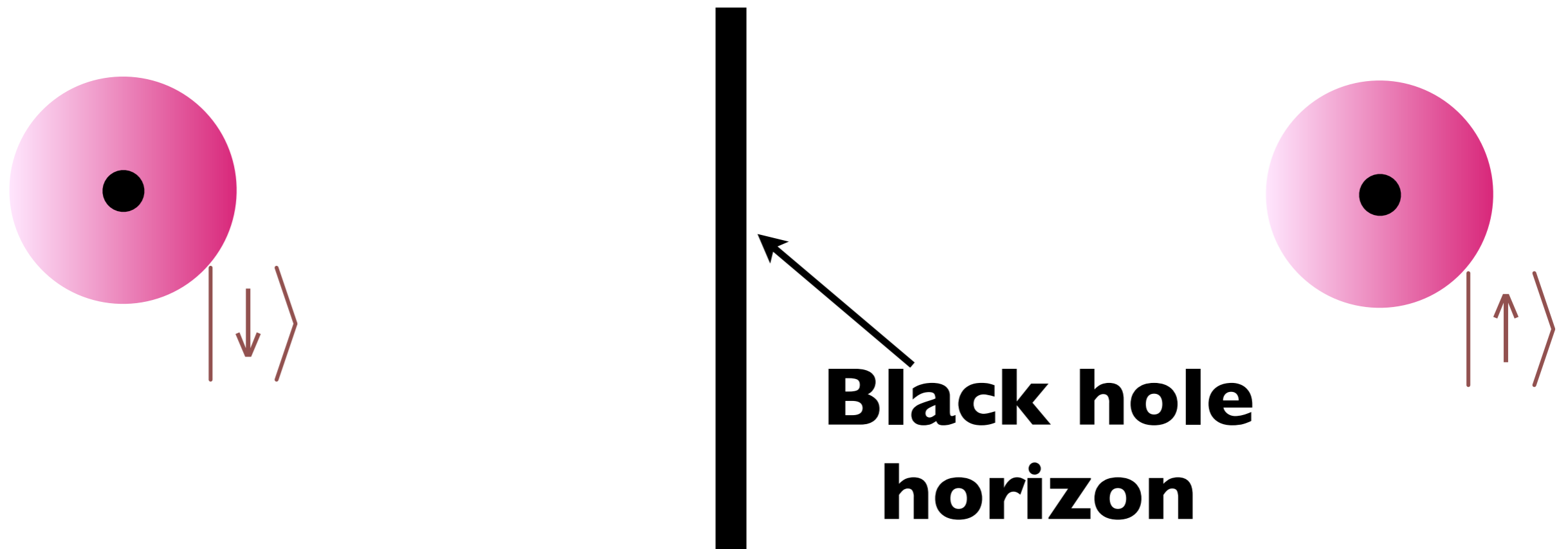
Quantum Entanglement across a black hole horizon

There is quantum entanglement between the inside and outside of a black hole



Quantum Entanglement across a black hole horizon

Hawking used this to show that black hole horizons have an entropy and a temperature (because to an outside observer, the state of the electron inside the black hole is an unknown)

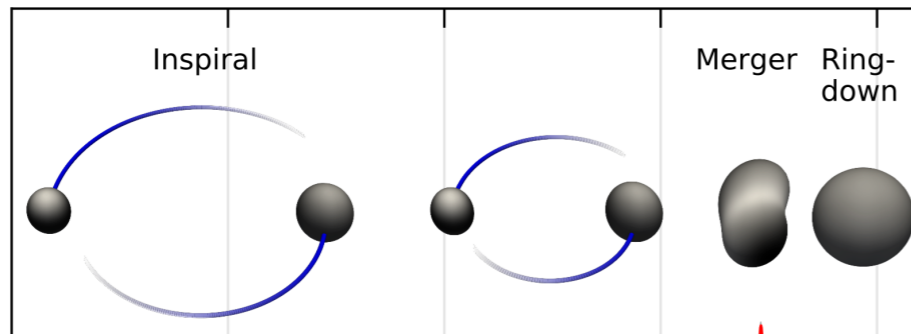
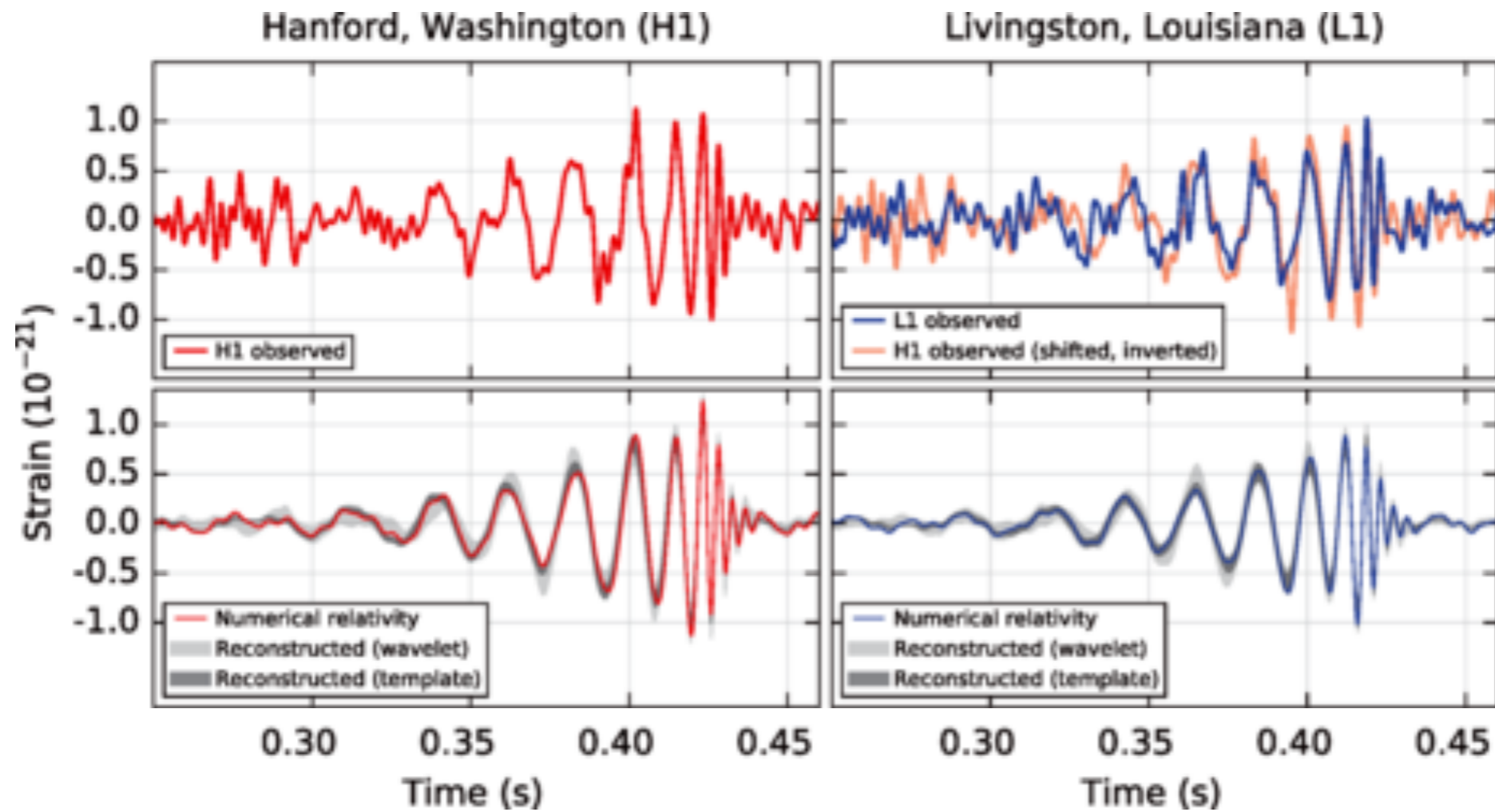


Black holes

- Black holes have an entropy and a temperature, T_H
- The entropy is proportional to their surface area.

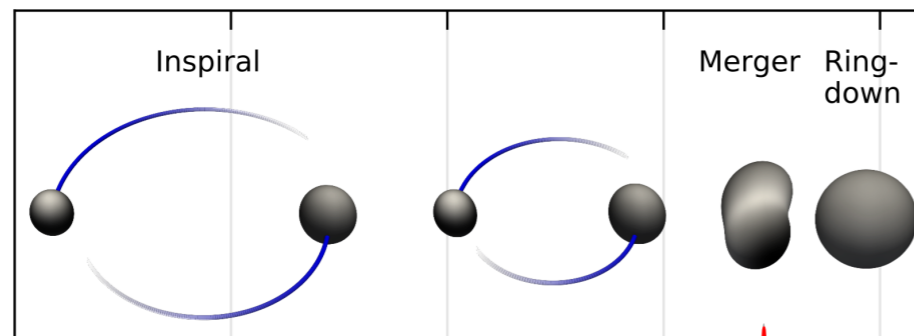
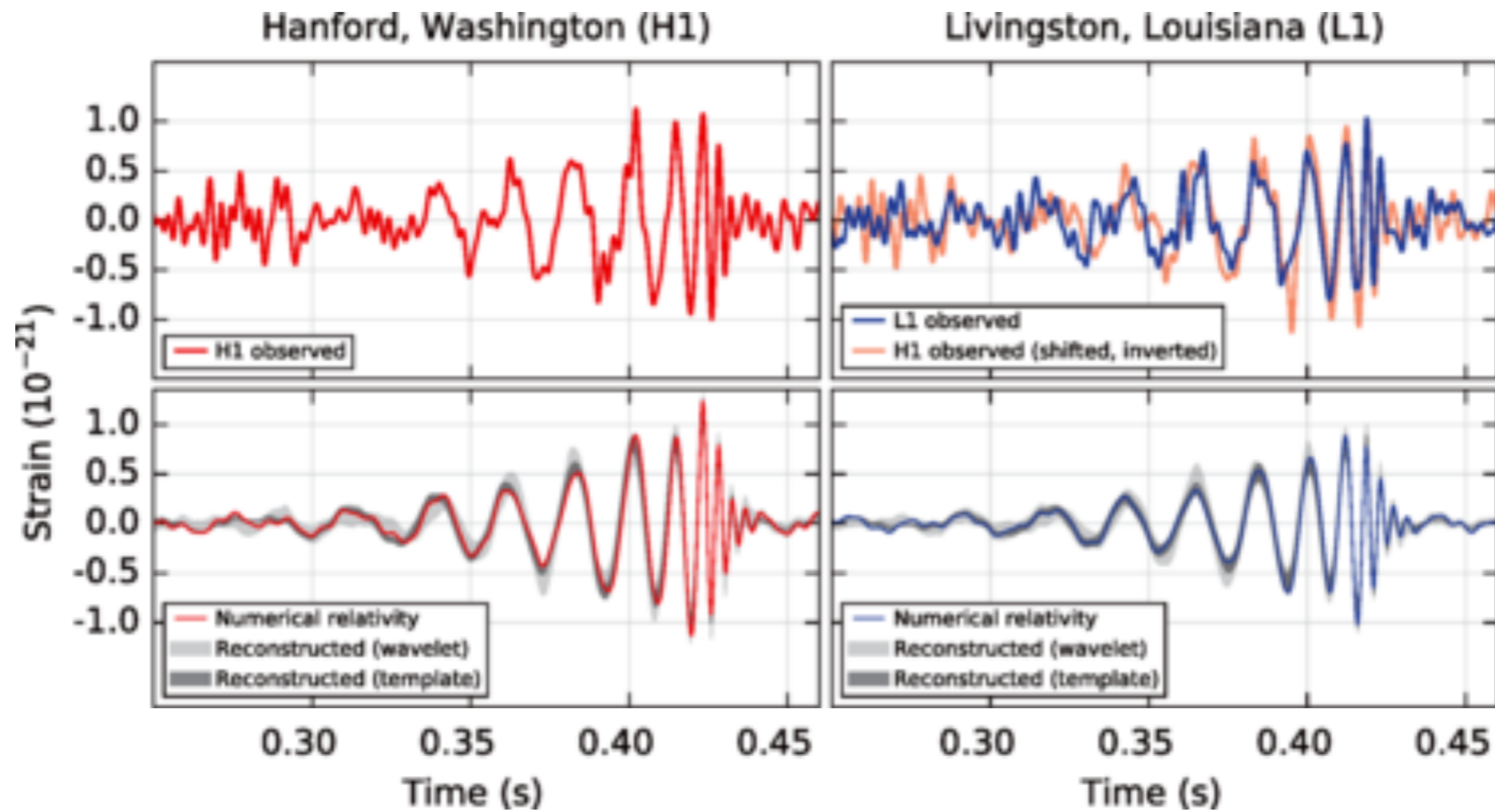
J. D. Bekenstein, PRD **7**, 2333 (1973)
S.W. Hawking, Nature **248**, 30 (1974)





LIGO
September 14, 2015

- The ring-down is predicted by General Relativity to happen in a time $\frac{8\pi GM}{c^3} \sim 8$ milliseconds.



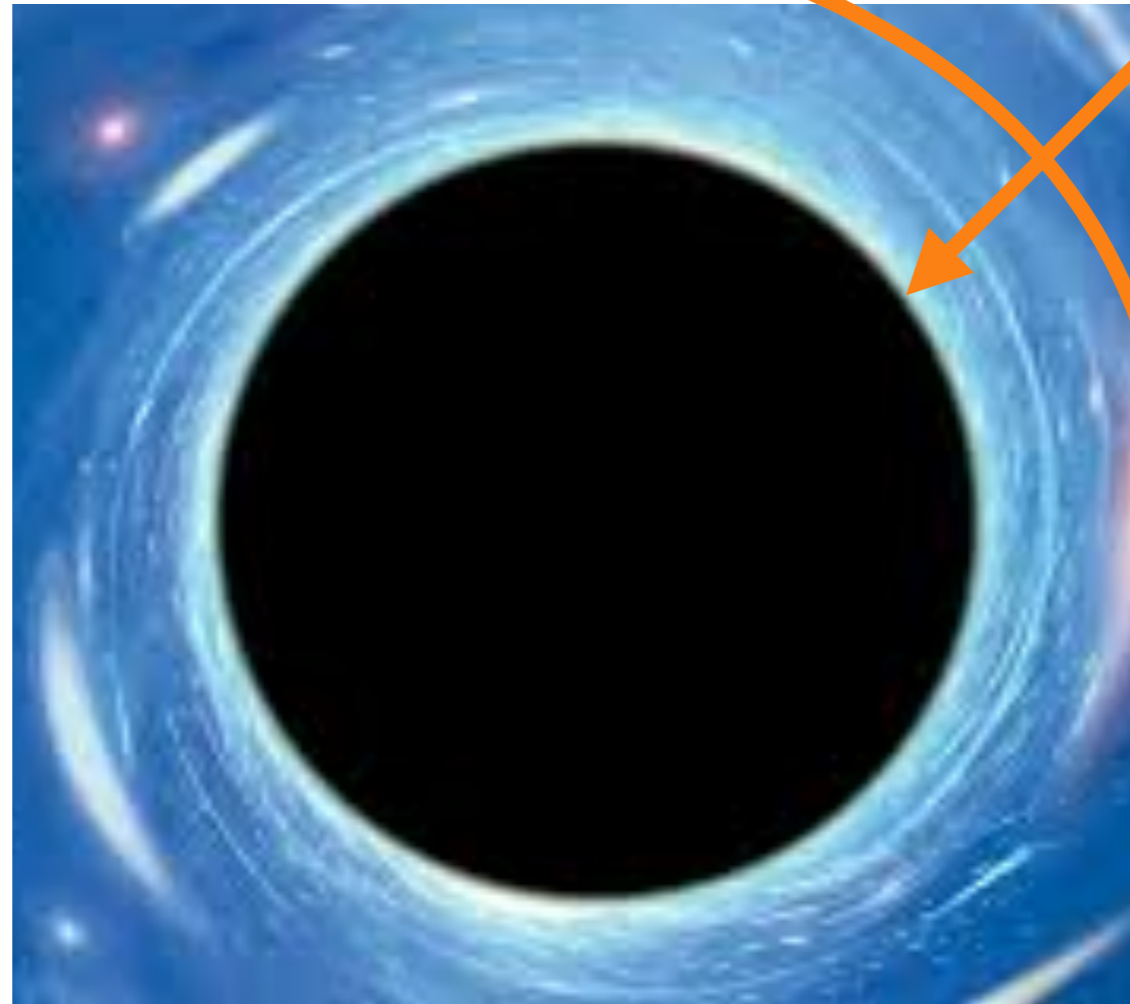
LIGO
September 14, 2015

- The ring-down is predicted by General Relativity to happen in a time $\frac{8\pi GM}{c^3} \sim 8$ milliseconds. Curiously this happens to equal $\frac{\hbar}{k_B T_H}$; so the ring down can also be viewed as the approach of a quantum system to thermal equilibrium at the fastest possible rate!

Black holes

- Black holes have an entropy and a temperature, T_H
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a Planckian time $\sim \hbar/(k_B T_H)$.





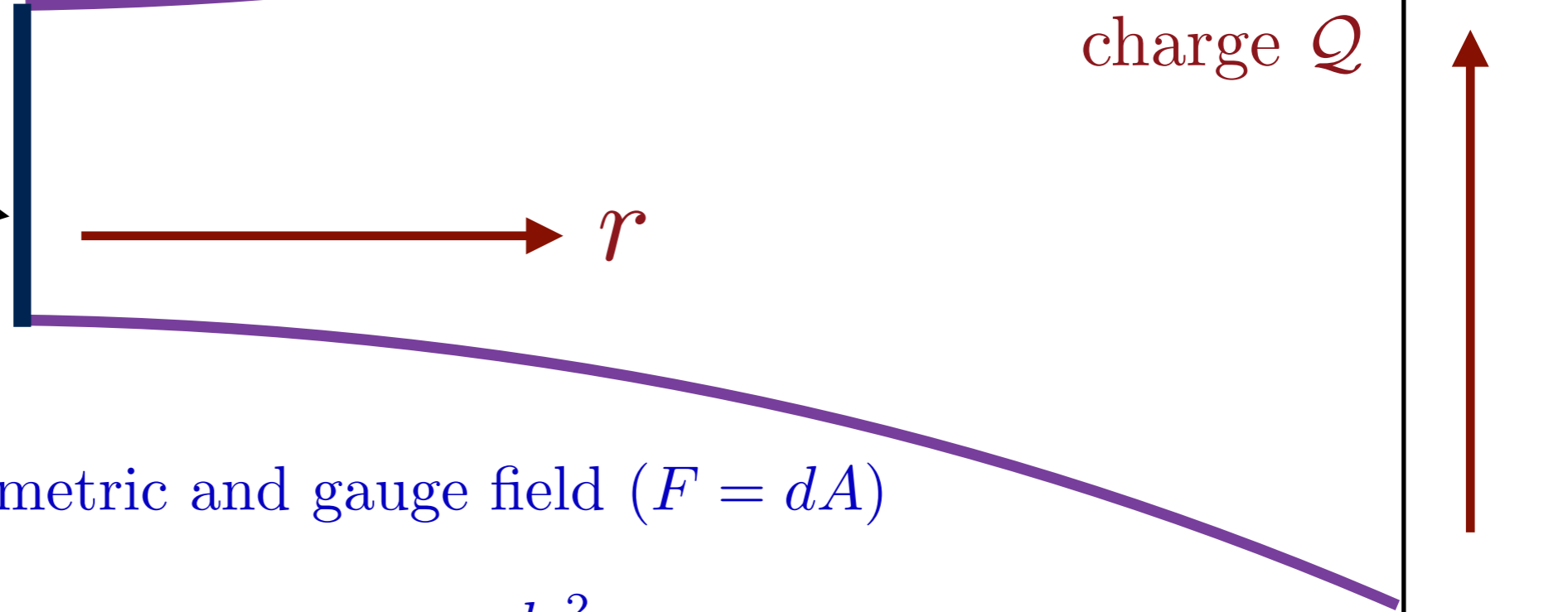
Zooming into the near-horizon region of a charged black hole, zooming at low temperature, yields a quantum theory in one space (ζ) and one time dimension



Charged black holes

$$S_{EM} = \frac{1}{2\kappa^2} \int d^4x \sqrt{-g} \left(\mathcal{R} + \frac{6}{L^2} - \frac{L^2}{g_F^2} F^2 \right)$$

Black hole horizon of radius r_0



Solutions of S_{EM} have metric and gauge field ($F = dA$)

$$ds^2 = -V(r)dt^2 + r^2 d\Omega_2^2 + \frac{dr^2}{V(r)} \quad , \quad A = \mu \left(1 - \frac{r_0}{r} \right) dt$$

$$V(r) = 1 + \frac{r^2}{L^2} + \frac{\Theta^2}{r^2} - \frac{M}{r}$$

where $d\Omega_2^2$ is the metric of the 2-sphere. All parameters of the solution, and the thermodynamics are determined in terms of the chemical potential μ , and the Hawking temperature of horizon, T .

Charged black holes

In the $T \rightarrow 0$ limit, at fixed μ , we obtain a charged black hole solution with radius $r_0(T \rightarrow 0, \mu) = R_h$. All properties of this black hole can be expressed in terms of R_h

- In the near-horizon region, we change co-ordinates from r to ζ so that

$$r - R_h = \frac{R_2^2}{\zeta} \quad , \quad R_2 = \frac{LR_h}{\sqrt{6R_h^2 + L^2}}.$$

Then the near-horizon metric becomes $\text{AdS}_2 \times \text{S}_2$, with

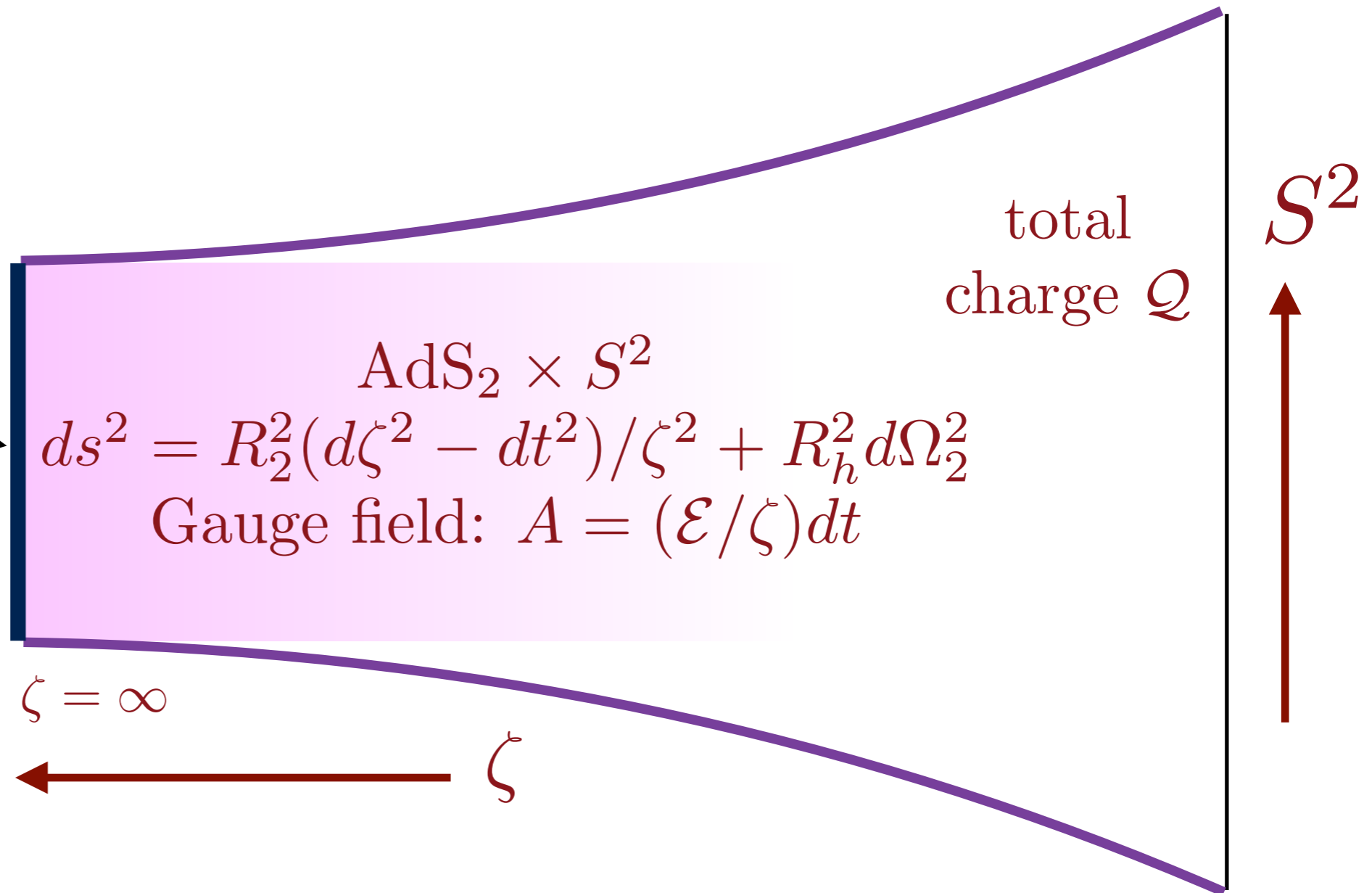
$$ds^2 = R_2^2 \left[\frac{-dt^2 + d\zeta^2}{\zeta^2} \right] + R_h^2 d\Omega_2^2 \quad , \quad A = \frac{\mathcal{E}}{\zeta} dt.$$

where the dimensionless electric field \mathcal{E} is

$$\mathcal{E} = \frac{g_F R_h \sqrt{3R_h^2 + L^2}}{6R_h^2 + L^2}.$$

Charged black holes

Black hole horizon of radius R_h and entropy s_0



- The entropy s_0 , the charge Q , and the dimensionless electric field \mathcal{E} obey

$$\frac{ds_0}{dQ} = 2\pi\mathcal{E}$$

Charged black holes

Probe fermion in the AdS₂ near horizon

- A probe fermion has a near-horizon Green's function at $T = 0$ of the form

$$G(\tau) \sim \begin{cases} -\tau^{-2\Delta} & \tau > 0 \\ e^{-2\pi\mathcal{E}}(-\tau)^{-2\Delta} & \tau < 0 \end{cases}, \quad T = 0$$

where Δ is the fermion scaling dimension. The *particle-hole asymmetry* is determined by the dimensionless electric field \mathcal{E} at the surface of the black hole.

- At $T > 0$ this has the conformal form

$$G(\tau) = -A \frac{e^{-2\pi\mathcal{E}T\tau}}{\sqrt{1 + e^{-4\pi\mathcal{E}}}} \left(\frac{T}{\sin(\pi T\tau)} \right)^{2\Delta}, \quad 0 < \tau < 1/T.$$

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random matrix model

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the complex SYK model

2. Einstein-Maxwell theory of charged black holes in AdS space

3. Fluctuations, and the Schwarzian

The SYK model

The theory of fluctuations about the large N saddle point are expressed in terms of a bilocal time Green's function, $G(\tau_1, \tau_2)$. This theory is invariant at low energies under a time reparameterization f , and an emergent gauge transformation g

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1)f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{G}(\sigma_1, \sigma_2)$$

where $f(\sigma)$ and $g(\sigma)$ are arbitrary functions.

The SYK model

The large N saddle point is

$$G_s(\tau_1 - \tau_2) \sim (\tau_1 - \tau_2)^{-1/2}.$$

The saddle point will be invariant under a reparamaterization $f(\tau)$ when choosing $G(\tau_1, \tau_2) = G_s(\tau_1 - \tau_2)$ leads to a transformed $\tilde{G}(\sigma_1, \sigma_2) = G_s(\sigma_1 - \sigma_2)$ (and similarly for Σ). It turns out this is true only for the $SL(2, \mathbb{R})$ transformations under which

$$f(\tau) = \frac{a\tau + b}{c\tau + d}, \quad ad - bc = 1.$$

So the (approximate) reparametrization symmetry is spontaneously broken down to $SL(2, \mathbb{R})$ by the saddle point.

The Schwarzian theory of the SYK model

The effective theory of time reparameterizations $f(\tau)$ broken down to $SL(2, \mathbb{R})$, and gauge transformations $g(\tau) = e^{i\phi(\tau)}$ is

$$S_{\text{eff}}[f, \phi] = \frac{K}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi \mathcal{E} T) \partial_\tau f)^2 - \frac{\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

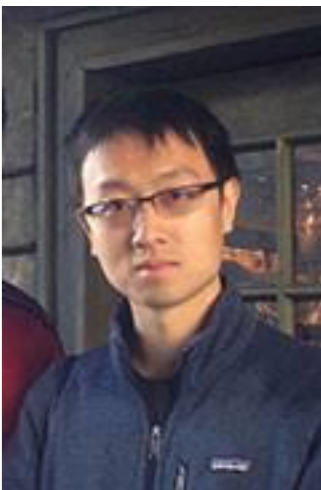
where $f(\tau)$ is a monotonic map from $[0, 1/T]$ to $[0, 1/T]$, the couplings K , γ , and \mathcal{E} can be related to thermodynamic derivatives and we have used the Schwarzian:

$$\{g, \tau\} \equiv \frac{g'''}{g'} - \frac{3}{2} \left(\frac{g''}{g'} \right)^2.$$

Specifically, an argument constraining the effective action at $T = 0$ is

$$S_{\text{eff}} \left[f(\tau) = \frac{a\tau + b}{c\tau + d}, \phi(\tau) = 0 \right] = 0,$$

and this is origin of the Schwarzian.

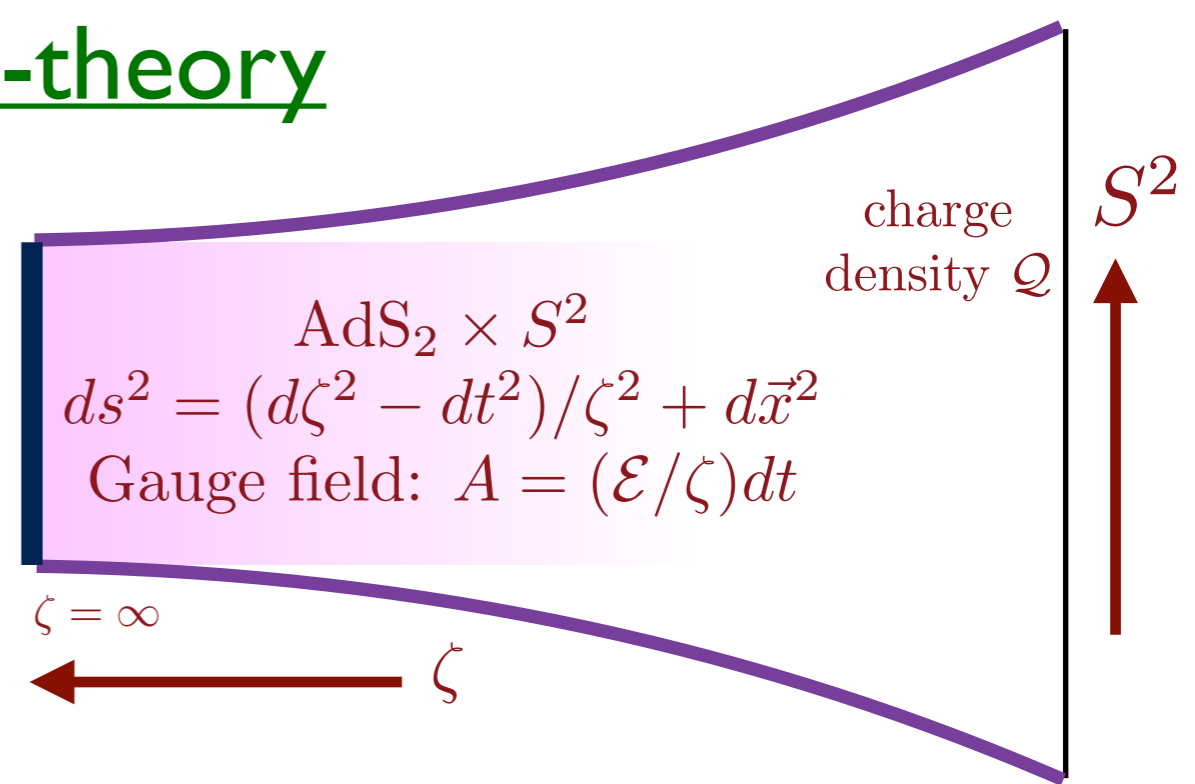


Yingfei Gu

J. Maldacena and D. Stanford, arXiv:1604.07818;
R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, arXiv:1612.00849;
A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746;
Yingfei Gu and S. Sachdev, unpublished



Einstein-Maxwell-theory



$$S_{4D} = \int d^4x \sqrt{-\hat{g}} \left(\hat{\mathcal{R}} + 6/L^2 - \frac{1}{4} \hat{F}_{\mu\nu} \hat{F}^{\mu\nu} \right),$$

- Near-extremal black hole of radius R and $T \ll 1/R$ from 4D Einstein-Maxwell
- \implies 2D Einstein-Maxwell with a scalar field (fluctuations of black hole radius)
- \implies Absence of 2D quantum gravity fluctuations: reduction to (0+1)D
- \implies (0 + 1)D theory is precisely the Schwarzian theory of the SYK model.

$$S_{\text{eff}}[f, \phi] = \frac{K}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi \mathcal{E} T) \partial_\tau f)^2 - \frac{\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857; K. Jensen, arXiv:1605.06098; J. Engelsoy, T.G. Mertens, and H. Verlinde, arXiv:1606.03438; P. Nayak, A. Shukla, R.M. Soni, S.P. Trivedi, and V. Vishal, arXiv:1802.09547; U. Moitra, S.P. Trivedi, and V. Vishal, arXiv:1808.08239; P. Chaturvedi, Yingfei Gu, Wei Song, Boyang Yu, arXiv:1808.08062; A. Gaikwad, L.K. Joshi, G. Mandal, and S.R. Wadia, arXiv:1802.07746

SYK models and black holes

- Reparameterization and gauge invariance are defining properties of Einstein-Maxwell theory
- In imaginary time, AdS_2 is the homogeneous hyperbolic space: two-dimensional surface of constant negative curvature. Its metric is invariant under $\text{SL}(2, \mathbb{R})$

$ds^2 = (d\tau^2 + d\zeta^2)/\zeta^2$ is invariant under

$$\tau' + i\zeta' = \frac{a(\tau + i\zeta) + b}{c(\tau + i\zeta) + d} \quad \text{with } ad - bc = 1.$$



Quantum matter without quasiparticles

- Planckian dynamics (*i.e.* fastest possible local thermalization in a time $\hbar/(k_B T)$) is realized in the ‘solvable’ SYK models.

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 - the SYK models
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- Lattices of SYK islands have led to a partial understanding of strange metals.