

Bekenstein-Hawking entropy and strange metals

CMSA Colloquium
Harvard University
September 16, 2015

Subir Sachdev

Talk online: sachdev.physics.harvard.edu



Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Infinite-range model of a Fermi liquid

$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j + \dots$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$\frac{1}{N} \sum_i c_i^\dagger c_i = Q$$

t_{ij} are independent random variables with $\overline{t_{ij}} = 0$ and $\overline{|t_{ij}|^2} = t^2$

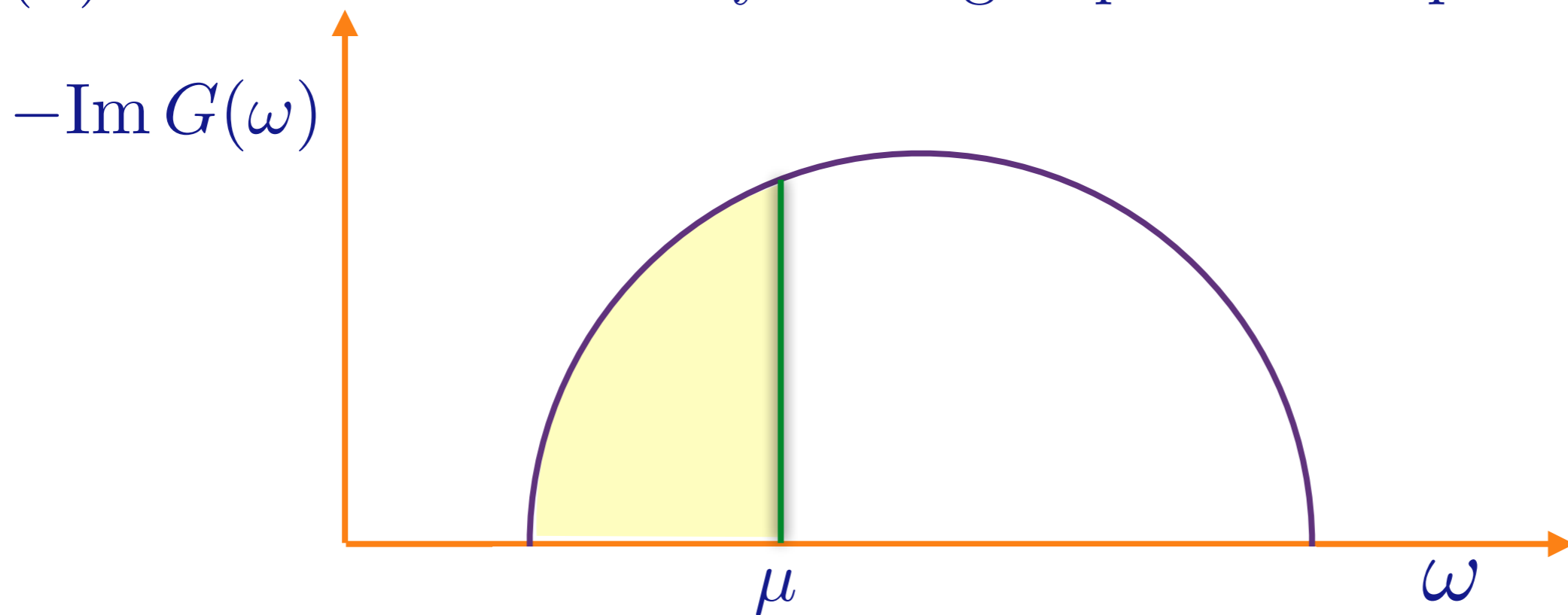
Fermions occupying the eigenstates of a
 $N \times N$ random matrix

Infinite-range model of a Fermi liquid

Feynman graph expansion in $t_{ij..}$, and graph-by-graph average, yields exact equations in the large N limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = t^2 G(\tau)$$
$$G(\tau = 0^-) = Q.$$

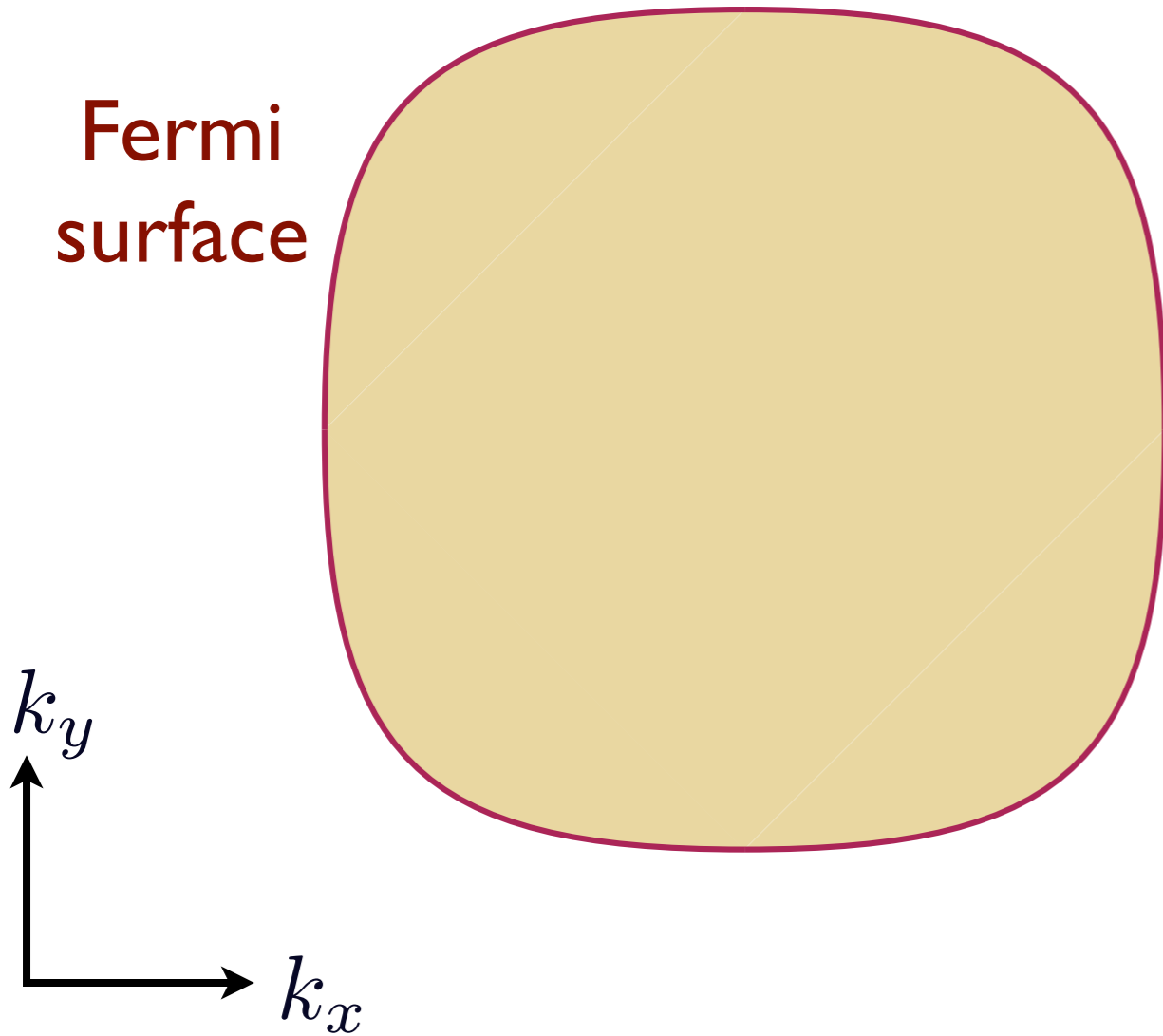
$G(\omega)$ can be determined by solving a quadratic equation.



Fermions occupying eigenstates with a “semi-circular” density of states

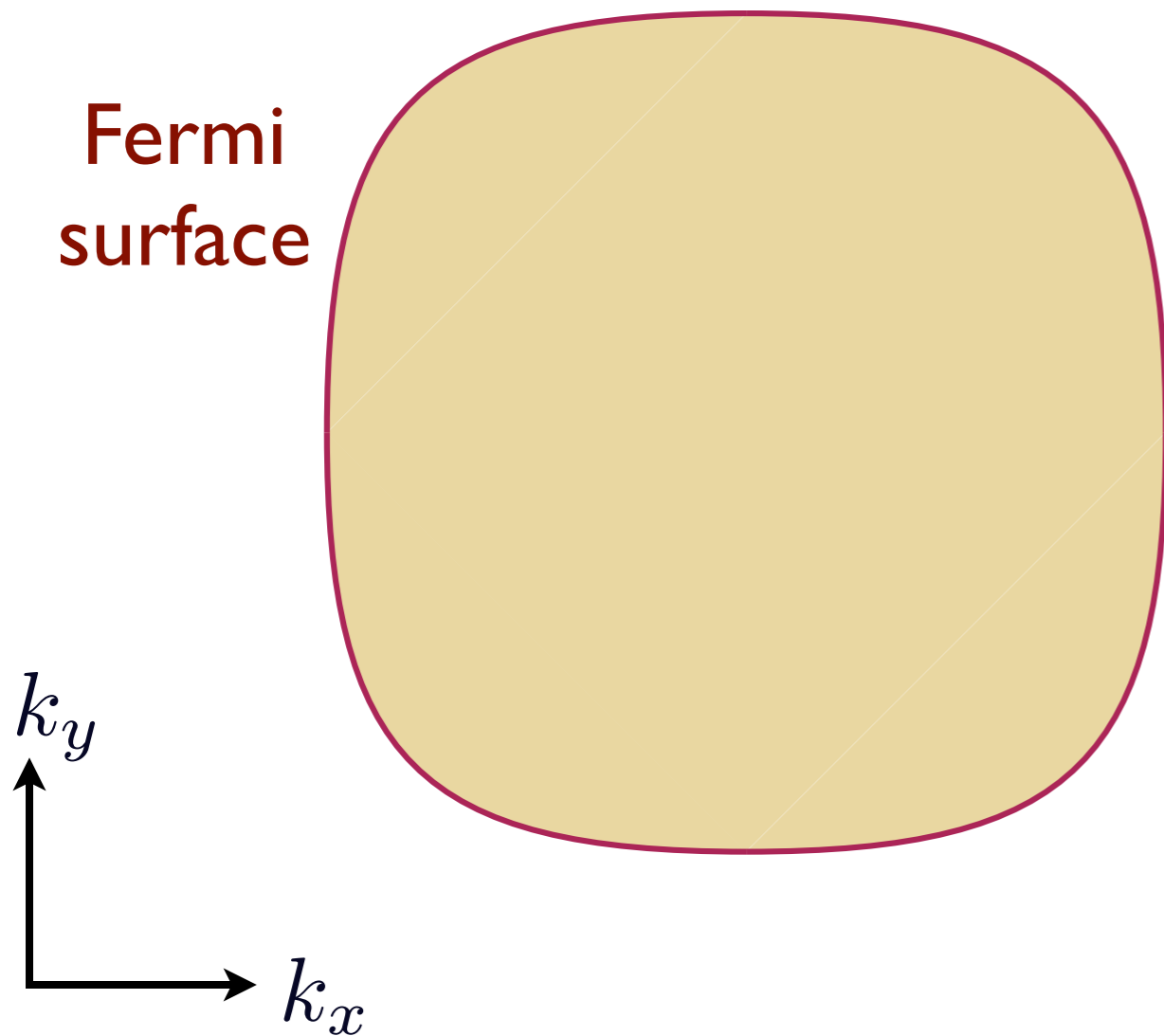
Ordinary quantum matter: the Fermi liquid (FL)

Fermi
surface



- Fermi surface separates empty and occupied states in momentum space.
- Area enclosed by Fermi surface = Q . Momenta of low energy excitations fixed by density of *all* electrons.
- Long-lived electron-like quasi-particle excitations near the Fermi surface: lifetime of quasi-particles $\sim 1/T^2$.

Ordinary quantum matter: the Fermi liquid (FL)

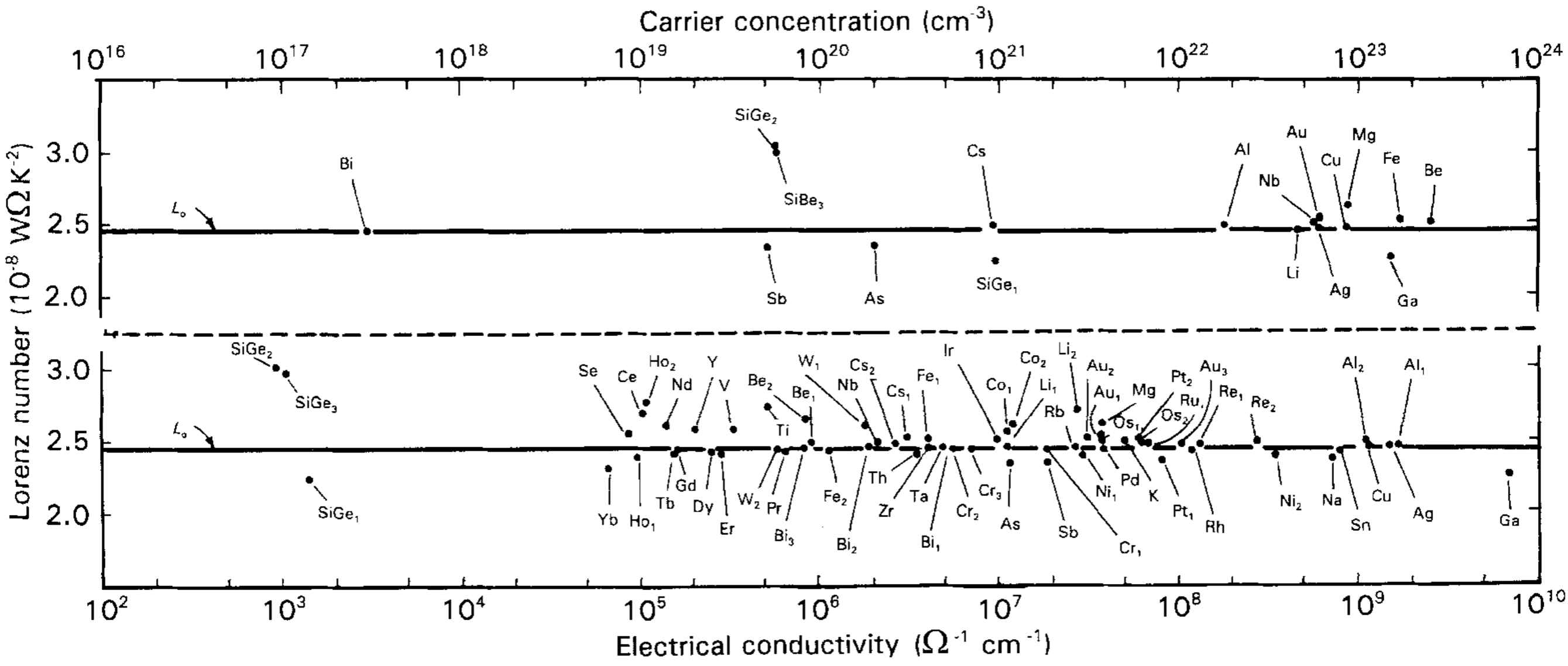


- Fermi surface separates empty and occupied states in momentum space.
- Area enclosed by Fermi surface = \mathcal{Q} . Momenta of low energy excitations fixed by density of *all* electrons.
- Long-lived electron-like quasi-particle excitations near the Fermi surface: lifetime of quasi-particles $\sim 1/T^2$.

- $$\frac{(\text{Thermal conductivity})}{T (\text{Electrical conductivity})} = \frac{\pi^2 k_B^2}{3e^2}$$

► Wiedemann-Franz law in a Fermi liquid:

$$\frac{\kappa}{\sigma T} \approx \frac{\pi^2 k_B^2}{3e^2} \approx 2.45 \times 10^{-8} \frac{W \cdot \Omega}{K^2}.$$



Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Infinite-range model of a strange metal

$$H = \frac{1}{(NM)^{1/2}} \sum_{i,j=1}^N \sum_{\alpha,\beta=1}^M J_{ij} c_{i\alpha}^\dagger c_{i\beta} c_{j\beta}^\dagger c_{j\alpha}$$

$$c_{i\alpha} c_{j\beta} + c_{j\beta} c_{i\alpha} = 0 \quad , \quad c_{i\alpha} c_{j\beta}^\dagger + c_{j\beta}^\dagger c_{i\alpha} = \delta_{ij} \delta_{\alpha\beta}$$

$$\frac{1}{M} \sum_{\alpha} c_{i\alpha}^\dagger c_{i\alpha} = Q$$

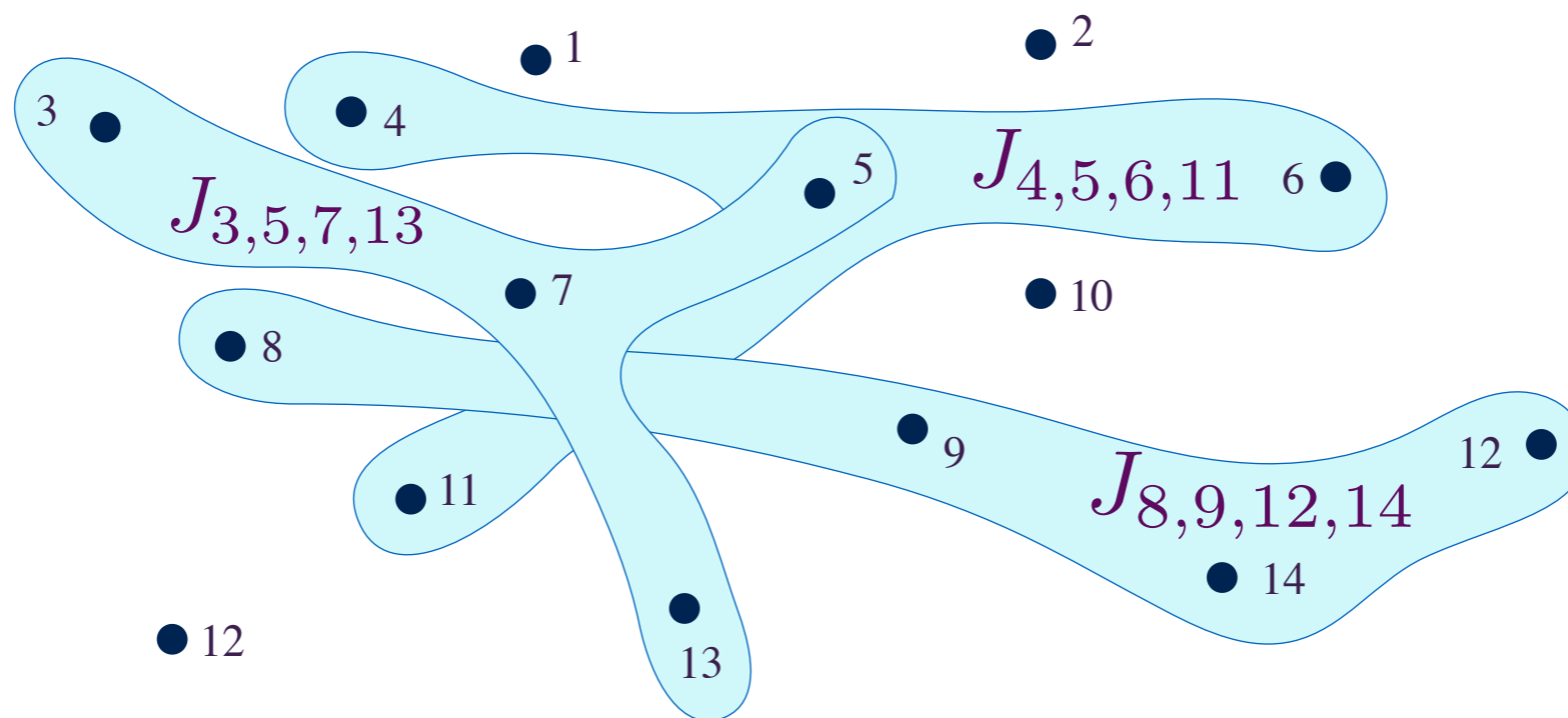
J_{ij} are independent random variables with $\overline{J_{ij}} = 0$ and $\overline{J_{ij}^2} = J^2$
 $N \rightarrow \infty$ at $M = 2$ yields spin-glass ground state.
 $N \rightarrow \infty$ and then $M \rightarrow \infty$ yields critical strange metal

Infinite-range model of a strange metal

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,l=1}^N J_{ij;kl} c_i^\dagger c_j^\dagger c_k c_l - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$Q = \frac{1}{N} \sum_i c_i^\dagger c_i$$



$J_{ij;kl}$ are independent random variables with $\overline{J_{ij;kl}} = 0$ and $\overline{|J_{ij;kl}|^2} = J^2$
 $N \rightarrow \infty$ yields same critical strange metal; simpler to study numerically

Infinite-range strange metals

Feynman graph expansion in $J_{ij..}$, and graph-by-graph average, yields exact equations in the large N limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -J^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = \mathcal{Q}.$$

Low frequency analysis shows that the solutions must be gapless and obey

$$\Sigma(z) = \mu - \frac{1}{A} \sqrt{z} + \dots \quad , \quad G(z) = \frac{A}{\sqrt{z}}$$

for some complex A . Let us also define $\tilde{\Sigma}(z) = \Sigma(z) - \mu$.

Infinite-range strange metals

At frequencies $\ll J$, the equations for G and Σ can be written as

$$\int d\tau_2 G(\tau_1, \tau_2) \tilde{\Sigma}(\tau_2, \tau_3) = -\delta(\tau_1 - \tau_3)$$
$$\tilde{\Sigma}(\tau_1, \tau_2) = -J^2 [G(\tau_1, \tau_2)]^2 G(\tau_2, \tau_1)$$

These equations are invariant under the reparametrization and gauge transformations

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} G(\sigma_1, \sigma_2)$$

$$\tilde{\Sigma}(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{\Sigma}(\sigma_1, \sigma_2)$$

where $f(\sigma)$ and $g(\sigma)$ are arbitrary functions.

A. Georges and O. Parcollet
PRB 59, 5341 (1999)
A. Kitaev, unpublished
S. Sachdev, arXiv:1506.05111

These equations and invariances have similarities to those of the large N limit of quantum spins at the spatial boundary of a CFT₂ (multi-channel Kondo problems)

O. Parcollet, A. Georges, G. Kotliar, and A. Sengupta
PRB 58, 3794 (1998)

Infinite-range strange metals

Local fermion density of states

$$\rho(\omega) = -\text{Im } G(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

\mathcal{E} encodes the particle-hole asymmetry

While \mathcal{E} determines the *low* energy spectrum, it is determined by the *total* fermion density Q :

$$Q = \frac{1}{4}(3 - \tanh(2\pi\mathcal{E})) - \frac{1}{\pi} \tan^{-1}(e^{2\pi\mathcal{E}}).$$

S. Sachdev and J. Ye, Phys. Rev. Lett. **70**, 3339 (1993)

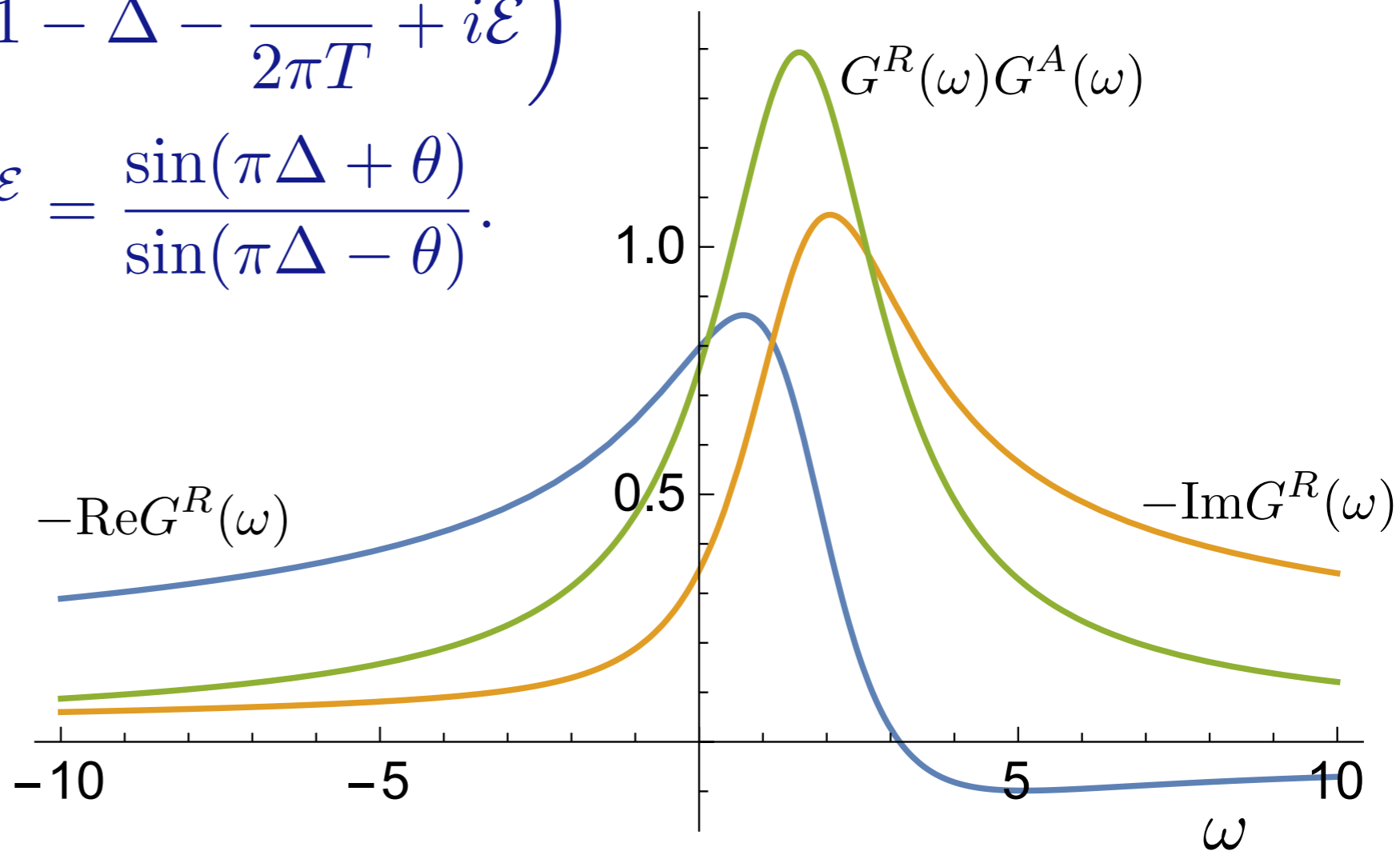
A. Georges, O. Parcollet, and S. Sachdev Phys. Rev. B **63**, 134406 (2001)

Infinite-range strange metals

At non-zero temperature, T , the Green's function also fully determined by \mathcal{E} .

$$G^R(\omega) = \frac{-iC e^{-i\theta}}{(2\pi T)^{1-2\Delta}} \frac{\Gamma\left(\Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}{\Gamma\left(1 - \Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}$$

where $\Delta = 1/4$ and $e^{2\pi\mathcal{E}} = \frac{\sin(\pi\Delta + \theta)}{\sin(\pi\Delta - \theta)}$.



S. Sachdev and J. Ye, Phys. Rev. Lett. **70**, 3339 (1993)

A. Georges and O. Parcollet PRB **59**, 5341 (1999)

A. Georges, O. Parcollet, and S. Sachdev Phys. Rev. B **63**, 134406 (2001)

Infinite-range strange metals

The entropy per site, \mathcal{S} , has a non-zero limit as $T \rightarrow 0$, and can be viewed as each site acquiring the universal boundary entropy of the multichannel Kondo problem.

Free spin has
entropy $\ln(2S + 1)$



Kondo-screened
spin has
no entropy

Metal



“Critically-screened”
spin has “irrational” entropy

N. Andrei and C. Destri, PRL **52**, 364 (1984).

A. M. Tsvelick, J. Phys. C **18**, 159 (1985).

I. Affleck and A. W. W. Ludwig, PRL **67**, 161 (1991).

S. Sachdev, C. Buragohain, and M. Vojta, Science **286**, 2479 (1999).

CFT



Infinite-range strange metals

The entropy per site, \mathcal{S} , has a non-zero limit as $T \rightarrow 0$, and can be viewed as each site acquiring the universal boundary entropy of the multichannel Kondo problem.

N. Andrei and C. Destri, PRL **52**, 364 (1984).

A. M. Tsvelick, J. Phys. C **18**, 159 (1985).

I. Affleck and A. W. W. Ludwig, PRL **67**, 161 (1991).

S. Sachdev, C. Buragohain, and M. Vojta, Science **286**, 2479 (1999).

“Critically-screened”
spin has “irrational” entropy

CFT



O. Parcollet, A. Georges, G. Kotliar, and A. Sengupta Phys. Rev. B **58**, 3794 (1998)

A. Georges, O. Parcollet, and S. Sachdev Phys. Rev. B **63**, 134406 (2001)

Infinite-range strange metals

The entropy per site, \mathcal{S} , has a non-zero limit as $T \rightarrow 0$, and can be viewed as each site acquiring the universal boundary entropy of the multichannel Kondo problem.

N. Andrei and C. Destri, PRL **52**, 364 (1984).

A. M. Tsvelick, J. Phys. C **18**, 159 (1985).

I. Affleck and A. W. W. Ludwig, PRL **67**, 161 (1991).

S. Sachdev, C. Buragohain, and M. Vojta, Science **286**, 2479 (1999).

This entropy obeys

$$\left(\frac{\partial \mathcal{S}}{\partial \mathcal{Q}} \right)_T = - \left(\frac{\partial \mu}{\partial T} \right)_\mathcal{Q} = 2\pi \mathcal{E}$$

Note that \mathcal{S} and \mathcal{E} involve low-lying states, while \mathcal{Q} depends upon *all* states, and details of the UV structure

O. Parcollet, A. Georges, G. Kotliar, and A. Sengupta Phys. Rev. B **58**, 3794 (1998)

A. Georges, O. Parcollet, and S. Sachdev Phys. Rev. B **63**, 134406 (2001)

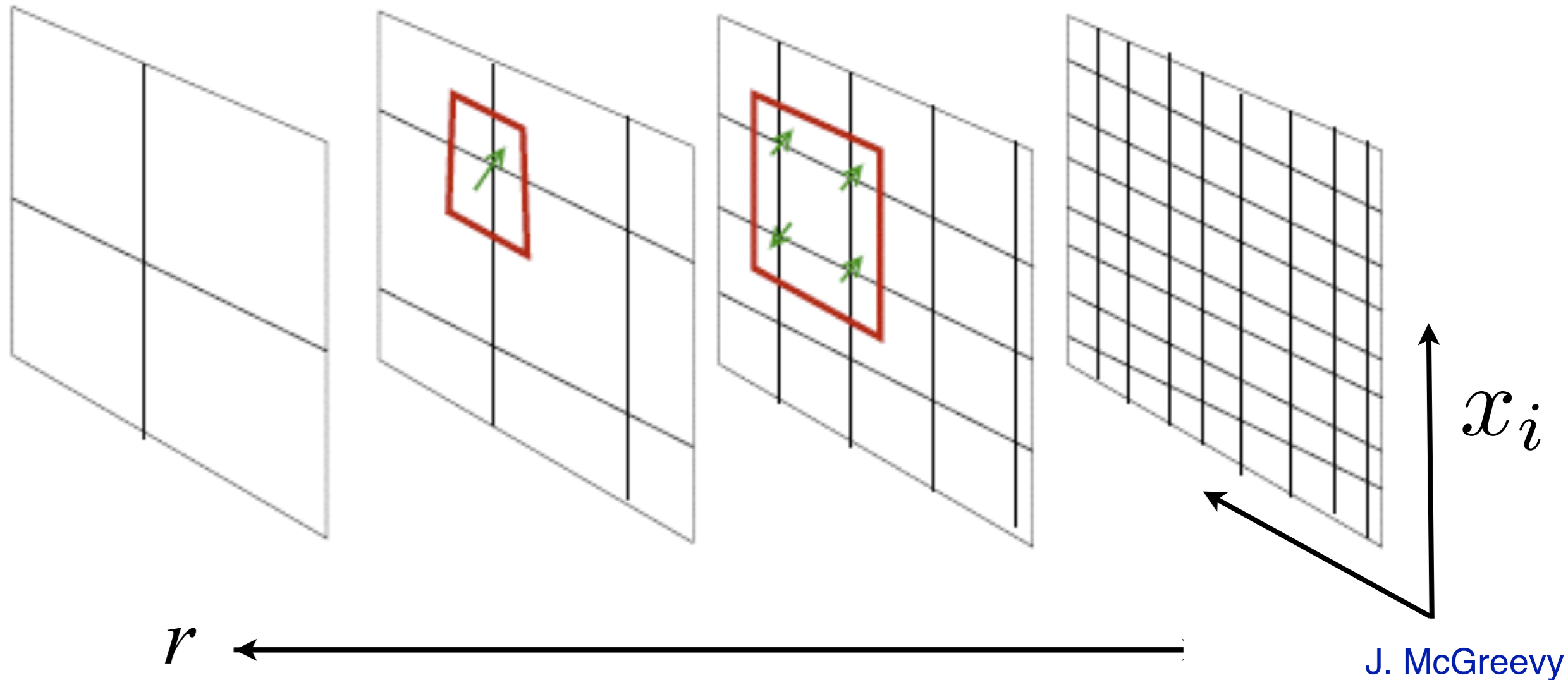
Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Quantum matter without quasiparticles

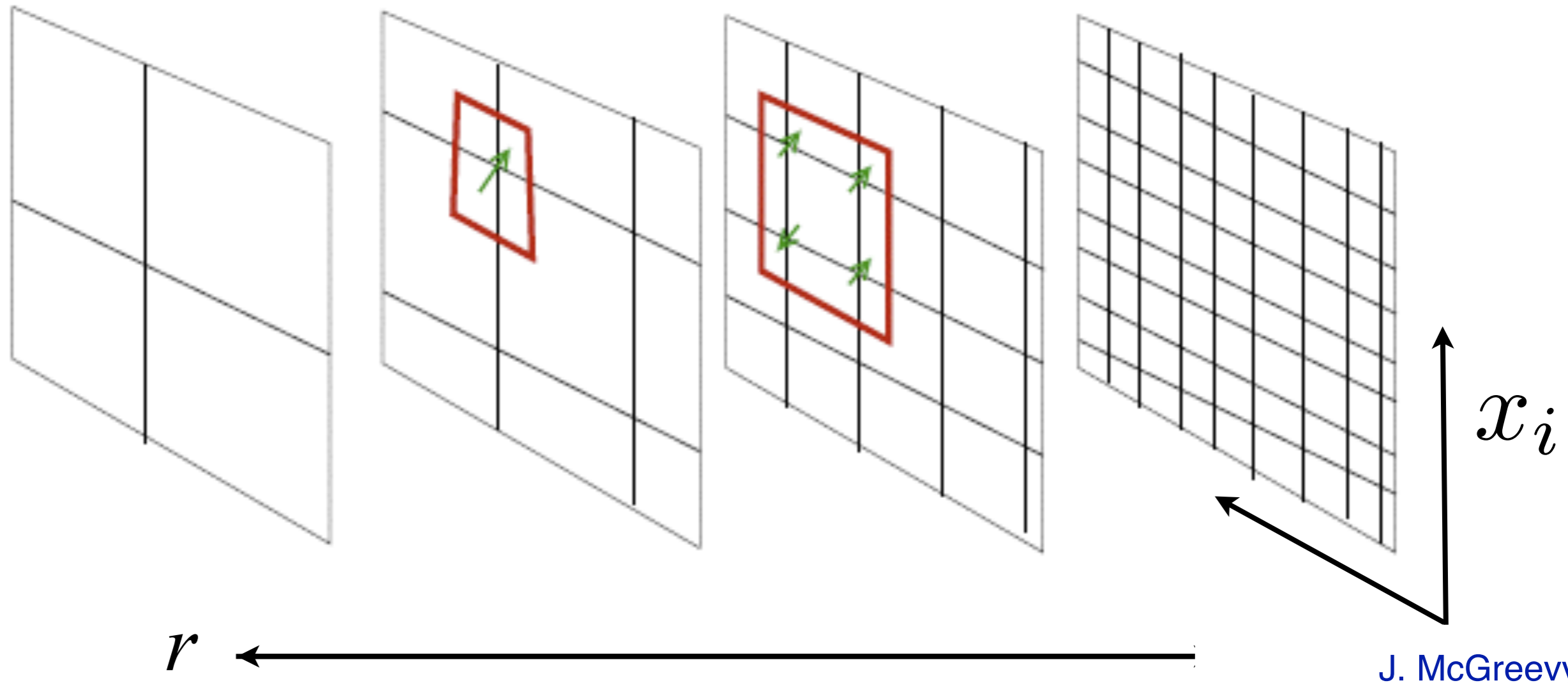
1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Holography



Key idea: \Rightarrow Implement r as an extra dimension, and map to a local theory in $d+2$ spacetime dimensions.

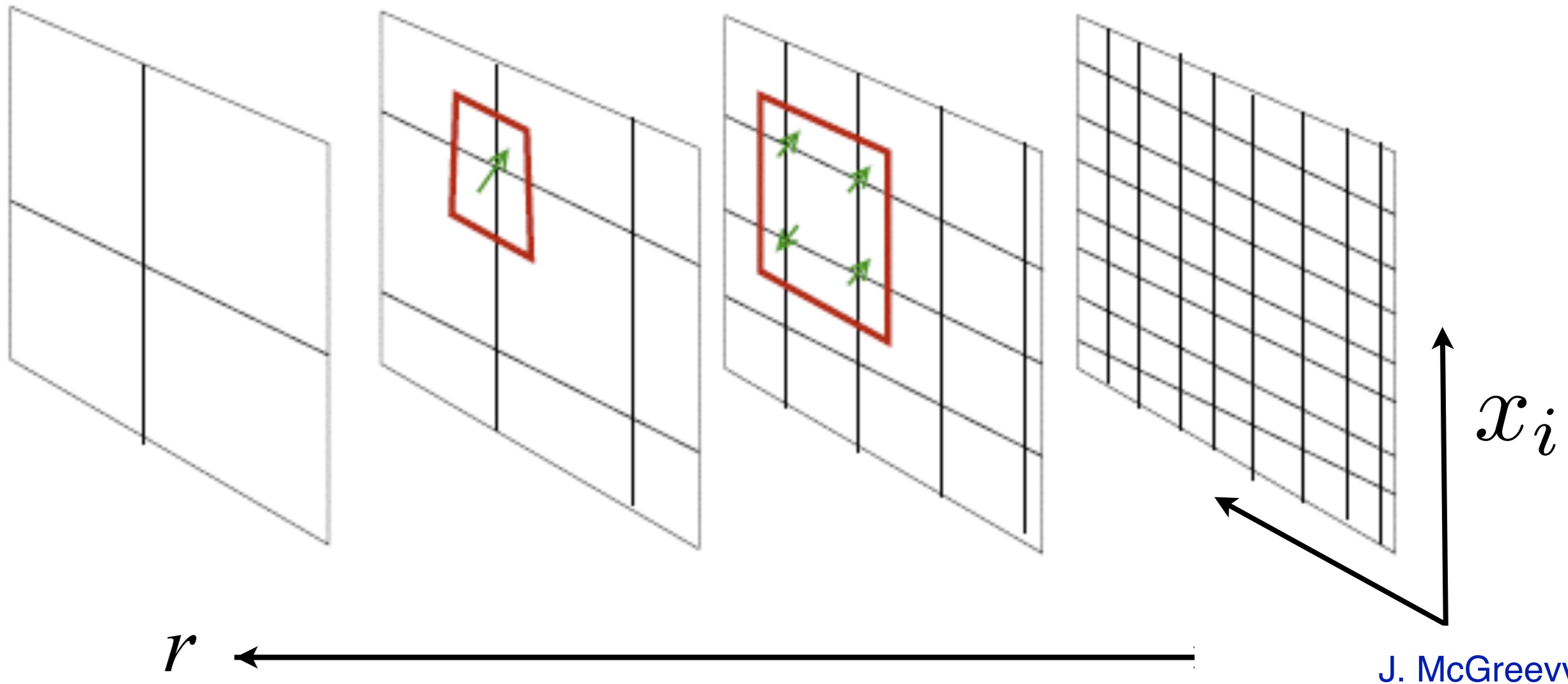
Holography



For a relativistic CFT in d spatial dimensions, the proper length, ds , in the holographic space is fixed by demanding the scale transformation ($i = 1 \dots d$)

$$x_i \rightarrow \zeta x_i \quad , \quad t \rightarrow \zeta t \quad , \quad ds \rightarrow ds$$

Holography



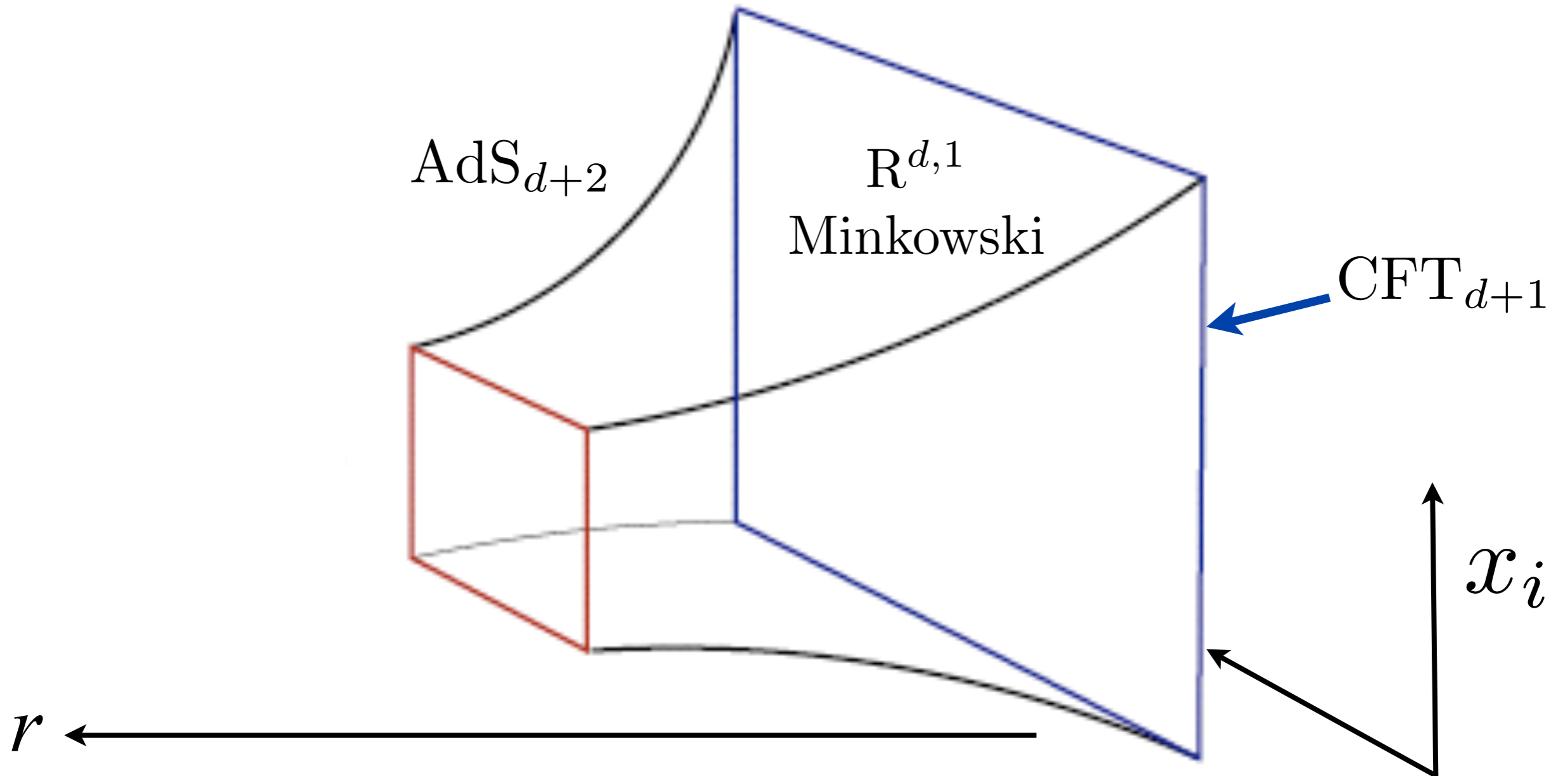
This gives the unique metric

$$ds^2 = \frac{1}{r^2} (-dt^2 + dr^2 + dx_i^2)$$

This is the metric of anti-de Sitter space AdS_{d+2} .

AdS/CFT correspondence at zero temperature

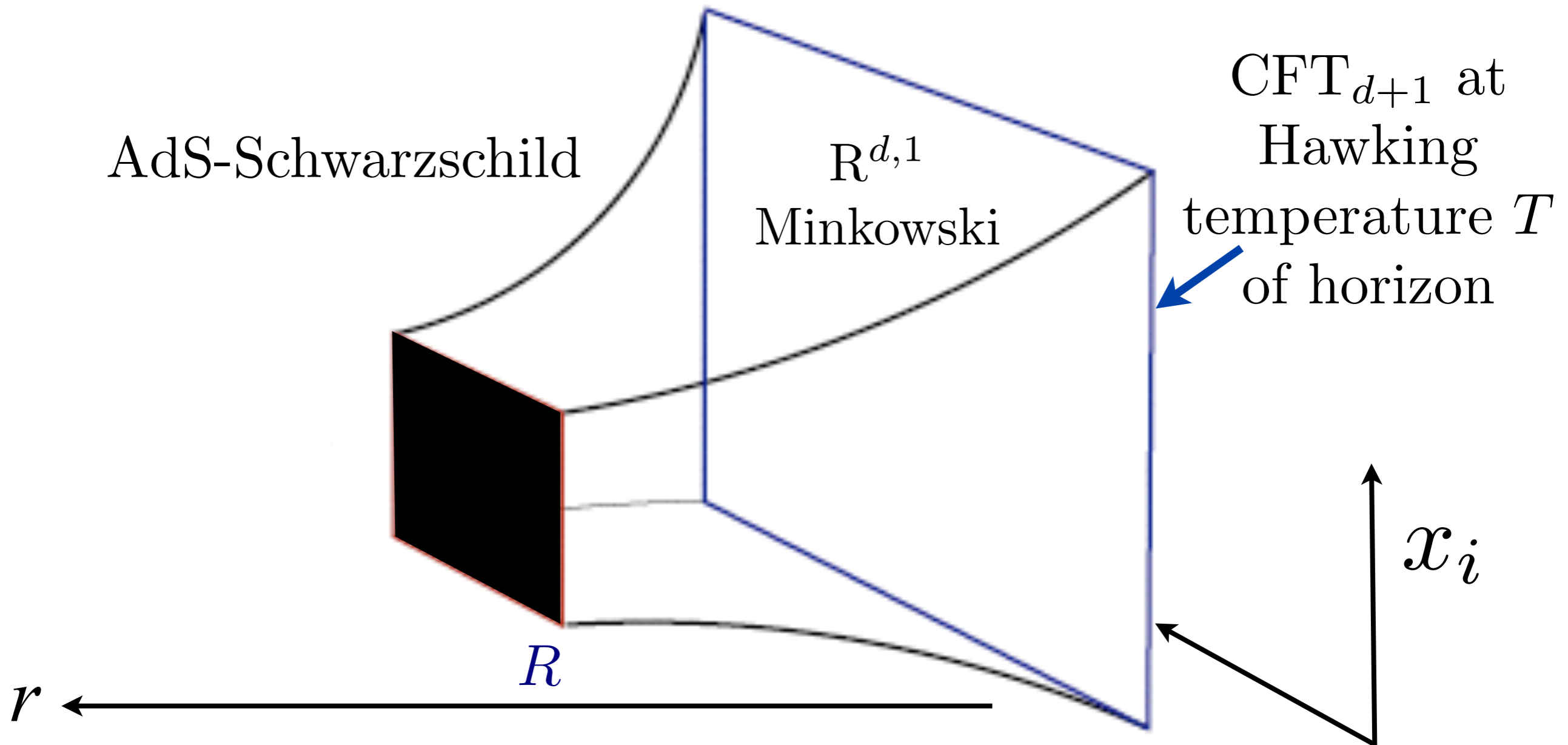
Einstein gravity $\mathcal{S}_E = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} \right) \right]$



$$ds^2 = \left(\frac{L}{r} \right)^2 [dr^2 - dt^2 + d\vec{x}^2]$$

AdS/CFT correspondence at non-zero temperature

Einstein gravity $\mathcal{S}_E = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} \right) \right]$

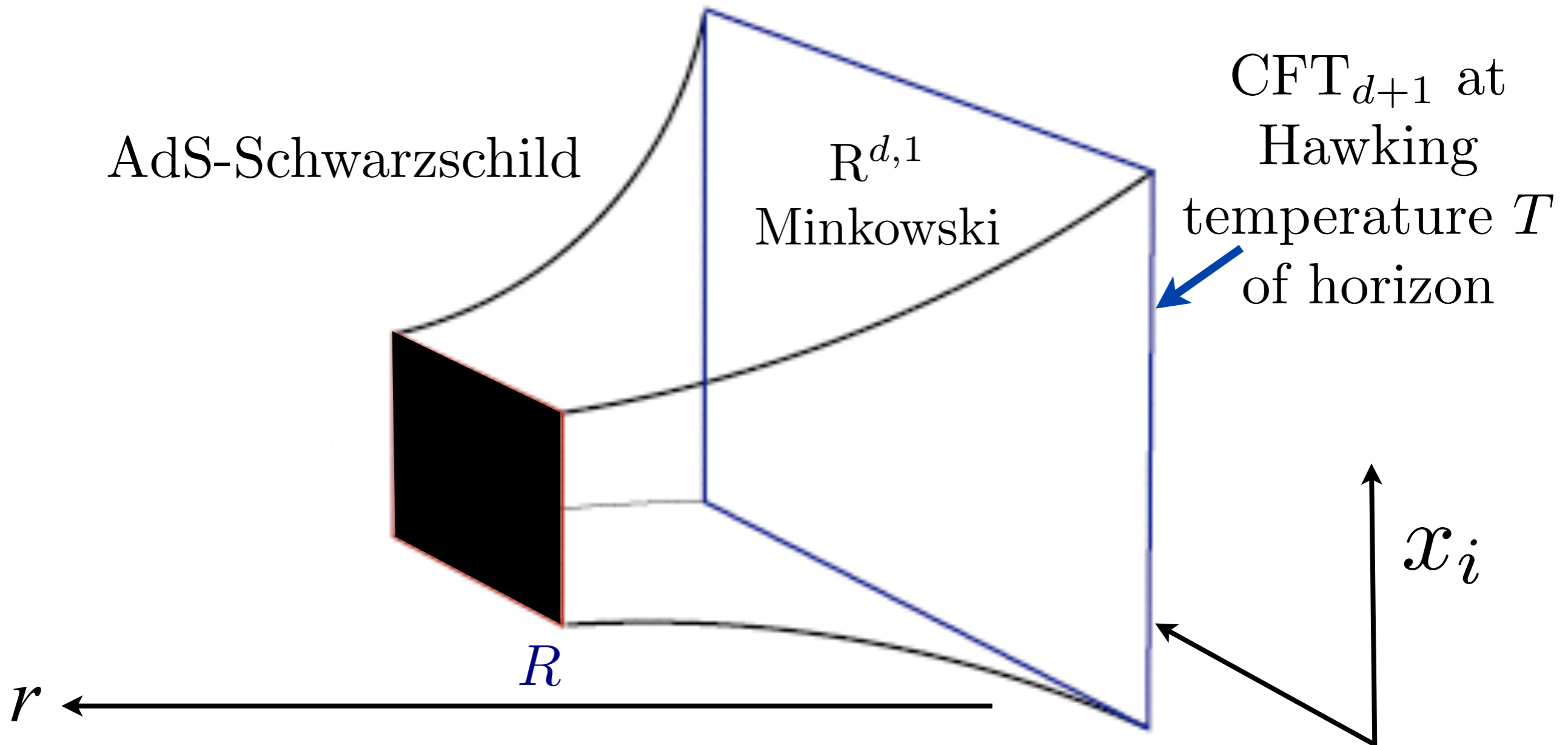


$$ds^2 = \left(\frac{L}{r} \right)^2 \left[\frac{dr^2}{f(r)} - f(r) dt^2 + d\vec{x}^2 \right]$$

with $f(r) = 1 - (r/R)^{d+1}$ and $T = (d+1)/(4\pi R)$.

AdS/CFT correspondence at non-zero temperature

Einstein gravity $\mathcal{S}_E = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} \right) \right]$

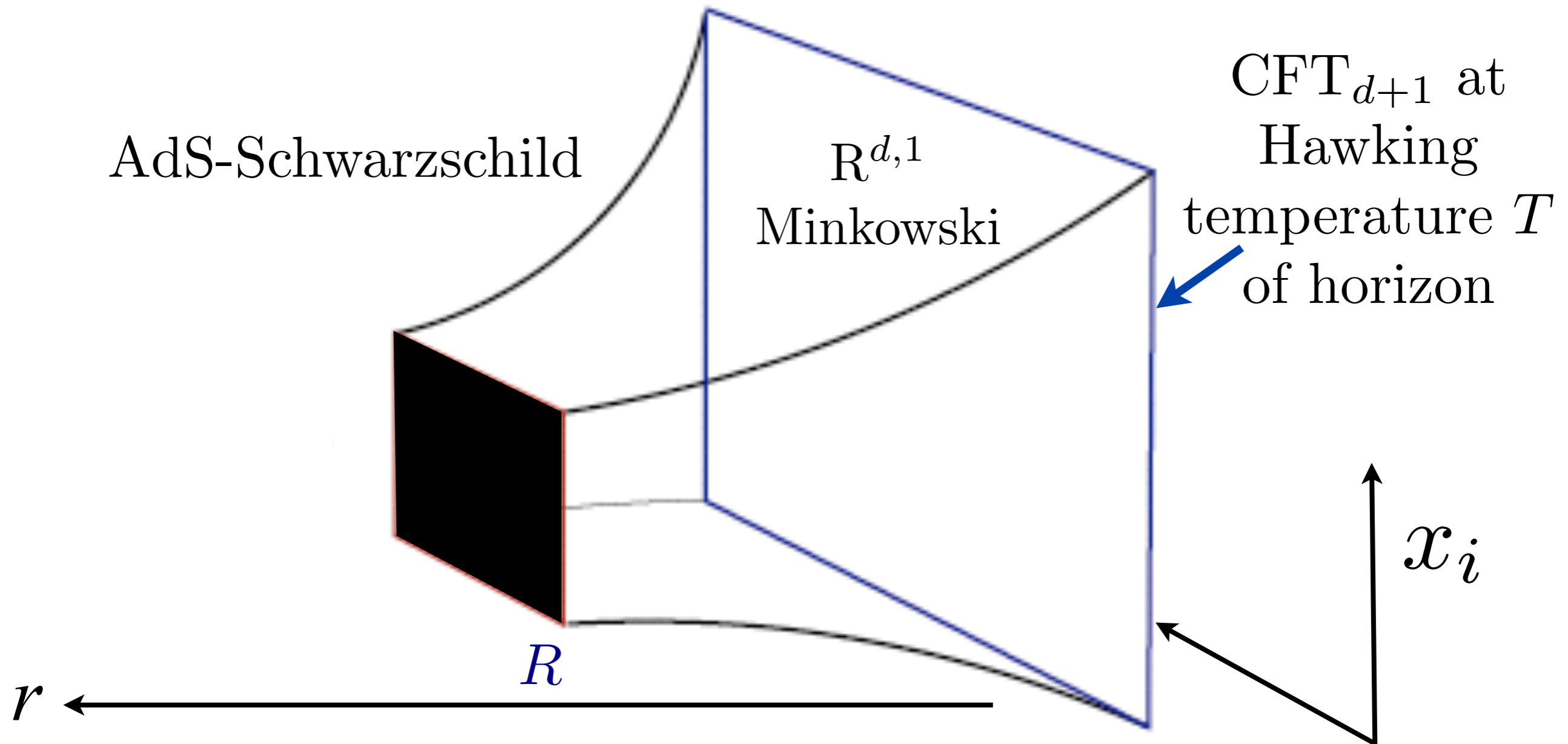


Entropy density of CFT $_{d+1}$, $\mathcal{S} \sim T^d$

Bekenstein-Hawking entropy density, $\mathcal{S}_{\text{BH}} \sim T^d$

AdS/CFT correspondence at non-zero temperature

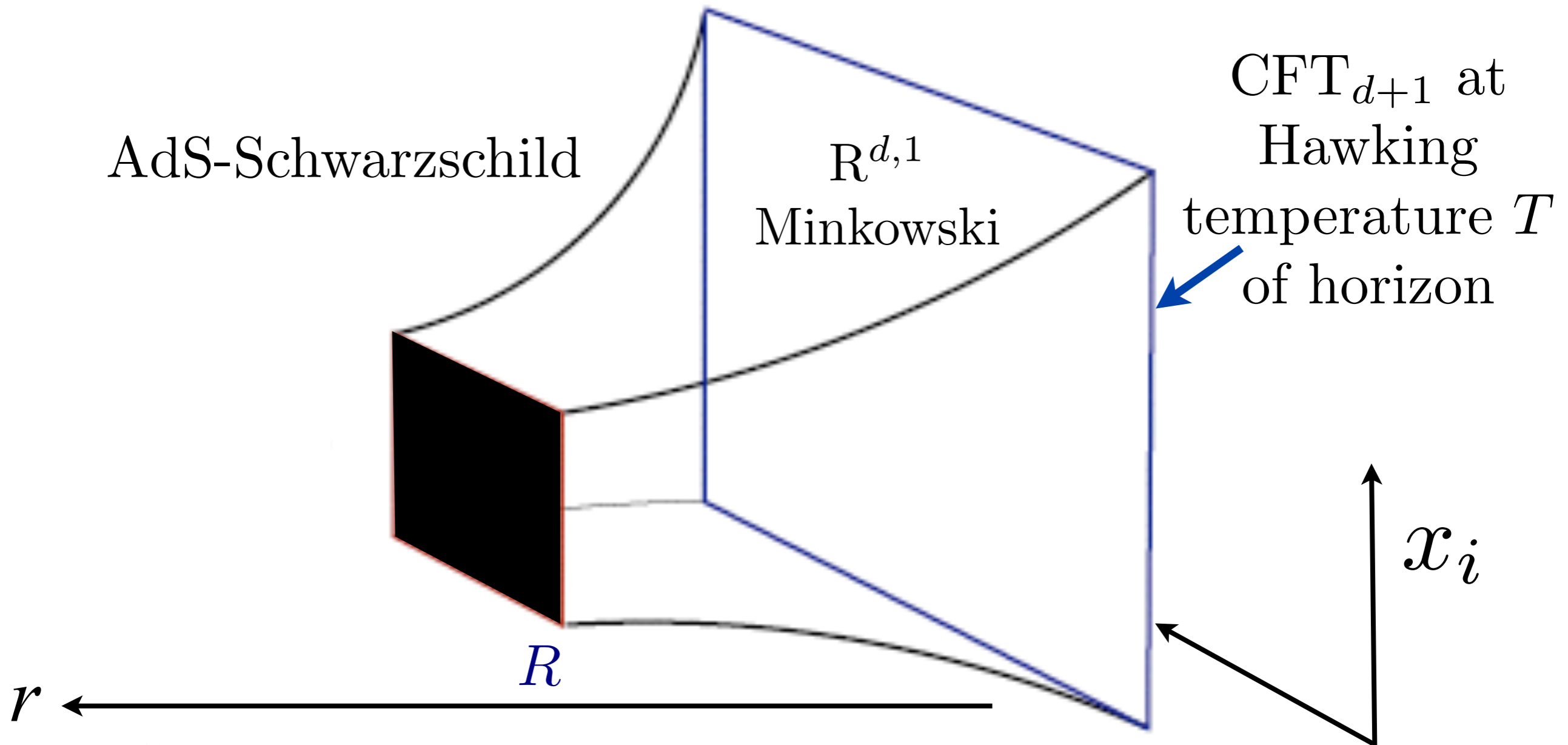
Einstein gravity $\mathcal{S}_E = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} \right) \right]$



For $\text{SU}(N)$ SYM in $d = 3$, $\mathcal{S}_{\text{BH}} = (\pi^2/2)N^2T^3$. But there is (still) no confirmation of this from a field-theory computation on SYM.

AdS/CFT correspondence at non-zero temperature

Einstein gravity $\mathcal{S}_E = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} \right) \right]$



Correspondence in $d = 1$:

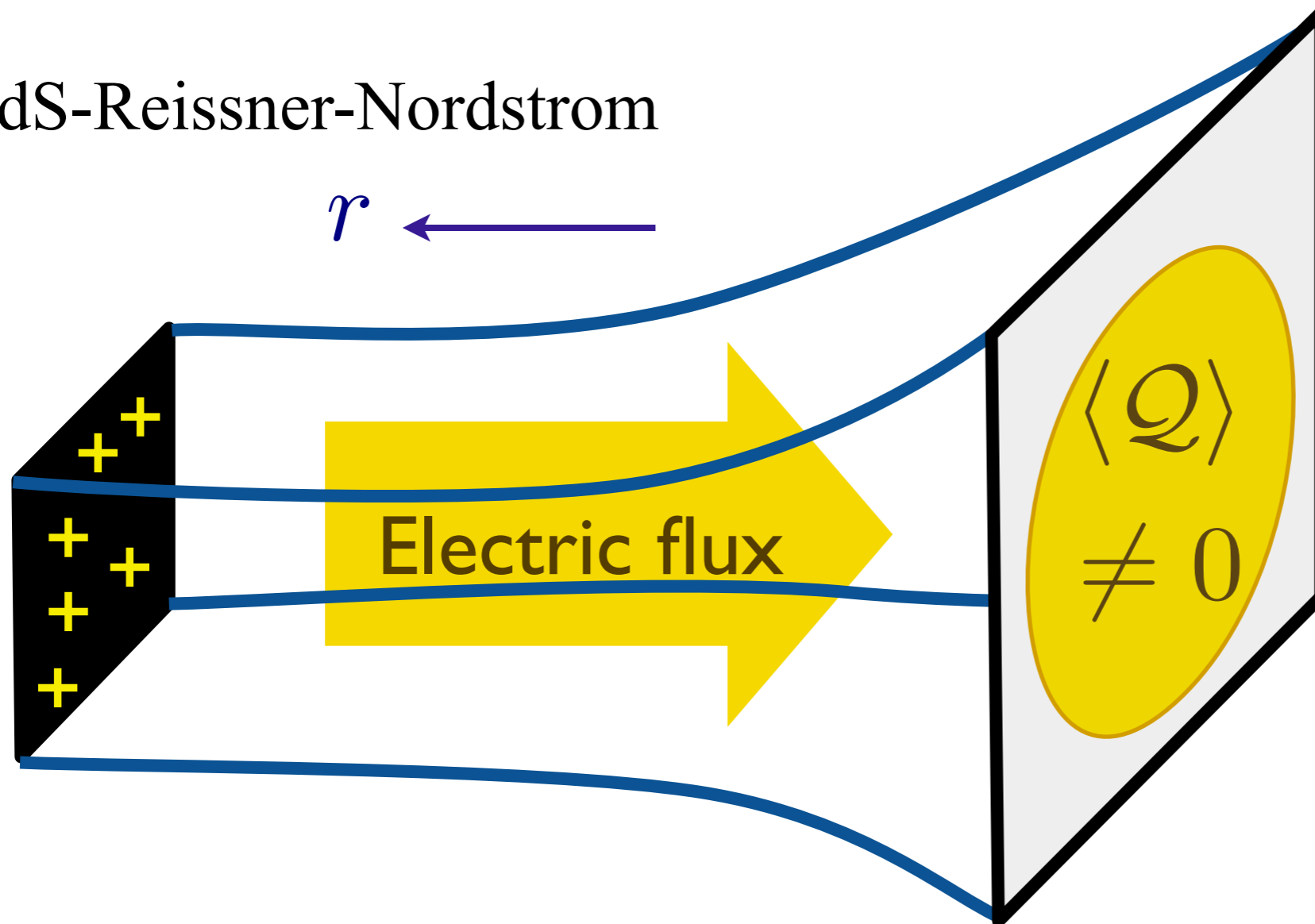
$$\mathcal{S} = \mathcal{S}_{\text{BH}} = \frac{\pi}{3} c T,$$

where $c = 12\pi L/\kappa^2$ is the central charge of the CFT₂.

Charged black branes

Einstein-Maxwell theory $\mathcal{S}_{EM} = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} - \frac{R^2}{g_F^2} F^2 \right) \right]$

AdS-Reissner-Nordstrom

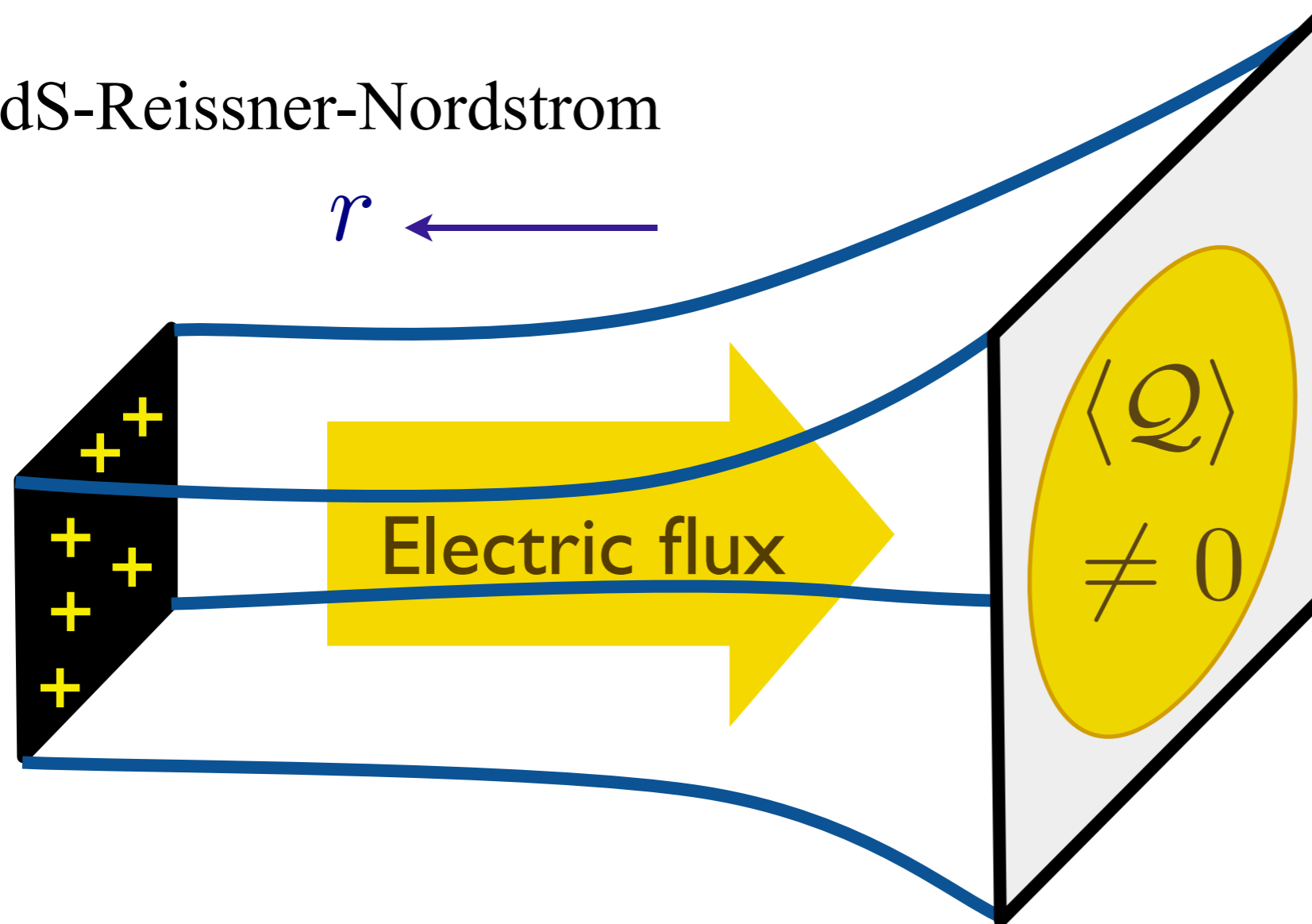


Quantum matter on the boundary with a variable charge density \mathcal{Q} of a global U(1) symmetry.

Charged black branes

Einstein-Maxwell theory $\mathcal{S}_{EM} = \int d^{d+2}x \sqrt{-g} \left[\frac{1}{2\kappa^2} \left(\mathcal{R} + \frac{d(d+1)}{L^2} - \frac{R^2}{g_F^2} F^2 \right) \right]$

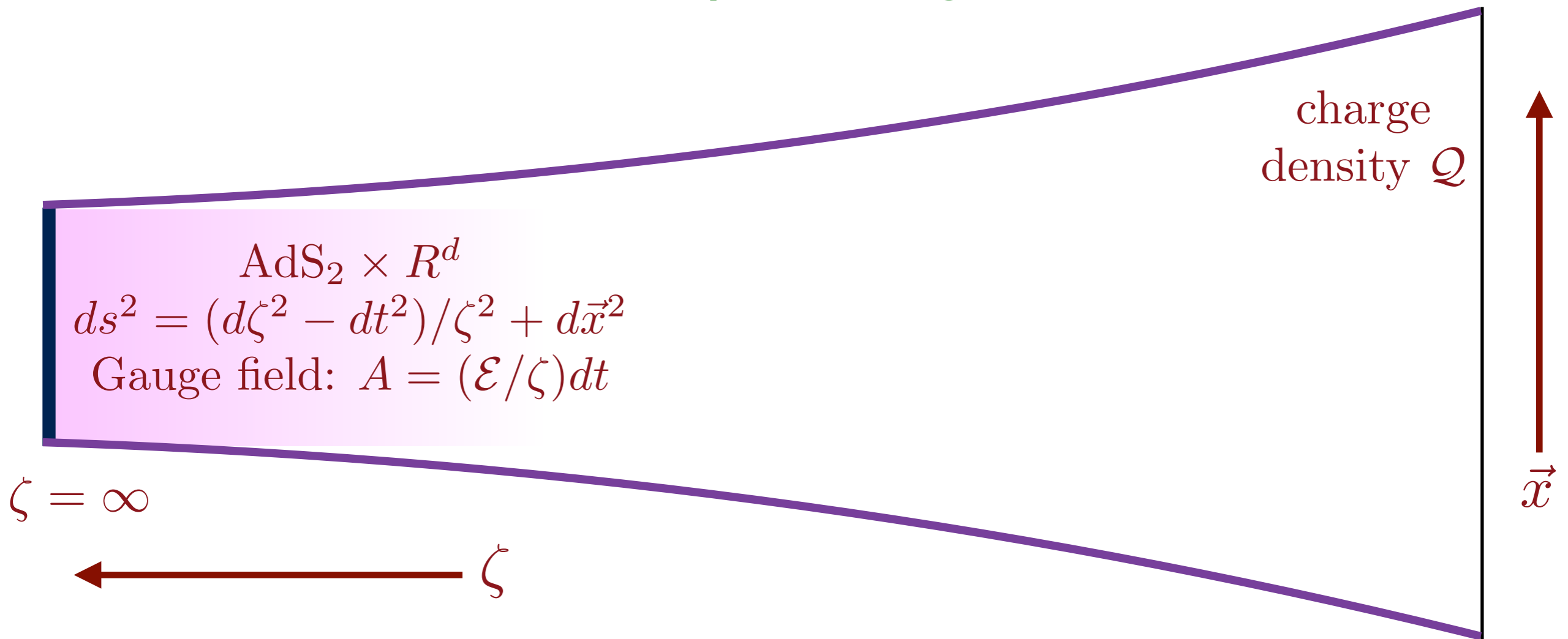
AdS-Reissner-Nordstrom



Quantum matter on the boundary with a variable charge density \mathcal{Q} of a global U(1) symmetry.

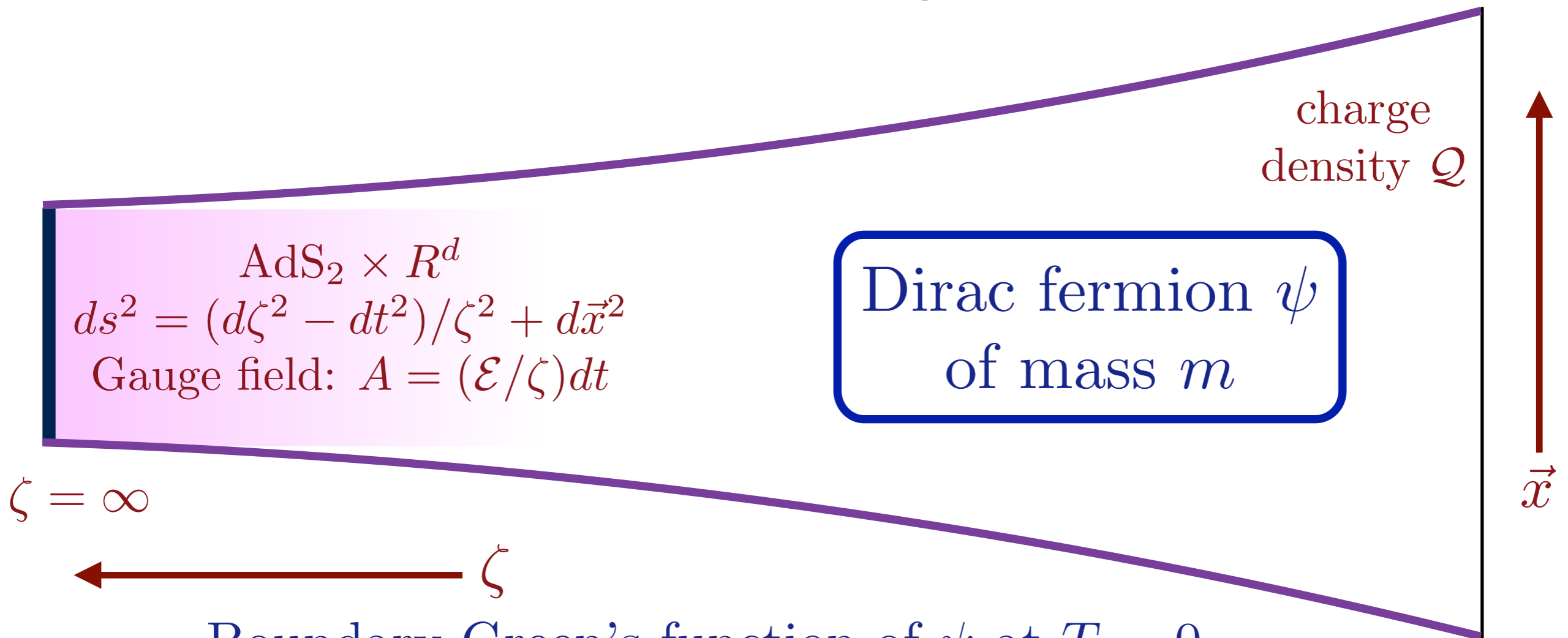
Realizes a strange metal: a state with an unbroken global U(1) symmetry with a continuously variable charge density, \mathcal{Q} , at $T = 0$ which does not have any quasiparticle excitations.

General Relativity of charged black branes



- Near-horizon metric is AdS_2 , with near-horizon electric field \mathcal{E} .

Quantum fields on charged black branes



Boundary Green's function of ψ at $T = 0$

$$\text{Im}G(\omega) \sim \begin{cases} \omega^{-(1-2\Delta)}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-(1-2\Delta)}, & \omega < 0. \end{cases}$$

where the fermion scaling dimension Δ is a function of m

\mathcal{E} encodes the particle-hole asymmetry

Quantum fields on charged black branes

Conformal mapping to $T > 0$

$$\zeta = \zeta_0$$

charge density \mathcal{Q}

$$ds^2 = [d\zeta^2 / (1 - \zeta^2 / \zeta_0^2) - (1 - \zeta^2 / \zeta_0^2) dt^2] / \zeta^2 + d\vec{x}^2$$

$$\text{Gauge field: } A = \mathcal{E}(1/\zeta - 1/\zeta_0) dt \text{ with } \zeta_0 = 1/(2\pi T)$$

Dirac fermion ψ
of mass m

$$\zeta = \infty$$

\vec{x}

Boundary Green's function of ψ at $T > 0$

is fully determined by \mathcal{E}

$$G^R(\omega) = \frac{-iC e^{-i\theta}}{(2\pi T)^{1-2\Delta}} \frac{\Gamma\left(\Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}{\Gamma\left(1 - \Delta - \frac{i\omega}{2\pi T} + i\mathcal{E}\right)}$$

$$\text{where } e^{2\pi\mathcal{E}} = \frac{\sin(\pi\Delta + \theta)}{\sin(\pi\Delta - \theta)}.$$

General Relativity of charged black branes

Conformal mapping to $T > 0$

$$\zeta = \zeta_0$$

charge
density Q

$$ds^2 = [d\zeta^2 / (1 - \zeta^2 / \zeta_0^2) - (1 - \zeta^2 / \zeta_0^2) dt^2] / \zeta^2 + d\vec{x}^2$$

$$\text{Gauge field: } A = \mathcal{E}(1/\zeta - 1/\zeta_0) dt \text{ with } \zeta_0 = 1/(2\pi T)$$

$$\zeta = \infty$$

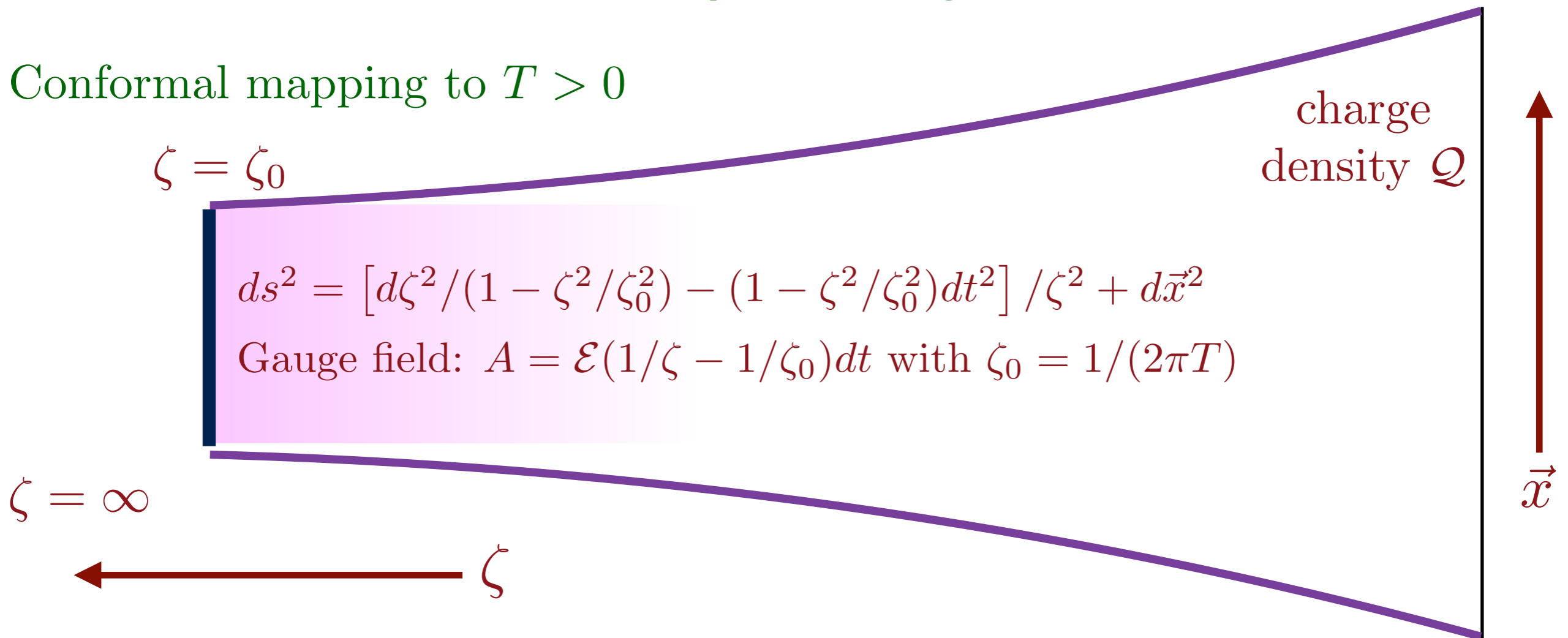
$$\zeta$$

$$\vec{x}$$

- As $T \rightarrow 0$, there is a non-zero Bekenstein-Hawking entropy, \mathcal{S}_{BH} .

General Relativity of charged black branes

Conformal mapping to $T > 0$



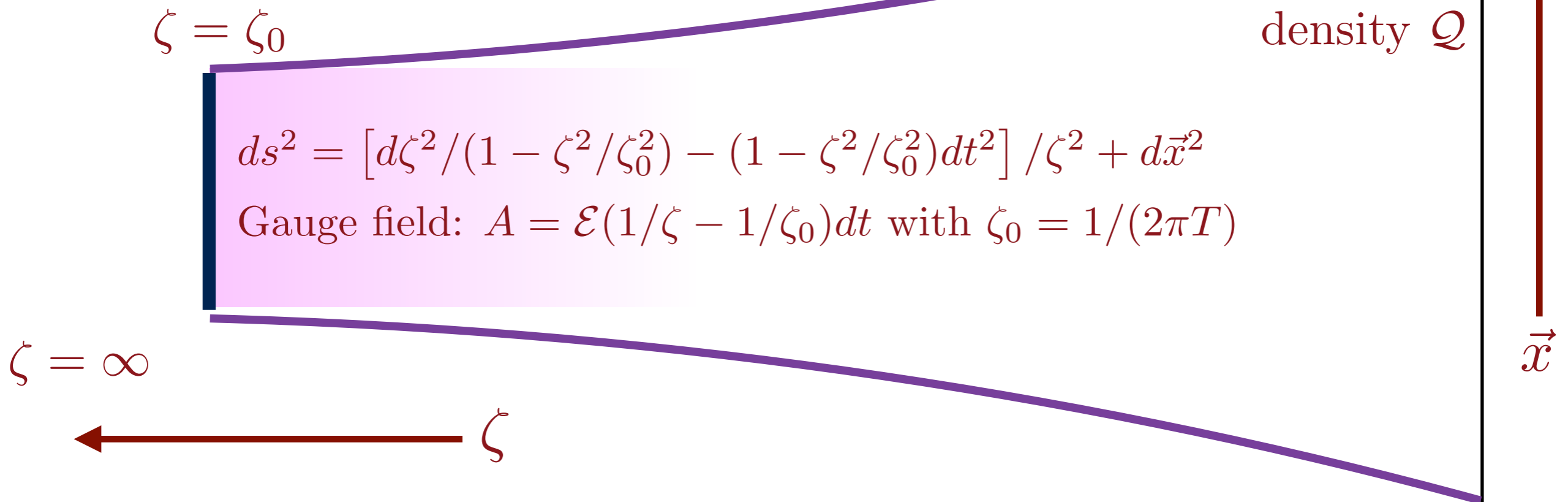
- As $T \rightarrow 0$, there is a non-zero Bekenstein-Hawking entropy, \mathcal{S}_{BH} .
- Using Gauss's Law, it can be shown that $\mu(T) = -2\pi\mathcal{E}T + \text{constant}$ as $T \rightarrow 0$.
- Using a thermodynamic Maxwell relation (also obeyed by gravity),

A. Sen
 hep-th/0506177
 S. Sachdev
 1506.05111

$$\left(\frac{\partial \mathcal{S}_{BH}}{\partial Q} \right)_T = - \left(\frac{\partial \mu}{\partial T} \right)_Q = 2\pi\mathcal{E}$$

General Relativity of charged black branes

Conformal mapping to $T > 0$



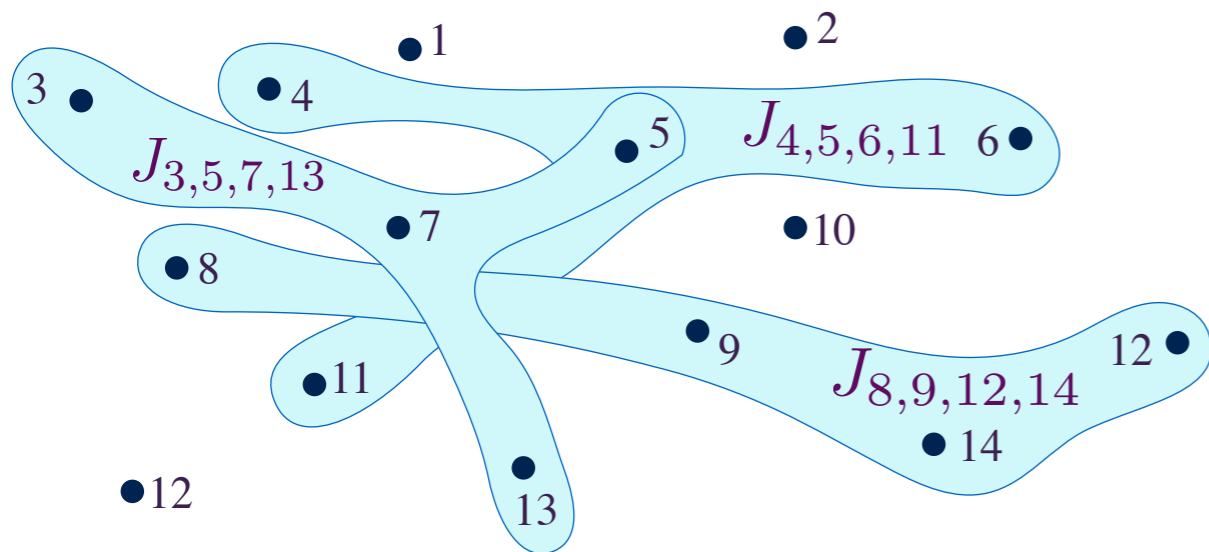
- As $T \rightarrow 0$, there is a non-zero Bekenstein-Hawking entropy, \mathcal{S}_{BH} .
- Using Gauss's Law, it can be shown that $\mu(T) = -2\pi\mathcal{E}T + \text{constant}$ as $T \rightarrow 0$.
- Using a thermodynamic Maxwell relation (also obeyed by gravity)

A. Sen
 hep-th/0506177
 S. Sachdev
 1506.05111

$$\left(\frac{\partial \mathcal{S}_{BH}}{\partial Q} \right)_T = - \left(\frac{\partial \mu}{\partial T} \right)_Q = 2\pi\mathcal{E}$$

Also obeyed by
 Wald entropy
 in higher-derivative
 gravity.

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N J_{ij;kl} c_i^\dagger c_j^\dagger c_k c_\ell$$



$$Q = \frac{1}{N} \sum_i \langle c_i^\dagger c_i \rangle.$$

Local fermion density of states

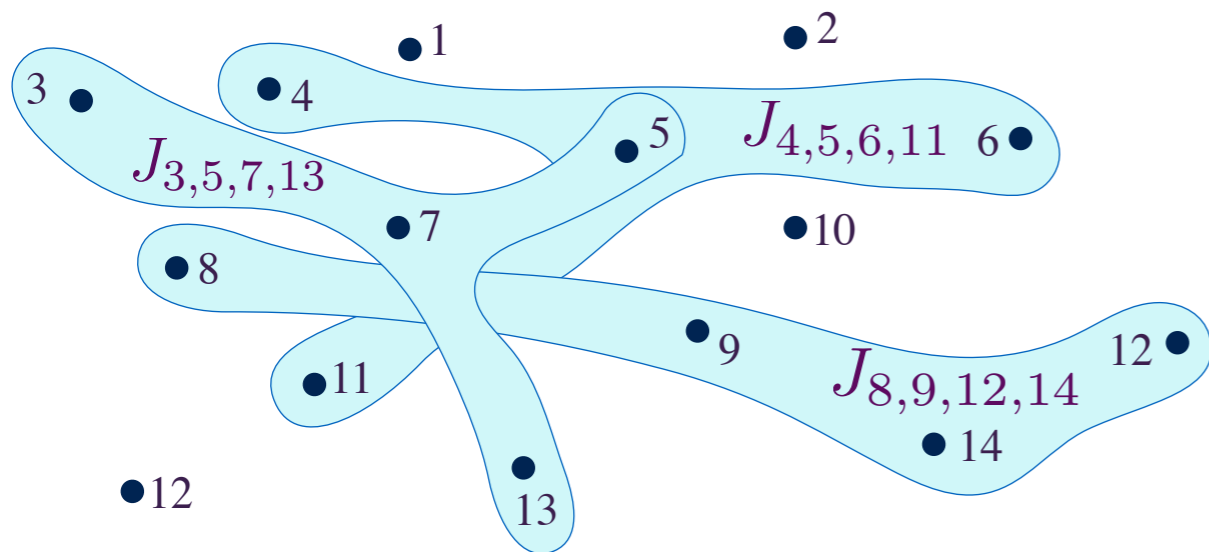
$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

Known 'equation of state'
determines \mathcal{E} as a function of Q

Microscopic zero temperature
entropy density, \mathcal{S} , obeys

$$\frac{\partial \mathcal{S}}{\partial Q} = 2\pi\mathcal{E}$$

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N J_{ij;kl} c_i^\dagger c_j^\dagger c_k c_\ell$$



$$Q = \frac{1}{N} \sum_i \langle c_i^\dagger c_i \rangle.$$

Local fermion density of states

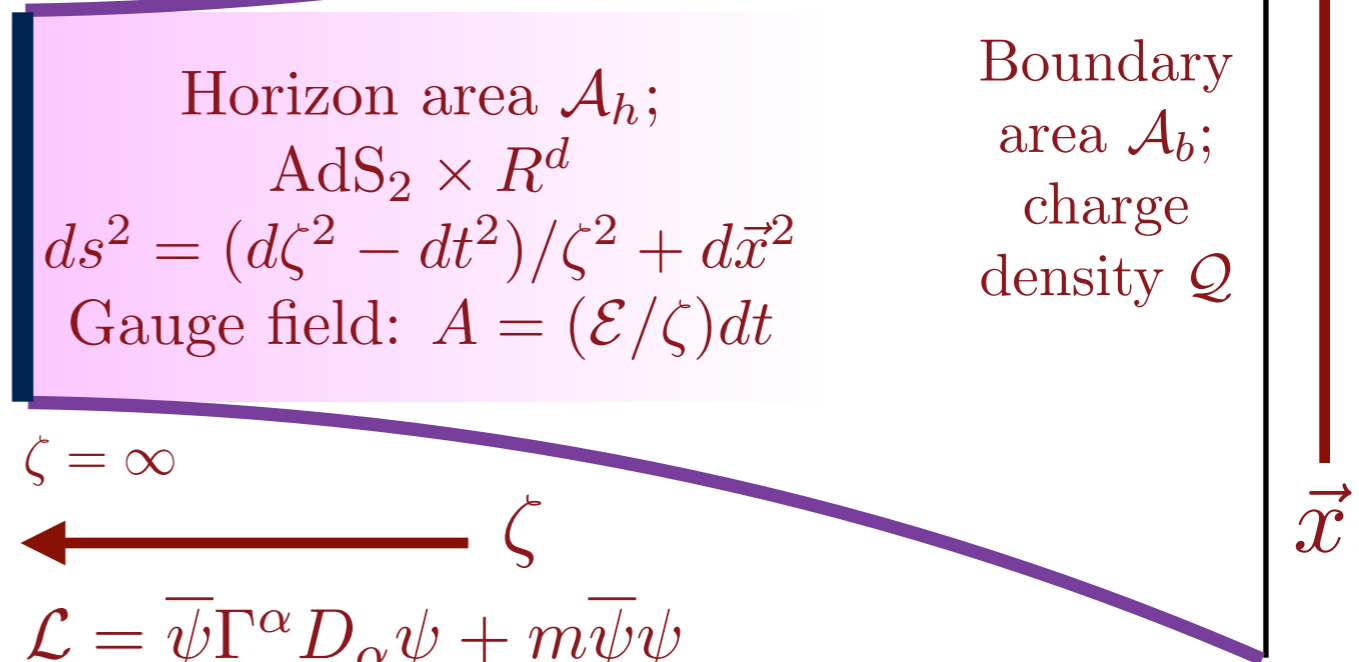
$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

Known ‘equation of state’ determines \mathcal{E} as a function of Q

Microscopic zero temperature entropy density, \mathcal{S} , obeys

$$\frac{\partial \mathcal{S}}{\partial Q} = 2\pi\mathcal{E}$$

Einstein-Maxwell theory
+ cosmological constant



$$\zeta = \infty$$

$$\zeta$$

$$\mathcal{L} = \bar{\psi} \Gamma^\alpha D_\alpha \psi + m \bar{\psi} \psi$$

Local fermion density of states

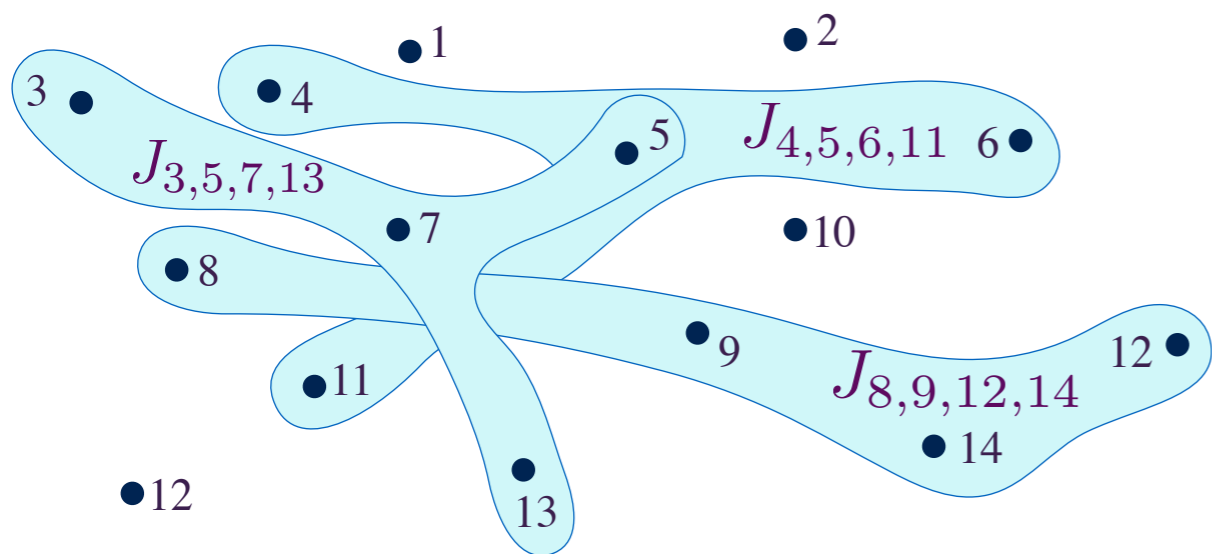
$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

‘Equation of state’ relating \mathcal{E} and Q depends upon the geometry of spacetime far from the AdS_2

Black hole thermodynamics (classical general relativity) yields

$$\frac{\partial \mathcal{S}_{\text{BH}}}{\partial Q} = 2\pi\mathcal{E}$$

$$H = \frac{1}{(2N)^{3/2}} \sum_{i,j,k,\ell=1}^N J_{ij;kl} c_i^\dagger c_j^\dagger c_k c_\ell$$



$$Q = \frac{1}{N} \sum_i \langle c_i^\dagger c_i \rangle.$$

Local fermion density of states

$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

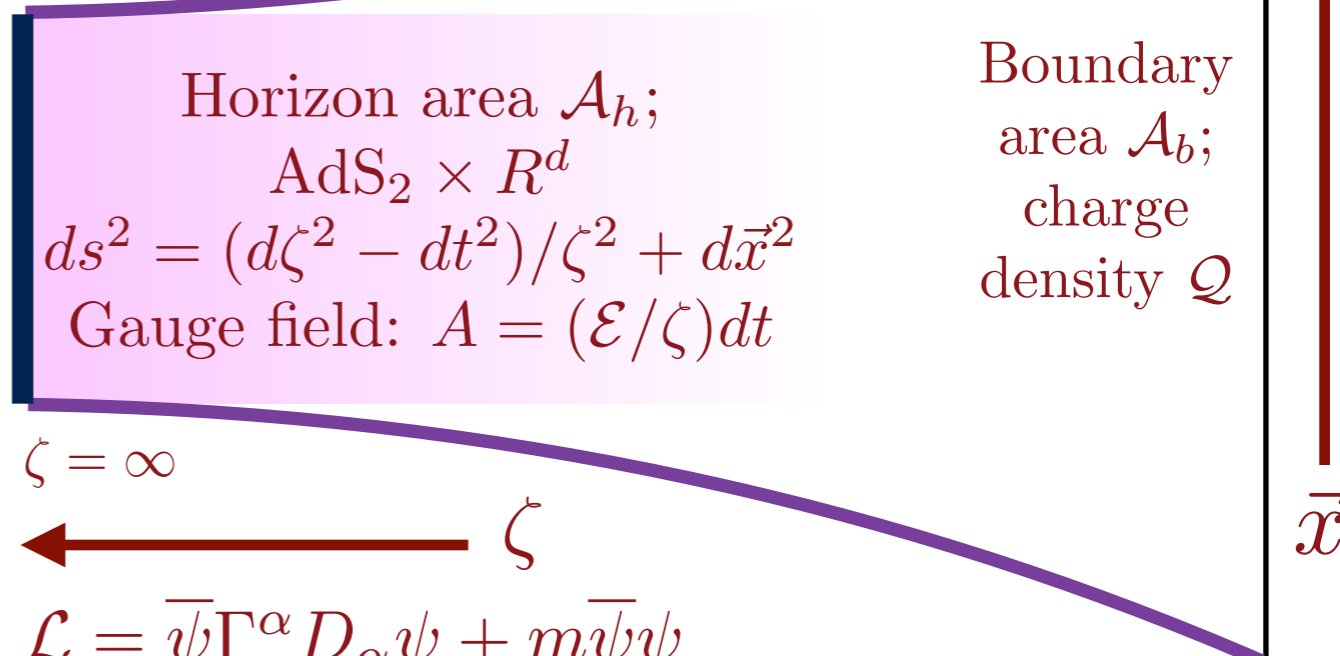
Known 'equation of state' determines \mathcal{E} as a function of Q

Microscopic zero temperature entropy density, \mathcal{S} , obeys

$$\frac{\partial \mathcal{S}}{\partial Q} = 2\pi\mathcal{E}$$

Evidence for AdS₂ gravity dual of H

Einstein-Maxwell theory + cosmological constant



Horizon area \mathcal{A}_h ;
AdS₂ × R^d
 $ds^2 = (d\zeta^2 - dt^2)/\zeta^2 + d\vec{x}^2$
Gauge field: $A = (\mathcal{E}/\zeta)dt$

Boundary area \mathcal{A}_b ;
charge density \mathcal{Q}

$\zeta = \infty$

ζ

$$\mathcal{L} = \bar{\psi} \Gamma^\alpha D_\alpha \psi + m \bar{\psi} \psi$$

Local fermion density of states

$$\rho(\omega) \sim \begin{cases} \omega^{-1/2}, & \omega > 0 \\ e^{-2\pi\mathcal{E}} |\omega|^{-1/2}, & \omega < 0. \end{cases}$$

'Equation of state' relating \mathcal{E} and \mathcal{Q} depends upon the geometry of spacetime far from the AdS₂

Black hole thermodynamics (classical general relativity) yields

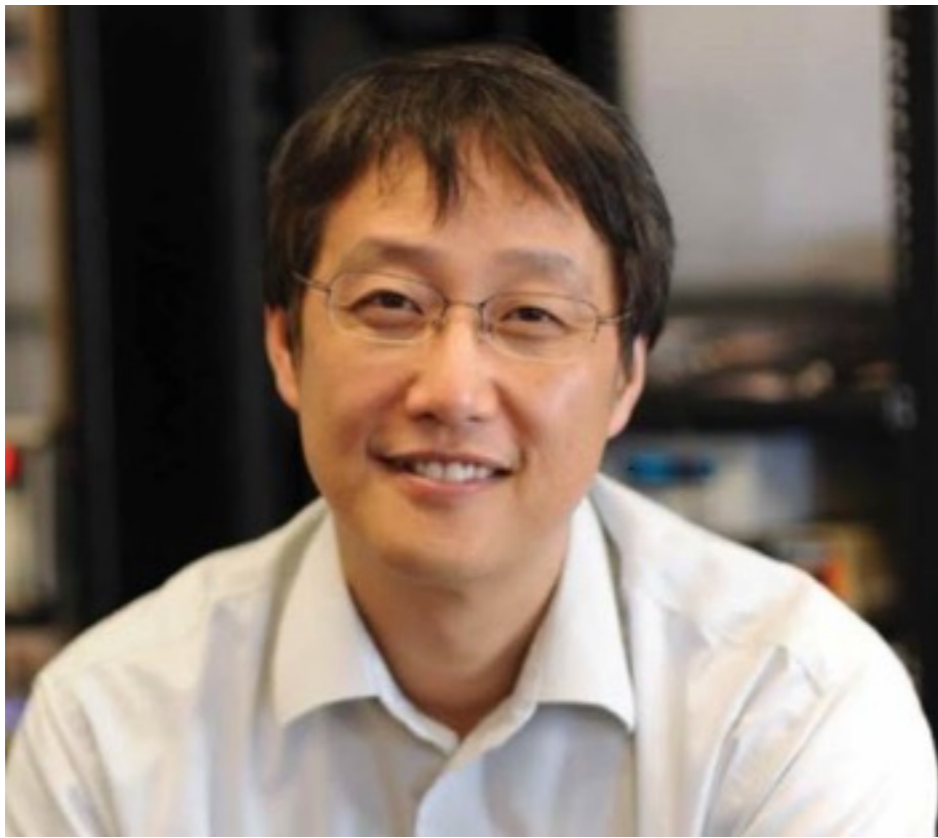
$$\frac{\partial \mathcal{S}_{\text{BH}}}{\partial Q} = 2\pi\mathcal{E}$$

Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene

Quantum matter without quasiparticles

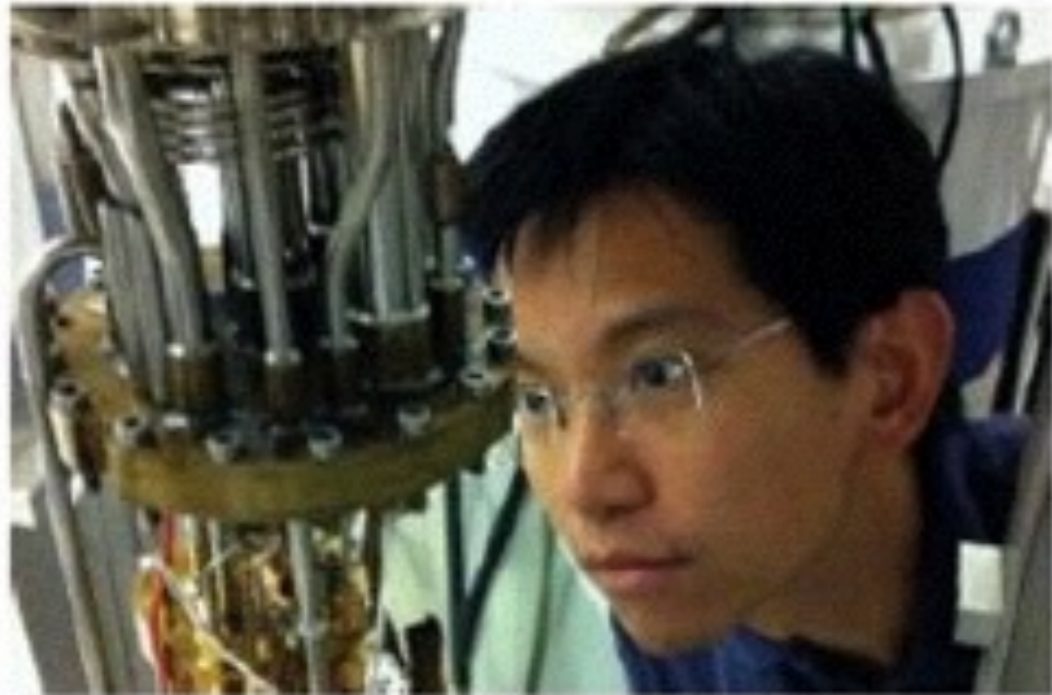
1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene



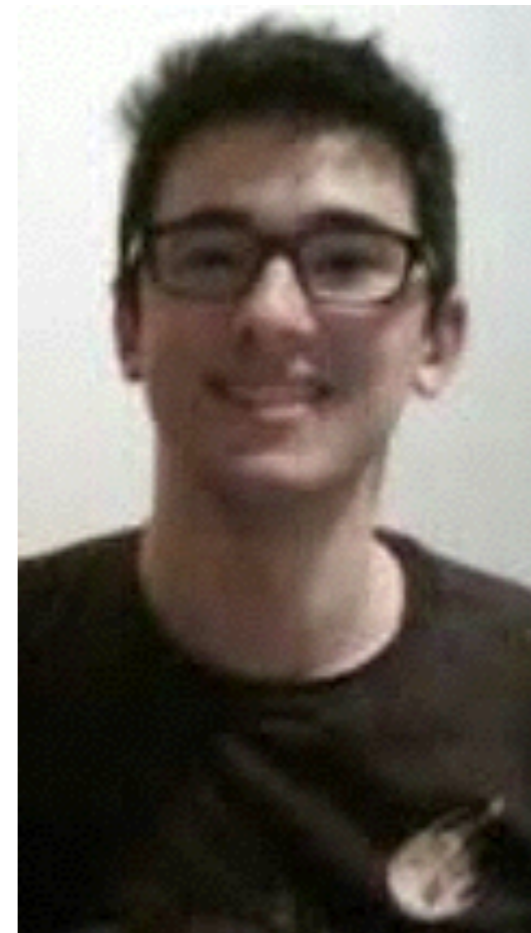
Philip Kim



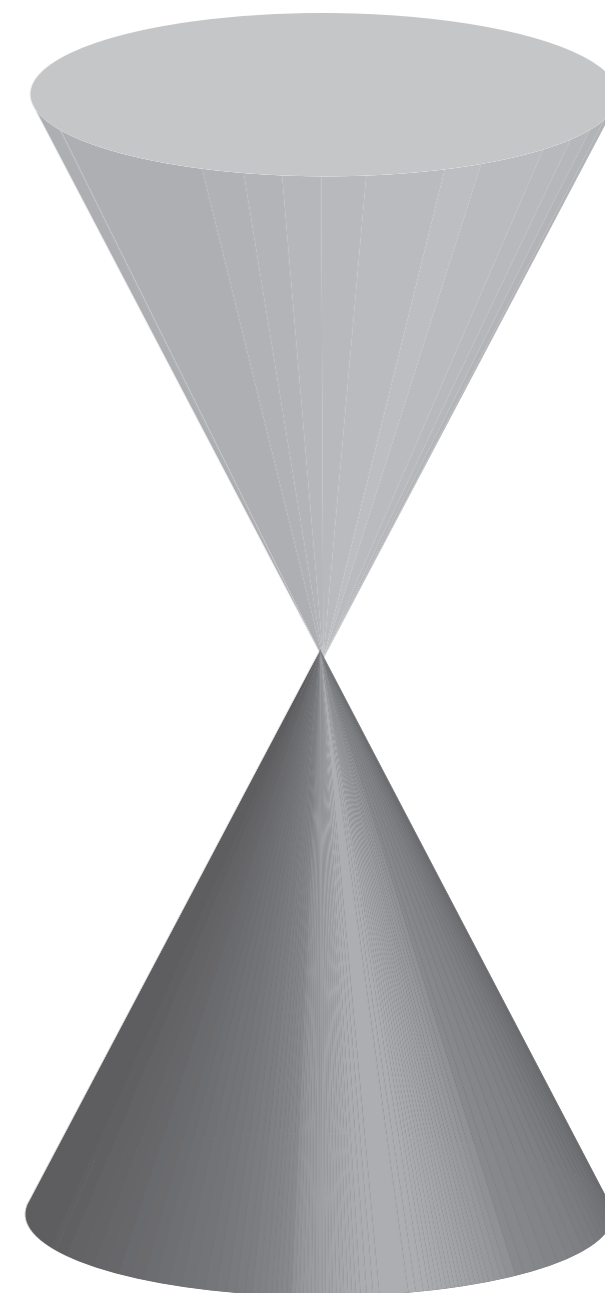
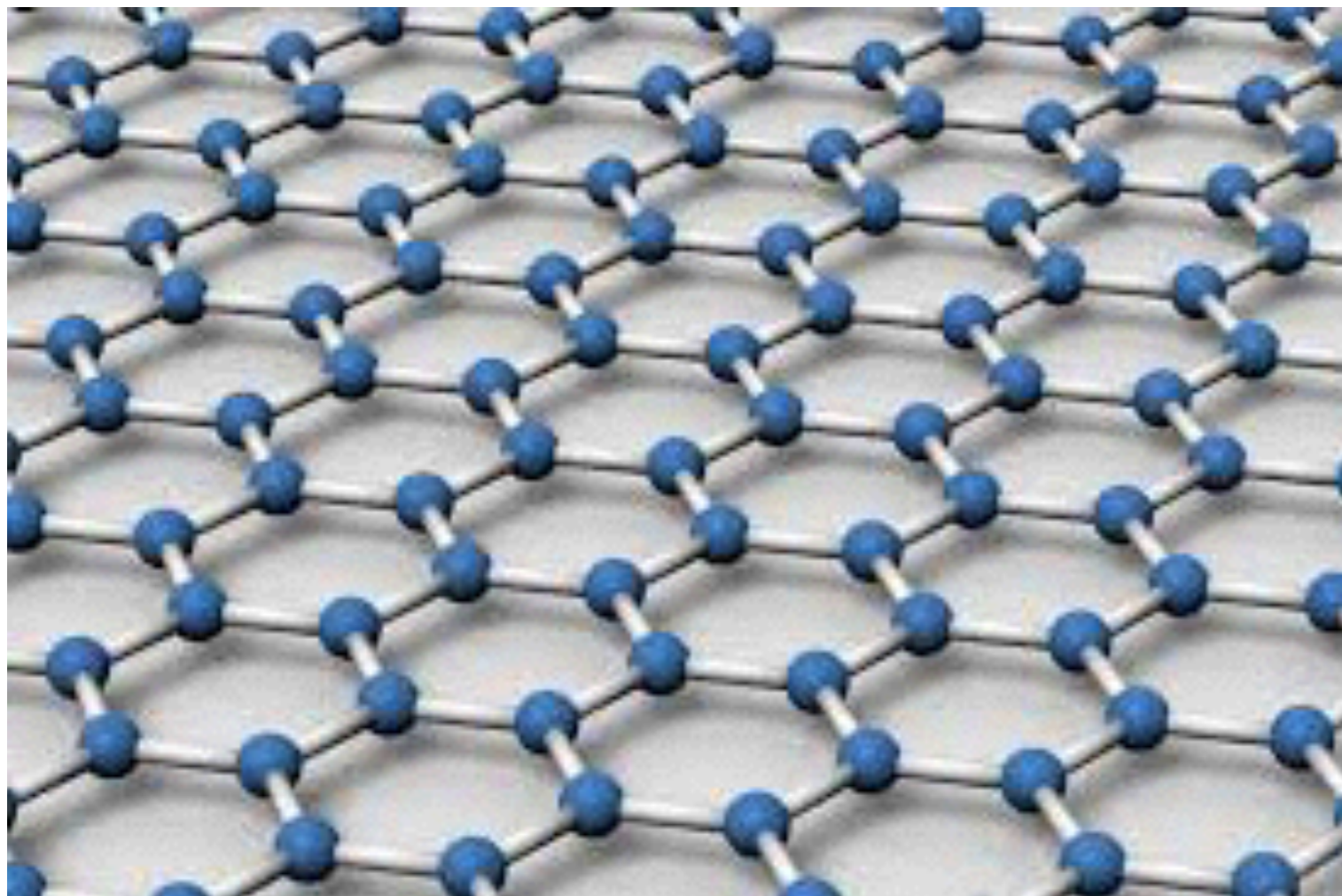
Jesse Crossno

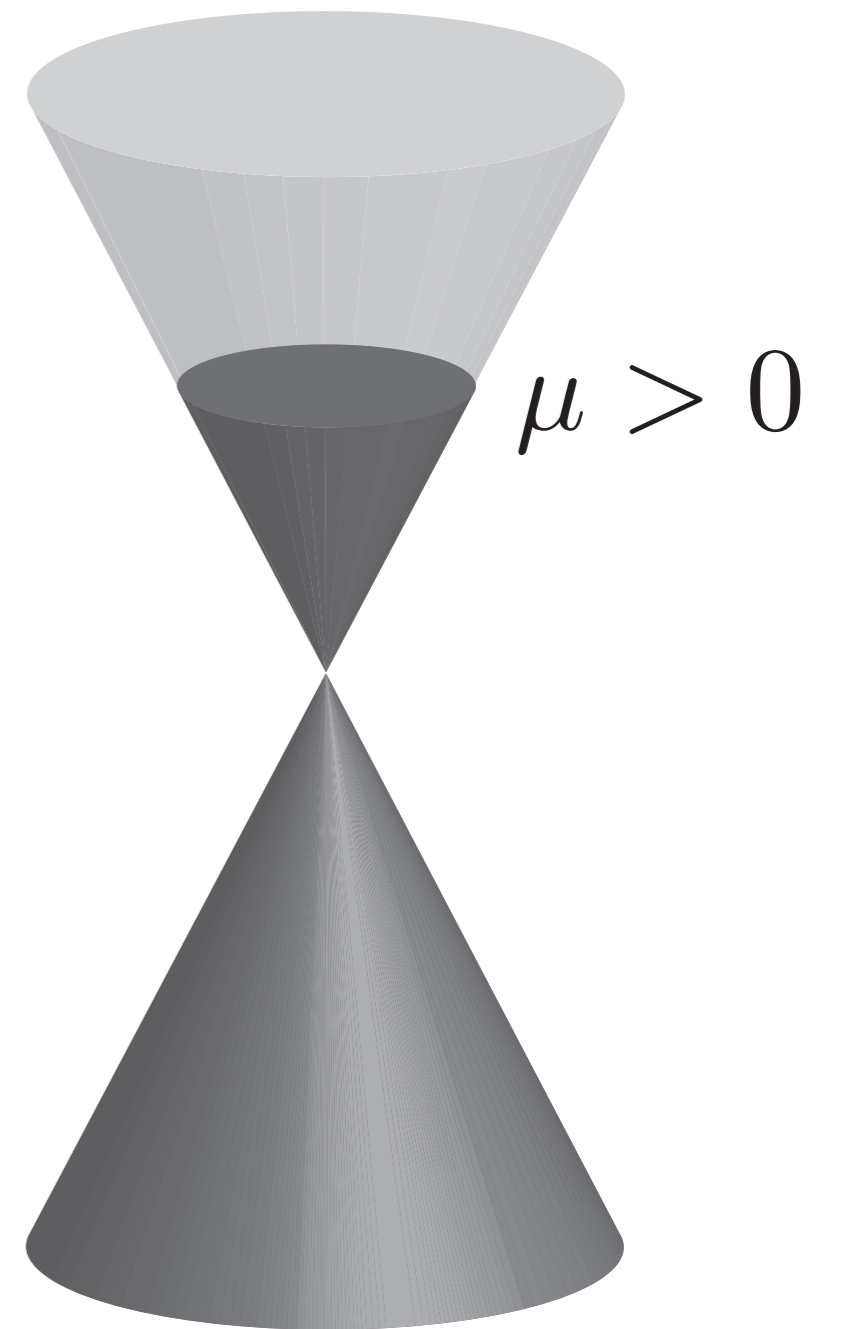


Kin Chung Fong

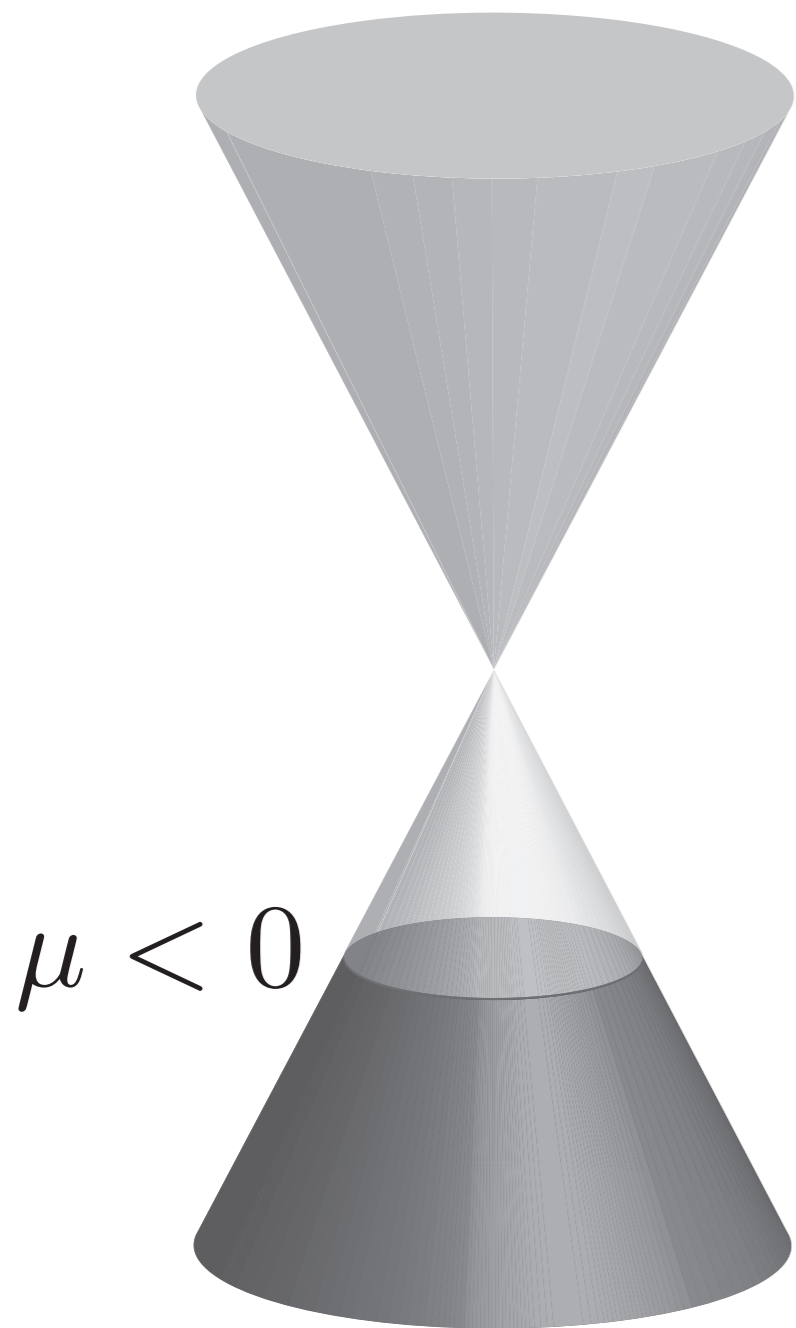


Andrew Lucas

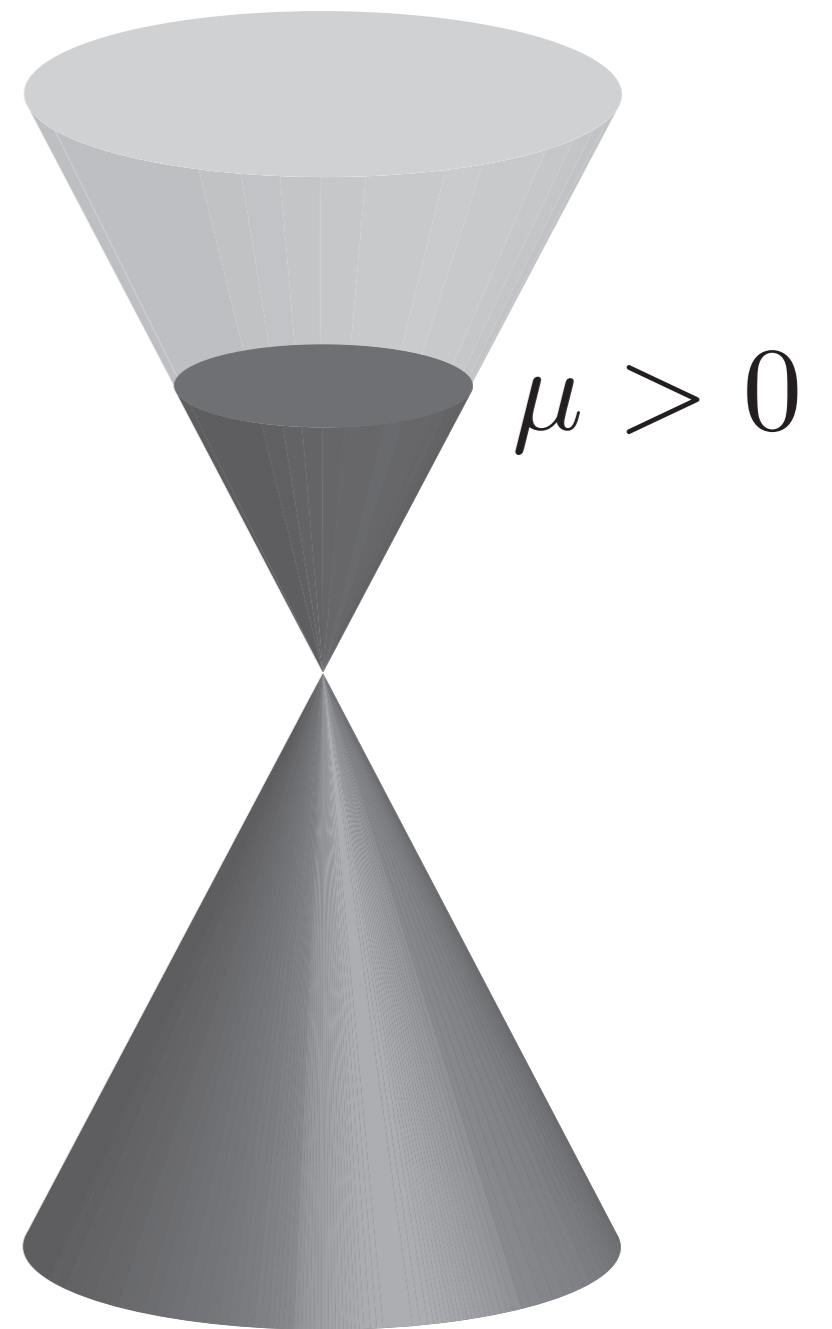




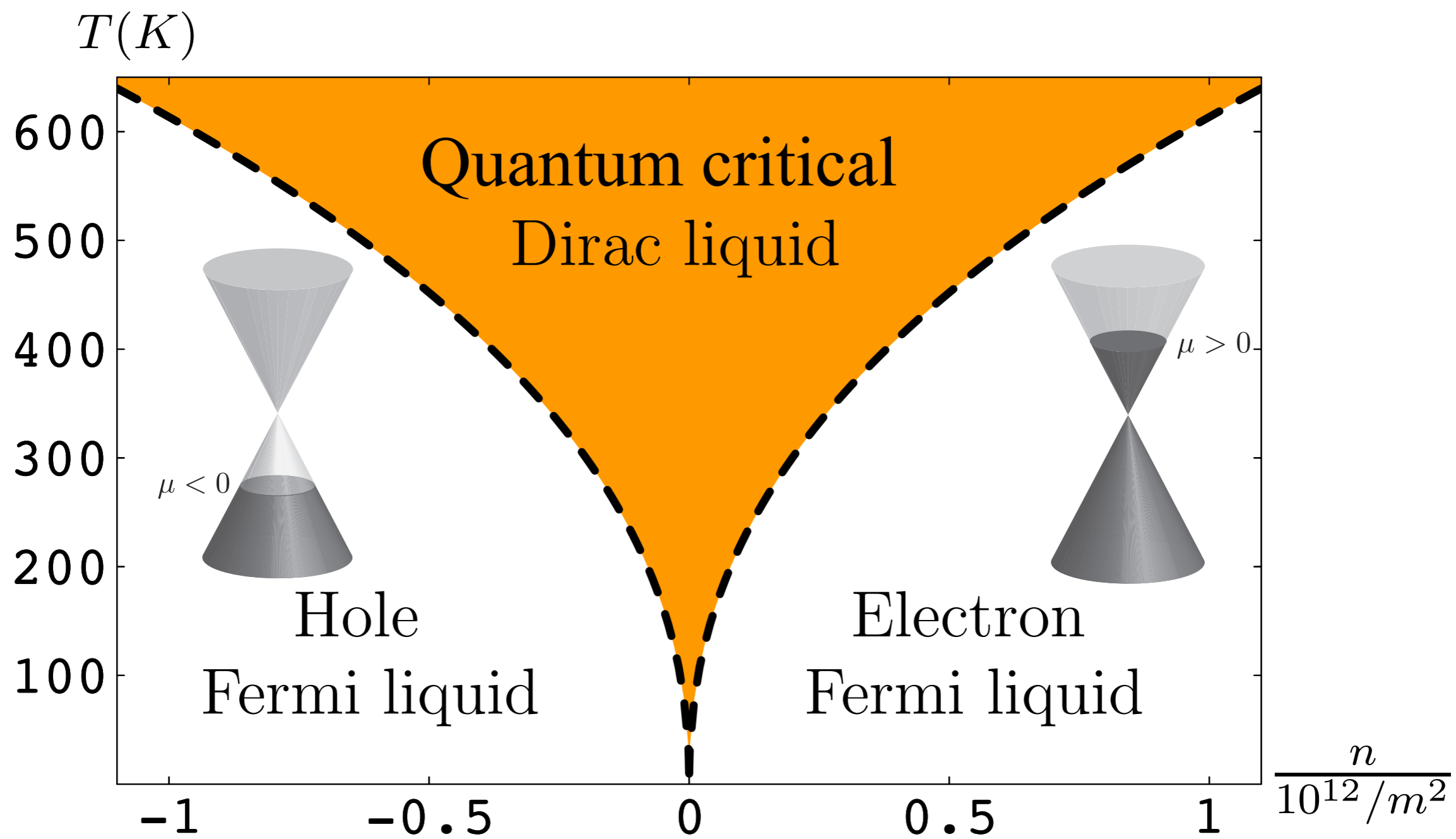
**Electron
Fermi surface**



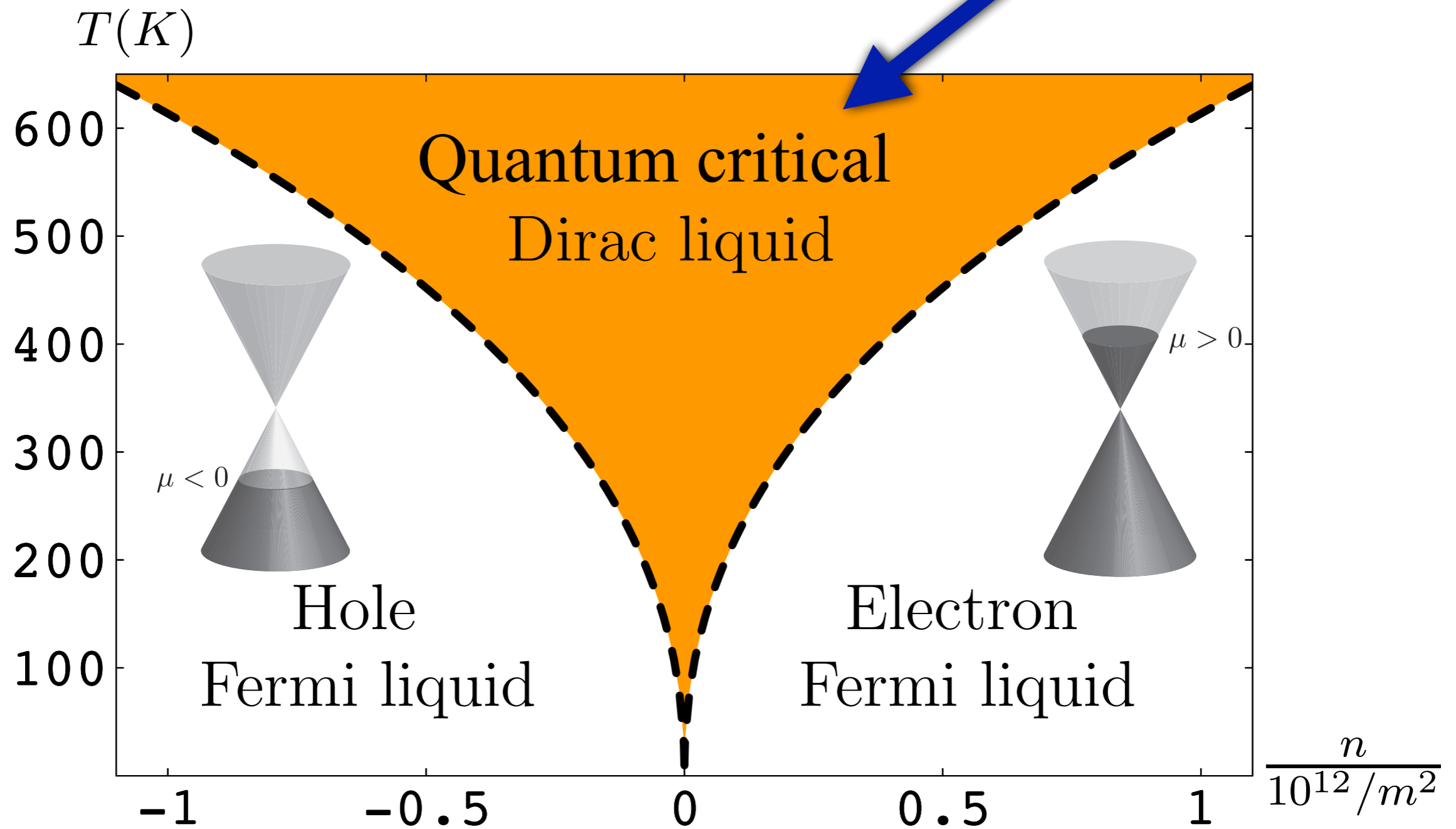
**Hole
Fermi surface**



**Electron
Fermi surface**



(slightly less)
strange metal



Transport in Strange Metals

universal constraints on transport

hydrodynamics

[Forster '70s]

[Hartnoll, others]

[Lucas, Sachdev PRB]

few conserved quantities

[Lucas 1506]

[Donos, Gauntlett 1506]

long time dynamics;
“renormalized IR fluid”
emerges

perturbative
limit

memory matrix

appropriate microscopics
for cuprates

[Lucas JHEP]

holography

matrix large N theory;
non-perturbative computations

figure from [Lucas, Sachdev, *Physical Review* **B91** 195122 (2015)]

S. A. Hartnoll, P. K. Kovtun, M. Müller, and S. Sachdev, PRB **76**, 144502 (2007)

Transport in Strange Metals

Recall that in a Fermi liquid, the Lorenz ratio $L = \kappa / (T\sigma)$, where κ is the thermal conductivity, and σ is the conductivity, is given by

$$L = \frac{\pi^2 k_B^2}{3e^2}$$

Transport in Strange Metals

Recall that in a Fermi liquid, the Lorenz ratio $L = \kappa/(T\sigma)$, where κ is the thermal conductivity, and σ is the conductivity, is given by

$$L = \frac{\pi^2 k_B^2}{3e^2}$$

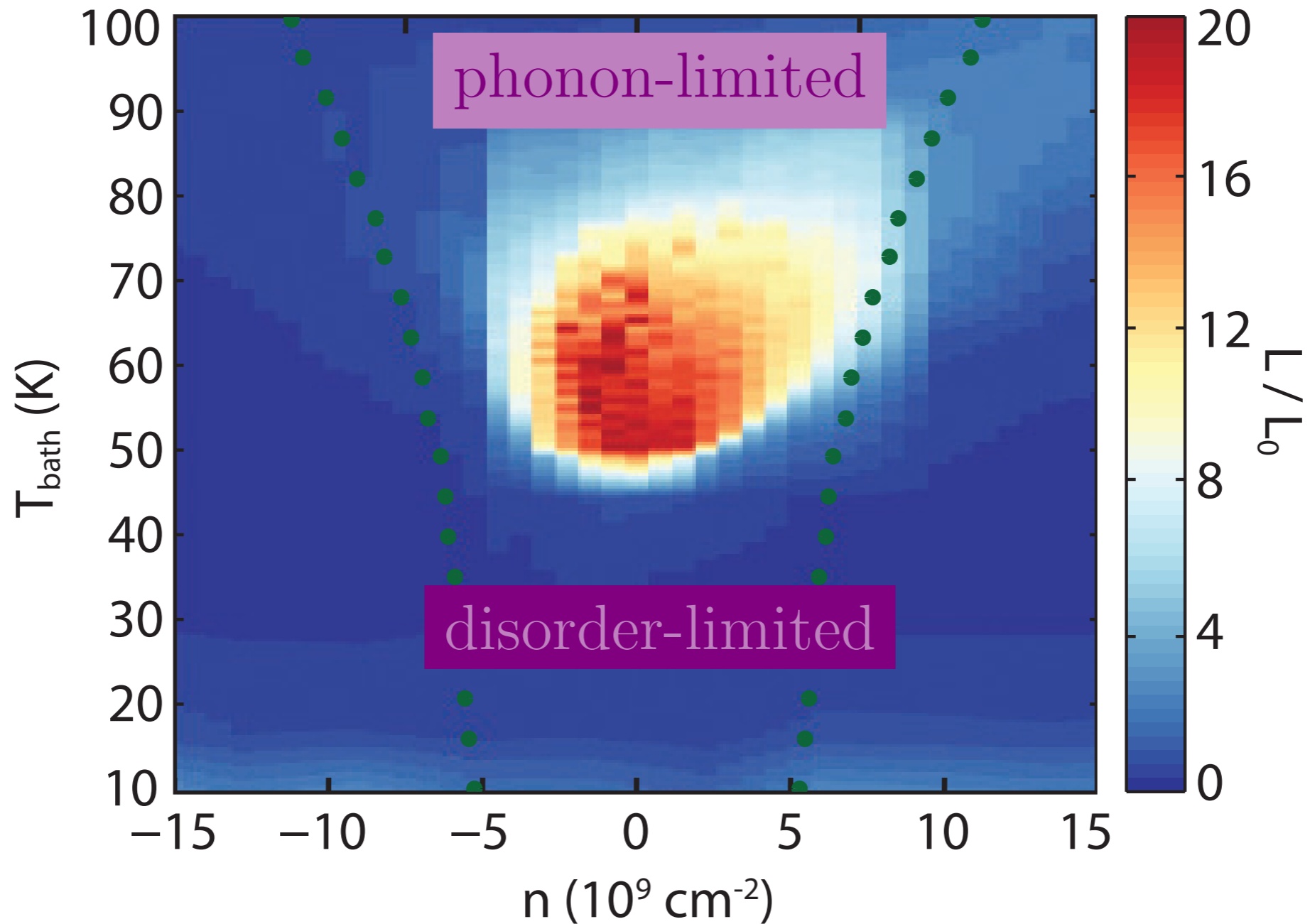
In contrast, for a strange metal with a “relativistic” Hamiltonian, L diverges as the charge density $Q \rightarrow 0$, and then the impurity momentum relaxation time $\tau_{\text{imp}} \rightarrow \infty$:

$$L = \frac{\mathcal{H}\tau_{\text{imp}}}{T^2\sigma_Q} \frac{1}{(1 + Q^2\tau_{\text{imp}}/(\mathcal{H}\sigma_Q))^2},$$

where \mathcal{H} is the enthalpy density. This divergence happens because at $Q = 0$, κ diverges as $\tau_{\text{imp}} \rightarrow \infty$, while $\sigma = \sigma_Q$, an intrinsic, finite, “quantum critical” conductivity.

Dirac Fluid in Graphene

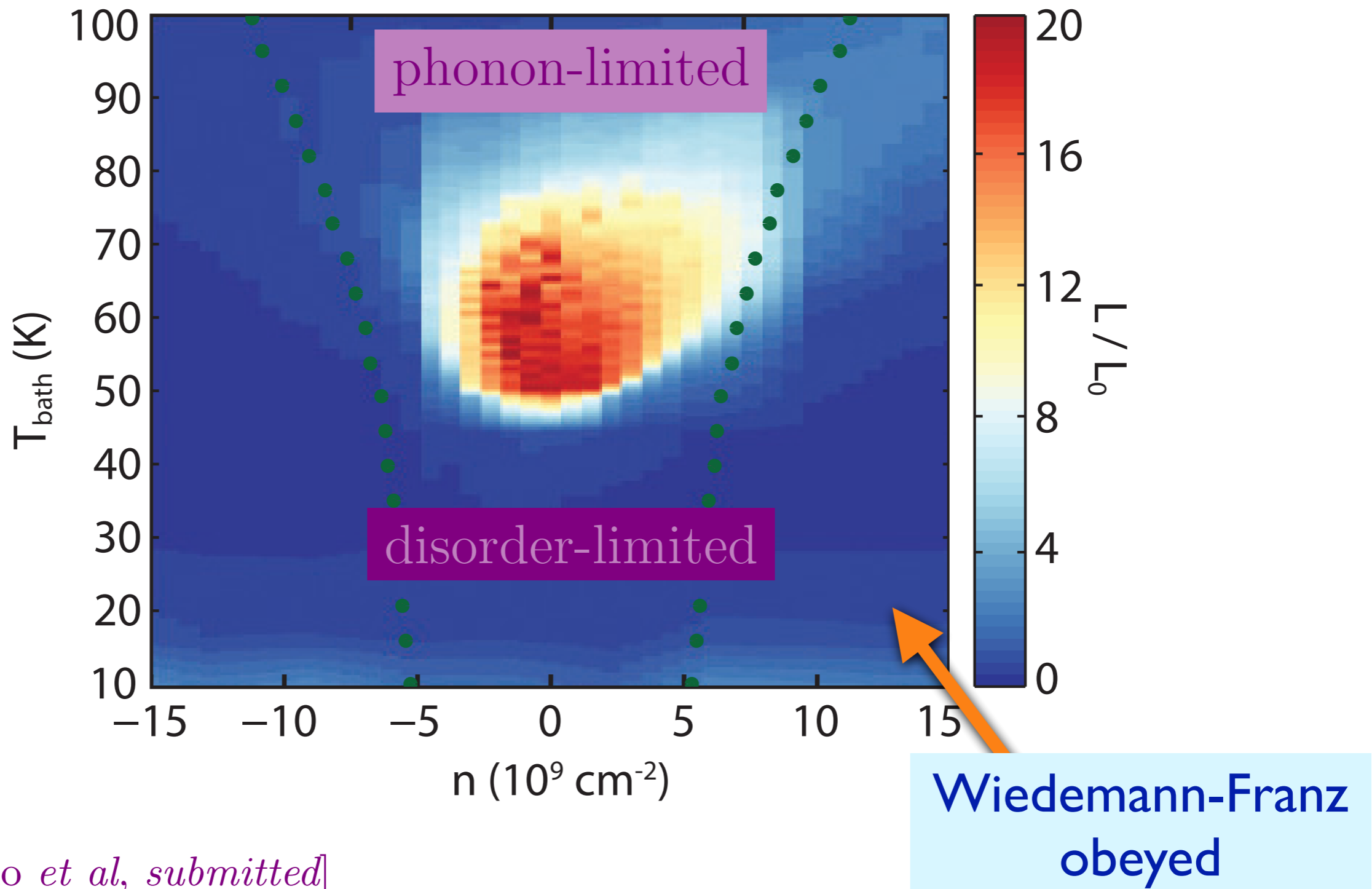
Wiedemann-Franz Law Violations in Experiment



[Crossno *et al*, *submitted*]

Dirac Fluid in Graphene

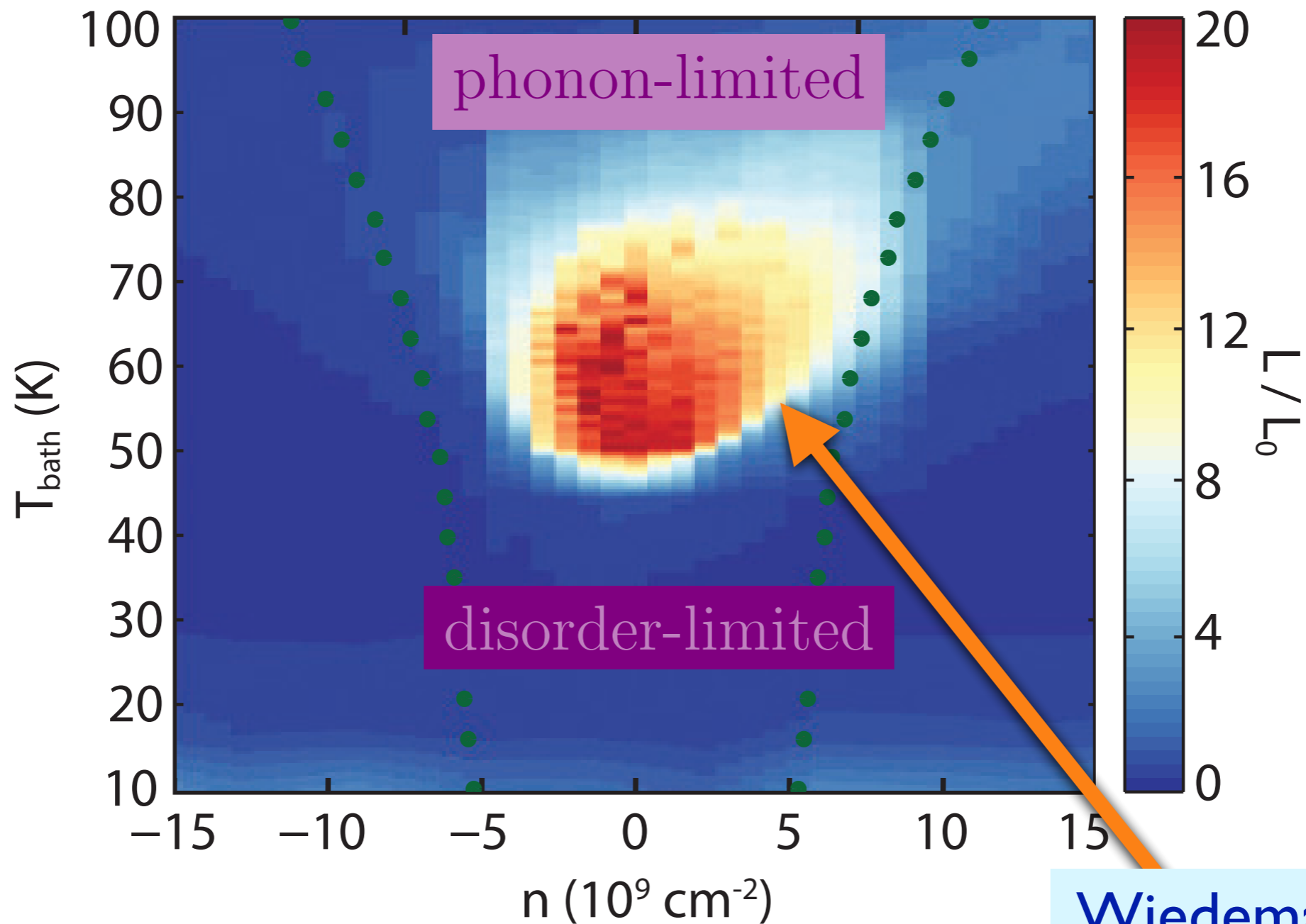
Wiedemann-Franz Law Violations in Experiment



[Crossno *et al*, *submitted*]

Dirac Fluid in Graphene

Wiedemann-Franz Law Violations in Experiment



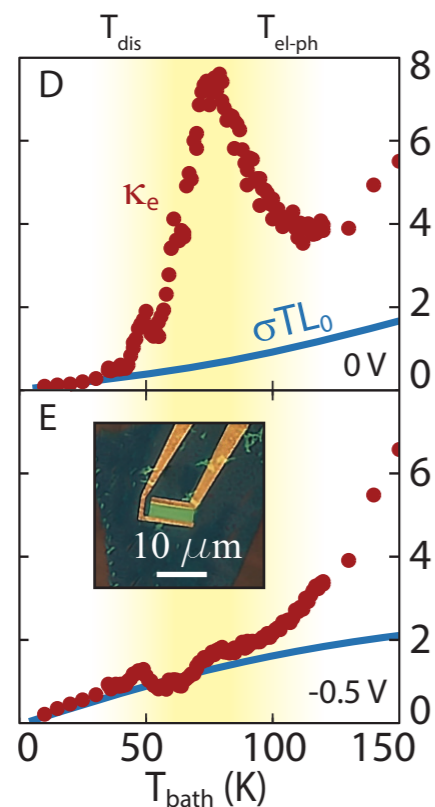
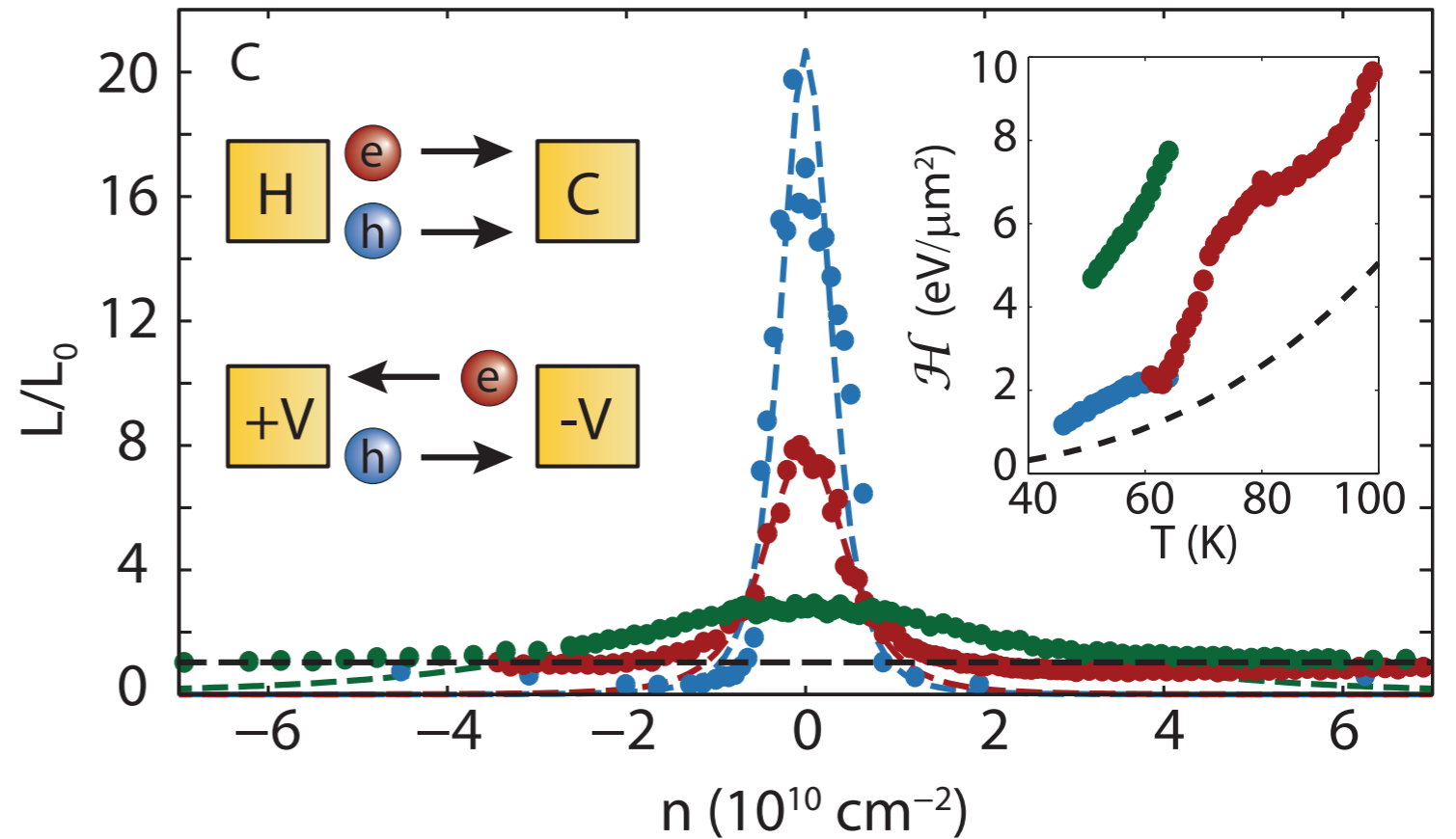
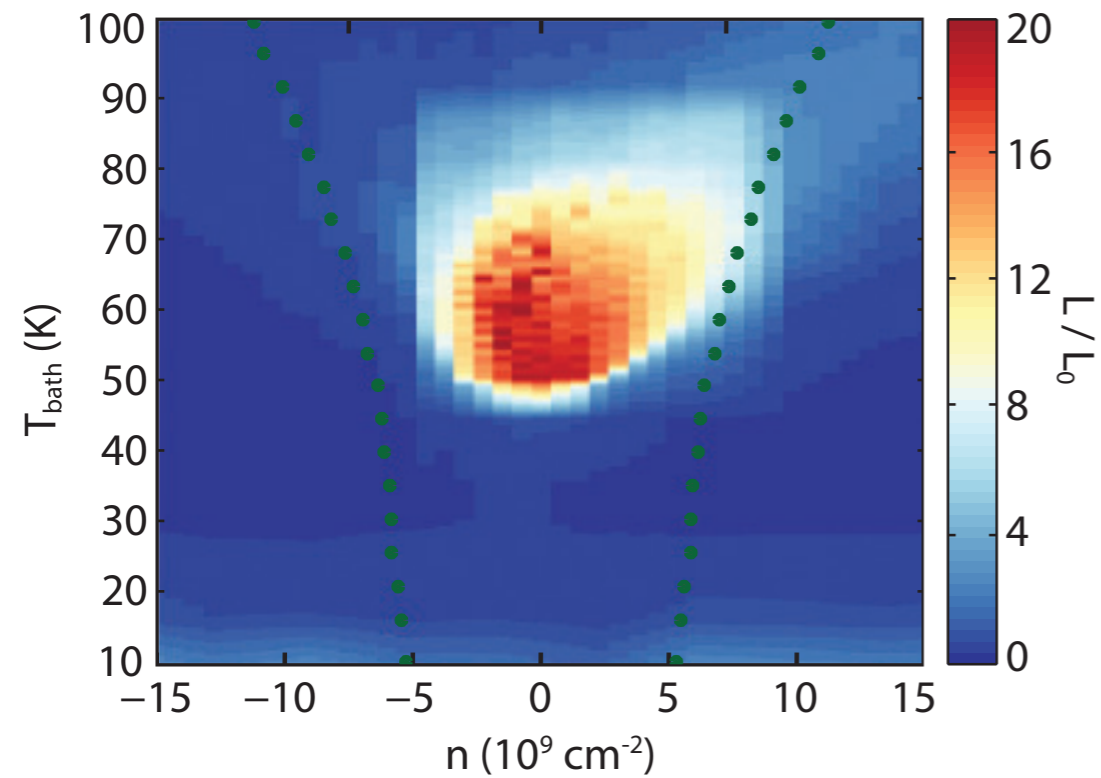
**Wiedemann-Franz
violated !**

[Crossno *et al*, *submitted*]

(submitted)

Observation of the Dirac fluid and the breakdown of the Wiedemann-Franz law in graphene

Jesse Crossno,^{1,2} Jing K. Shi,¹ Ke Wang,¹ Xiaomeng Liu,¹ Achim Harzheim,¹ Andrew Lucas,¹ Subir Sachdev,^{1,3}
Philip Kim,^{1,2,*} Takashi Taniguchi,⁴ Kenji Watanabe,⁴ Thomas A. Ohki,⁵ and Kin Chung Fong^{5,†}



$$\begin{aligned} \text{Lorentz ratio } L &= \kappa / (T\sigma) \\ &= \frac{\mathcal{H}\tau_{\text{imp}}}{T^2\sigma_Q} \frac{1}{(1 + Q^2\tau_{\text{imp}}/(\mathcal{H}\sigma_Q))^2} \end{aligned}$$

Quantum matter without quasiparticles

1. A solvable model of an ordinary metal
2. A solvable model of a strange metal
3. Holography and charged black holes
4. The (slightly less) strange metal in graphene