

The SYK model: a window into black holes and non-Fermi liquids

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INSTITUTE FOR
ADVANCED STUDY

PHYSICS



HARVARD

Talk online: sachdev.physics.harvard.edu

1. SYK models

2. Time reparameterization soft mode

3. Charged black holes

4. Spin glass of $S=1/2$ $SU(2)$ spins

5. Random t - J model

A simple model of a metal with quasiparticles

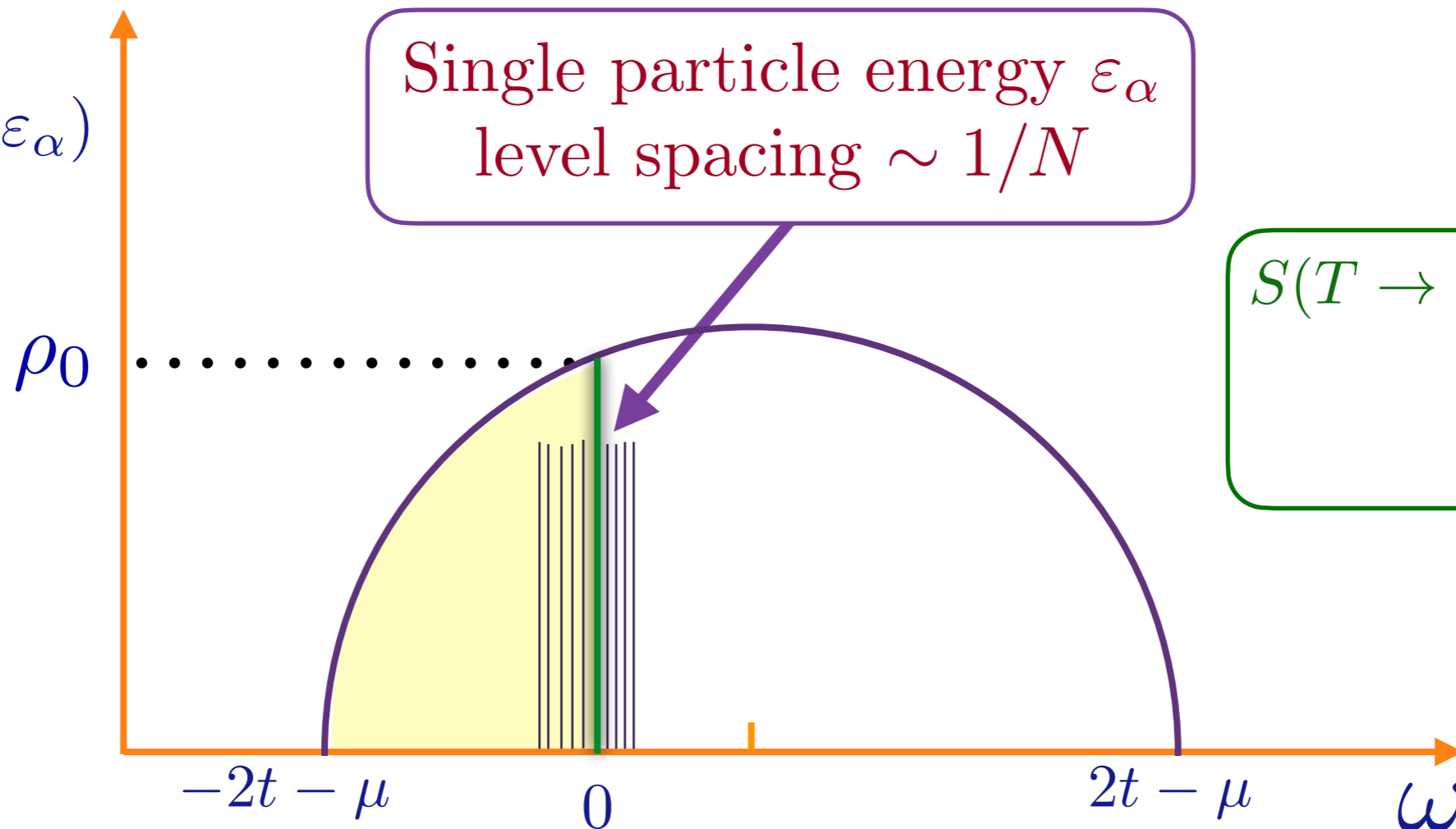
$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$\frac{1}{N} \sum_i c_i^\dagger c_i = Q$$

t_{ij} are independent random variables with $\overline{t_{ij}} = 0$ and $\overline{|t_{ij}|^2} = t^2$

$$\rho(\omega) = \frac{1}{N} \sum_\alpha \delta(\omega - \varepsilon_\alpha)$$



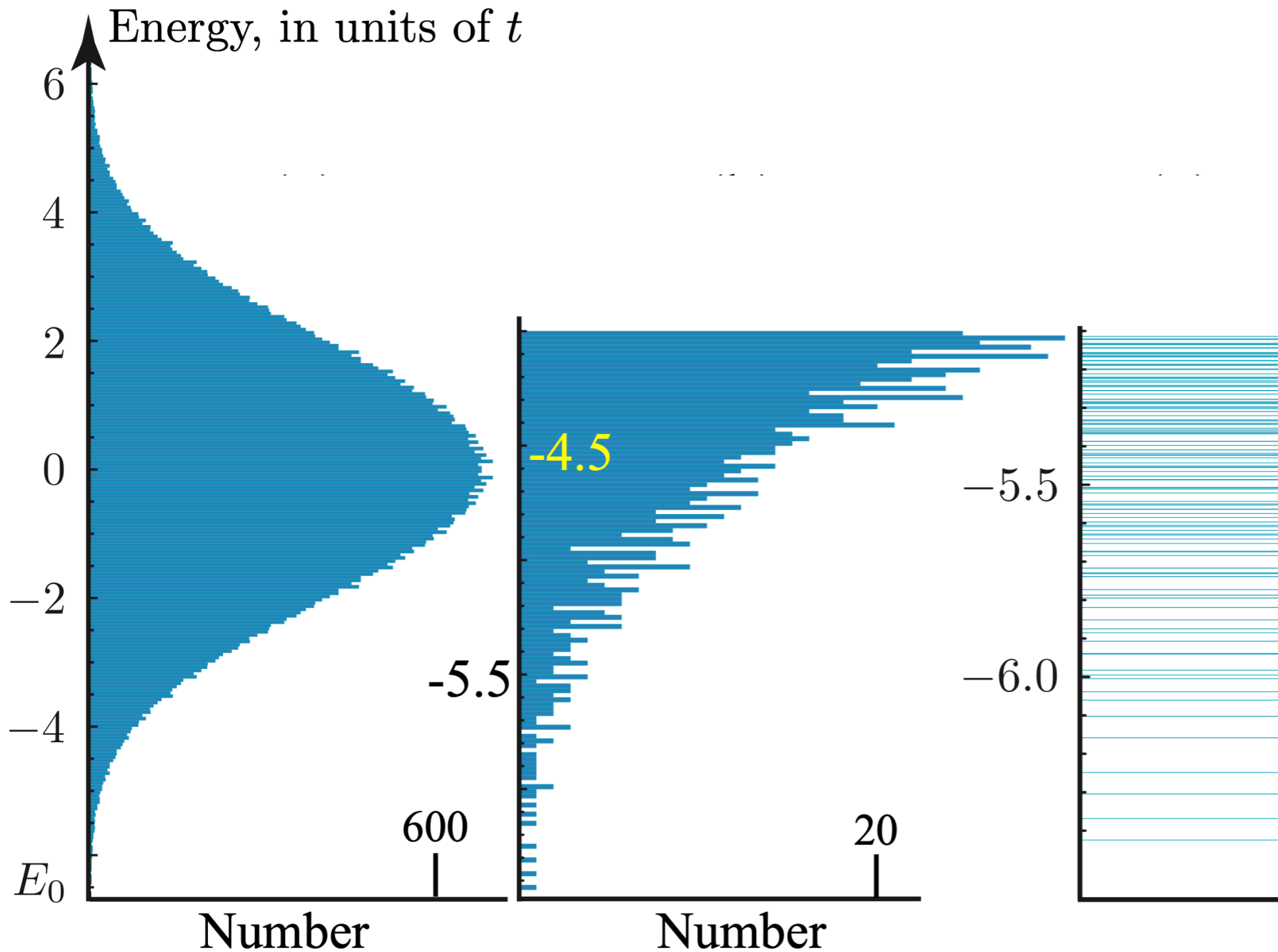
$$S(T \rightarrow 0) = N\gamma T$$

$$\gamma = \frac{\pi^2}{3} \rho_0$$

Many-body density of states

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$

For random matrix model:
 $E_0 + E_i = \sum_{\alpha} n_{\alpha} \epsilon_{\alpha}$
 $n_{\alpha} = 0, 1,$
occupation number

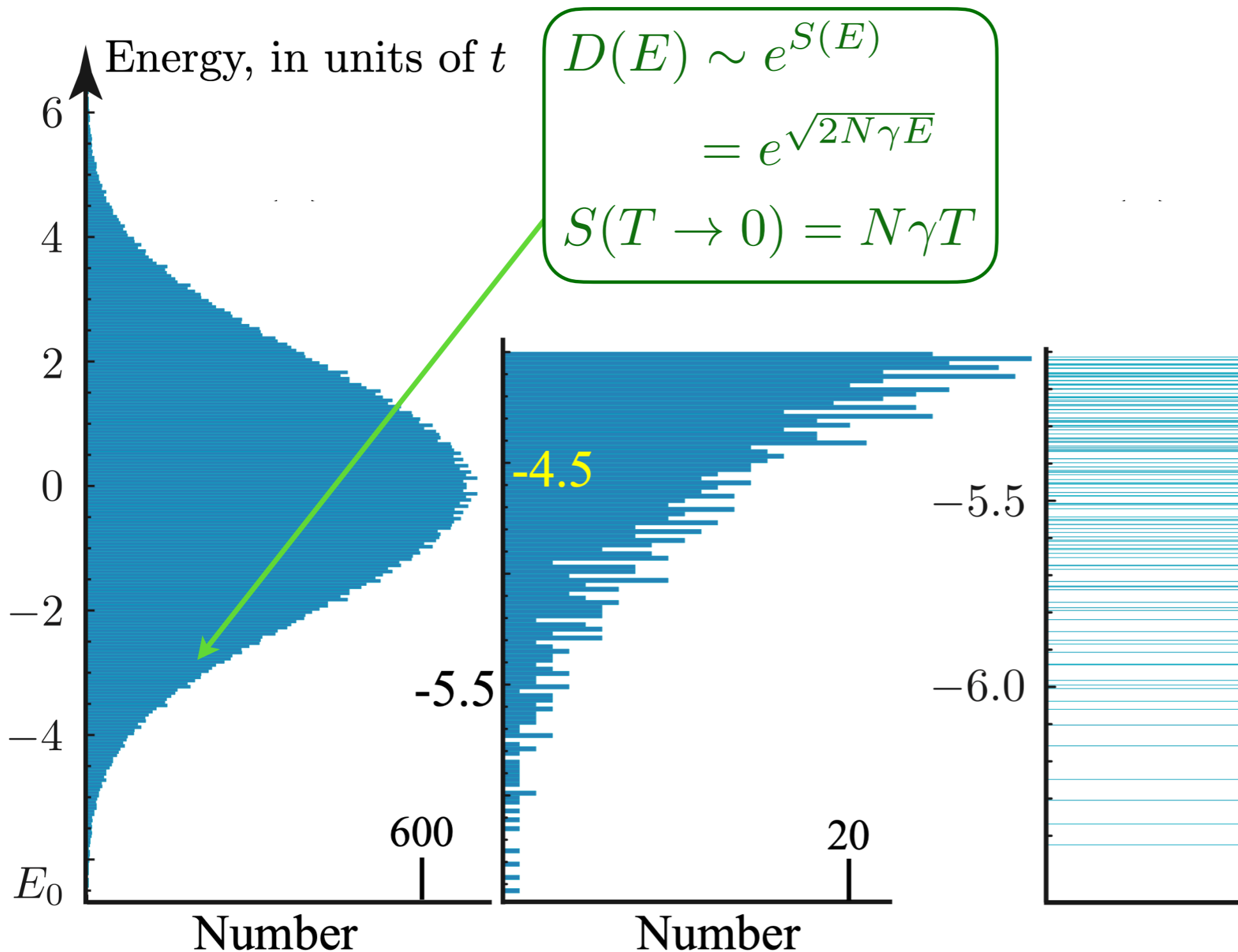


Random matrix model

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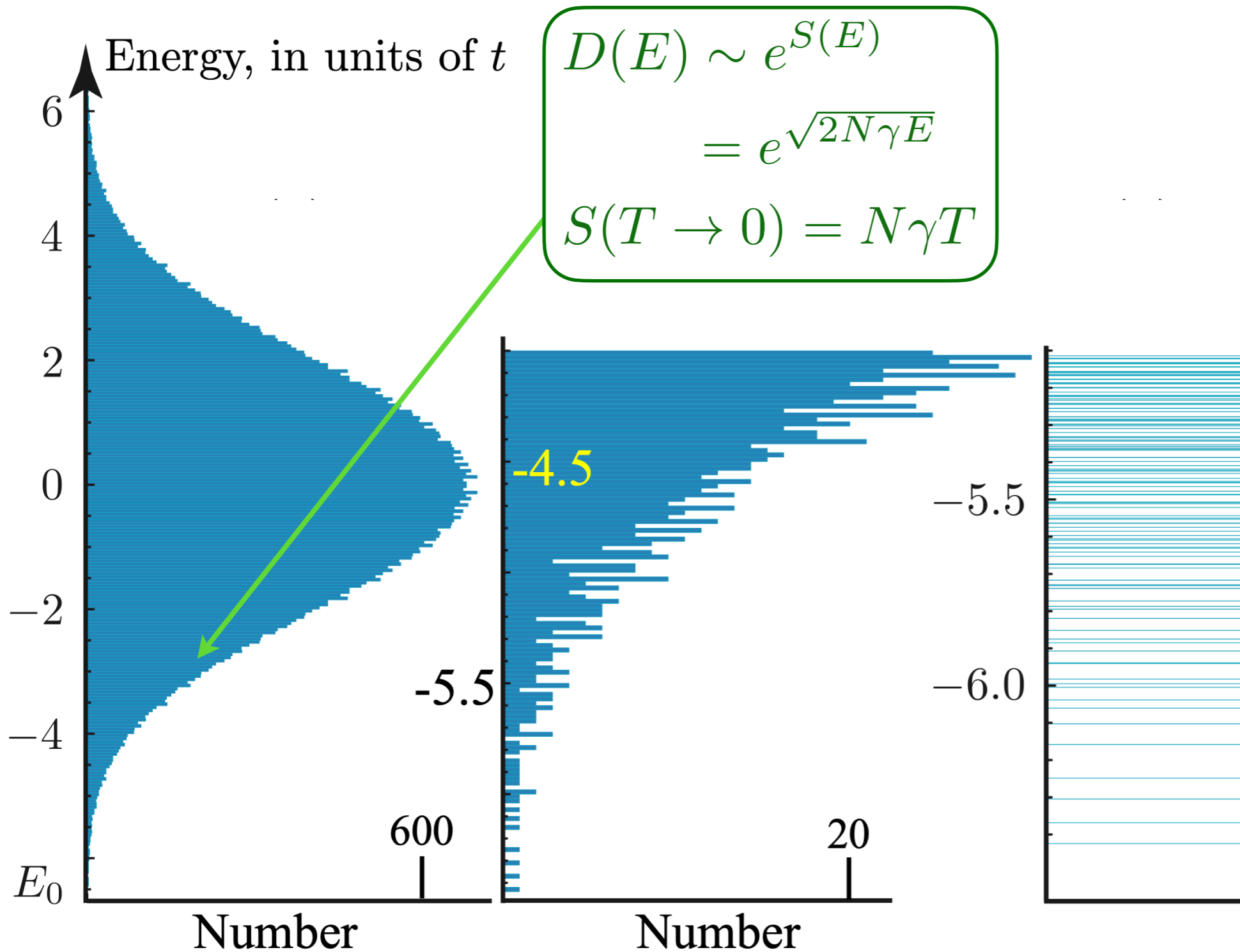


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Random matrix model

The Sachdev-Ye-Kitaev (SYK) model

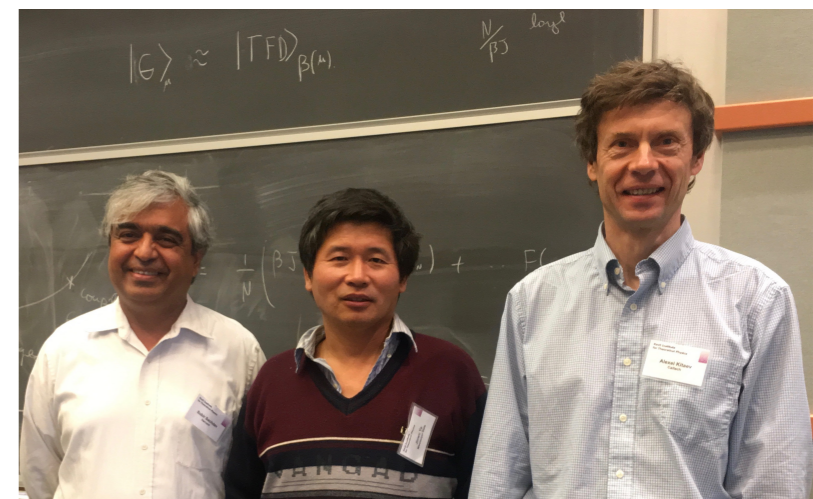
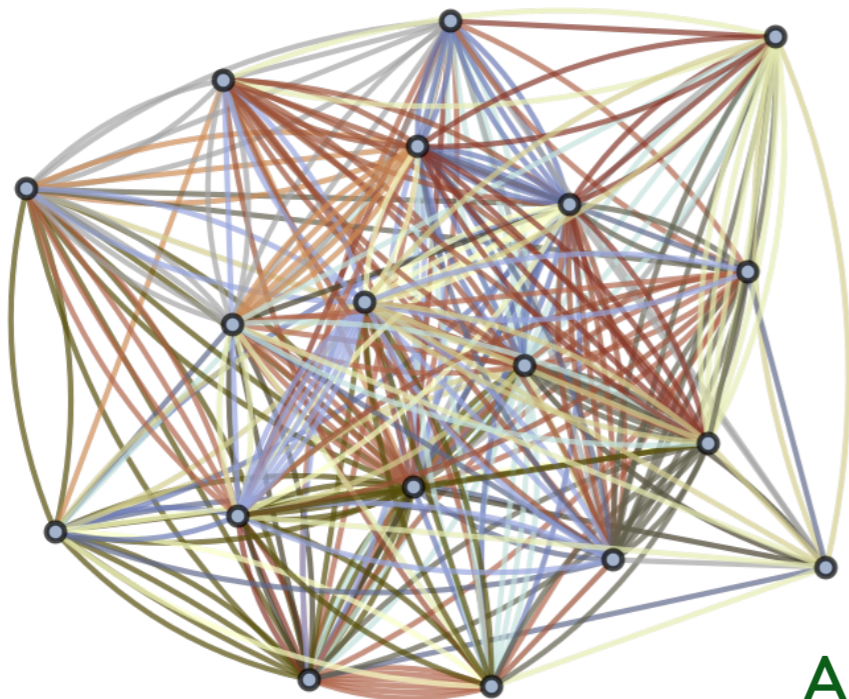
(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large N limit;
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} - \mu \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$$c_{\alpha} c_{\beta} + c_{\beta} c_{\alpha} = 0 \quad , \quad c_{\alpha} c_{\beta}^{\dagger} + c_{\beta}^{\dagger} c_{\alpha} = \delta_{\alpha\beta}$$

$$Q = \frac{1}{N} \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$U_{\alpha\beta;\gamma\delta}$ are independent random variables with $\overline{U_{\alpha\beta;\gamma\delta}} = 0$ and $\overline{|U_{\alpha\beta;\gamma\delta}|^2} = U^2$
 $N \rightarrow \infty$ yields critical strange metal.

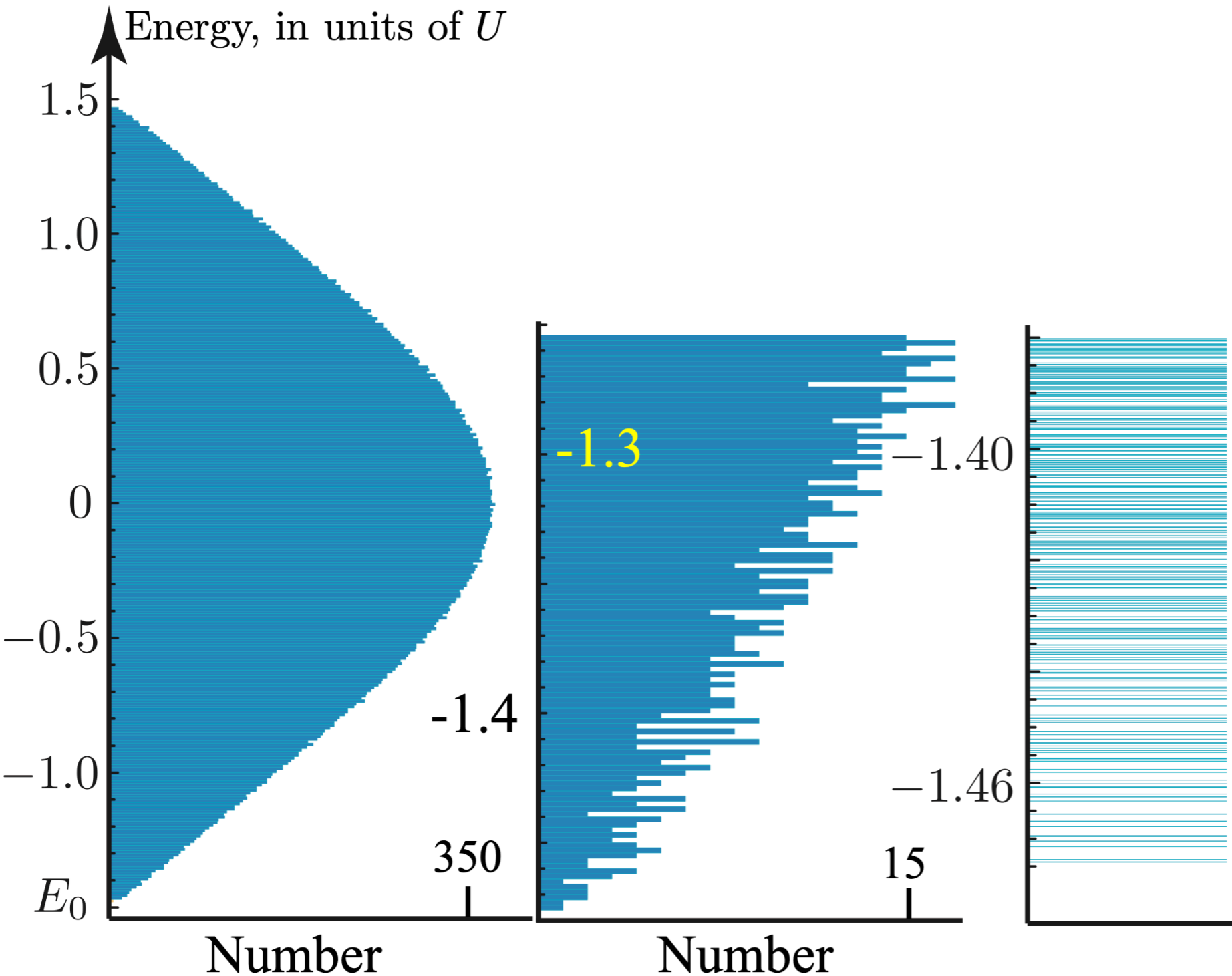


S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

Many-body density of states

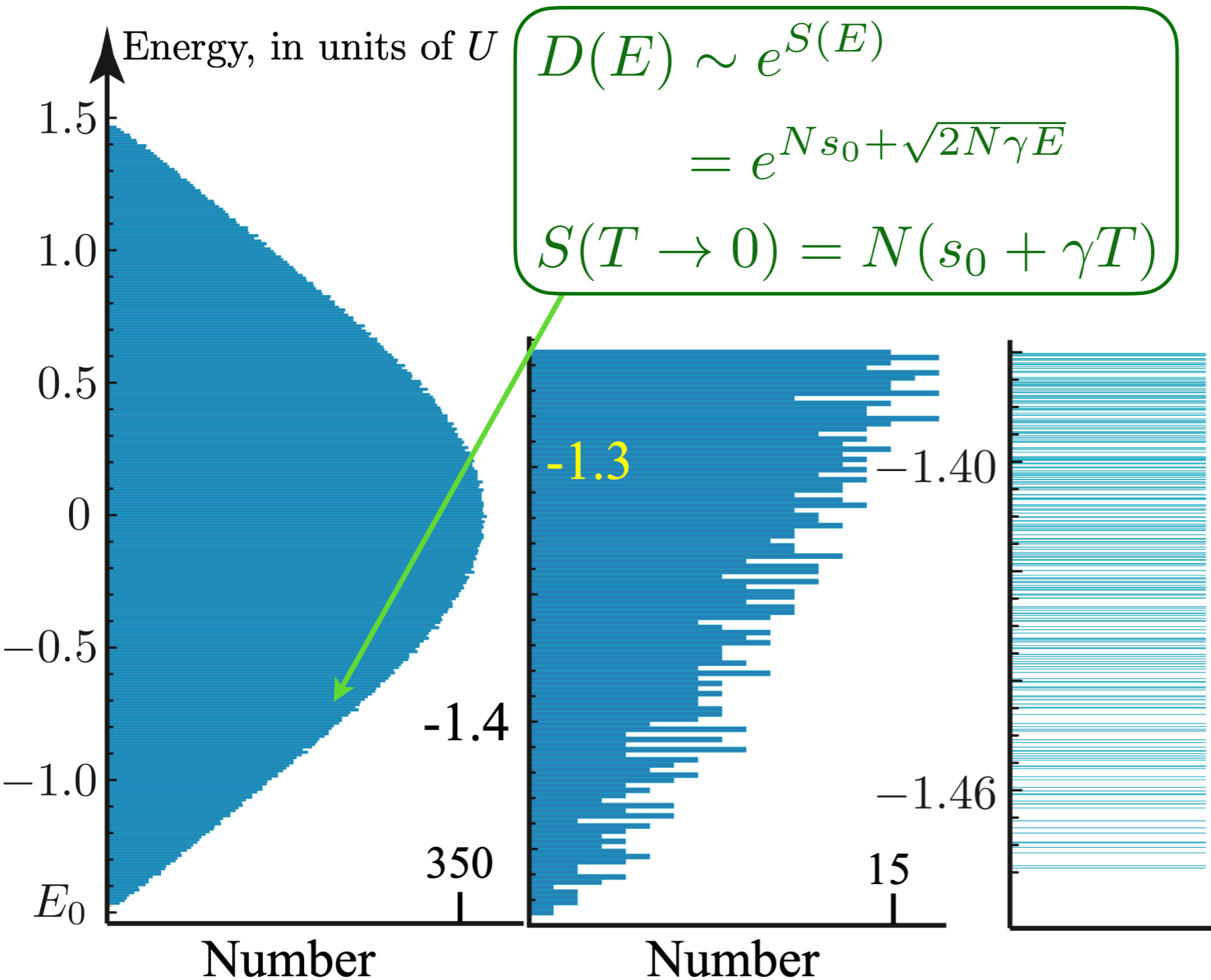
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Complex SYK model

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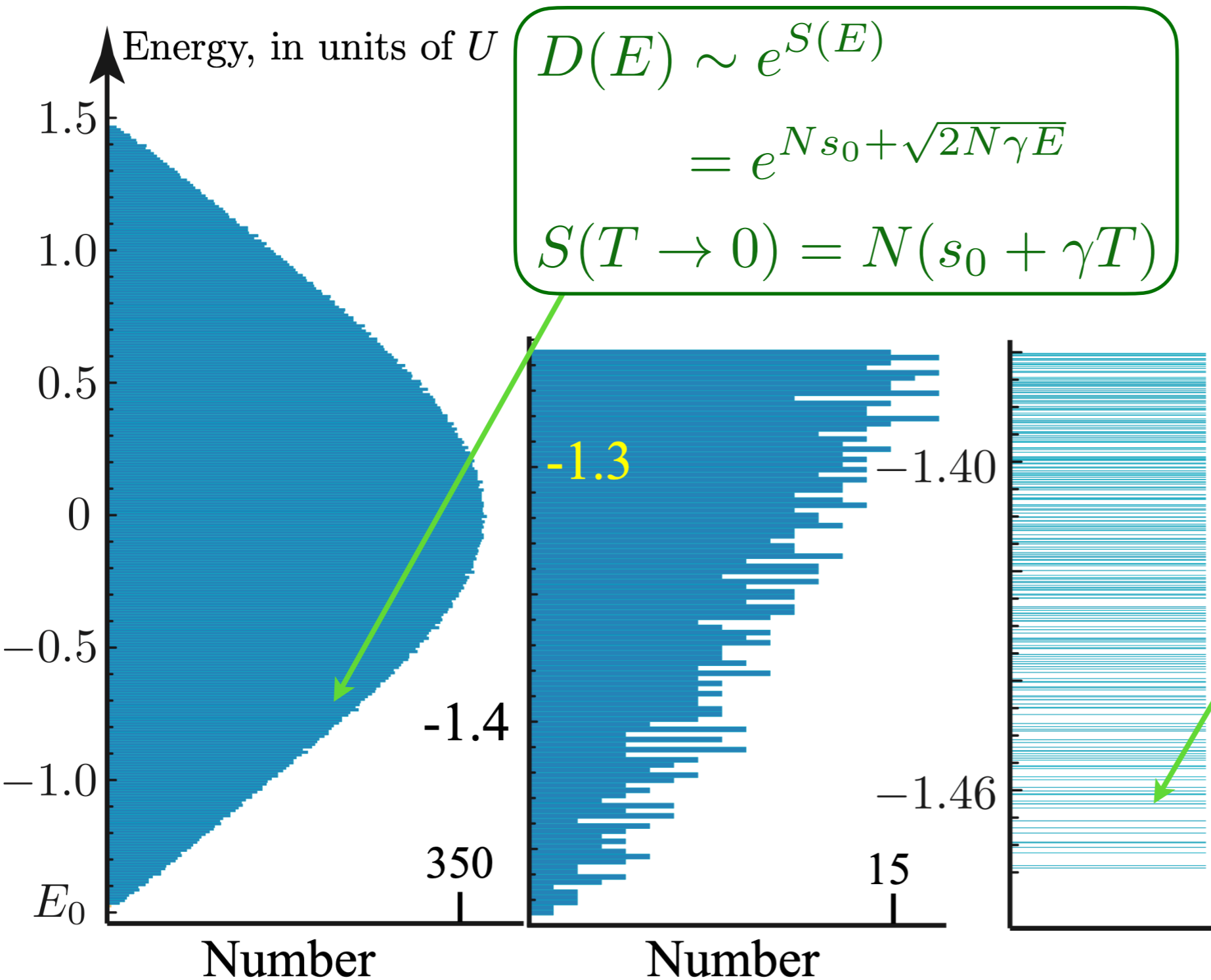
$$s_0 = 0.464848 \dots$$

A. Georges, O. Parcollet, and
S. Sachdev,
PRB **63**, 134406 (2001)

Complex SYK model

Many-body density of states

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim e^{S(E)}$$

$$= e^{Ns_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{Ns_0} \sqrt{2N\gamma E}$$

No quasiparticle decomposition of many-body states

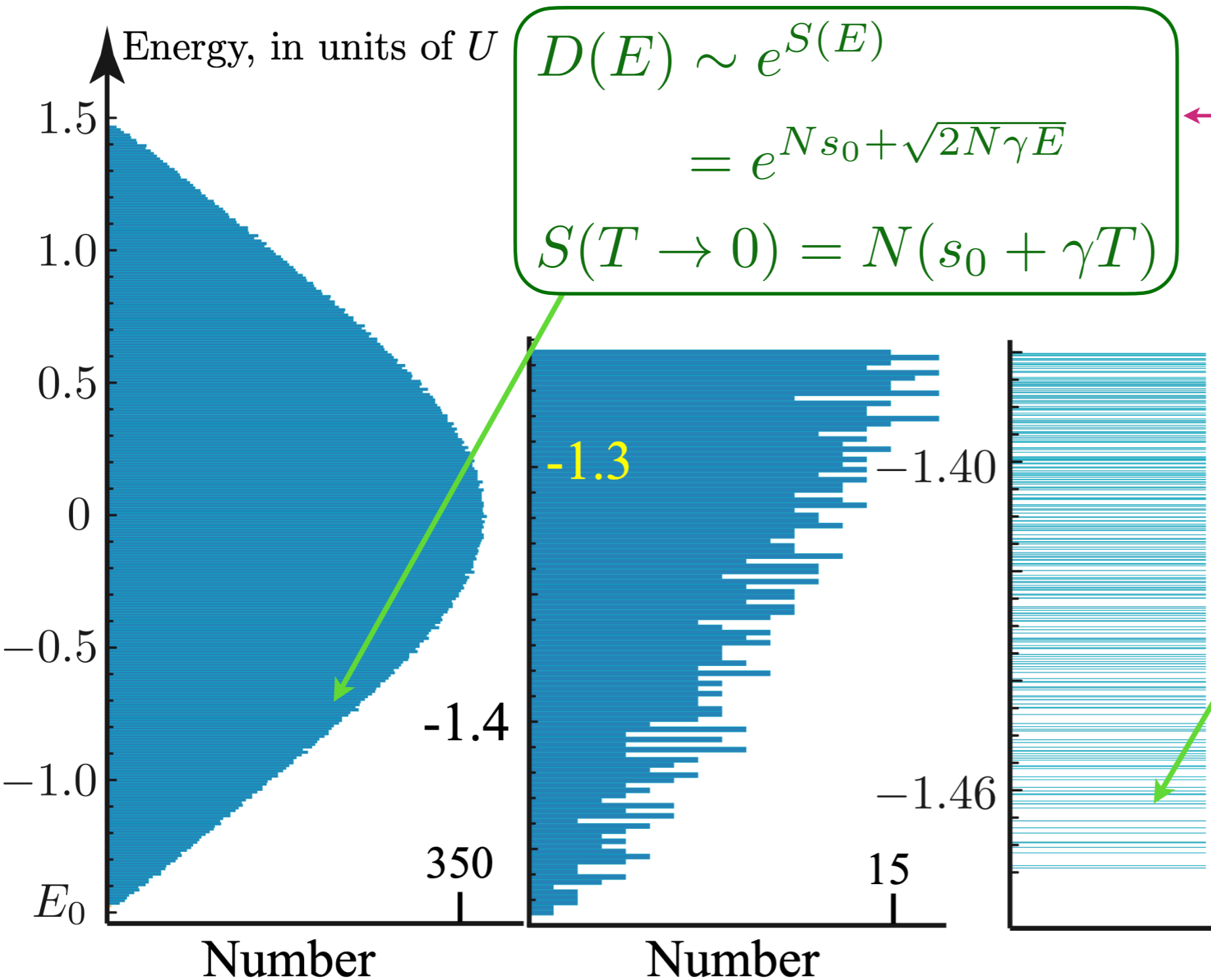
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$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim 2 e^{Ns_0} \sinh(\sqrt{2N\gamma E})$$

$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln \left(\frac{U}{T} \right)$$

$$D(E) \sim 2 e^{Ns_0} \sqrt{2N\gamma E}$$

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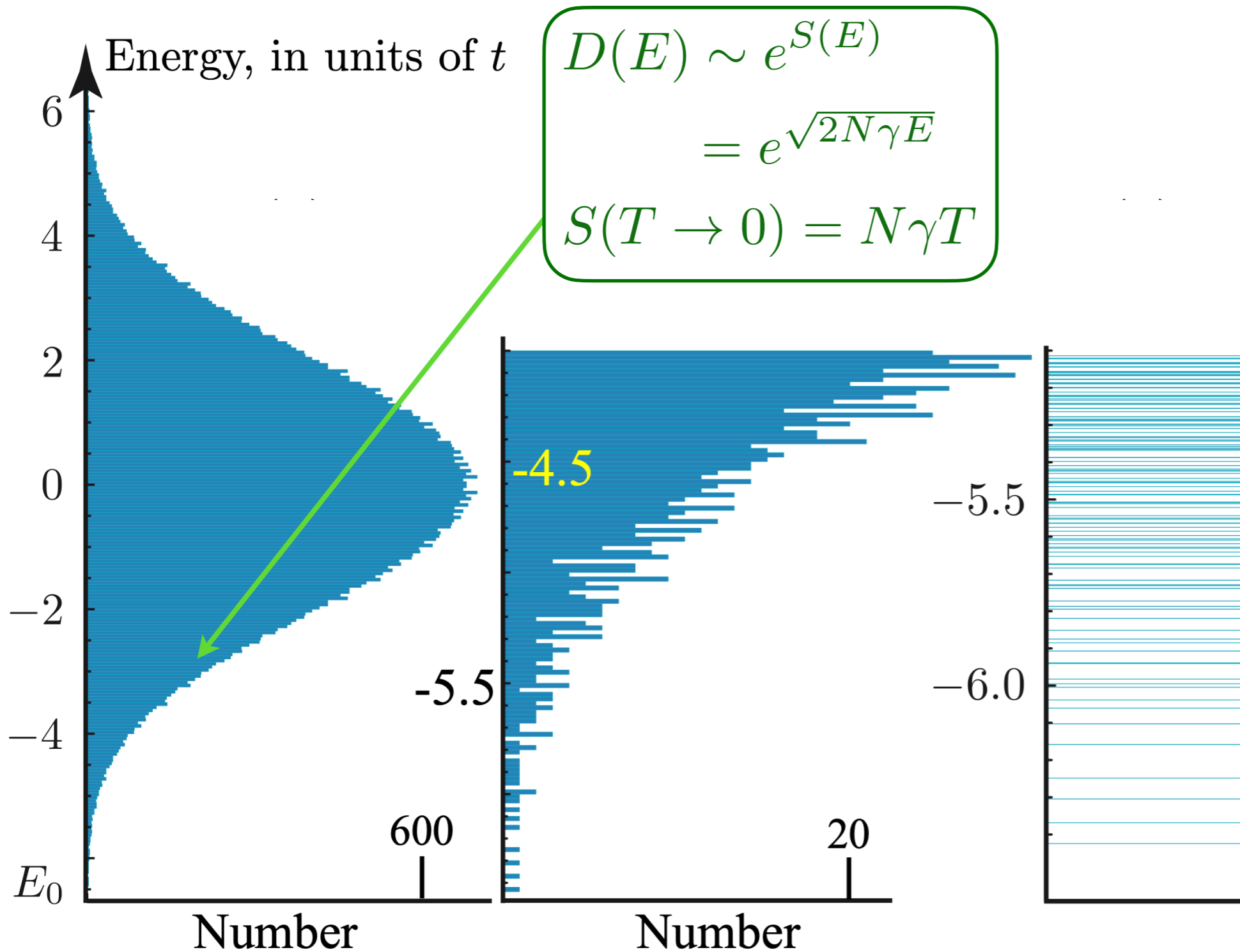
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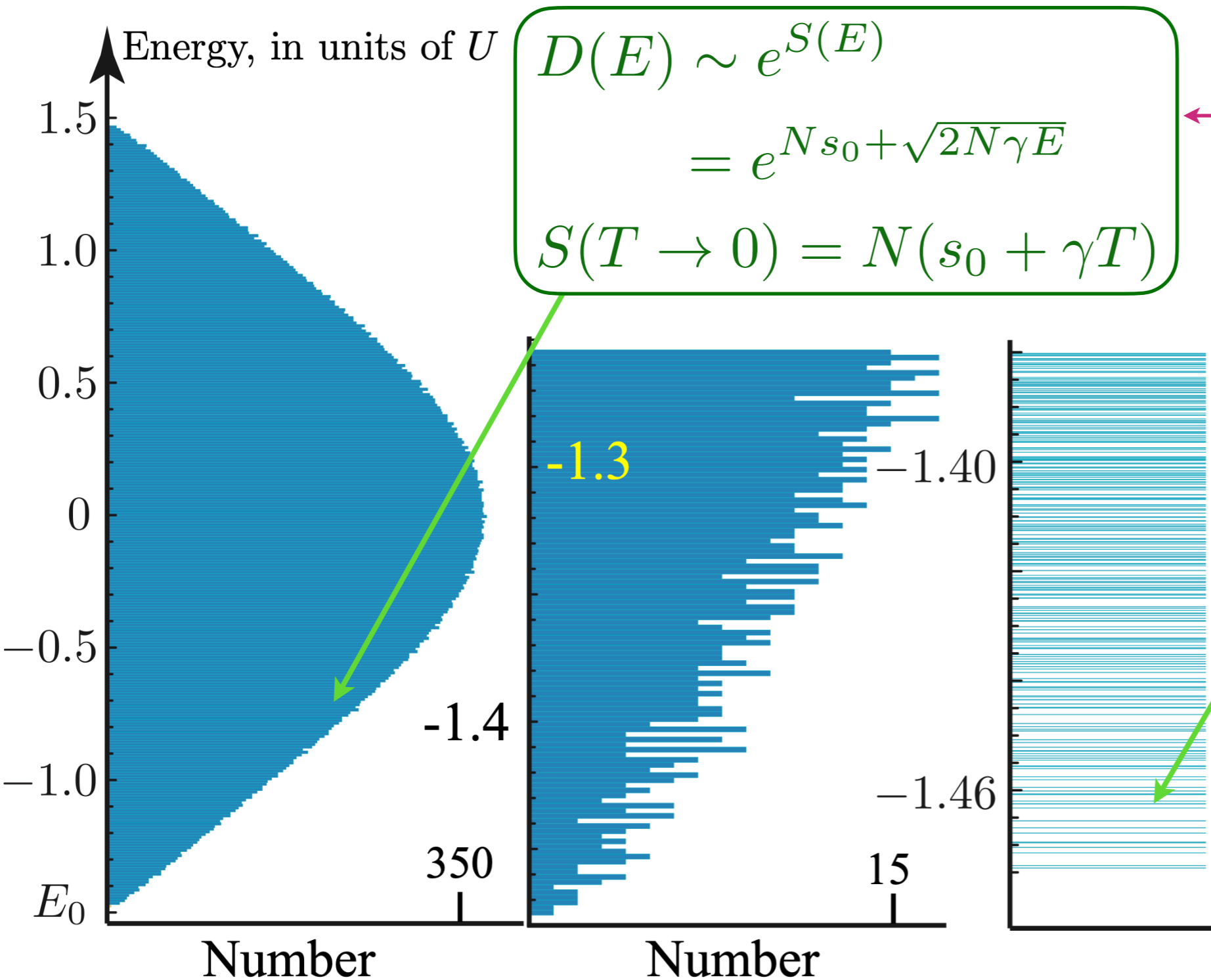


$D(E) \sim N$

Random matrix model

Many-body density of states

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim 2 e^{Ns_0} \sinh(\sqrt{2N\gamma E})$$

$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln\left(\frac{U}{T}\right)$$

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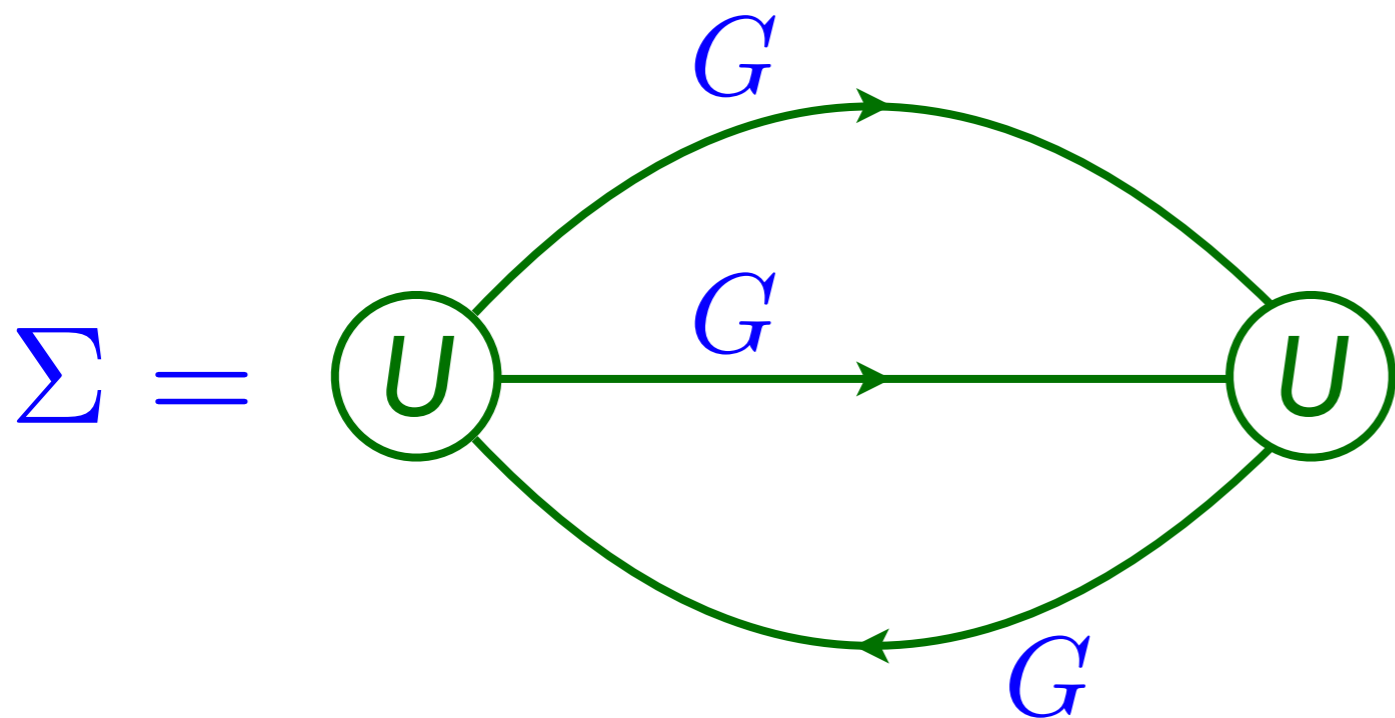
A. Georges, O. Parcollet, and S. Sachdev,
 PRB **63**, 134406 (2001)

Complex SYK model

The complex SYK model

Feynman graph expansion in $U_{\alpha\beta;\gamma\delta}$, and graph-by-graph average, yields exact equations in the large N limit:

$$G(i\omega) = \frac{1}{i\omega + \mu - \Sigma(i\omega)} \quad , \quad \Sigma(\tau) = -U^2 G^2(\tau) G(-\tau)$$
$$G(\tau = 0^-) = \mathcal{Q}.$$



S. Sachdev and J. Ye,
PRL **70**, 3339 (1993)

Conformal solution at $\mu = 0$, $G(\tau) \sim \frac{\text{sgn}(\tau)}{\sqrt{|\tau|}}$.

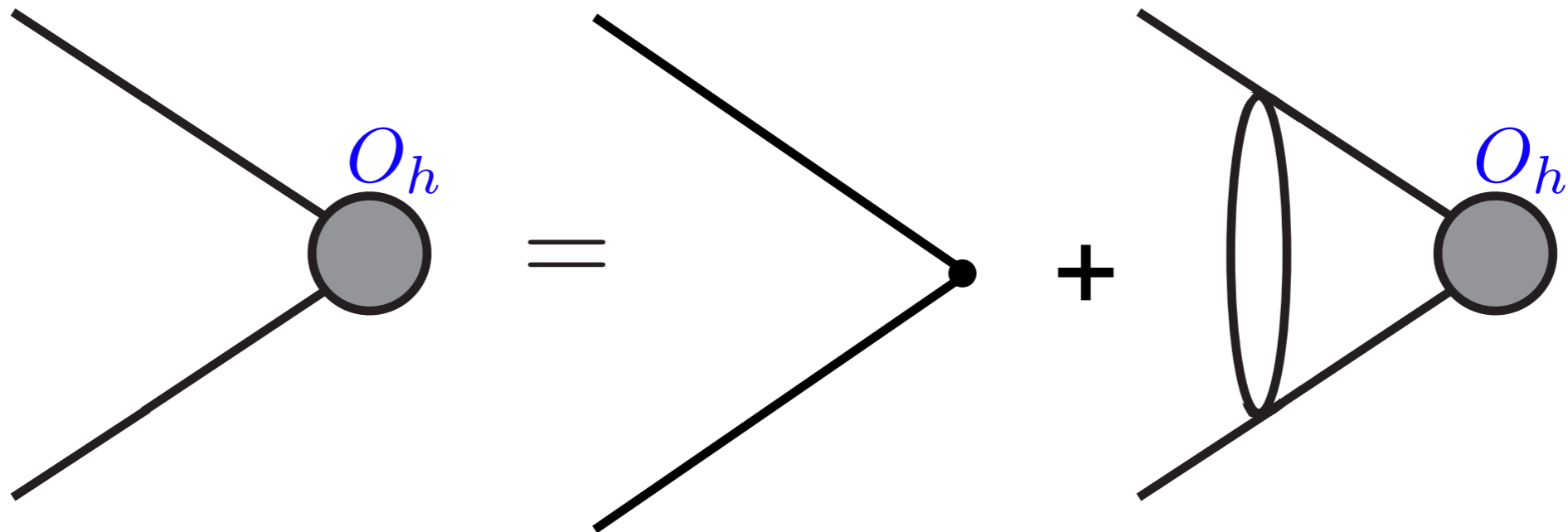


The SYK model

Conformal Perturbation theory

$$S = S_{\text{CFT}} + \sum_h g_h \int_0^\beta d\tau O_h(\tau)$$

where $G_{\text{CFT}} \sim \text{sgn}(\tau)/\sqrt{|\tau|}$ and $\langle O_h(\tau)O_h(0) \rangle \sim 1/|\tau|^{2h}$

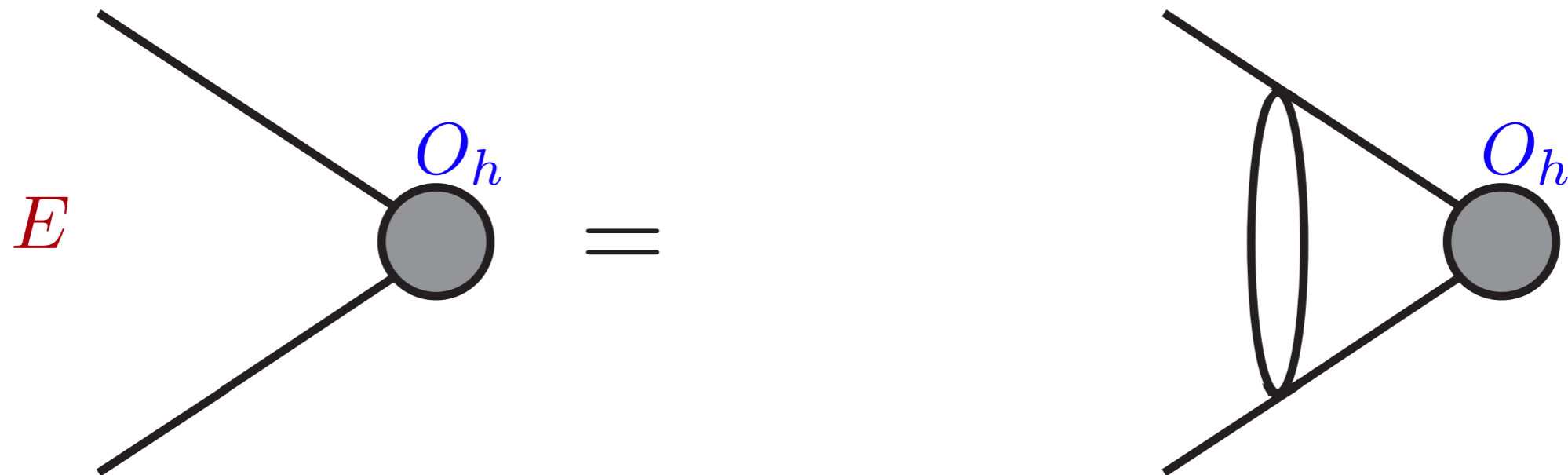


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where $G_{\text{CFT}} \sim \text{sgn}(\tau)/\sqrt{|\tau|}$ and $\langle O_h(\tau)O_h(0) \rangle \sim 1/|\tau|^{2h}$



Solution of eigenvalue equation with $E = 1$ yields a tower of operators O_h with scaling dimensions h . Smallest non-trivial value is $h = 2$, and O_2 is the ‘boundary graviton’.

$$G(\tau) \sim \frac{\text{sgn}(\tau)}{\sqrt{|\tau|}} \left(1 + \sum_h \frac{c_h g_h}{|\tau|^{h-1}} + \dots \right)$$

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Time reparameterization symmetry and 2D gravity

After introducing replicas $a = 1 \dots n$, and integrating out the disorder, the partition function can be written as

$$Z = \int \mathcal{D}c_{\alpha a}(\tau) \exp \left[- \sum_{ia} \int_0^\beta d\tau c_{\alpha a}^\dagger \left(\frac{\partial}{\partial \tau} - \mu \right) c_{\alpha a} - \frac{U^2}{4N^3} \sum_{ab} \int_0^\beta d\tau d\tau' \left| \sum_i c_{\alpha a}^\dagger(\tau) c_{\alpha b}(\tau') \right|^4 \right].$$

For simplicity, we neglect the replica indices, and introduce the identity

$$1 = \int \mathcal{D}G(\tau_1, \tau_2) \mathcal{D}\Sigma(\tau_1, \tau_2) \exp \left[-N \int_0^\beta d\tau_1 d\tau_2 \Sigma(\tau_1, \tau_2) \left(G(\tau_2, \tau_1) + \frac{1}{N} \sum_i c_\alpha(\tau_2) c_\alpha^\dagger(\tau_1) \right) \right].$$

Time reparameterization symmetry and 2D gravity

Then the partition function can be written as a path integral with an action S analogous to a Luttinger-Ward functional

$$Z = \int \mathcal{D}G(\tau_1, \tau_2) \mathcal{D}\Sigma(\tau_1, \tau_2) \exp(-NS)$$

$$S = \ln \det [\delta(\tau_1 - \tau_2)(\partial_{\tau_1} + \mu) - \Sigma(\tau_1, \tau_2)] \\ + \int d\tau_1 d\tau_2 [\Sigma(\tau_1, \tau_2)G(\tau_2, \tau_1) + (U^2/2)G^2(\tau_2, \tau_1)G^2(\tau_1, \tau_2)]$$

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$$+ \int d\tau_1 d\tau_2 [\Sigma(\tau_1, \tau_2) G(\tau_2, \tau_1) + (U^2/2) G^2(\tau_2, \tau_1) G^2(\tau_1, \tau_2)]$$

At frequencies $\ll U$, the time derivative in the determinant is less important, and without it the path integral is invariant under the reparametrization and gauge transformations

$$\tau = f(\sigma)$$

$$G(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-1/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{G}(\sigma_1, \sigma_2)$$

$$\Sigma(\tau_1, \tau_2) = [f'(\sigma_1) f'(\sigma_2)]^{-3/4} \frac{g(\sigma_1)}{g(\sigma_2)} \tilde{\Sigma}(\sigma_1, \sigma_2)$$

where $f(\sigma)$ and $g(\sigma)$ are arbitrary functions.

A. Georges and O. Parcollet
PRB **59**, 5341 (1999)

A. Kitaev, 2015

S. Sachdev, PRX **5**, 041025 (2015)

Time reparameterization symmetry and 2D gravity

Reparametrization and phase zero modes

We can write the path integral for the SYK model as

$$\mathcal{Z} = \int \mathcal{D}G(\tau_1, \tau_2) \mathcal{D}\Sigma(\tau_1, \tau_2) e^{-NS[G, \Sigma]}$$

for a known action $S[G, \Sigma]$. We find the saddle point, G_s, Σ_s , and only focus on the “Nambu-Goldstone” modes associated with breaking reparameterization and U(1) gauge symmetries by writing

$$G(\tau_1, \tau_2) = [f'(\tau_1)f'(\tau_2)]^{1/4} G_s(f(\tau_1) - f(\tau_2)) e^{i\phi(\tau_1) - i\phi(\tau_2)}$$

(and similarly for Σ). Then the path integral is approximated by

$$\mathcal{Z} = \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) e^{-E_0/T + NS(E_0) - NS_{\text{eff}}[f, \phi]},$$

where $E_0 \propto N$ is the ground state energy.

J. Maldacena and D. Stanford, arXiv:1604.07818;
R. Davison, Wenbo Fu, A. Georges, Yingfei Gu, K. Jensen, S. Sachdev, arXiv:1612.00849;
S. Sachdev, PRX **5**, 041025 (2015); J. Maldacena, D. Stanford, and Zhenbin Yang, arXiv:1606.01857;
K. Jensen, arXiv:1605.06098; J. Engelsoy, T.G. Mertens, and H. Verlinde, arXiv:1606.03438

Time reparameterization symmetry and 2D gravity

Symmetry arguments, and explicit computations, show that the effective action is

$$S_{\text{eff}}[f, \phi] = \frac{NK}{2} \int_0^{1/T} d\tau (\partial_\tau \phi + i(2\pi\mathcal{E}T)\partial_\tau f)^2 - \frac{N\gamma}{4\pi^2} \int_0^{1/T} d\tau \{ \tan(\pi T f(\tau)), \tau \},$$

where $f(\tau)$ is a monotonic map from $[0, 1/T]$ to $[0, 1/T]$, the couplings K , γ , and \mathcal{E} can be related to thermodynamic derivatives and we have used the Schwarzian:

$$\{g, \tau\} \equiv \frac{g'''}{g'} - \frac{3}{2} \left(\frac{g''}{g'} \right)^2.$$

Specifically, an argument constraining the effective at $T = 0$ is

$$S_{\text{eff}} \left[f(\tau) = \frac{a\tau + b}{c\tau + d}, \phi(\tau) = 0 \right] = 0,$$

and this is origin of the Schwarzian.

- Exact evaluation of the path integral over $f(\tau)$ and $\phi(\tau)$ leads to the many-body density of states

$$D(E) \sim 2e^{S_0} \sinh(\sqrt{2N\gamma E})$$

- Saddle-point shift leads to a correction to the Green's function:

$$G(\tau) \sim \frac{\text{sgn}(\tau)}{\sqrt{|\tau|}} \left(1 + \frac{\alpha_G}{|\tau|} + \dots \right)$$

There is a universal relationship between α_G and γ , the coefficient of the Schwarzian and the linear- T entropy. This is due to the connection between the $h = 2 O_h$ operator and the ‘boundary graviton’.

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Thermodynamics of quantum black holes:

$$\int \mathcal{D}g_{\mu\nu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity}}^{(3+1)} [g_{\mu\nu}] \right)$$



Metric of
spacetime

In general, this summation is not well defined, because to the uncontrollably large number of spacetime configurations.

Thermodynamics of quantum black holes:

$$\int \mathcal{D}g_{\mu\nu} \exp\left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity}}^{(3+1)}[g_{\mu\nu}]\right)$$
$$= \exp(S_{BH}) \times \left(\dots????\dots \right)$$

Metric of spacetime

Gibbons, Hawking (1977)

$$S_{BH} = \frac{A(T)c^3}{4G\hbar}$$

($\hbar/(k_B T)$ is the length of the Euclidean time circle)

$A(T)$ is the area of the black hole horizon at a temperature T .

Interpretation: Black holes have finite number of quantum degrees of freedom, and black hole entropy is their entanglement entropy across the horizon.

Thermodynamics of quantum black holes with charge Q :

$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right) \\ = \exp(S_{BH}) \times \left(\dots????\dots \right)$$

Gibbons, Hawking (1977)

$$S_{BH}(T \rightarrow 0, Q) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2}T}{\hbar c} \right)$$

A_0 is the area of the charged black hole horizon at $T = 0$, Q is the black hole charge.

Thermodynamics of quantum black holes with charge Q :

$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right) \\ = \exp(S_{BH}) \times \left(\dots????\dots \right)$$

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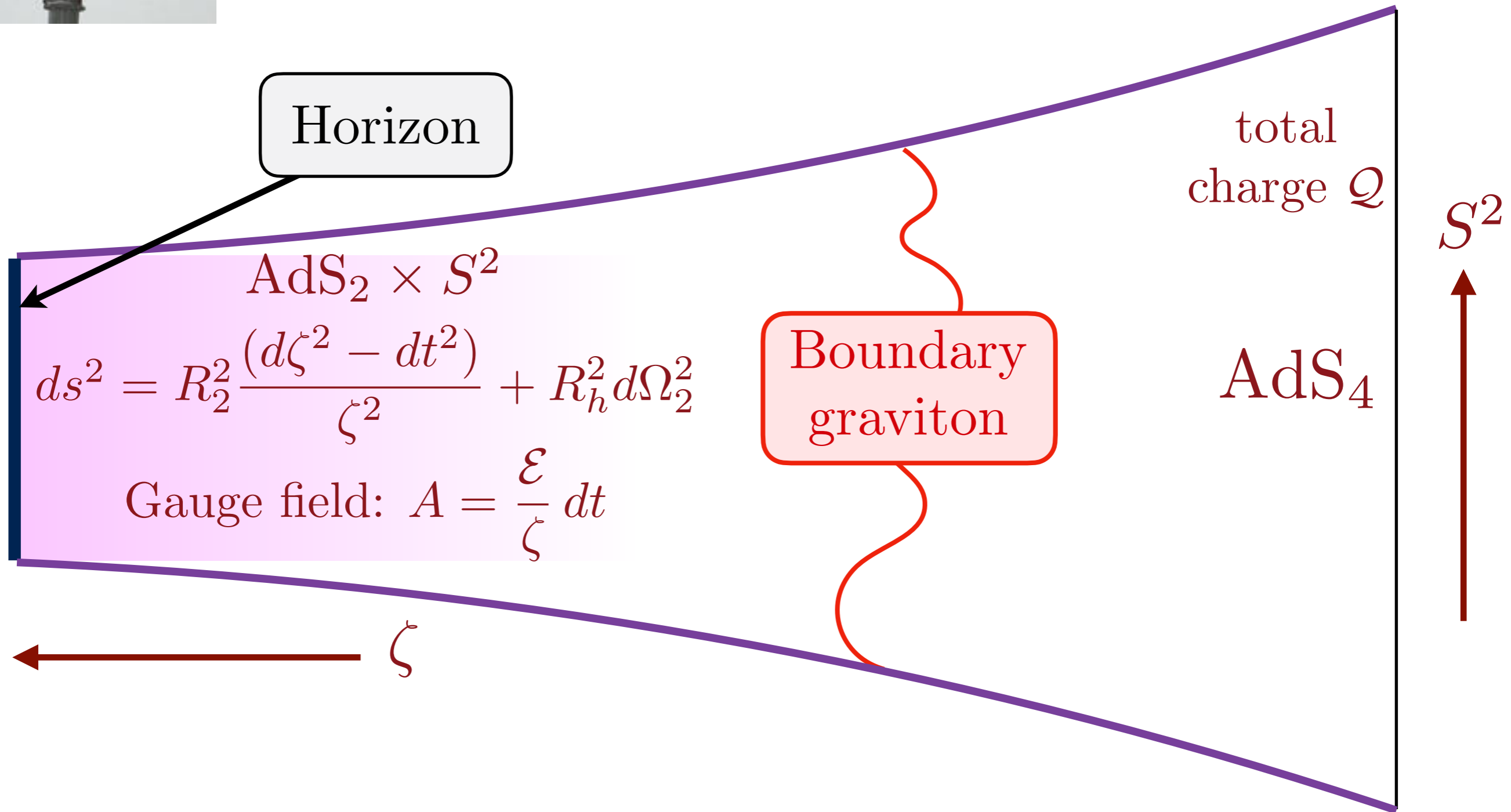
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Note the similarity to the large N entropy of the SYK model !

Sachdev PRL 2010



Reissner-Nordstrom black hole of Einstein-Maxwell theory



Thermodynamics of quantum black holes with charge Q :

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Sachdev PRL 2010

Thermodynamics of quantum black holes with charge \mathcal{Q} :

$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_\mu] \right) \\ \approx \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{JTgravity of AdS}_2 \text{ and boundary}}^{(1+1)}[g_{\mu\nu}, A_\mu] \right)$$

$$S_{BH}(T \rightarrow 0, \mathcal{Q}) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0 c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2} T}{\hbar c} \right)$$

A_0 is the area of the charged black hole horizon at $T = 0$, \mathcal{Q} is the black hole charge.

Thermodynamics of quantum black holes with charge \mathcal{Q} :

$$\begin{aligned}
 & \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_\mu] \right) \\
 & \approx \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{JTgravity of AdS}_2 \text{ and boundary}}^{(1+1)}[g_{\mu\nu}, A_\mu] \right) \\
 & = \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) \exp \left(-\text{Schwarzian boundary graviton} + \text{rotor action}[f, \phi] \right)
 \end{aligned}$$

$$S_{BH}(T \rightarrow 0, \mathcal{Q}) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0 c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2} T}{\hbar c} \right)$$

A_0 is the area of the charged black hole horizon at $T = 0$, \mathcal{Q} is the black hole charge.

Thermodynamics of quantum black holes with charge Q :

$$\begin{aligned} & \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity} + \text{Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_\mu] \right) \\ & \approx \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_\mu \exp \left(-\frac{1}{\hbar} \mathcal{S}_{\text{JTgravity of AdS}_2 \text{ and boundary}}^{(1+1)}[g_{\mu\nu}, A_\mu] \right) \\ & = \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) \exp \left(-\text{Schwarzian boundary graviton} + \text{rotor action}[f, \phi] \right) \end{aligned}$$

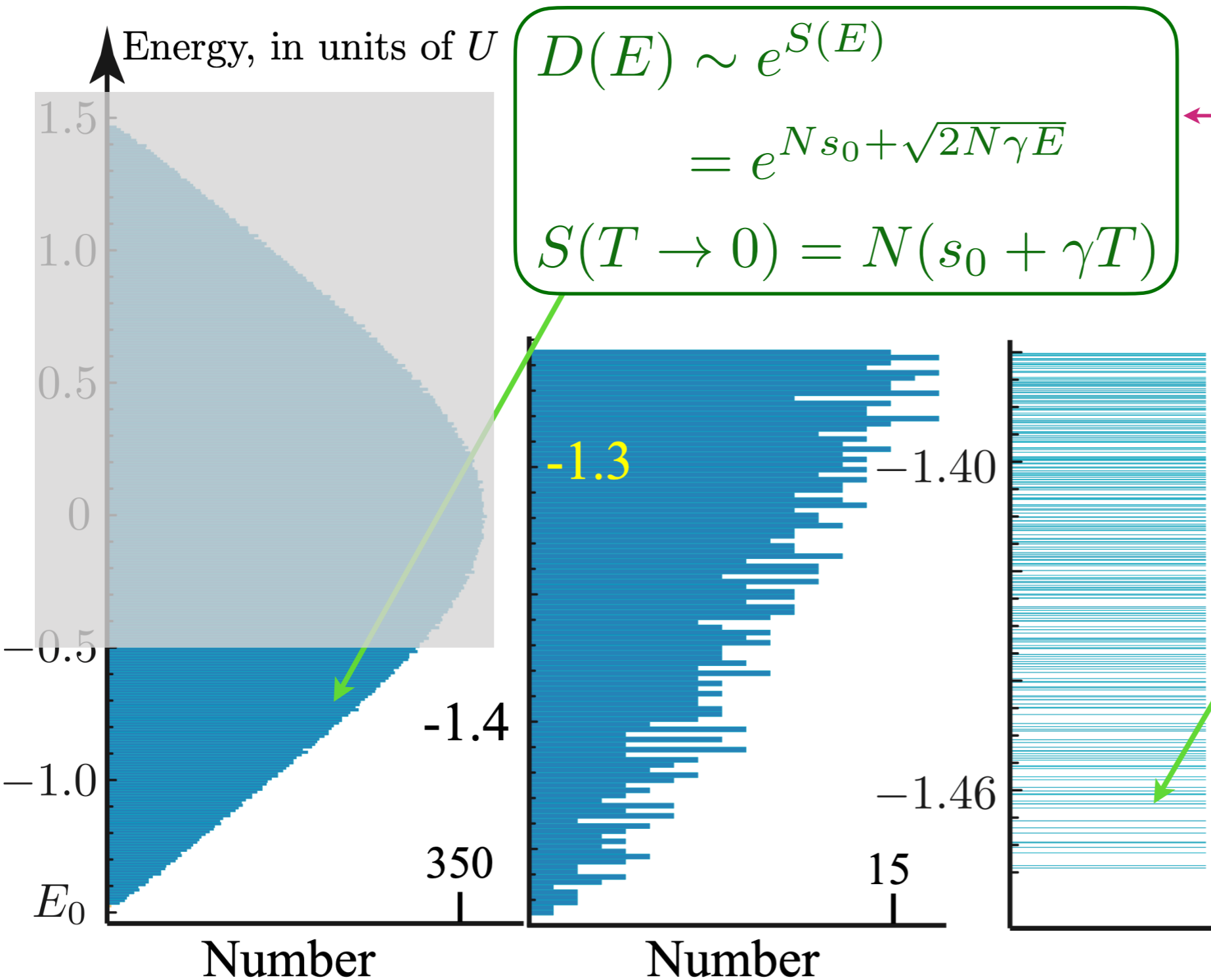
$$S(T \rightarrow 0, Q) = S_{BH} - \frac{3}{4} \ln \left(\frac{\hbar c^5}{GT^2} \right)$$

$$S_{BH} = \frac{A(T)c^3}{4G\hbar} = \frac{A_0 c^3}{4G\hbar} \left(1 + \frac{2(\pi A_0)^{1/2} T}{\hbar c} \right)$$

A_0 is the area of the charged black hole horizon at $T = 0$, Q is the black hole charge. The $\ln T$ term is the contribution of the boundary graviton.

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$$D(E) \sim e^{S(E)}$$

$$= e^{N s_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{N s_0} \sinh(\sqrt{2N\gamma E})$$

$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln \left(\frac{U}{T} \right)$$

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No quasiparticle decomposition of many-body states

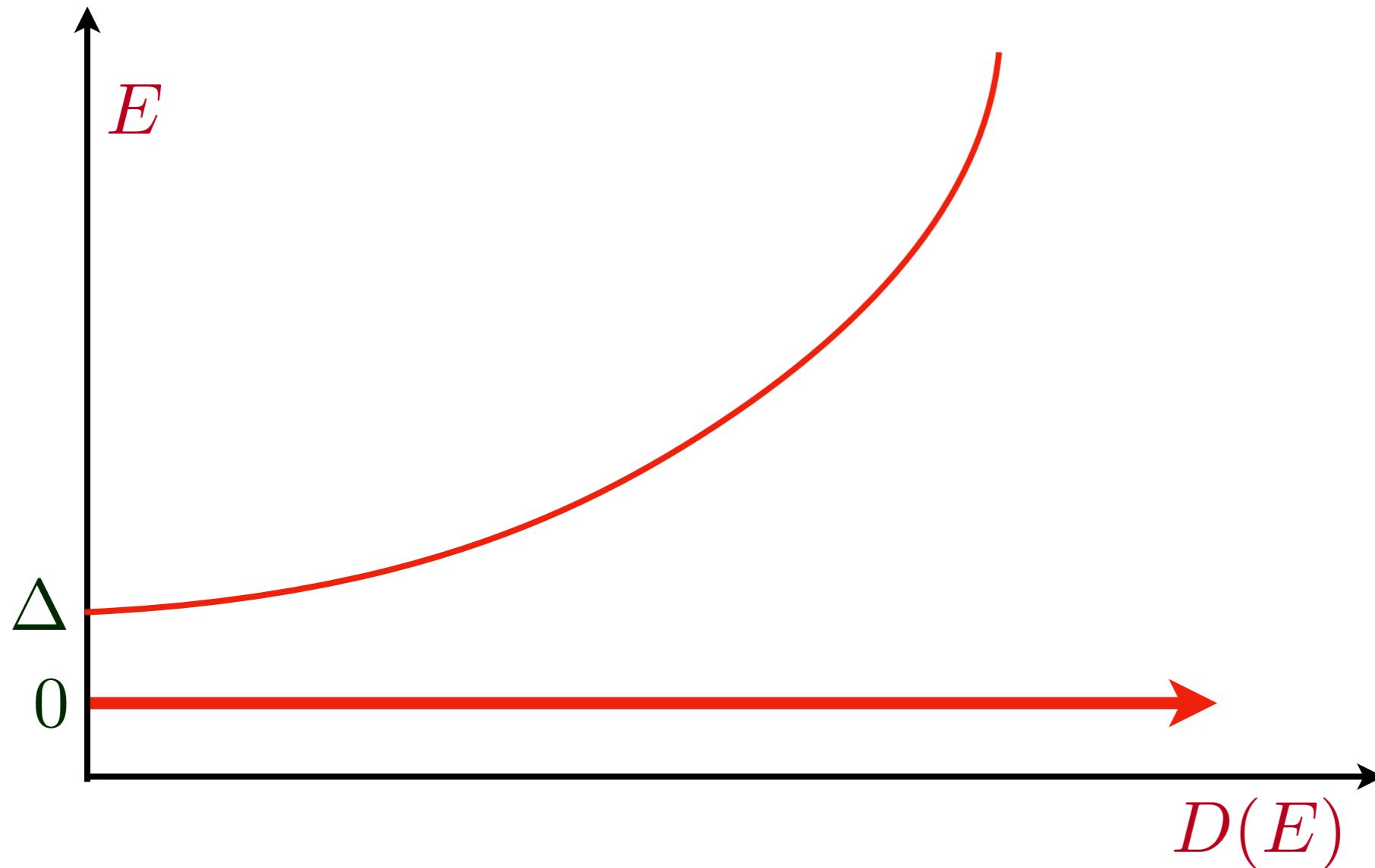
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Complex SYK model

Many-body density of states

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim \exp\left(\frac{A_0}{4G} + \dots\right) \delta(E) + f_{\text{reg}}(E - \Delta), \quad \Delta \sim R_h^{-1}$$

Supersymmetric black holes and SYK models

1. SYK models

2. Time reparameterization soft mode

3. Charged black holes

4. Spin glass of $S=1/2$ $SU(2)$ spins

5. Random t - J model

Quantum generalization of the Sherrington-Kirkpatrick model to $S = 1/2$ spins with $SU(2)$ symmetry

$$H = \sum_{i < j} J_{ij} \mathbf{S}_i \cdot \mathbf{S}_j$$

$$[S_{i\mu}, S_{j\nu}] = i\delta_{ij}\epsilon_{\mu\nu\lambda}S_{i\lambda} \quad , \quad \mathbf{S}_i^2 = 3/4$$

$$\overline{J_{ij}} = 0, \quad \overline{J_{ij}^2} = J^2, \quad \text{Different } J_{ij} \text{ uncorrelated.}$$

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Two possible ground states

I. Spin glass order

$$\lim_{\tau \rightarrow \infty} \langle \mathbf{S}_i(\tau) \cdot \mathbf{S}_i(0) \rangle = q > 0$$

where q is the spin glass order parameter.

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Two possible ground states

II. SY spin liquid

$$\chi_L(\tau) = \langle \mathbf{S}_i(\tau) \cdot \mathbf{S}_i(0) \rangle \sim 1/\tau \quad , \quad \tau \rightarrow \infty$$
$$= -G(\tau)G(-\tau) \text{ at } M = \infty$$

Obtained as $M \rightarrow \infty$ in a model with $SU(M)$ symmetry

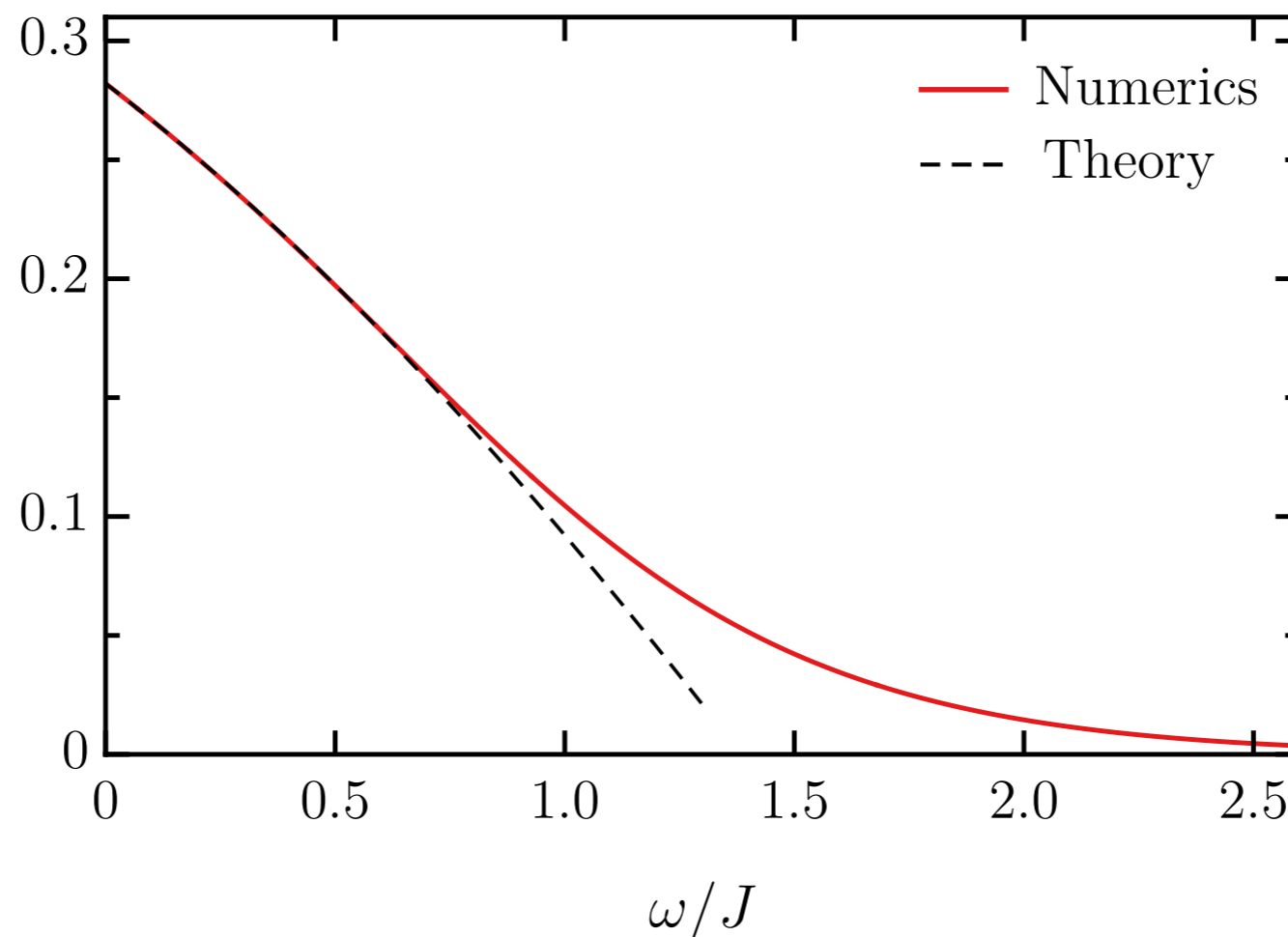
Write $\mathbf{S}_i \Rightarrow f_{i\alpha}^\dagger f_{i\beta}$ with $\alpha, \beta = 1 \dots M$ and $\sum_\alpha f_{i\alpha}^\dagger f_{i\alpha} = M/2$. Then the large M saddle point equations are identical to those of SY(K).

Dynamic spin susceptibility of SY spin liquid at $M = \infty$

$$\chi_L(\tau) = -G(\tau)G(-\tau)$$

$$\text{Im}\chi_L(\omega) \sim \text{sgn}(\omega) \left[1 - \mathcal{C}\gamma|\omega| - \frac{7}{16}(\mathcal{C}\gamma)^2|\omega|^2 - \mathcal{C}'|\omega|^{2.77354\dots} + \frac{37}{48}(\mathcal{C}\gamma)^3|\omega|^3 - \dots \right]$$

Numerical solution of SYK equations (SY, PRL 1993), compared with conformal perturbation theory. \mathcal{C} is a known number, and γ is the co-efficient of the action for the ‘boundary graviton’ in holographic dual.

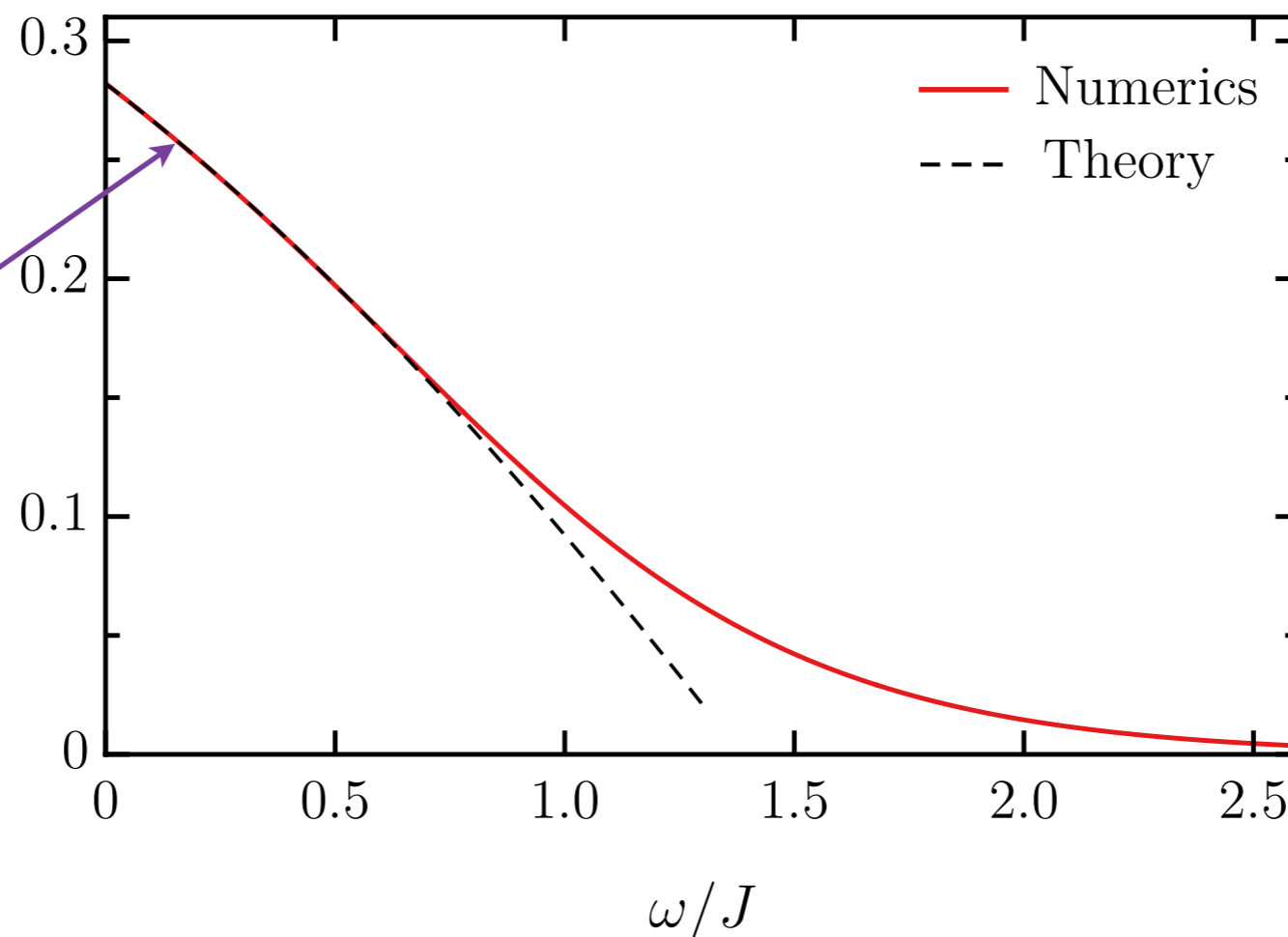


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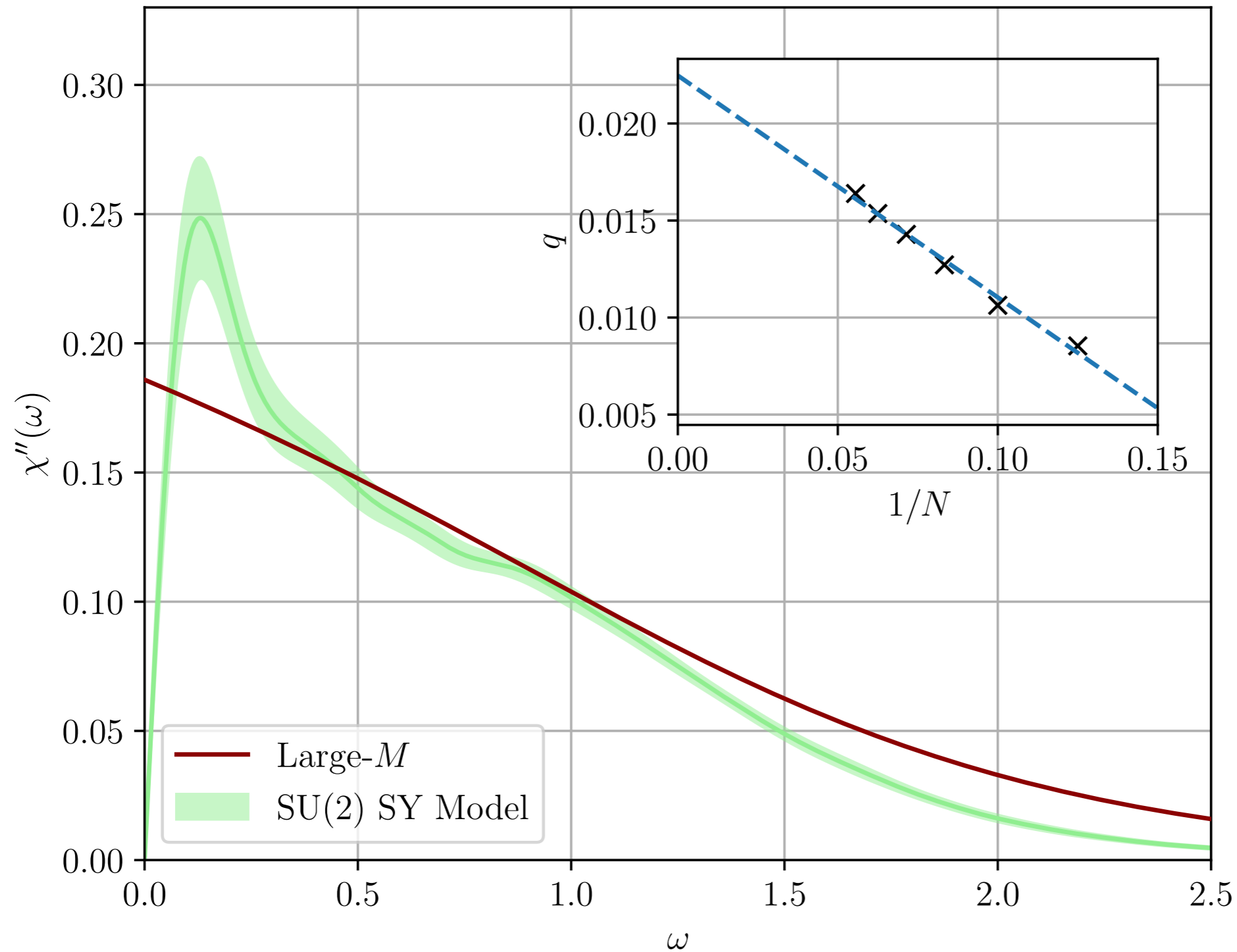
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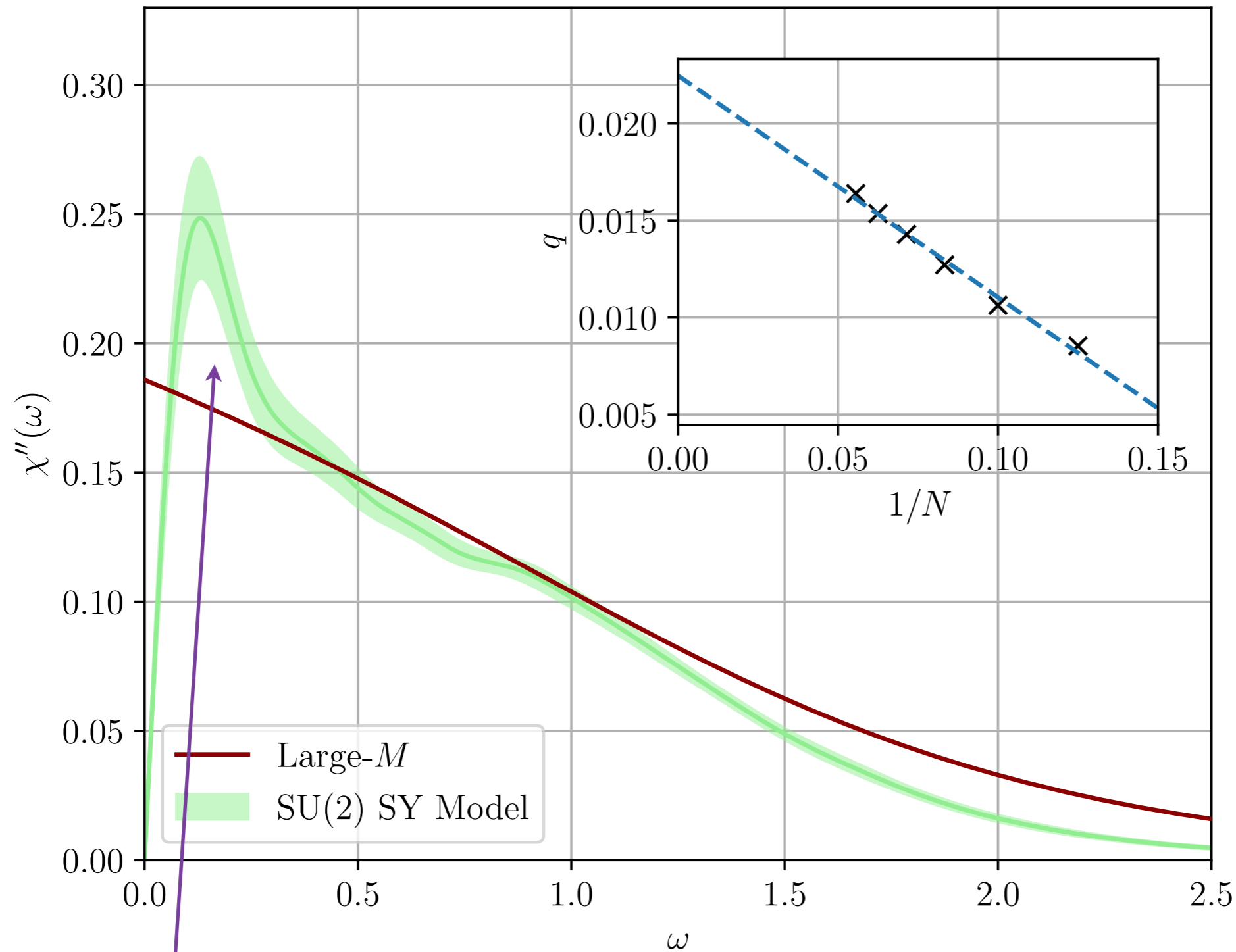
Correction
from the
boundary
graviton



Exact diagonalization of clusters of SU(2) spins



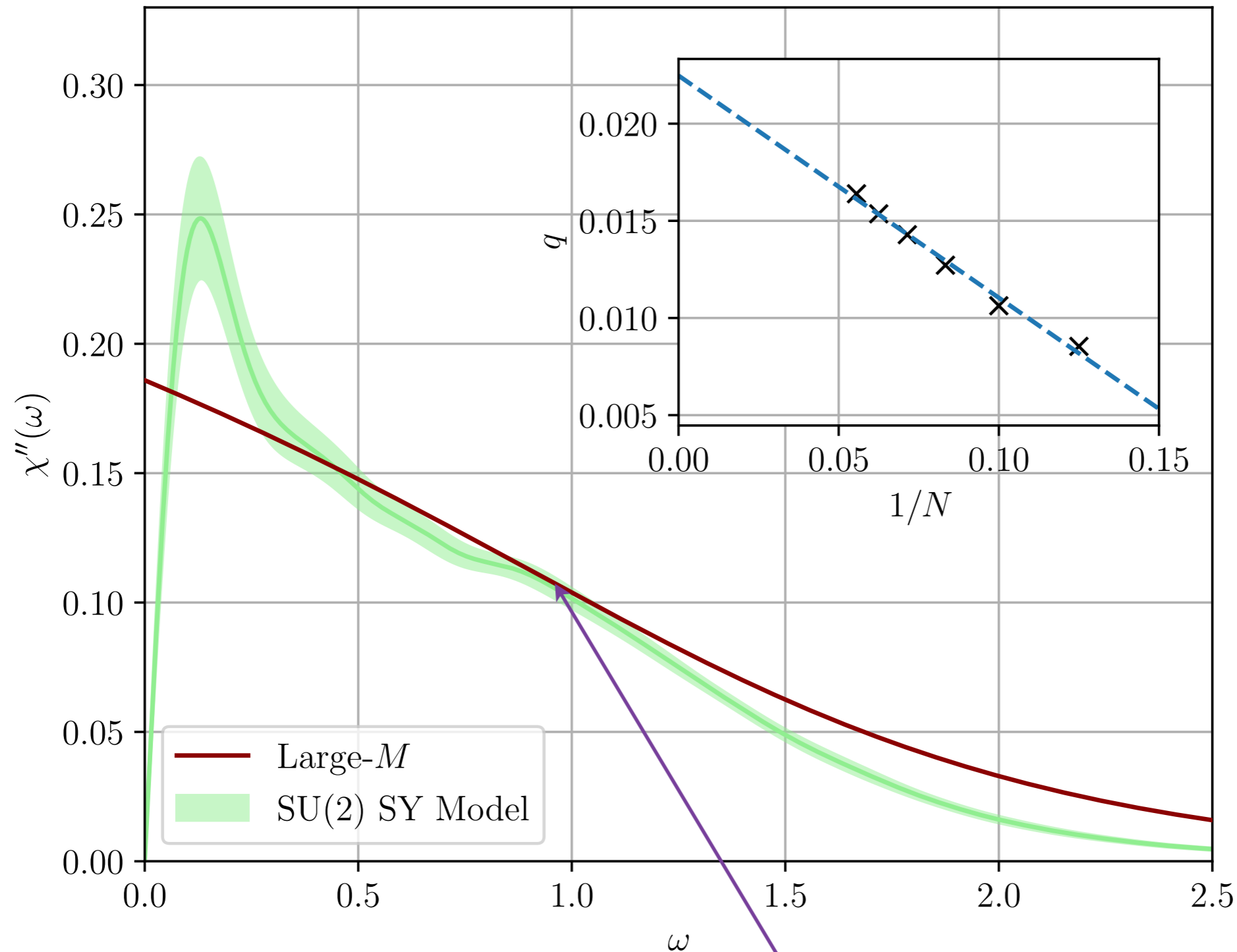
Exact diagonalization of clusters of SU(2) spins



Spin glass



Exact diagonalization of clusters of SU(2) spins



SY spin liquid (boundary graviton)



G-Σ-Q theory for co-existence of spin glass and SY spin liquid

Quantum spin glass order parameter: $Q_{ab}(\tau)$, $a, b = 1 \dots n$, $n \rightarrow 0$ replicas.
 Q_{ab} is τ independent for $a \neq b$.

$$Q_{ab}(\tau, \tau') = \frac{1}{N} \sum_i \mathbf{S}_{ia}(\tau) \cdot \mathbf{S}_{ib}(\tau')$$

$$q_{EA} = \overline{\langle \mathbf{S}_i(\tau) \rangle \cdot \langle \mathbf{S}_i(0) \rangle} = \lim_{n \rightarrow 0} \frac{1}{n(n-1)} \sum_{a \neq b} Q_{ab}$$

$$q = \lim_{\tau \rightarrow \infty} \overline{\langle \mathbf{S}_i(\tau) \cdot \mathbf{S}_i(0) \rangle} = \lim_{n \rightarrow 0} \lim_{\tau \rightarrow \infty} \frac{1}{n} \sum_a Q_{aa}(\tau)$$

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$$\frac{\mathcal{F}}{N} = \frac{MJ^2}{4} \int d\tau [Q_{ab}(\tau)]^2 - \ln \mathcal{Z}_f[Q]$$

$$\mathcal{Z}_f[Q] = \int \mathcal{D}G_{ab}(\tau, \tau') \mathcal{D}\Sigma_{ab}(\tau, \tau') \mathcal{D}\lambda_a(\tau) \exp[-MI[Q]]$$

$$I[Q] = -\ln \det \left[-\delta'(\tau - \tau') \delta_{ab} - i\lambda_a(\tau) \delta(\tau - \tau') \delta_{ab} - \Sigma_{ab}(\tau, \tau') \right] - \frac{i}{2} \int d\tau \lambda_a(\tau) \\ + \int d\tau d\tau' \left[-\Sigma_{ab}(\tau, \tau') G_{ba}(\tau', \tau) + \frac{J^2}{2} Q_{ab}(\tau - \tau') G_{ab}(\tau, \tau') G_{ba}(\tau', \tau) \right]$$

1. SYK models

2. Time reparameterization soft mode

3. Charged black holes

4. Spin glass of $S=1/2$ $SU(2)$ spins

5. Random t - J model

Dope the quantum Sherrington-Kirkpatrick model with mobile electrons

$$H = \sum_{i < j} \left[-t_{ij} c_{i\alpha}^\dagger c_{j\alpha} + \text{H.c.} + J_{ij} \mathbf{S}_i \cdot \mathbf{S}_j \right]$$

$$\mathbf{S}_i = \frac{1}{2} c_{i\alpha}^\dagger \boldsymbol{\sigma}_{\alpha\beta} c_{i\beta}$$

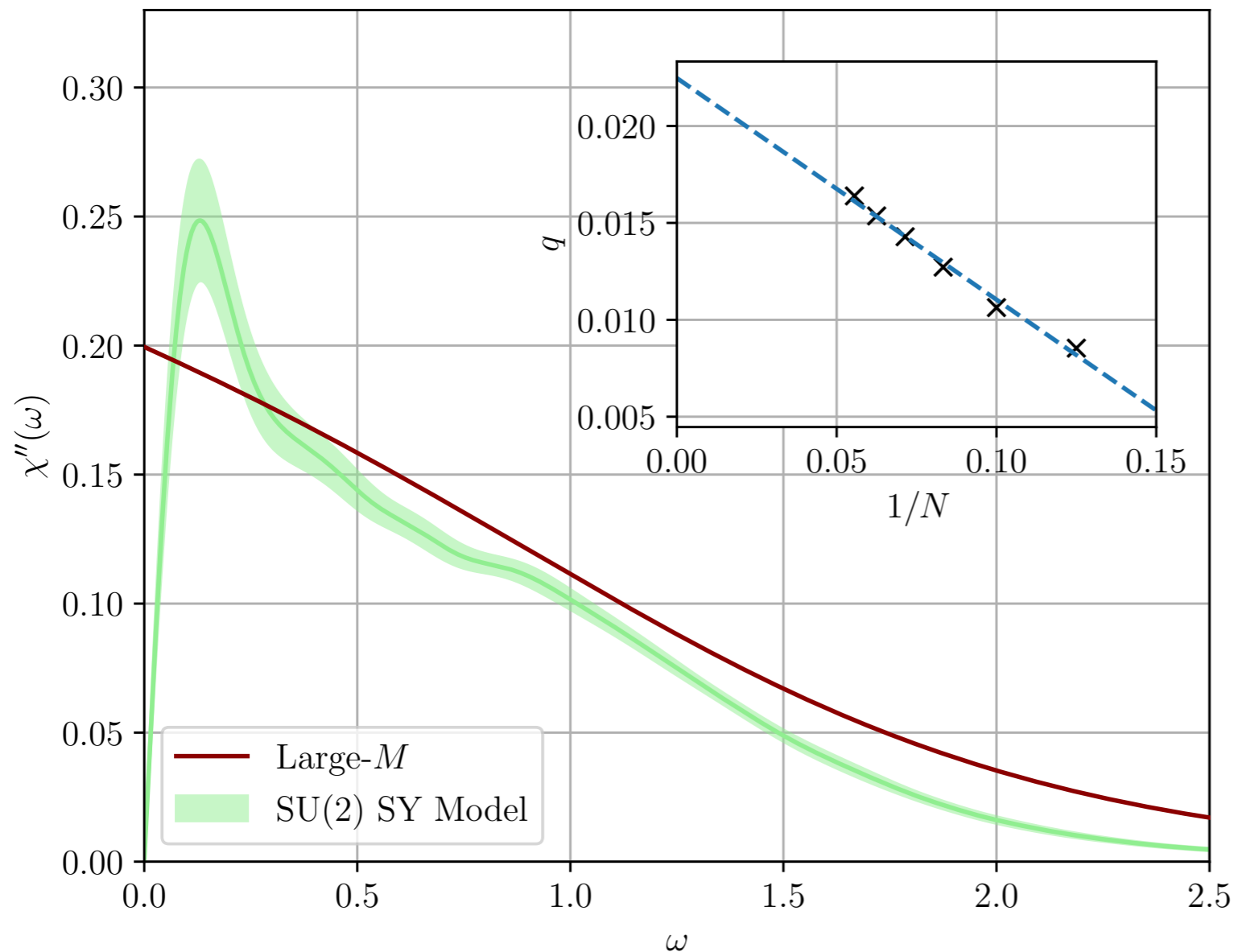
$$[c_{i\alpha}, c_{j\beta}^\dagger]_+ = \delta_{ij} \delta_{\alpha\beta} \quad , \quad \sum_{\alpha} c_{i\alpha}^\dagger c_{i\alpha} \leq 1$$

$$\frac{1}{N} \sum_{i\alpha} c_{i\alpha}^\dagger c_{i\alpha} = 1 - p$$

$$\overline{J_{ij}} = 0, \quad \overline{J_{ij}^2} = J^2, \quad \text{Different } J_{ij} \text{ uncorrelated.}$$

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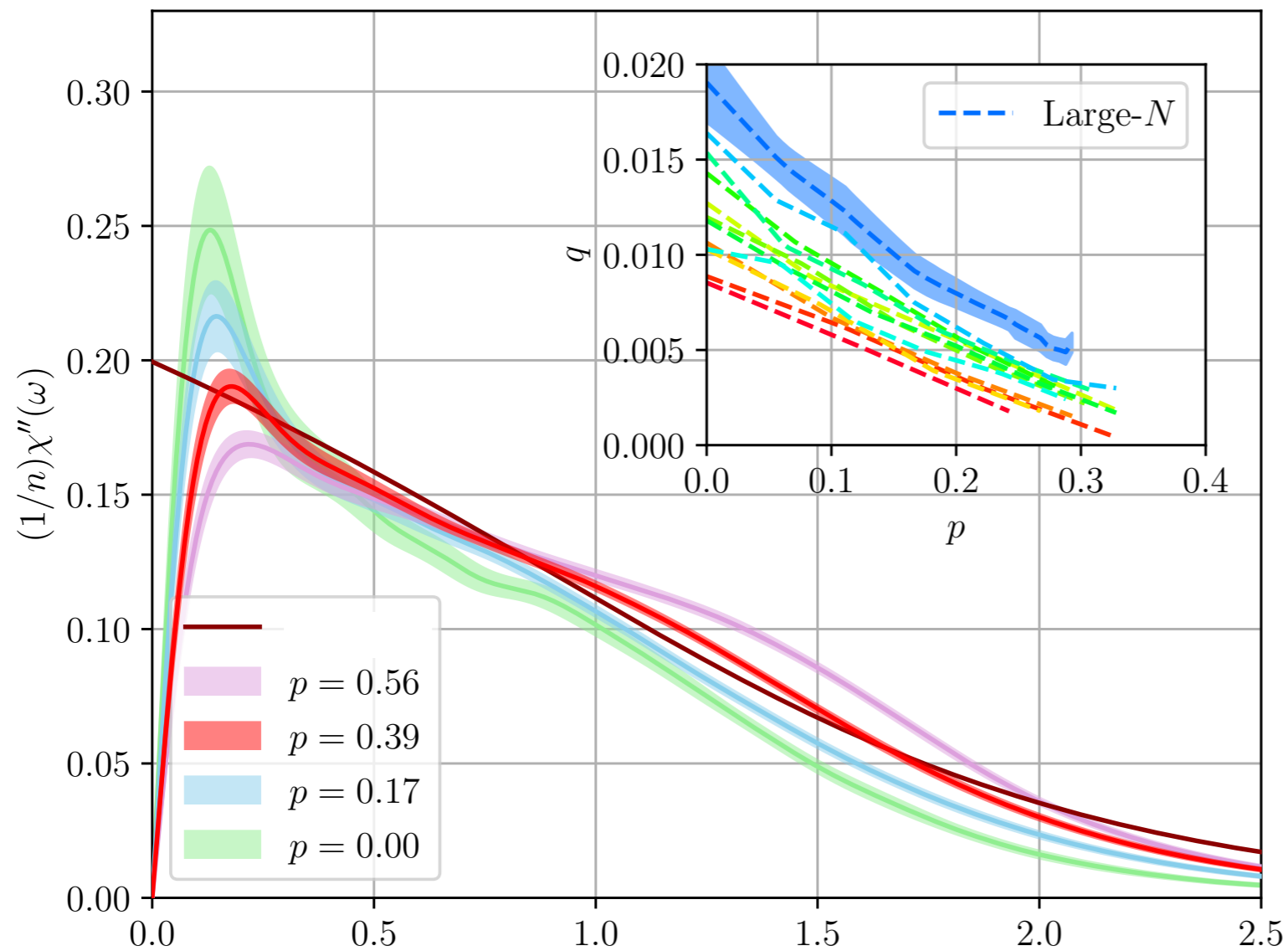
Exact diagonalization of clusters of SU(2) spins



The peak at small ω indicates the presence of spin glass order.

The large M SYK theory predicts $\chi''(\omega) \sim \text{sgn}(\omega) [1 - c|\omega| + \dots]$

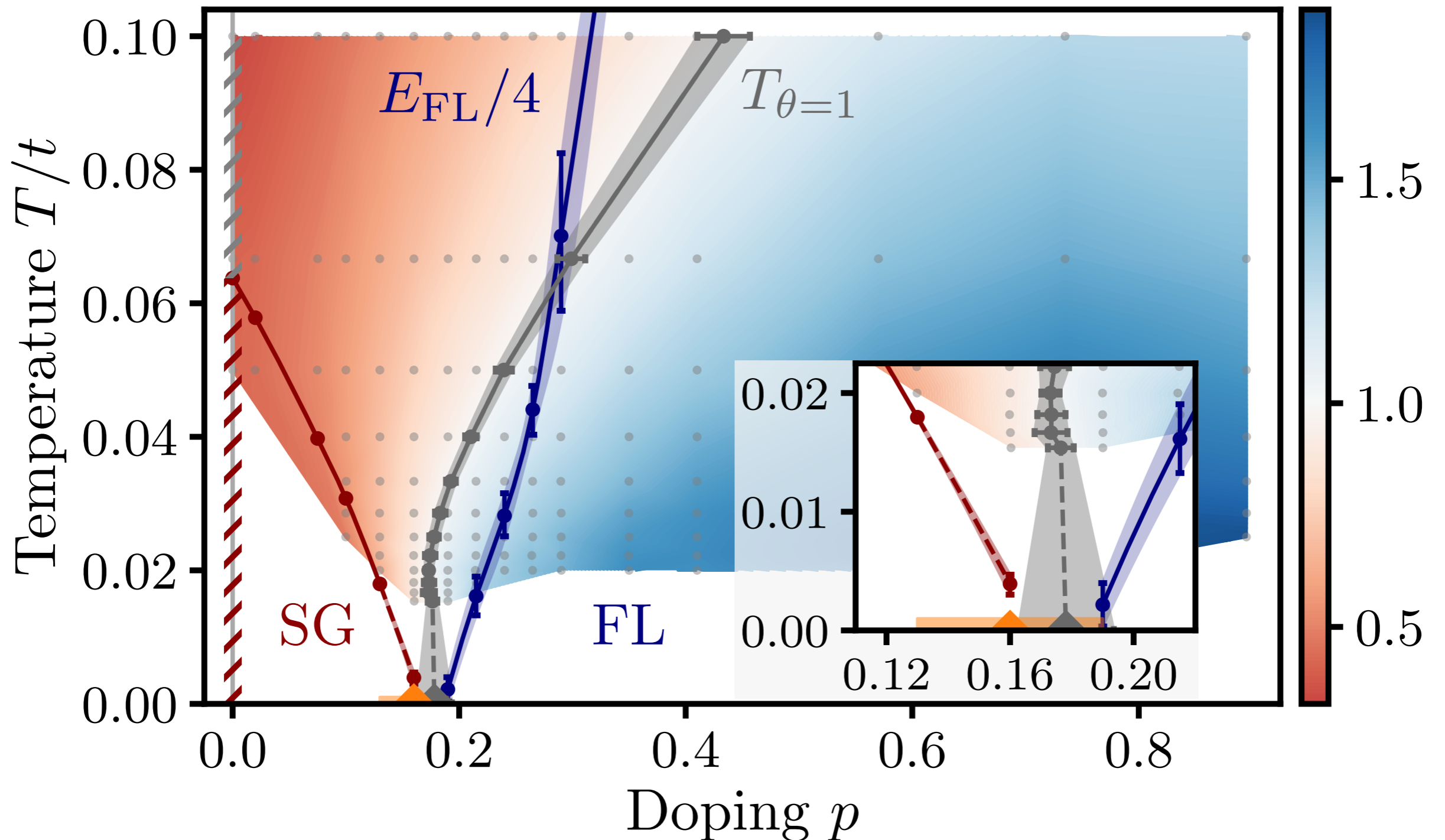
Exact diagonalization of t - J clusters



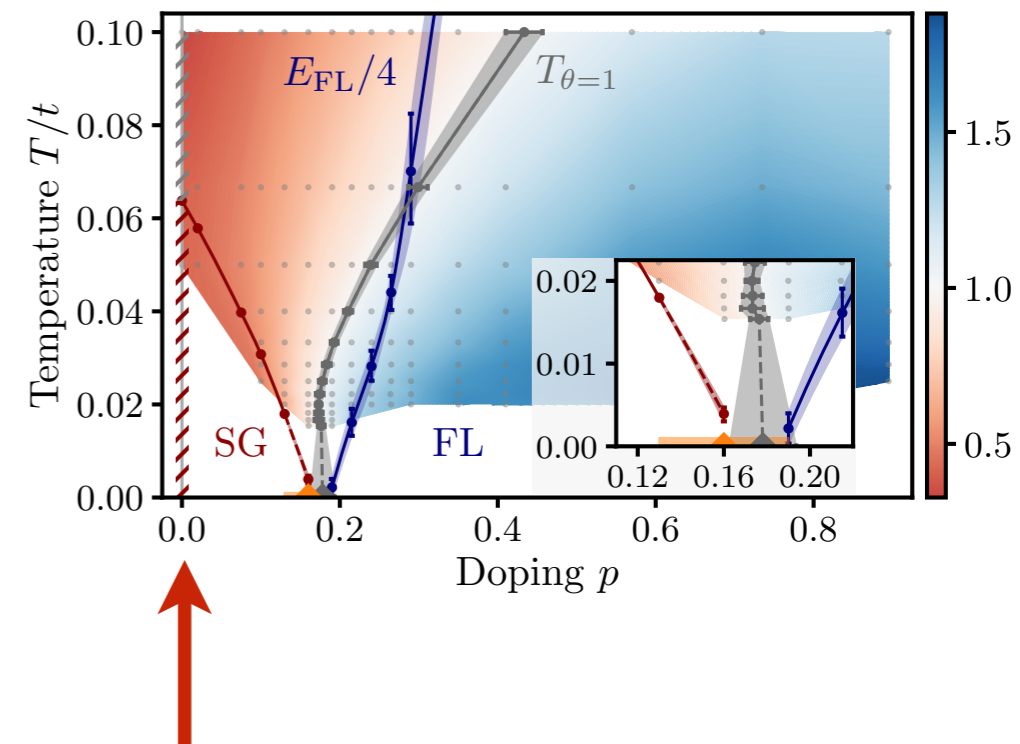
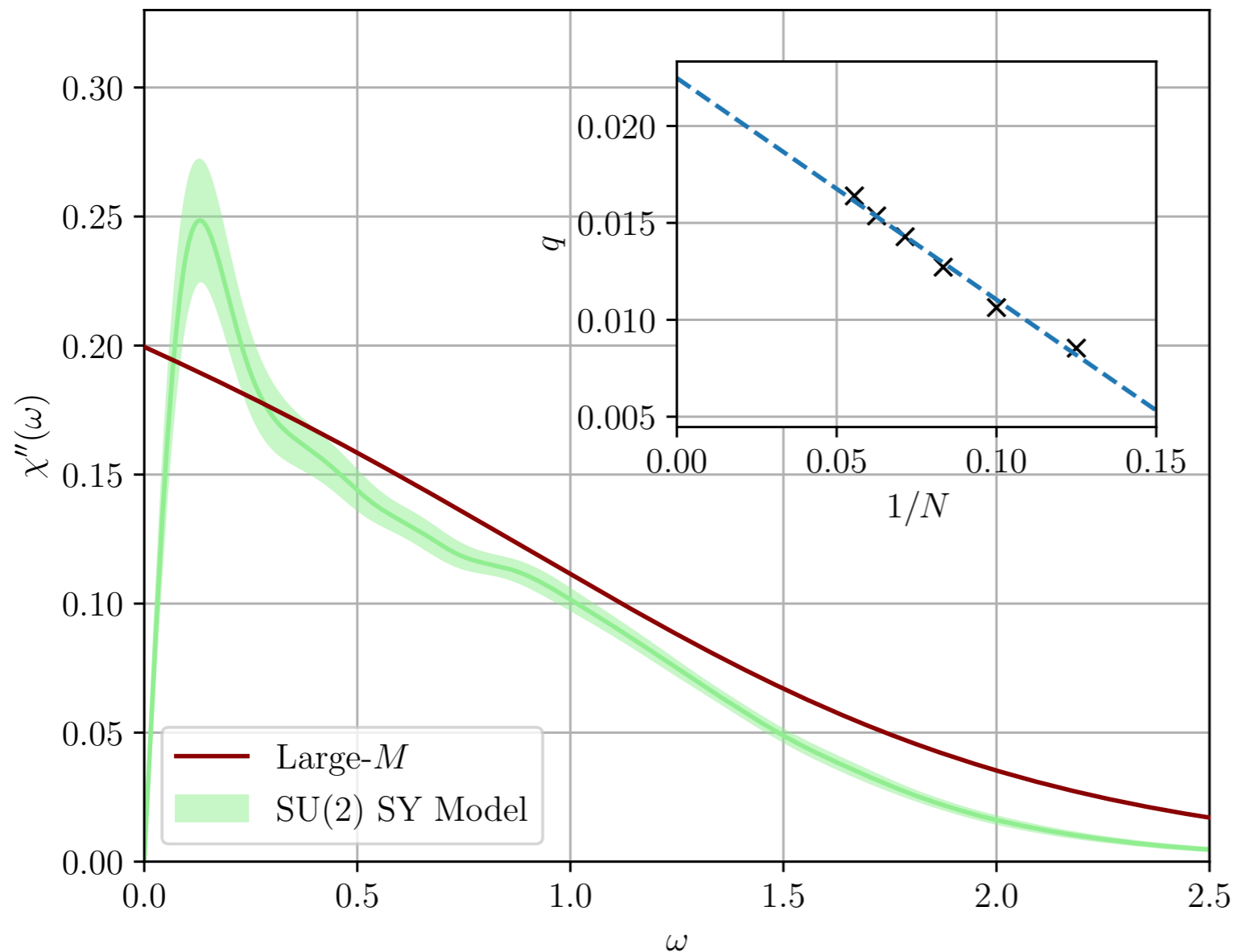
Spin glass order disappears above $p \approx 1/3$.

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and this fits well near criticality

QMC solution of DMFT equations



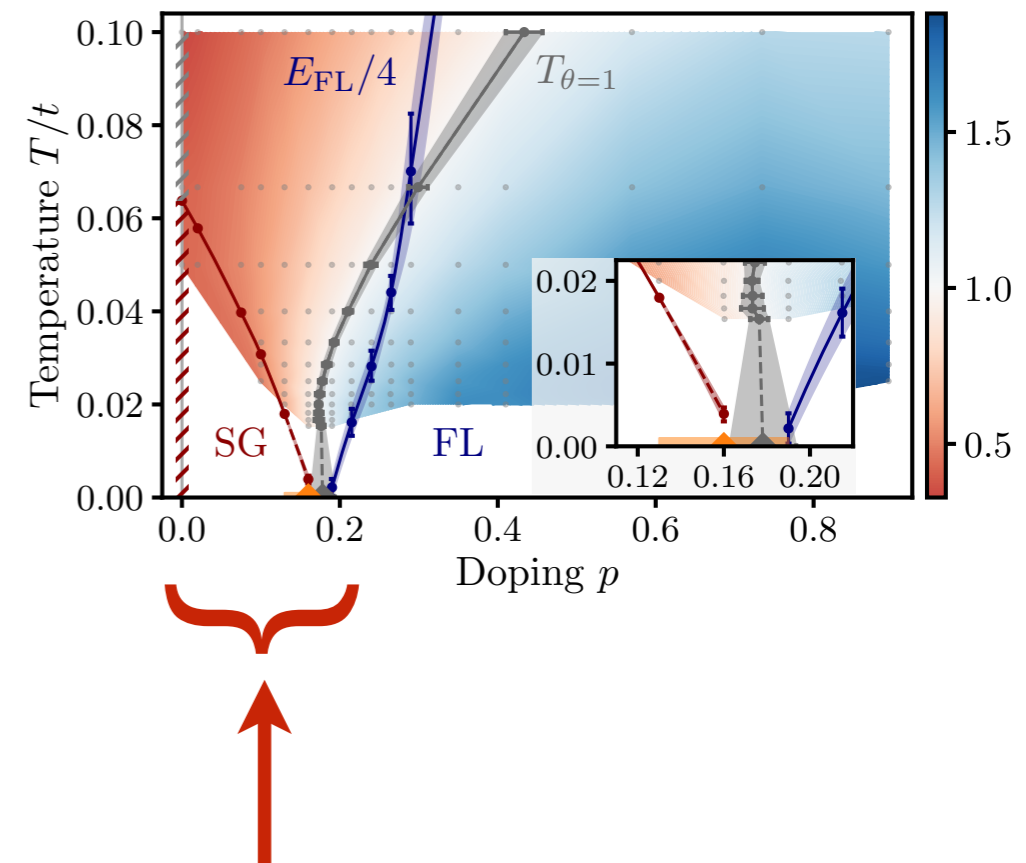
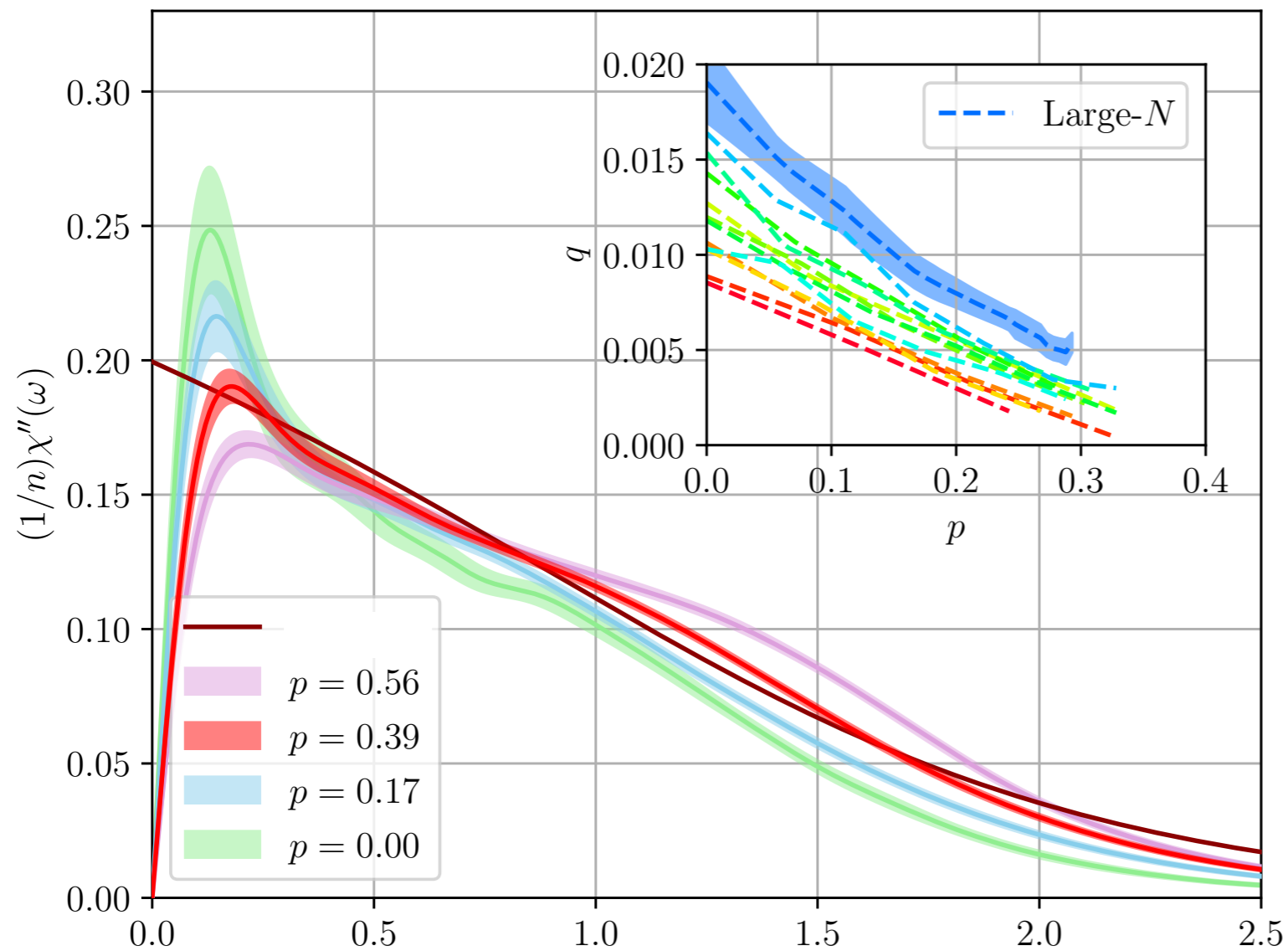
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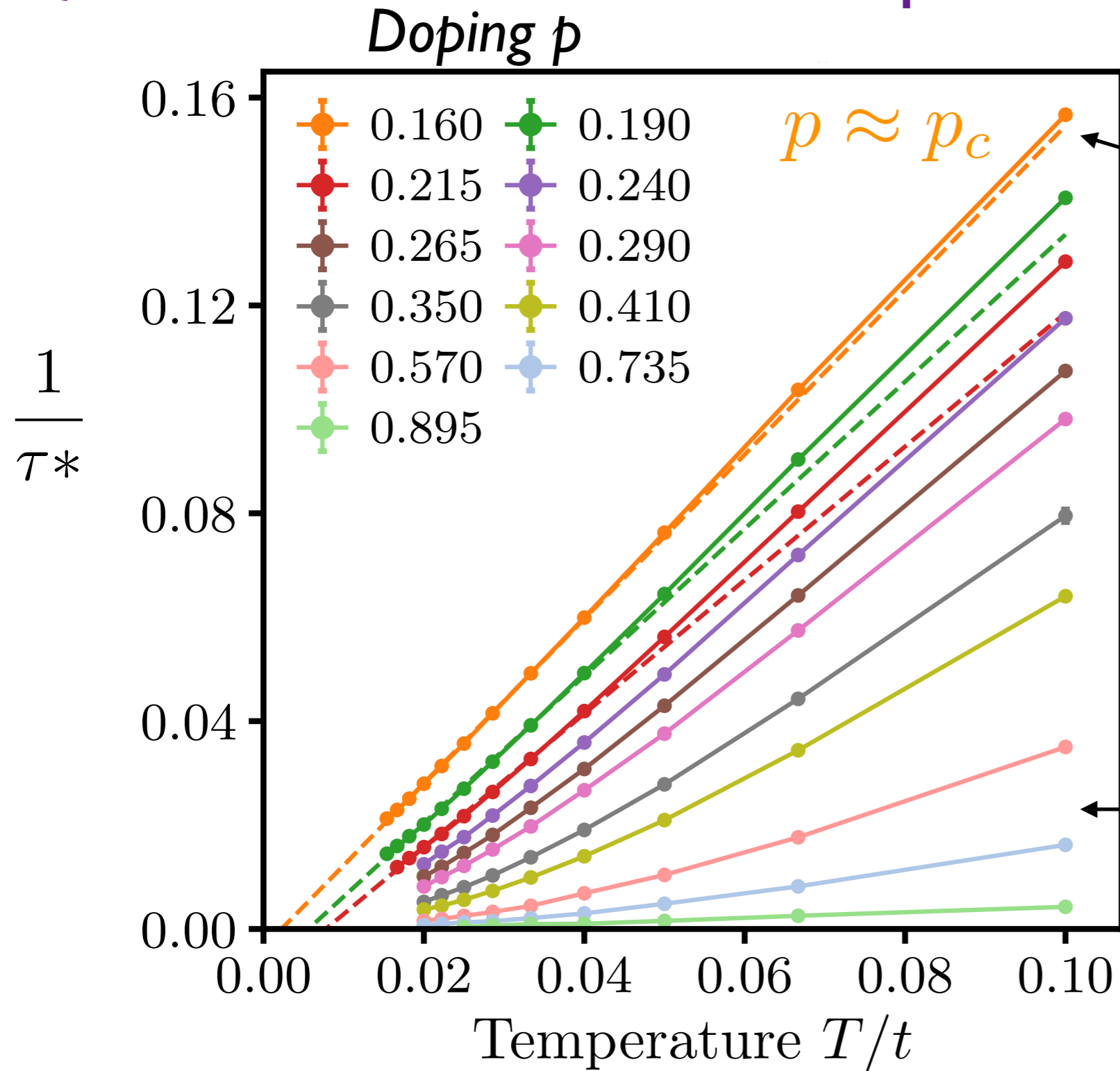
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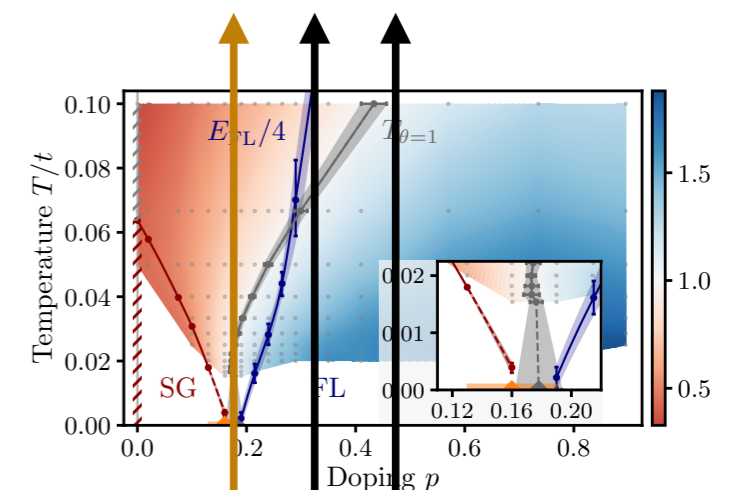
QMC solution of DMFT equations



$$\frac{1}{\tau^*} \simeq c \frac{k_B T}{\hbar}$$

Planckian metal
for $p \approx p_c$

$$\frac{1}{\tau^*} \propto T^2$$



P. T. Dumitrescu, N. Wentzell, A. Georges,
O. Parcollet, arXiv:2103.08607

Summary

- SYK: a solvable model without quasiparticle excitations, exhibiting thermalization and many-body chaos in a time of order $\hbar/(k_B T)$, independent of microscopic energy scales.

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- Low energy theory of time reparameterizations is the theory of the boundary graviton in 2D quantum gravity on AdS_2 .
- Boundary graviton leads to:
 - Dynamic spin susceptibility $\sim \text{sgn}(\omega) [1 - c|\omega| + \dots]$
 - Universal $-3/2 \ln(1/T)$ correction to Bekenstein-Hawking entropy of low T charged black holes in Einstein gravity.

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- Spin $S = 1/2$ $SU(2)$ quantum Sherrington-Kirkpatrick model display co-existence of spin glass order and $SY(K)$ spin liquid behavior
- Random t - J model has a quantum phase transition where spin glass order vanishes, with Planckian metal criticality.