

# Planckian metals and black holes

S. N. Bose Colloquium Series  
on Quantum Materials and Devices  
S. N. Bose National Centre for Basic Sciences  
Kolkata, November 10, 2021

Subir Sachdev

Talk online: [sachdev.physics.harvard.edu](http://sachdev.physics.harvard.edu)



INSTITUTE FOR  
ADVANCED STUDY

PHYSICS



HARVARD

1. Introduction to Planckian metals

2. Introduction to black holes

3. The SYK model

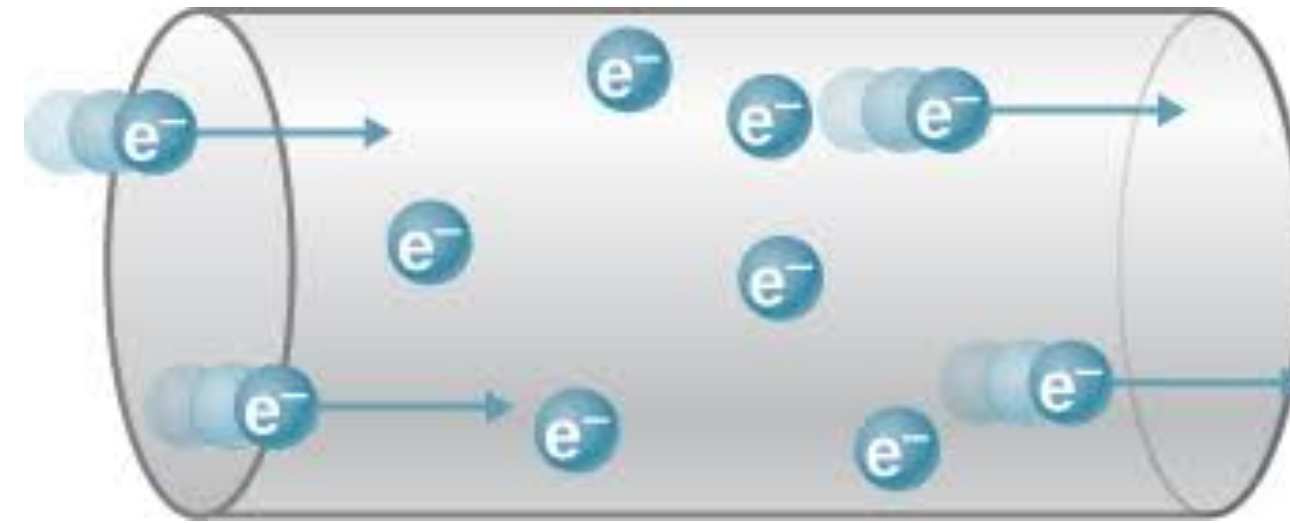
4. Progress on the theory of black holes

5. Progress on the theory of Planckian metals

*A. Random  $t$ - $J$  model*

*B. Fermi surface coupled to a critical boson*

## Current flow with quasiparticles in Copper

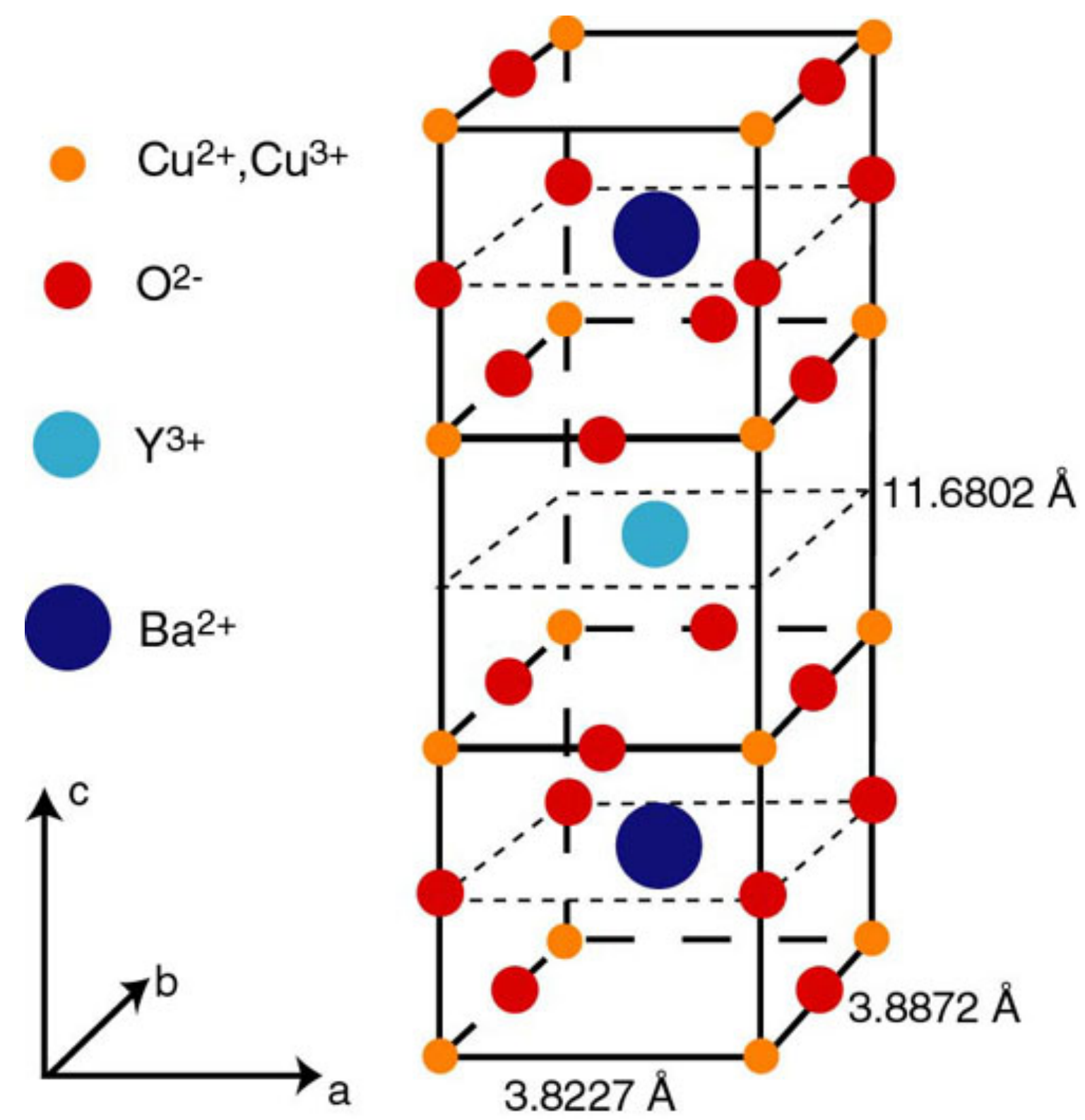
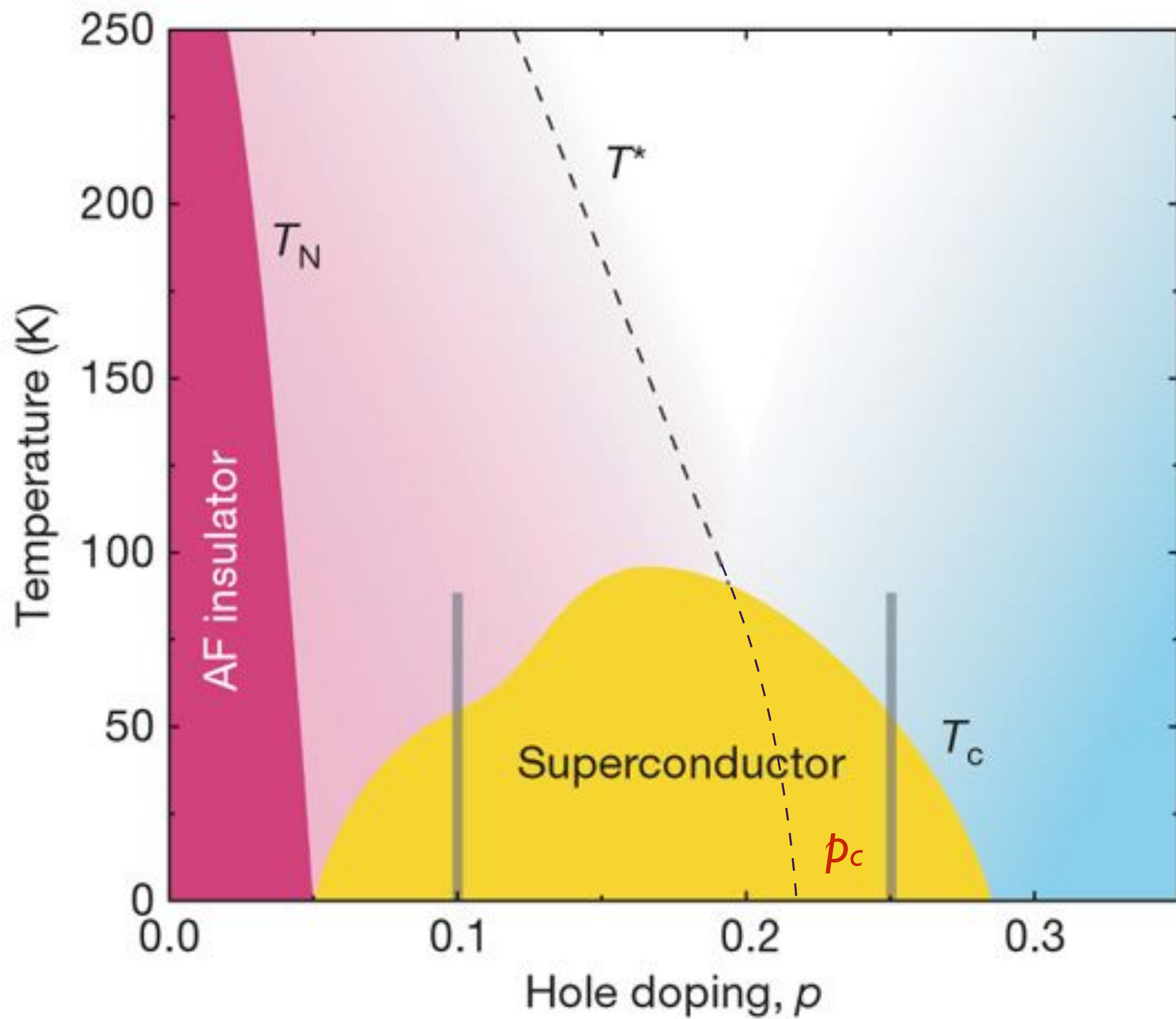


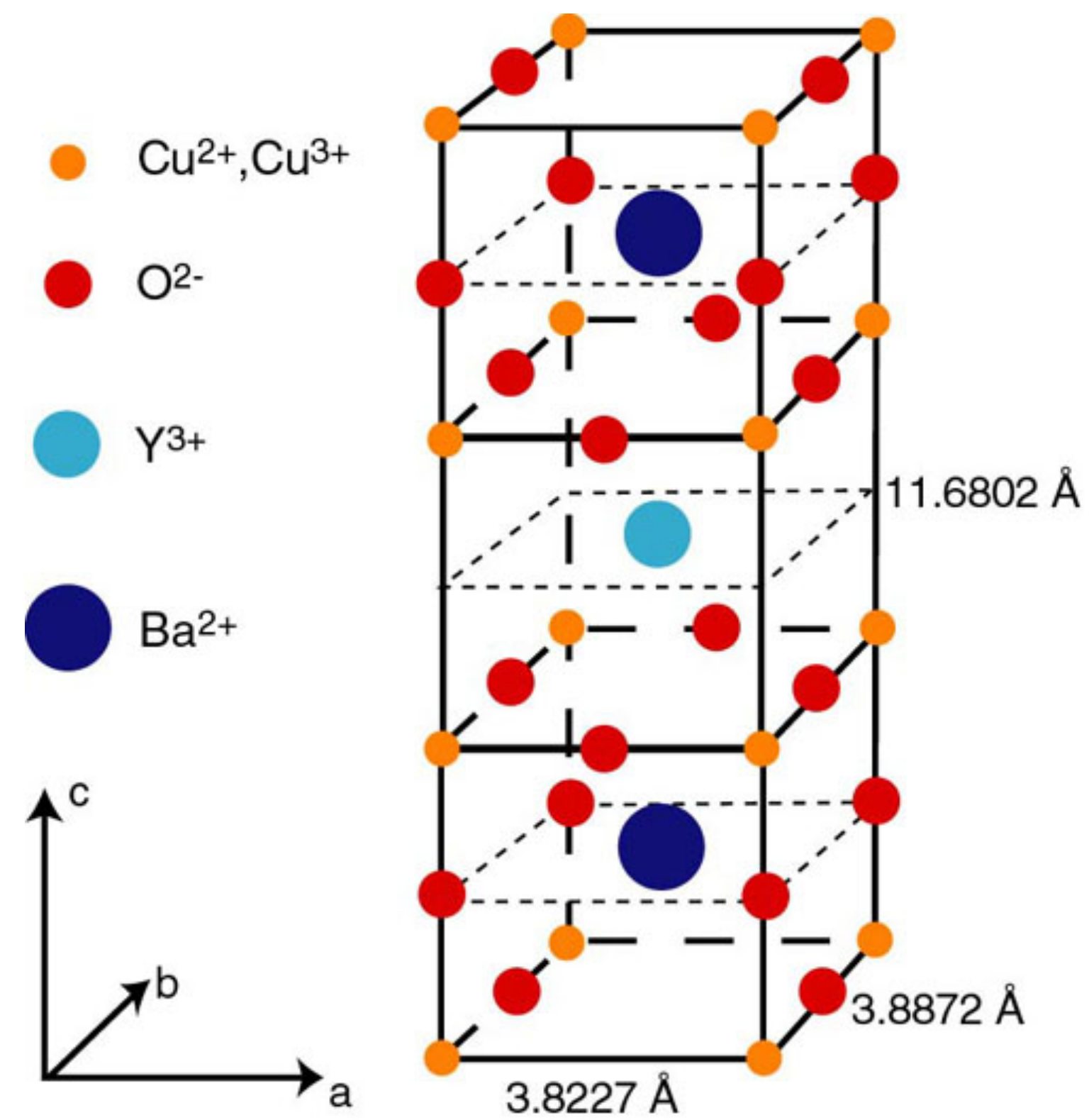
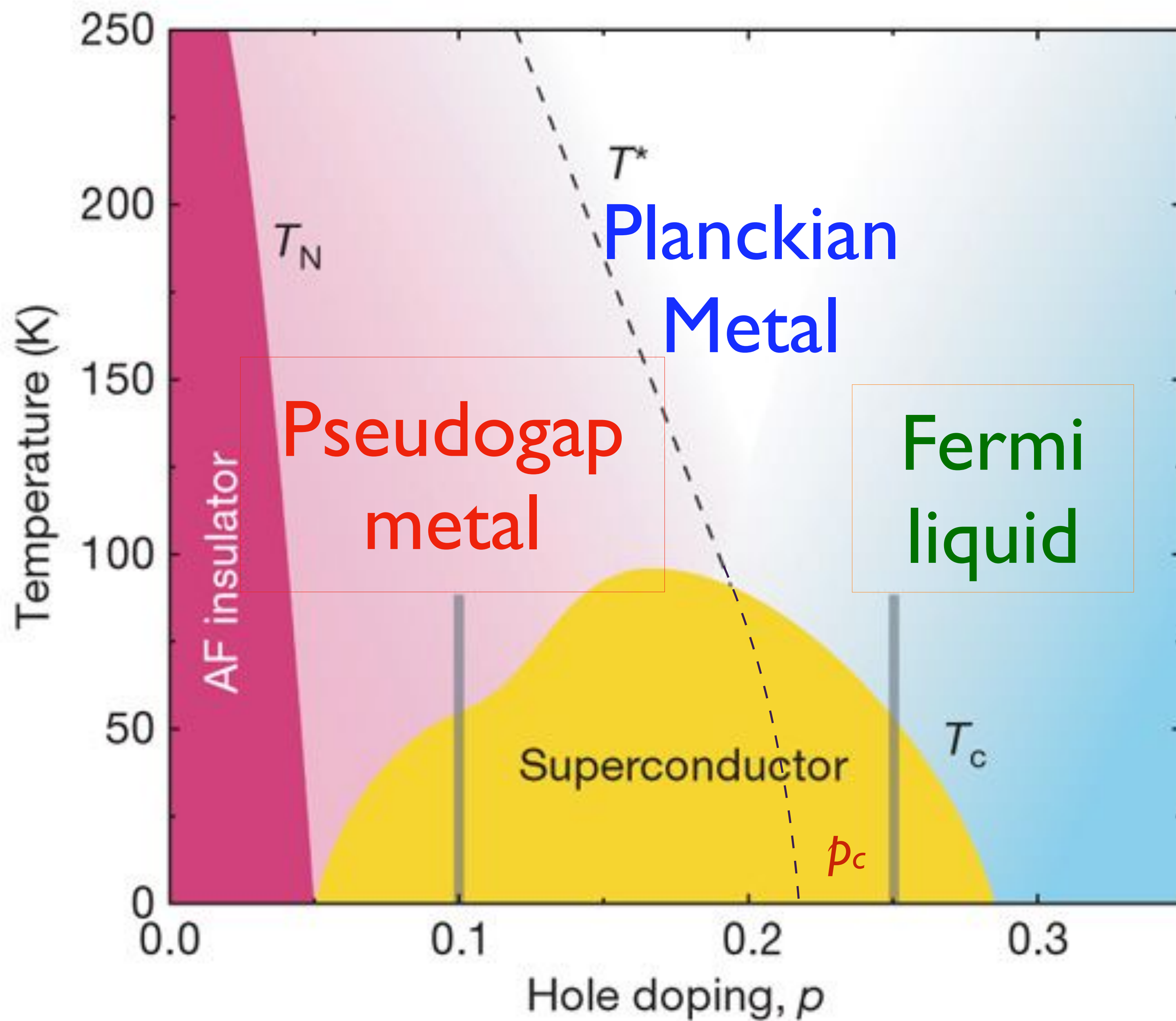
Flowing quasiparticles scatter off each other  
in a typical scattering time  $\tau \sim 1/T^2$

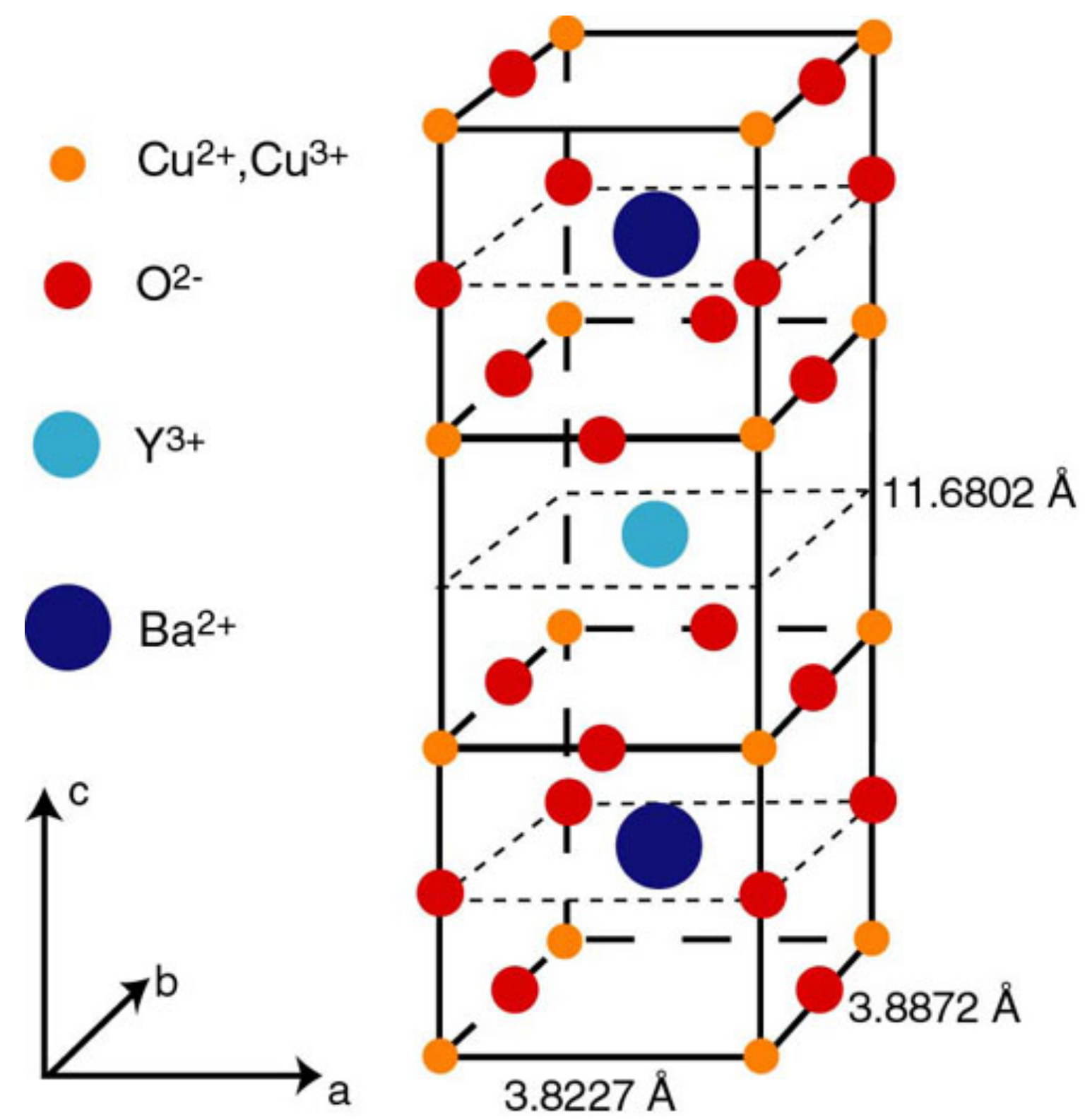
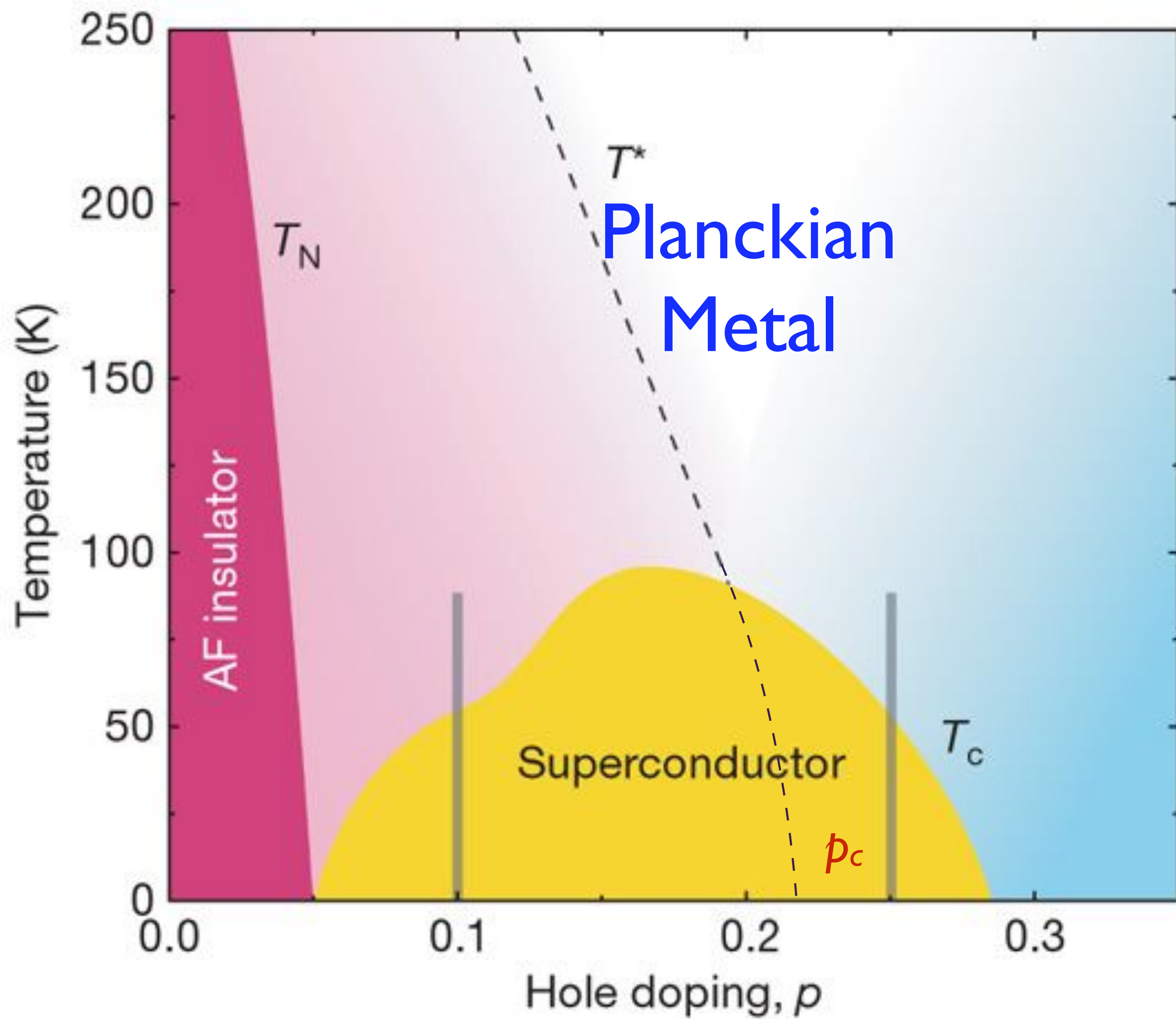
This time is much longer than a limiting ‘Planckian time’  $\frac{\hbar}{k_B T}$ .

The long scattering time implies that quasiparticles are well-defined.

The motion of quasiparticles is ‘ballistic’ or ‘integrable’  
up to the long time  $\tau$ , after which it is chaotic.



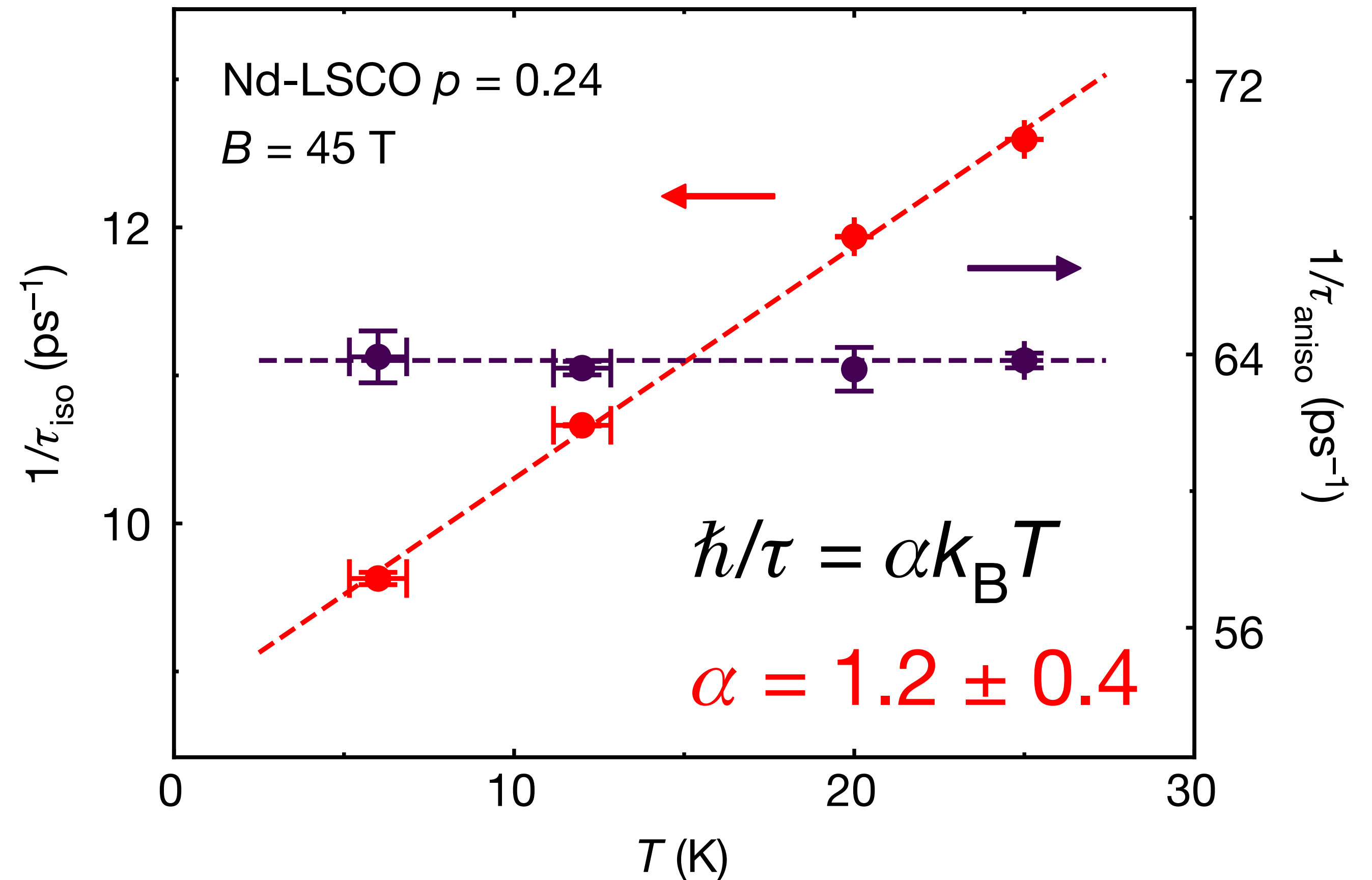
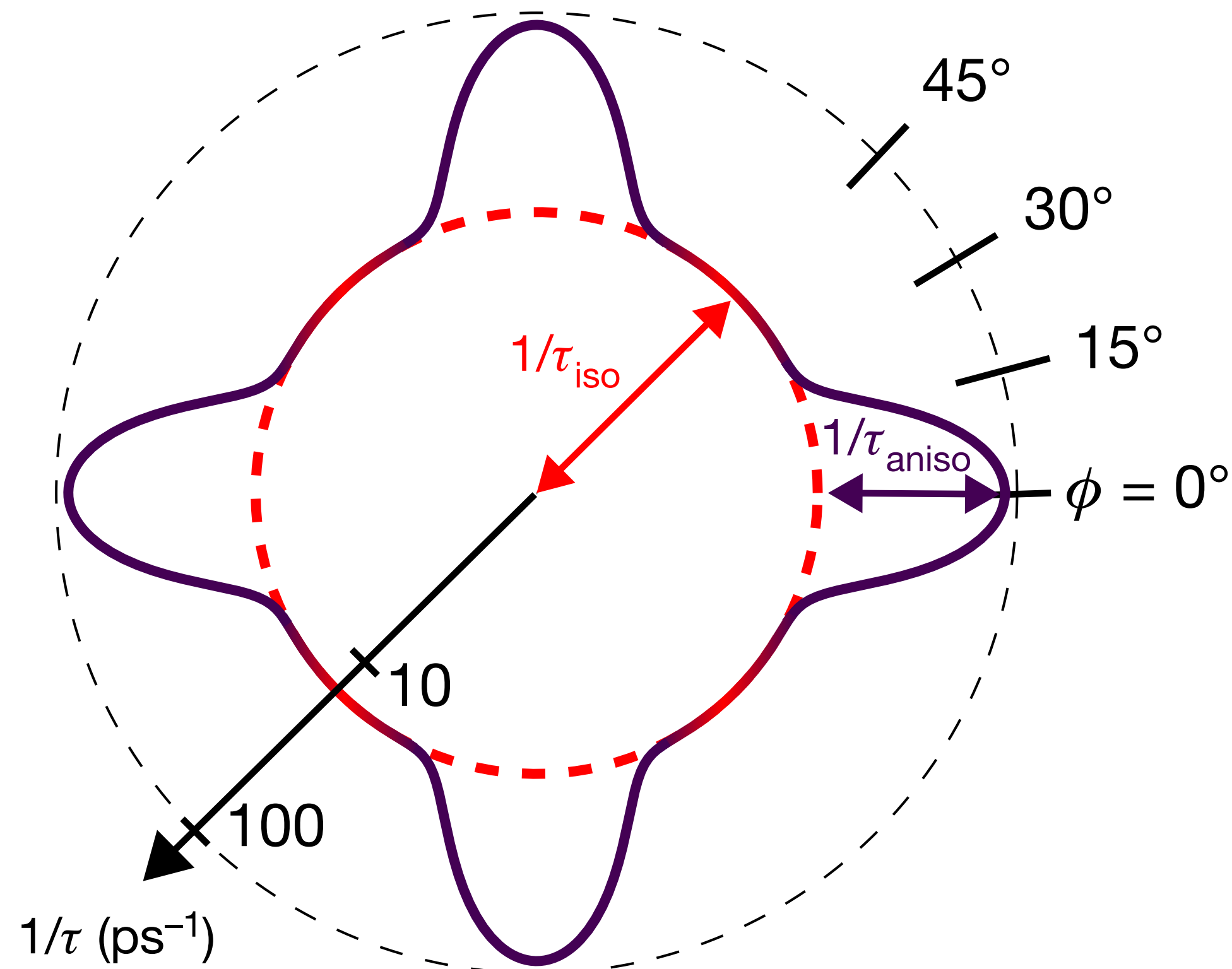




# Linear-in temperature resistivity from an isotropic Planckian scattering rate

Nature **595**, 667-672 (2021)

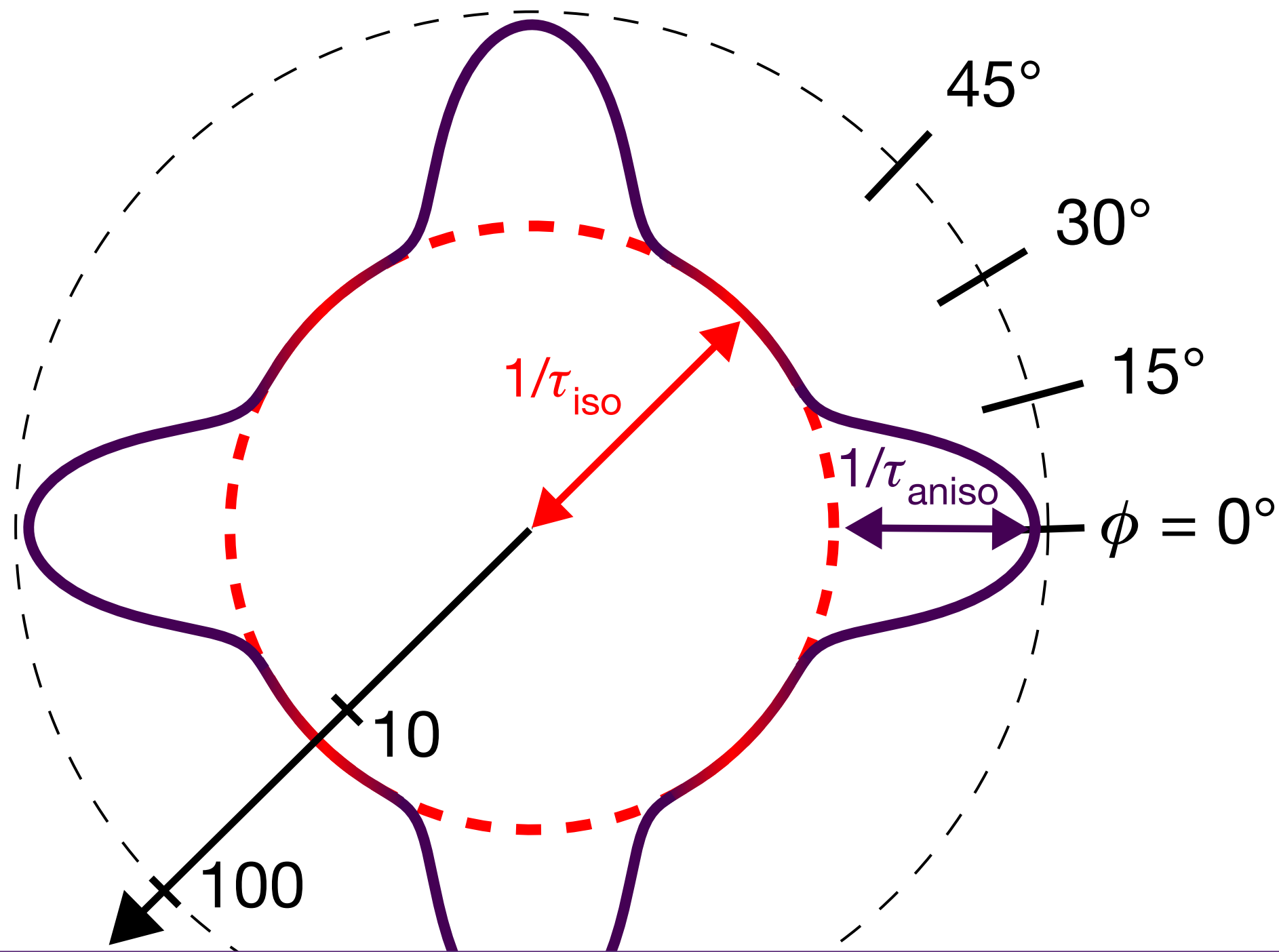
G. Grissonnanche, Y. Fang, A. Legros, S. Verret, F. Laliberté, C. Collignon, J. Zhou, D. Graf, P. Goddard, L. Taillefer, B. J. Ramshaw



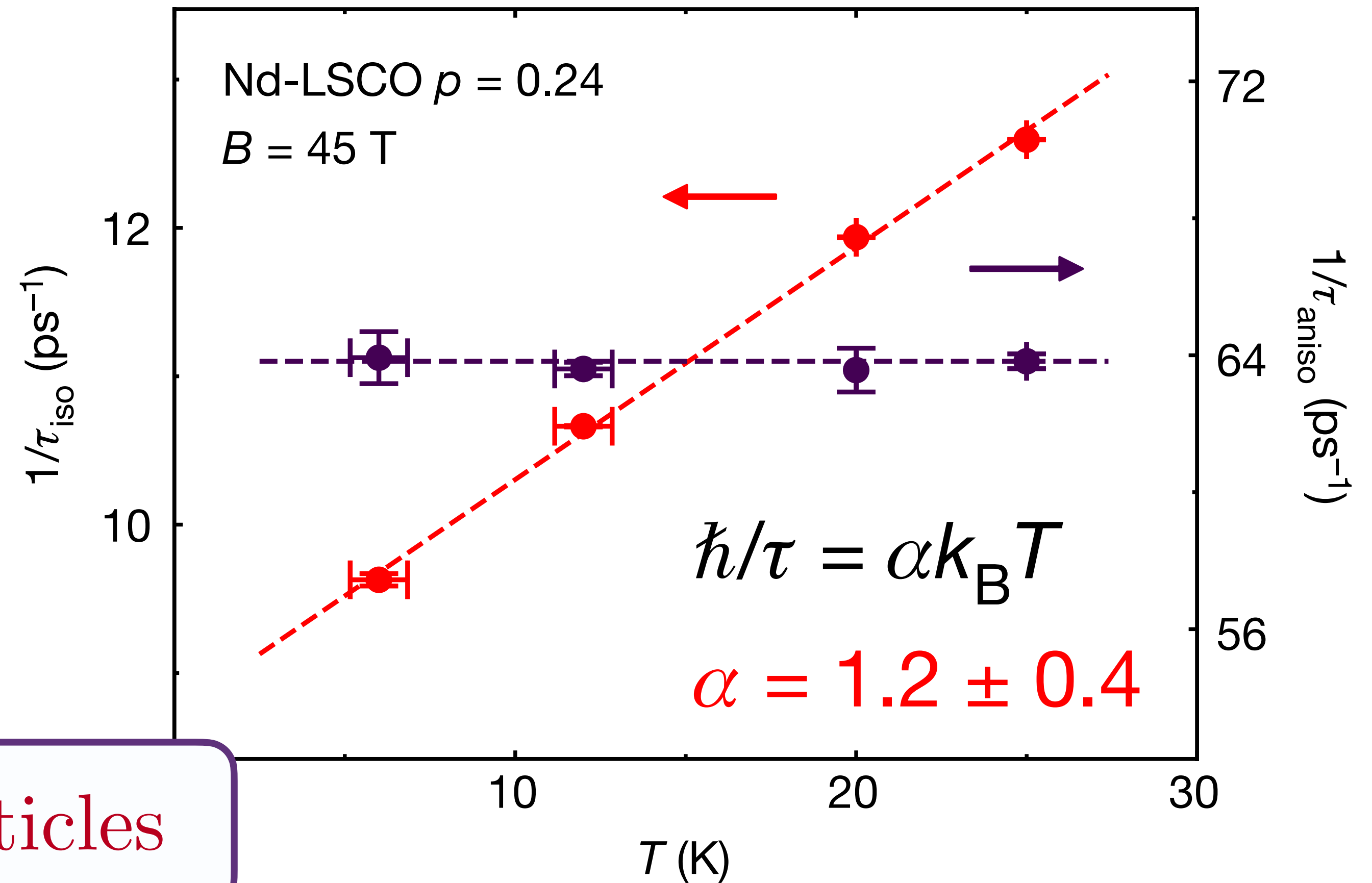
# Linear-in temperature resistivity from an isotropic Planckian scattering rate

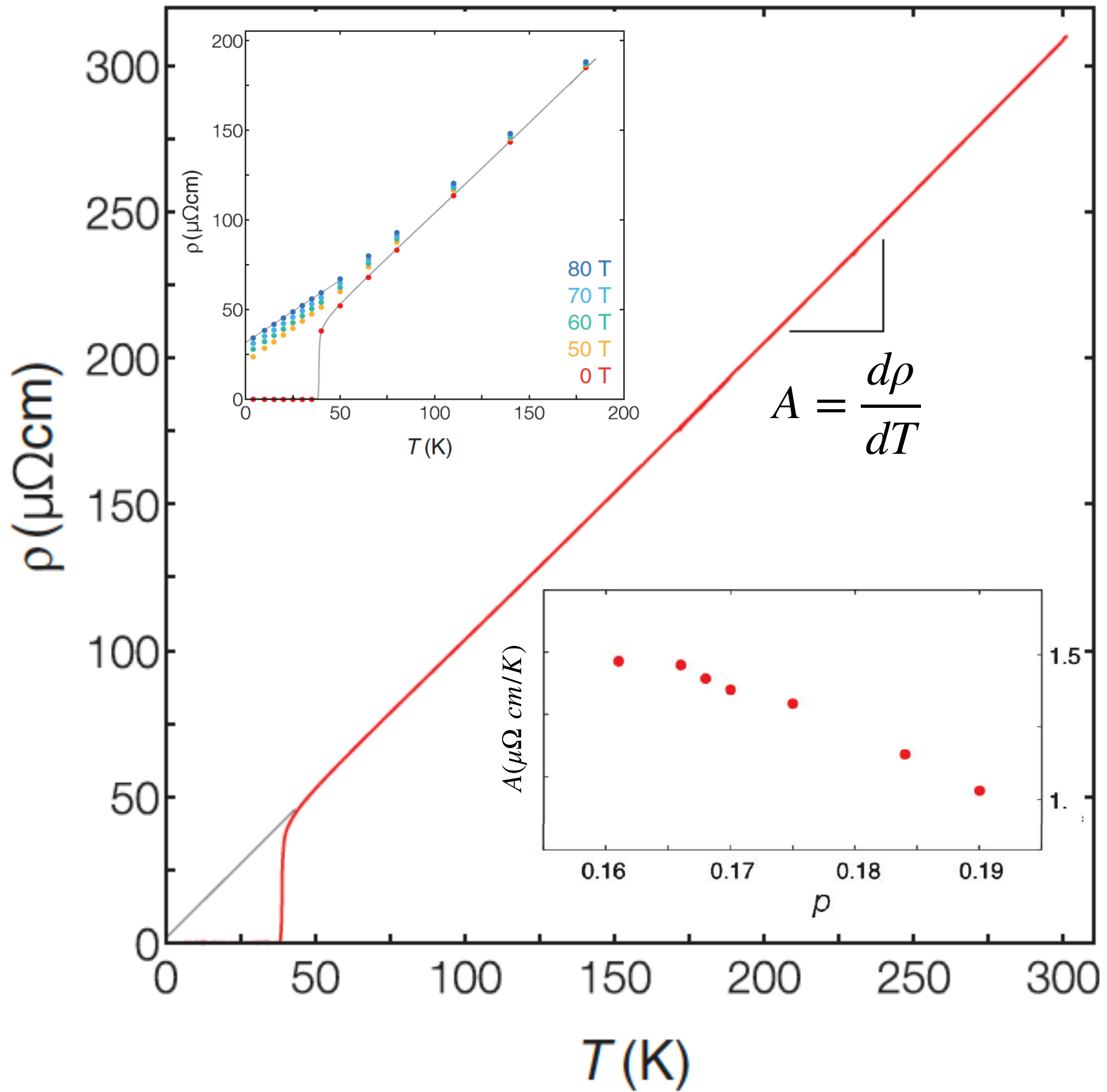
Nature **595**, 667-672 (2021)

G. Grissonnanche, Y. Fang, A. Legros, S. Verret, F. Laliberté, C. Collignon, J. Zhou, D. Graf, P. Goddard, L. Taillefer, B. J. Ramshaw

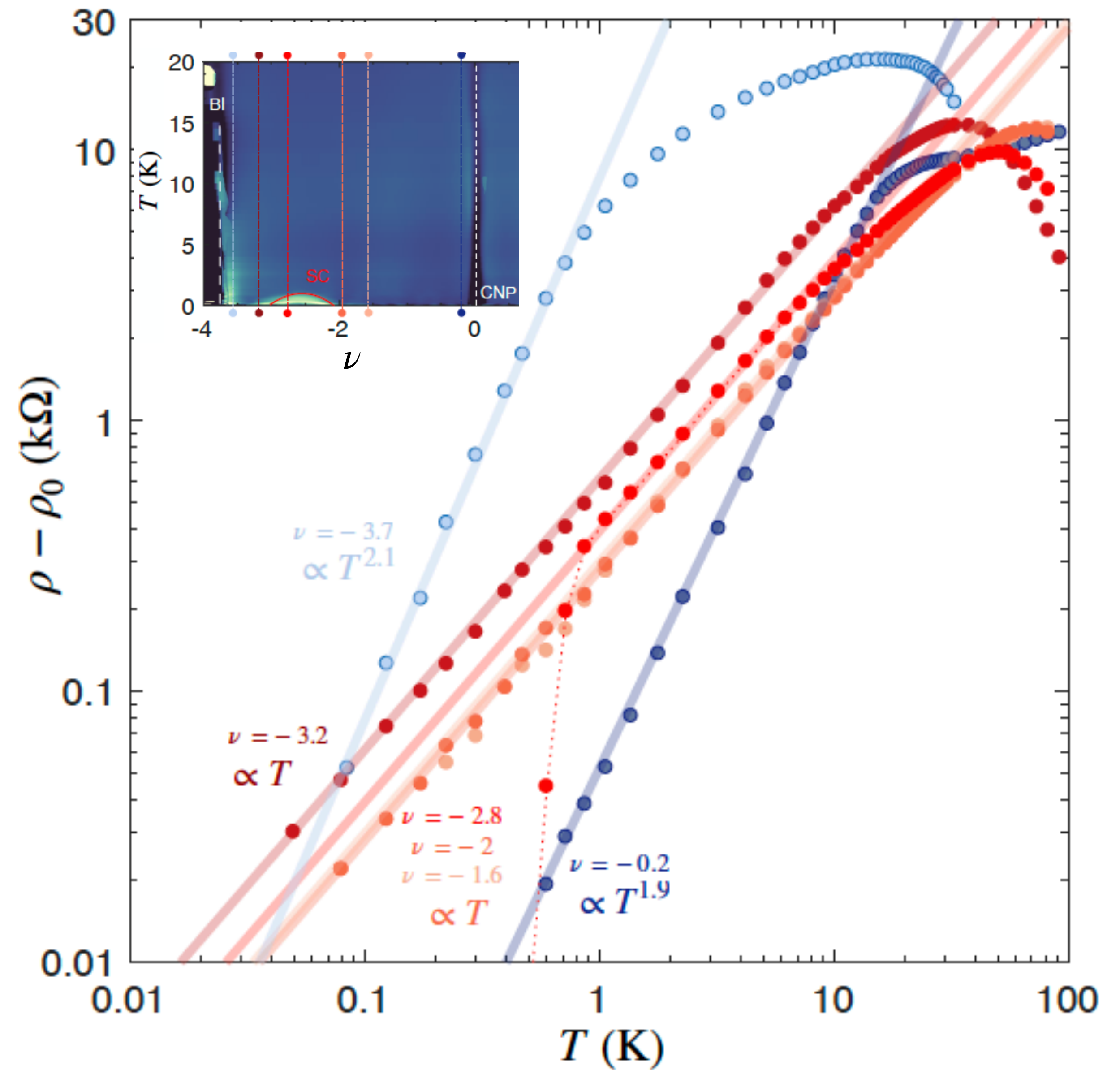


Current flow without quasiparticles





LSCO: Giraldo-Gallo et al. 2018



MATBG: Jaoui et al. 2021

# Questions

- Theory for a fermion system with variable density without quasiparticles, and relaxation time  $\sim \hbar/(k_B T)$ .
- Needed: theory for collision time in resistivity  $\sim \hbar/(k_B T)$ .
- Needed: theory for the appearance of superconductivity (and other broken symmetries) in such a ‘Planckian metal’.

1. Introduction to Planckian metals

2. Introduction to black holes

3. The SYK model

4. Progress on the theory of black holes

5. Progress on the theory of Planckian metals

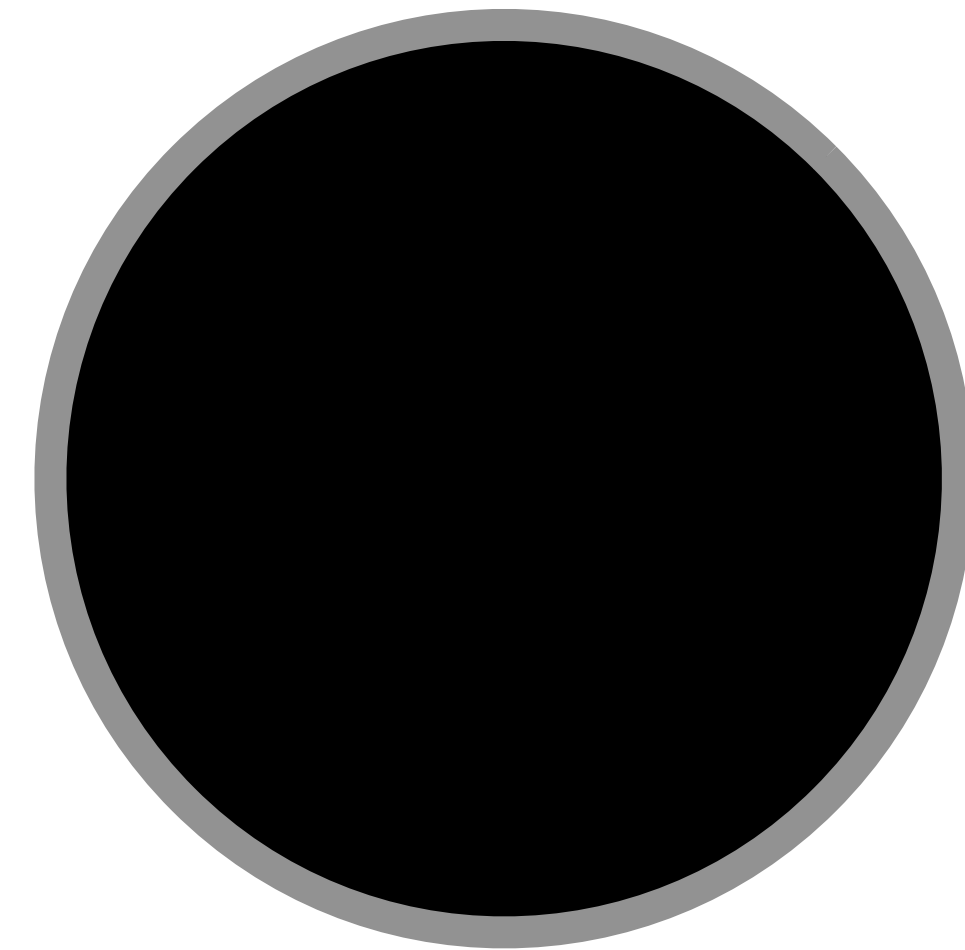
*A. Random  $t$ - $J$  model*

*B. Fermi surface coupled to a critical boson*

# Black Holes

Objects so dense that light is gravitationally bound to them.

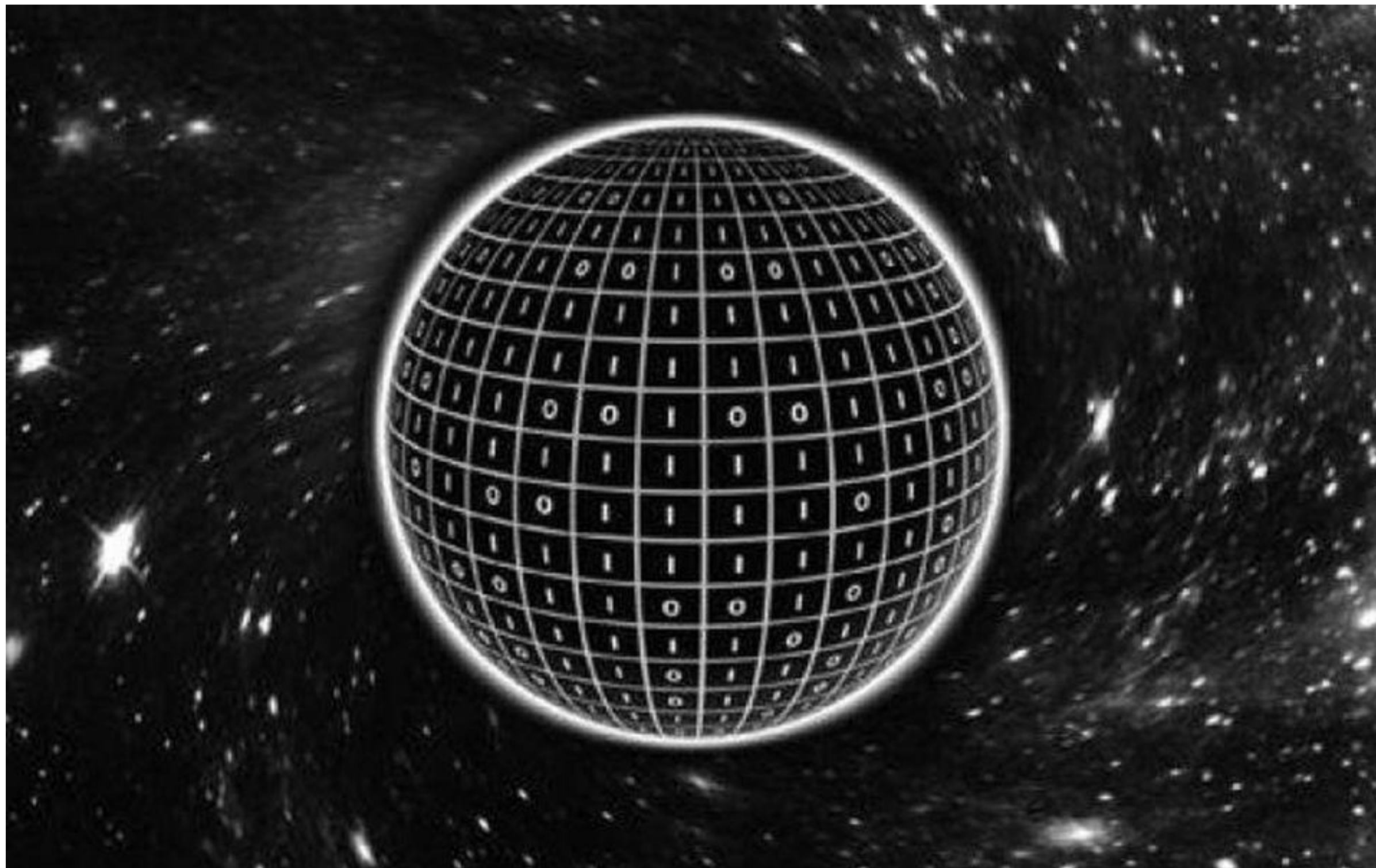
Horizon radius  $R = \frac{2GM}{c^2}$



$G$  Newton's constant,  $c$  velocity of light,  $M$  mass of black hole  
For  $M = \text{earth's mass}$ ,  $R \approx 9 \text{ mm!}$

# Quantum Black holes

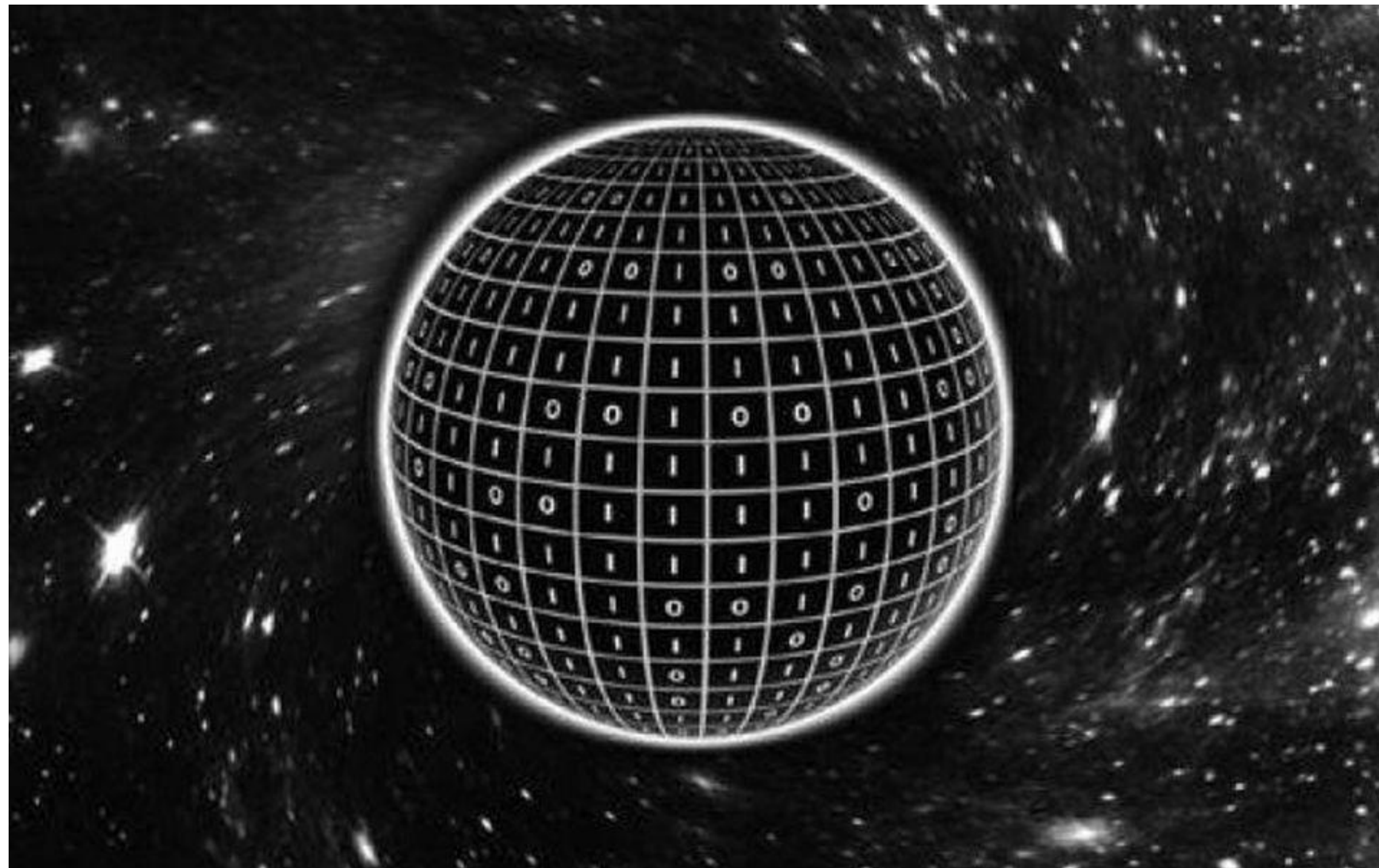
- Black holes have an entropy and a temperature,  $T_H = \hbar c^3 / (8\pi G M k_B)$ .
- The entropy is proportional to their surface area.



J. D. Bekenstein, PRD **7**, 2333 (1973)  
S.W. Hawking, Nature **248**, 30 (1974)

# Quantum Black holes

- Black holes have an entropy and a temperature,  $T_H = \hbar c^3 / (8\pi G M k_B)$ .
- The entropy is proportional to their surface area.



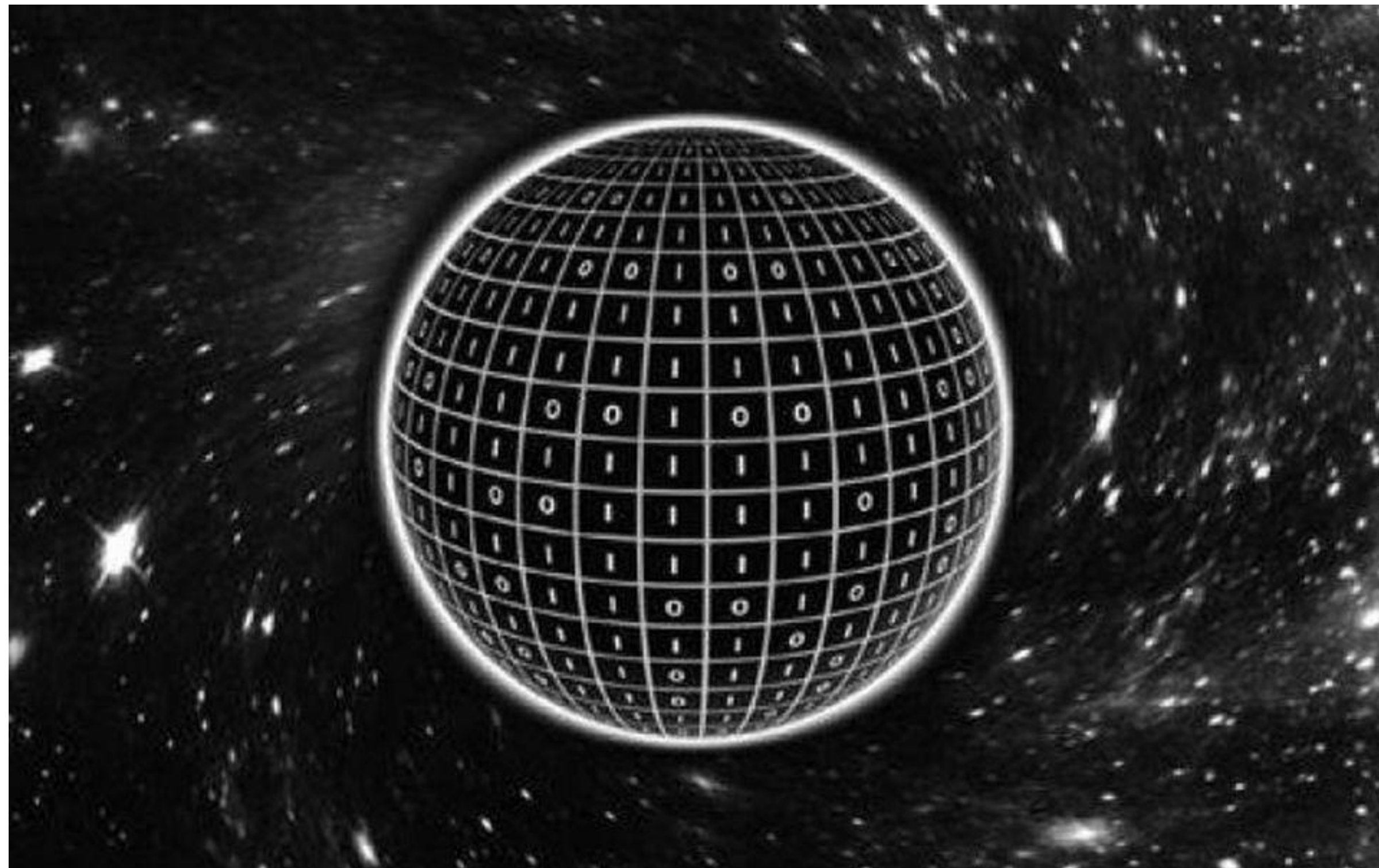
J. D. Bekenstein, PRD **7**, 2333 (1973)  
S.W. Hawking, Nature **248**, 30 (1974)

## Remarkable features:

- Entropy is finite.
- Entropy is not proportional to volume

# Quantum Black holes

- Black holes have an entropy and a temperature,  $T_H = \hbar c^3 / (8\pi G M k_B)$ .
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a Planckian time  $\sim 8\pi G M / c^3$



J. D. Bekenstein, PRD **7**, 2333 (1973)  
S.W. Hawking, Nature **248**, 30 (1974)

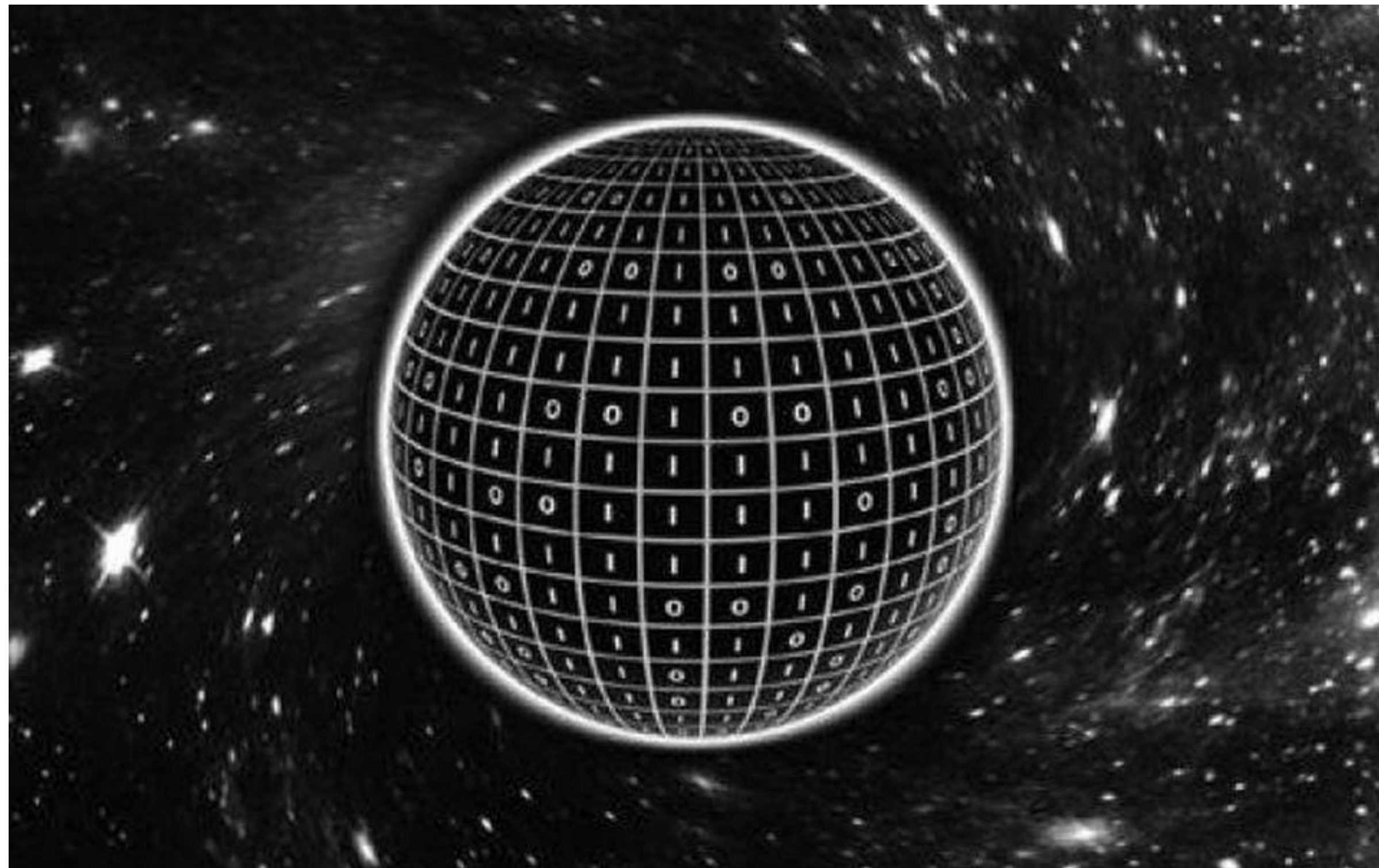
C.V. Vishveshwara, Nature **227**, 936 (1970)

## Remarkable features:

- Entropy is finite.
- Entropy is not proportional to volume

# Quantum Black holes

- Black holes have an entropy and a temperature,  $T_H = \hbar c^3 / (8\pi G M k_B)$ .
- The entropy is proportional to their surface area.
- They relax to thermal equilibrium in a Planckian time  $\sim 8\pi G M / c^3 = \hbar / (k_B T_H)$ .



J. D. Bekenstein, PRD **7**, 2333 (1973)

S.W. Hawking, Nature **248**, 30 (1974)

C.V. Vishveshwara, Nature **227**, 936 (1970)

## Remarkable features:

- Entropy is finite.
- Entropy is not proportional to volume

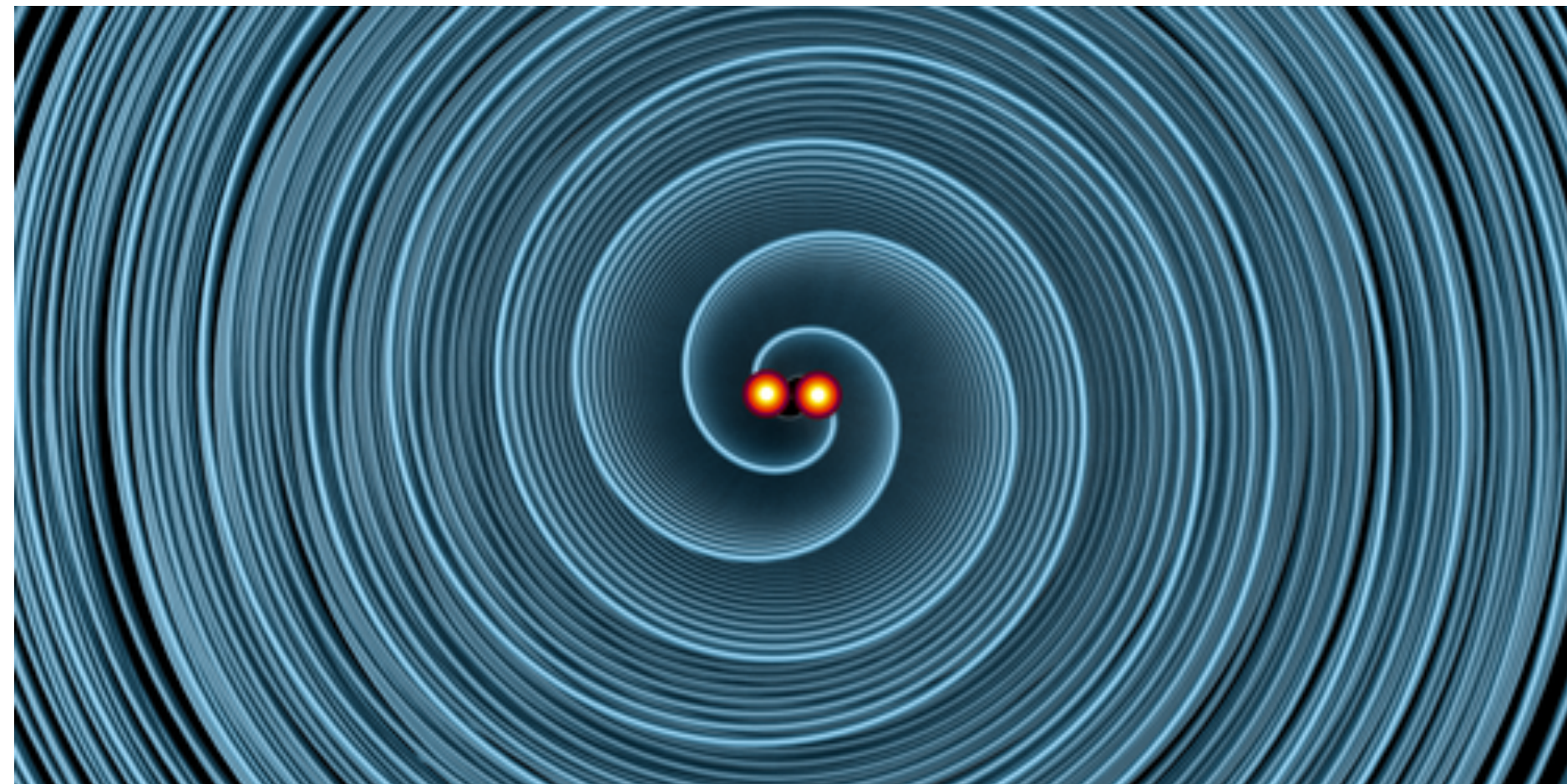
# Black Holes Obey Information-Emission Limits

## Limits

April 22, 2021 • *Physics 14, s47* –Christopher Crockett

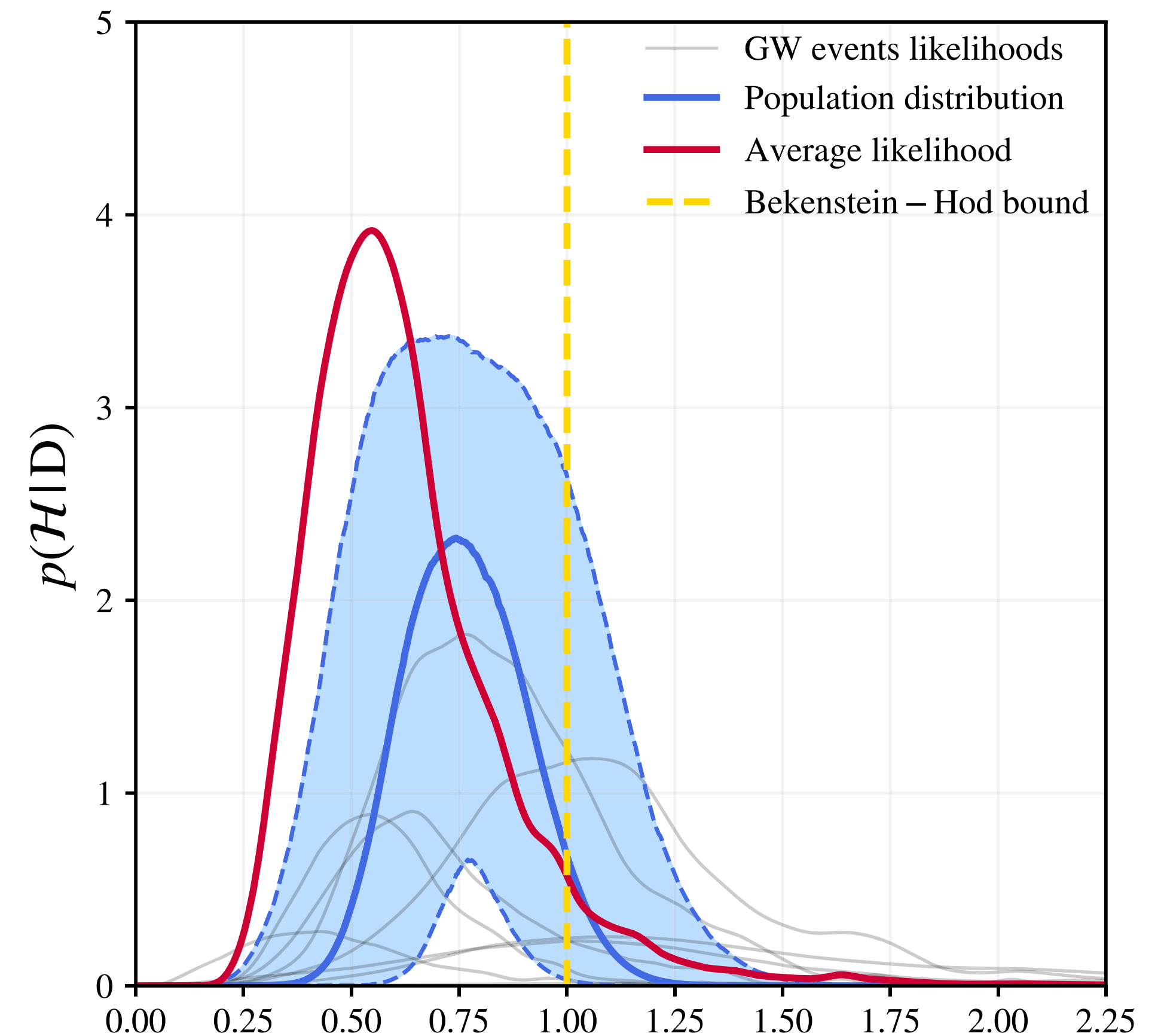
G. Carullo, D. Laghi, J. Veitch, W. Del Pozzo, *Phys. Rev. Lett.* **126**, 161102 (2021)

An analysis of the gravitational waves emitted from black hole mergers confirms that black holes are the fastest known information dissipaters.



Gravity wave observations of 8 different black holes show a relaxation time

$$\tau \sim \frac{\hbar}{k_B T}$$



$$\mathcal{H} = \frac{1}{\pi} \frac{\hbar/\tau}{k_B T}$$

# Thermodynamics of quantum black holes with charge $Q$ :



$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right)$$

Metric of  
spacetime

Electromagnetic  
gauge field

In general, this integral is not well defined, because of an uncontrollably large number of spacetime configurations.

# Thermodynamics of quantum black holes with charge $Q$ :



$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right)$$
$$= \exp(S_{BH}) \times \left( \dots????\dots \right)$$

Gibbons, Hawking (1977)

Chambin, Emparan, Johnson, Myers (1999)

$$S_{BH}(T \rightarrow 0, Q) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left( 1 + \frac{2(\pi A_0)^{1/2}T}{\hbar c} \right)$$

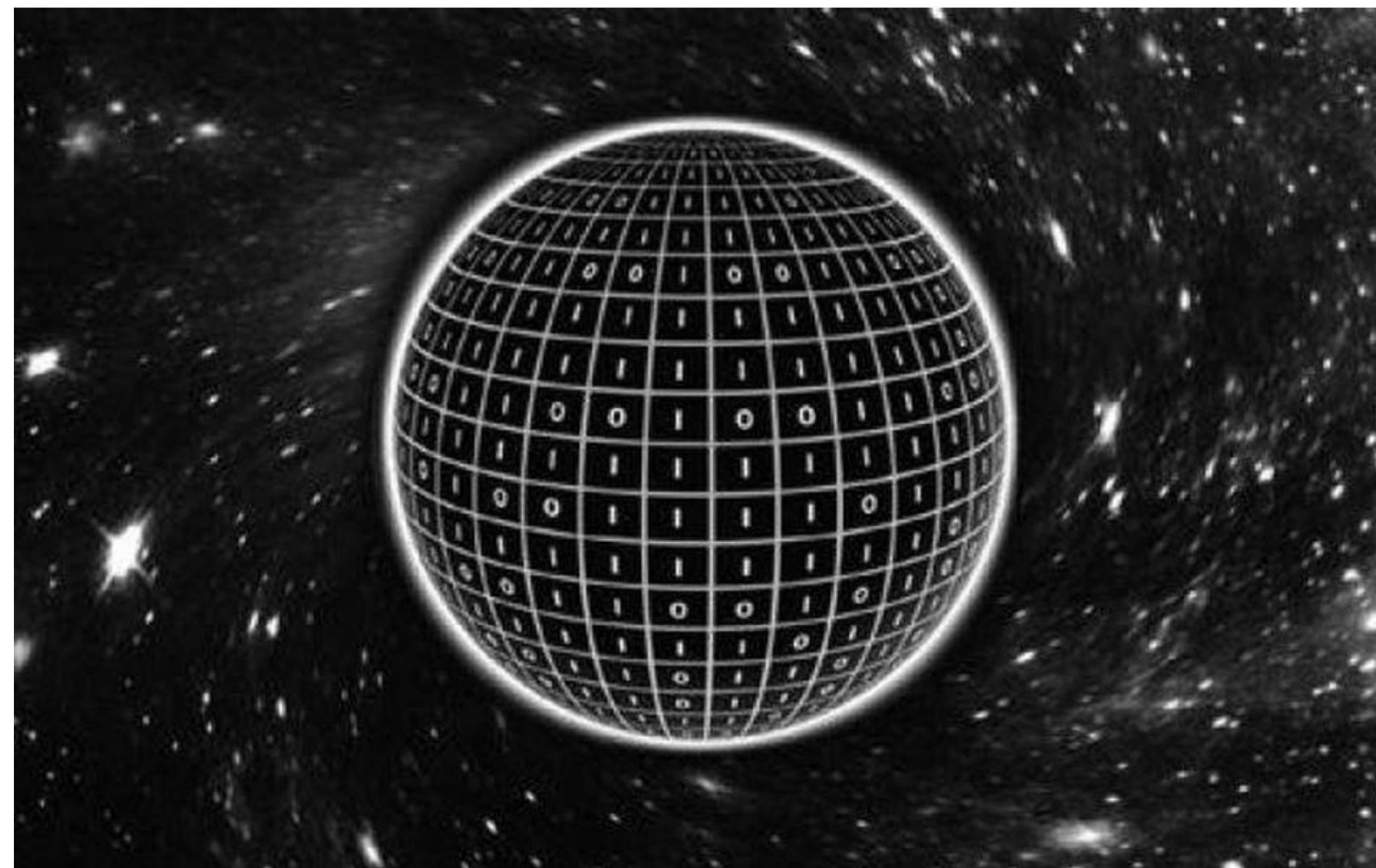
$A_0$  is the area of the charged black hole horizon at  $T = 0$ .

$Q$  is the black hole charge.

$A_0$  is a function of  $Q$ .

# Questions

- Is Einstein-Maxwell theory meaningful beyond the saddle point, and can we compute quantum fluctuation corrections to  $S_{BH}$ ?
- Can the resulting entropy be understood as that of a unitary quantum system with a discrete spectrum ?
- Can we compute the evolution of the entropy as the black hole evaporates? Is it that of an evaporating unitary quantum system?



1. Introduction to Planckian metals

2. Introduction to black holes

3. The SYK model

4. Progress on the theory of black holes

5. Progress on the theory of Planckian metals

*A. Random  $t$ - $J$  model*

*B. Fermi surface coupled to a critical boson*

Needed:

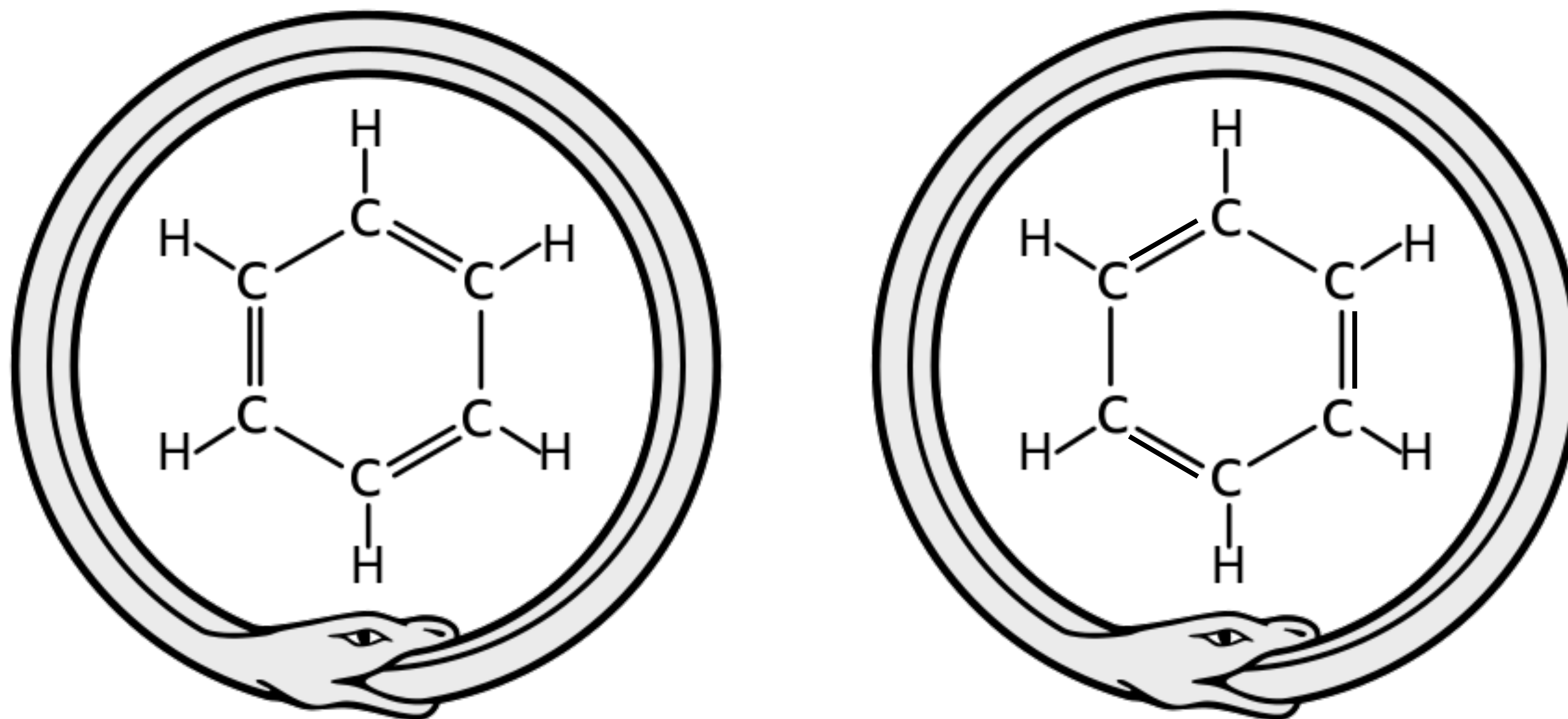
A solvable model of multi-particle entanglement leading to a compressible quantum state with no quasiparticle excitations

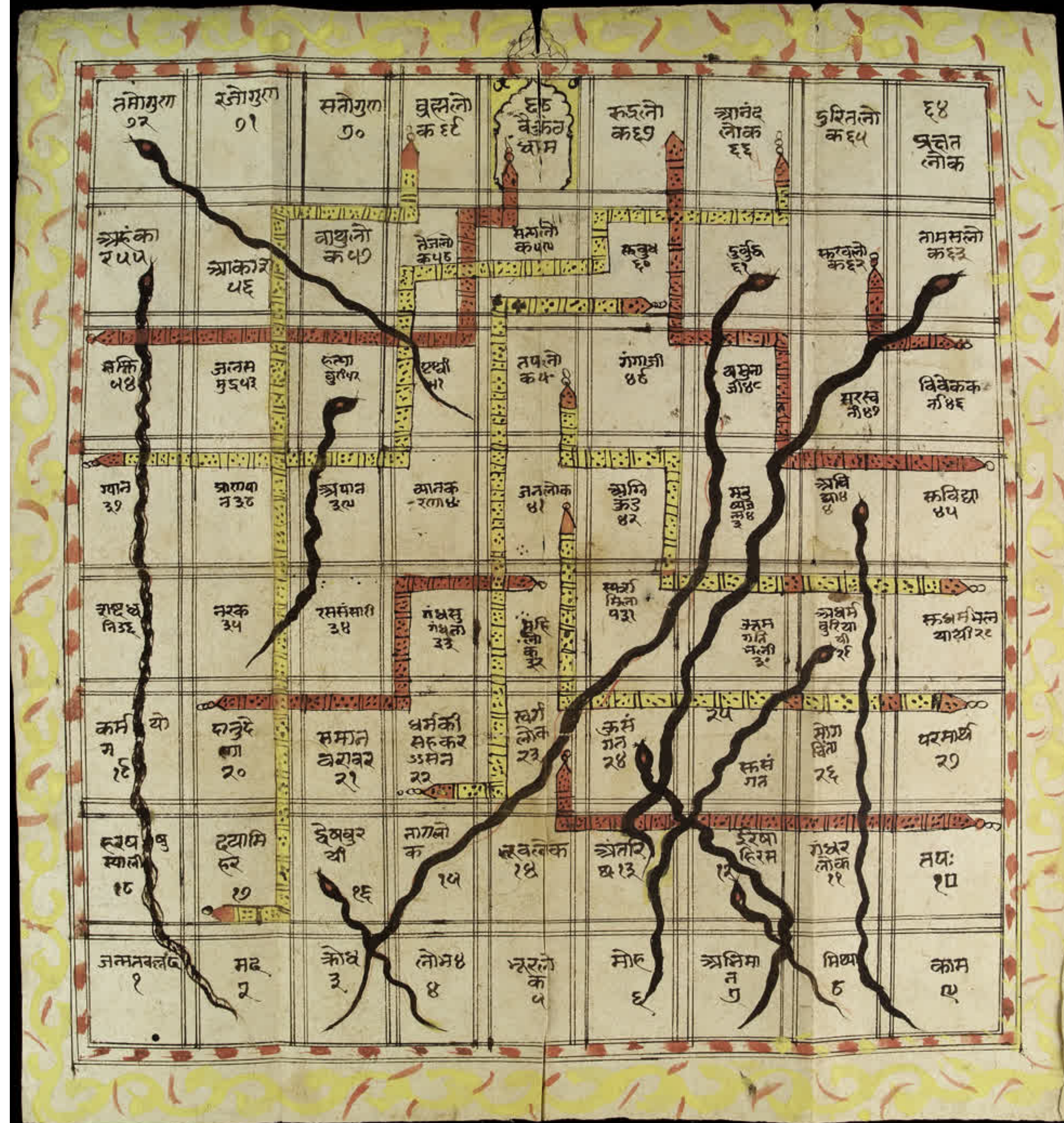


August Kekule, theory of the benzene molecule, 1865

# Kekule's dream

Kekulé spoke of the creation of the theory. He said that he had discovered the ring shape of the benzene molecule after having a reverie or day-dream of a snake seizing its own tail\*





My  
dream\*

Snakes  
and  
ladders

\*Not true

# The Sachdev-Ye-Kitaev (SYK) model

(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large  $N$  limit;  
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} - \mu \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

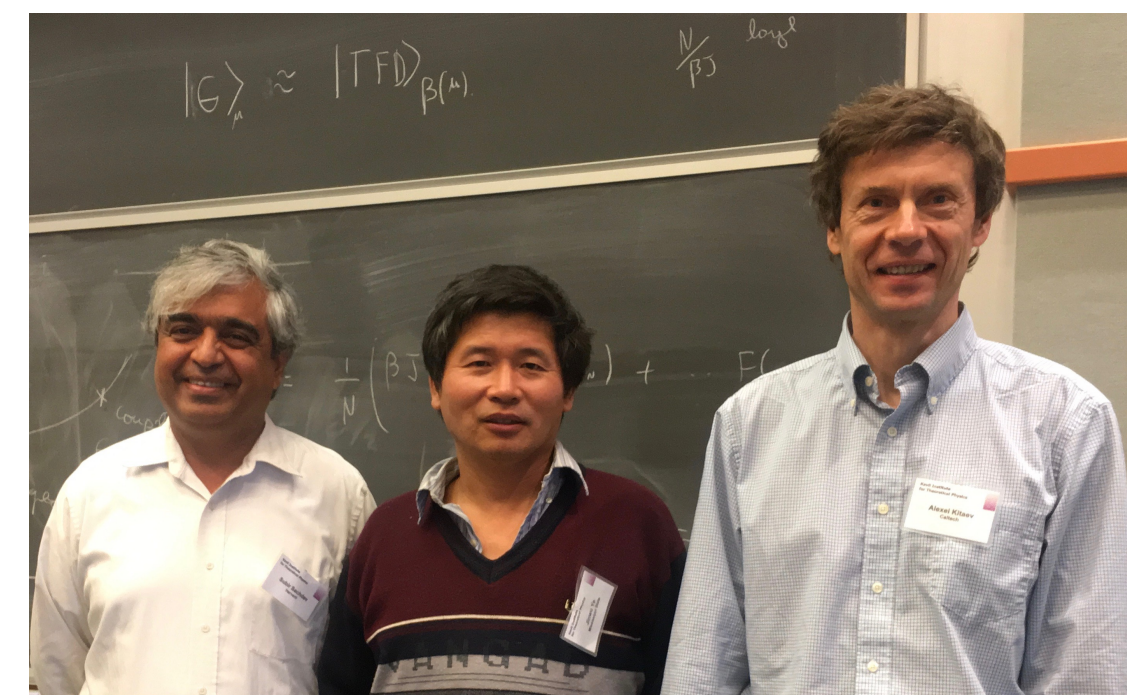
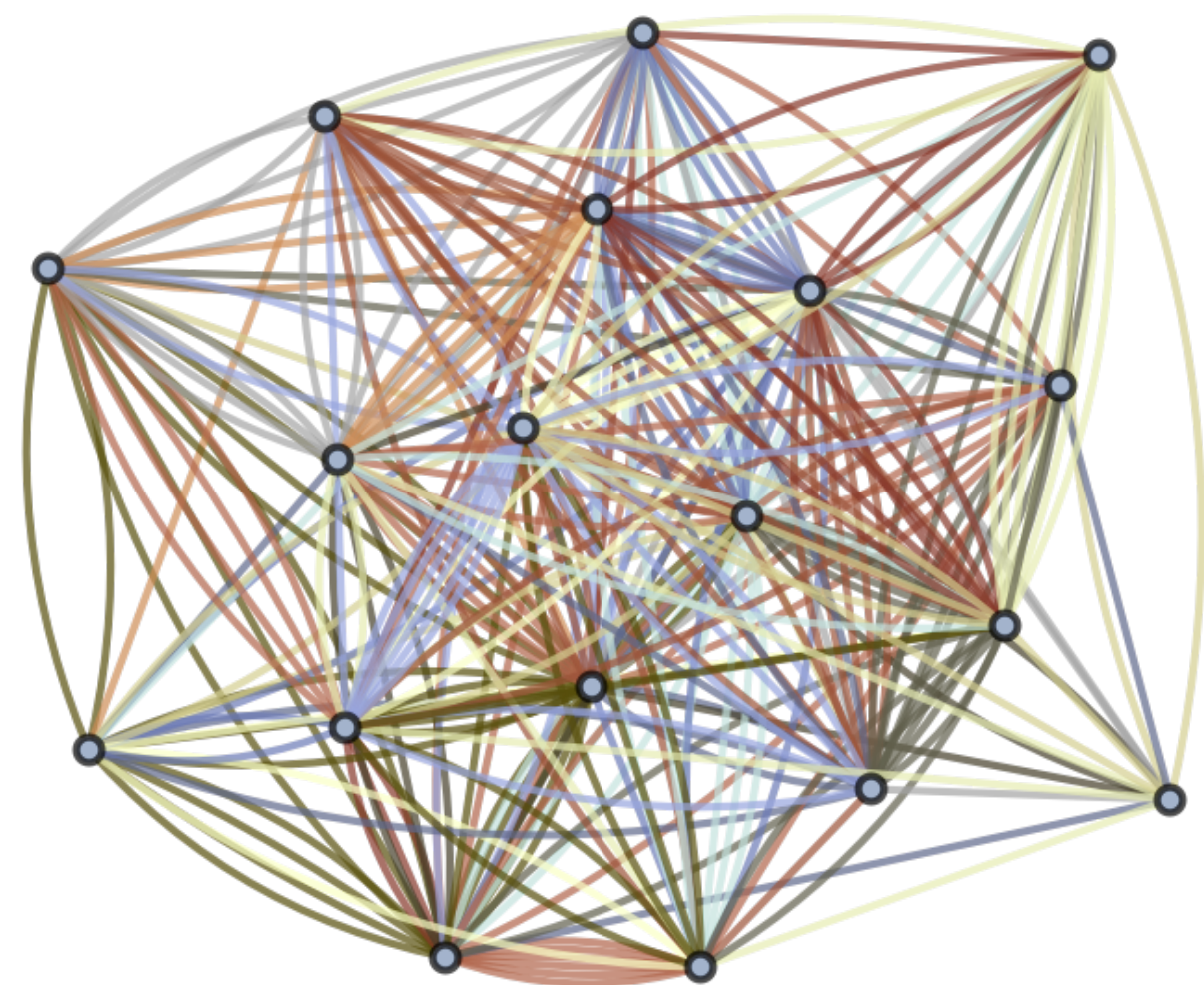
$$c_{\alpha} c_{\beta} + c_{\beta} c_{\alpha} = 0 \quad , \quad c_{\alpha} c_{\beta}^{\dagger} + c_{\beta}^{\dagger} c_{\alpha} = \delta_{\alpha\beta}$$

$$Q = \frac{1}{N} \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$U_{\alpha\beta;\gamma\delta}$  are independent random variables with  $\overline{U_{\alpha\beta;\gamma\delta}} = 0$  and  $\overline{|U_{\alpha\beta;\gamma\delta}|^2} = U^2$   
 $N \rightarrow \infty$  yields critical strange metal.

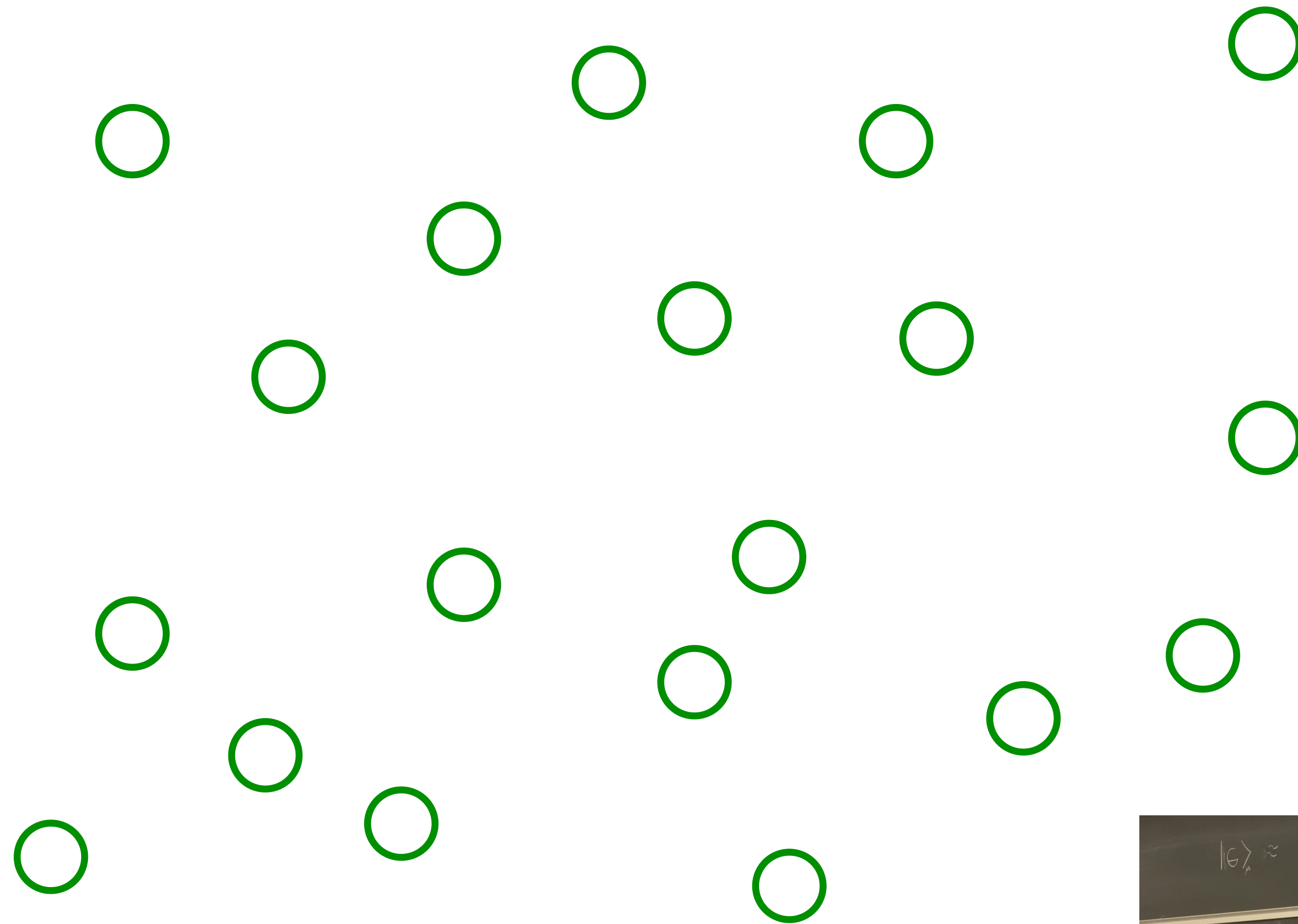
S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)

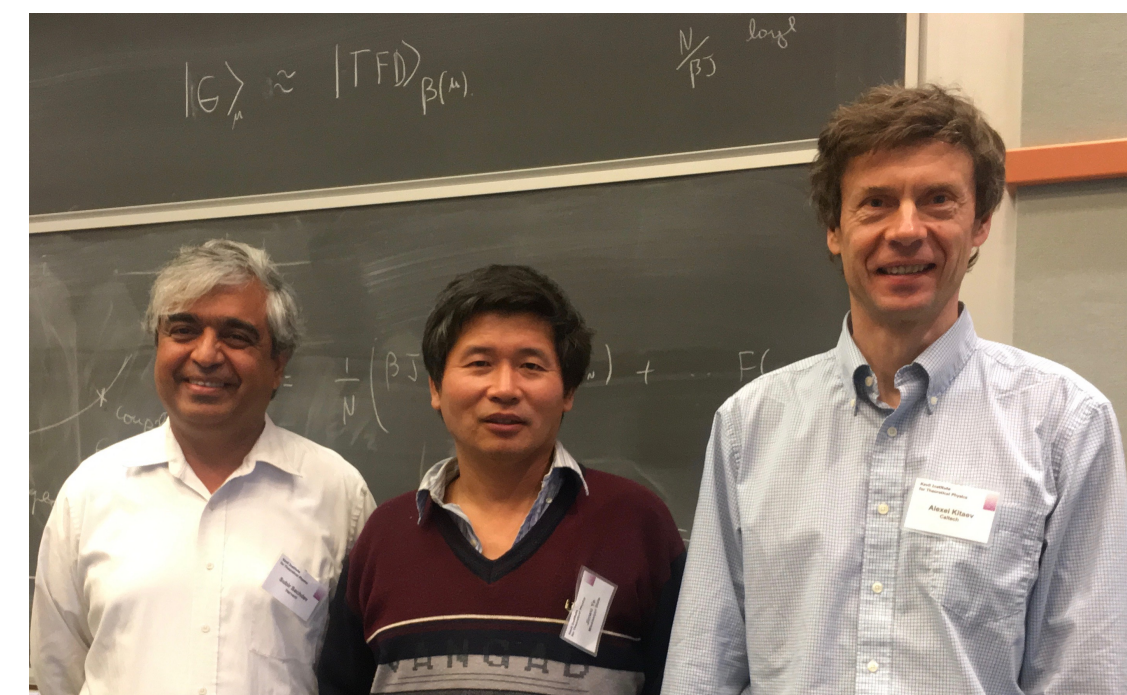


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

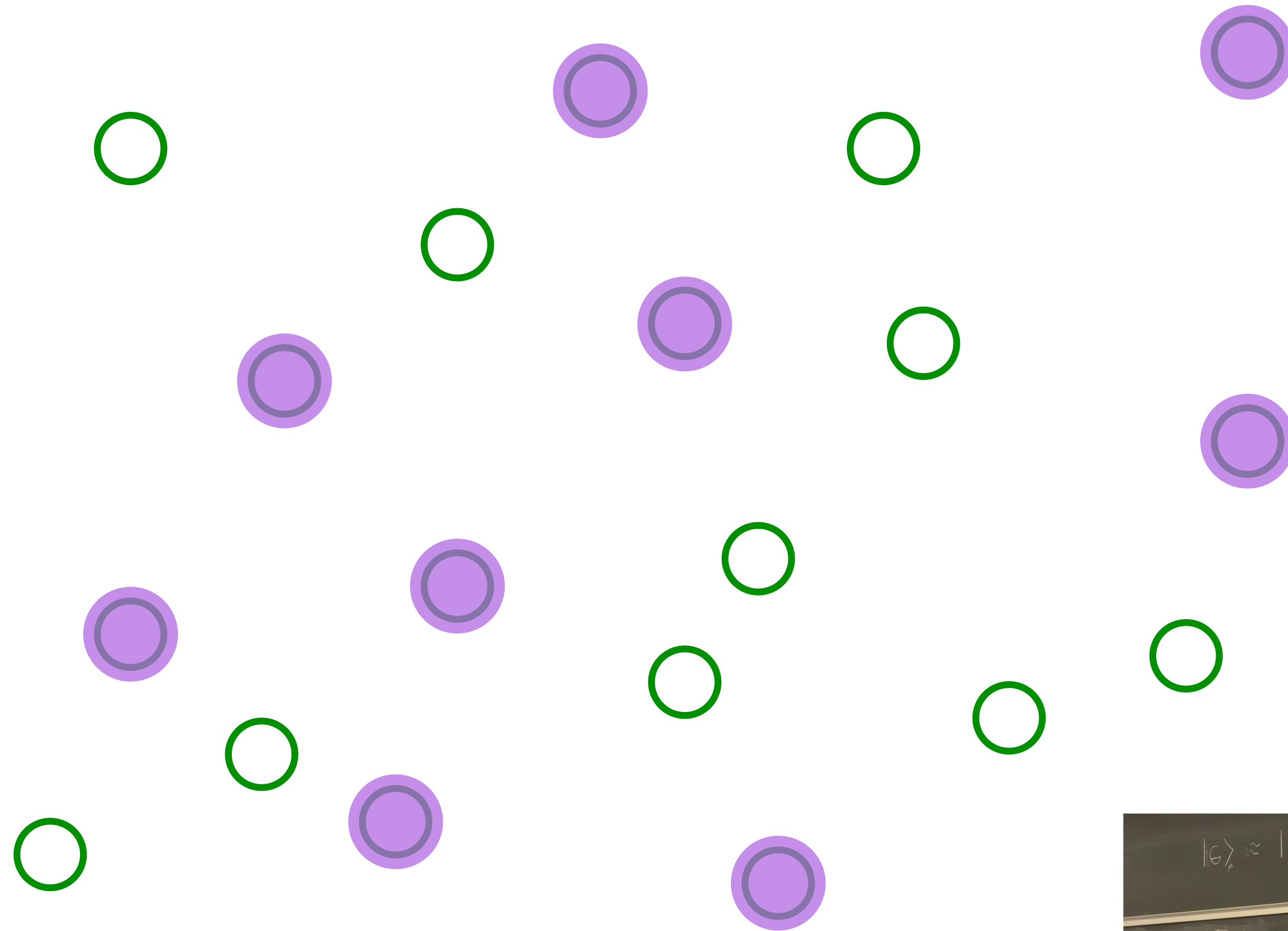


Pick a set of random positions

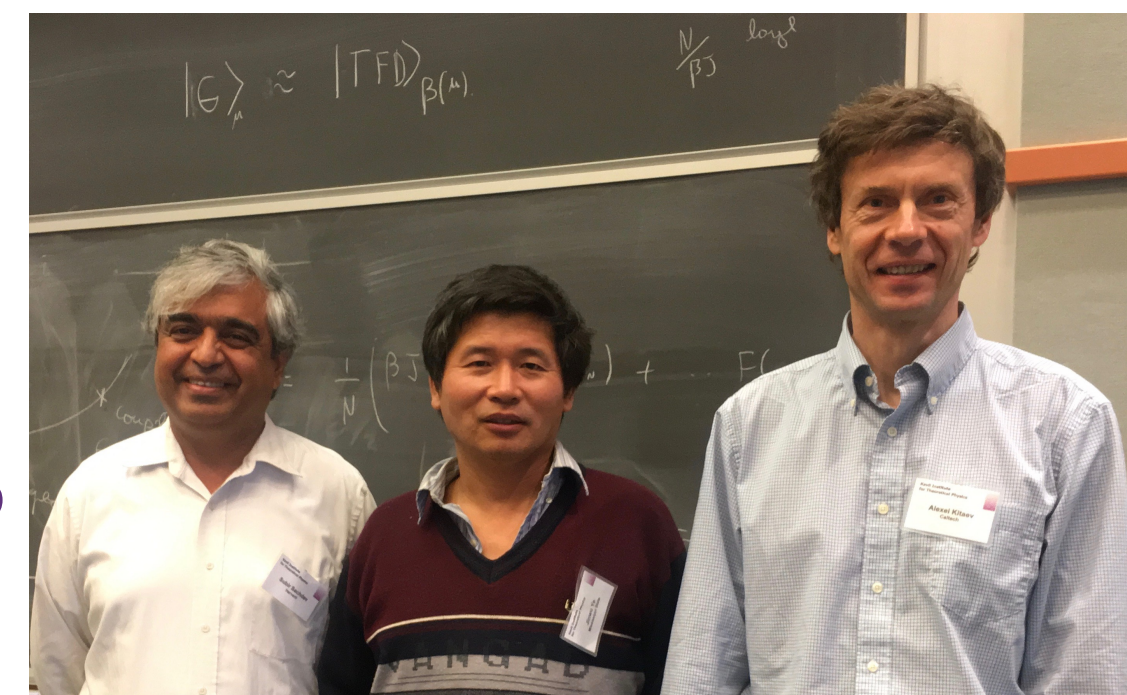


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

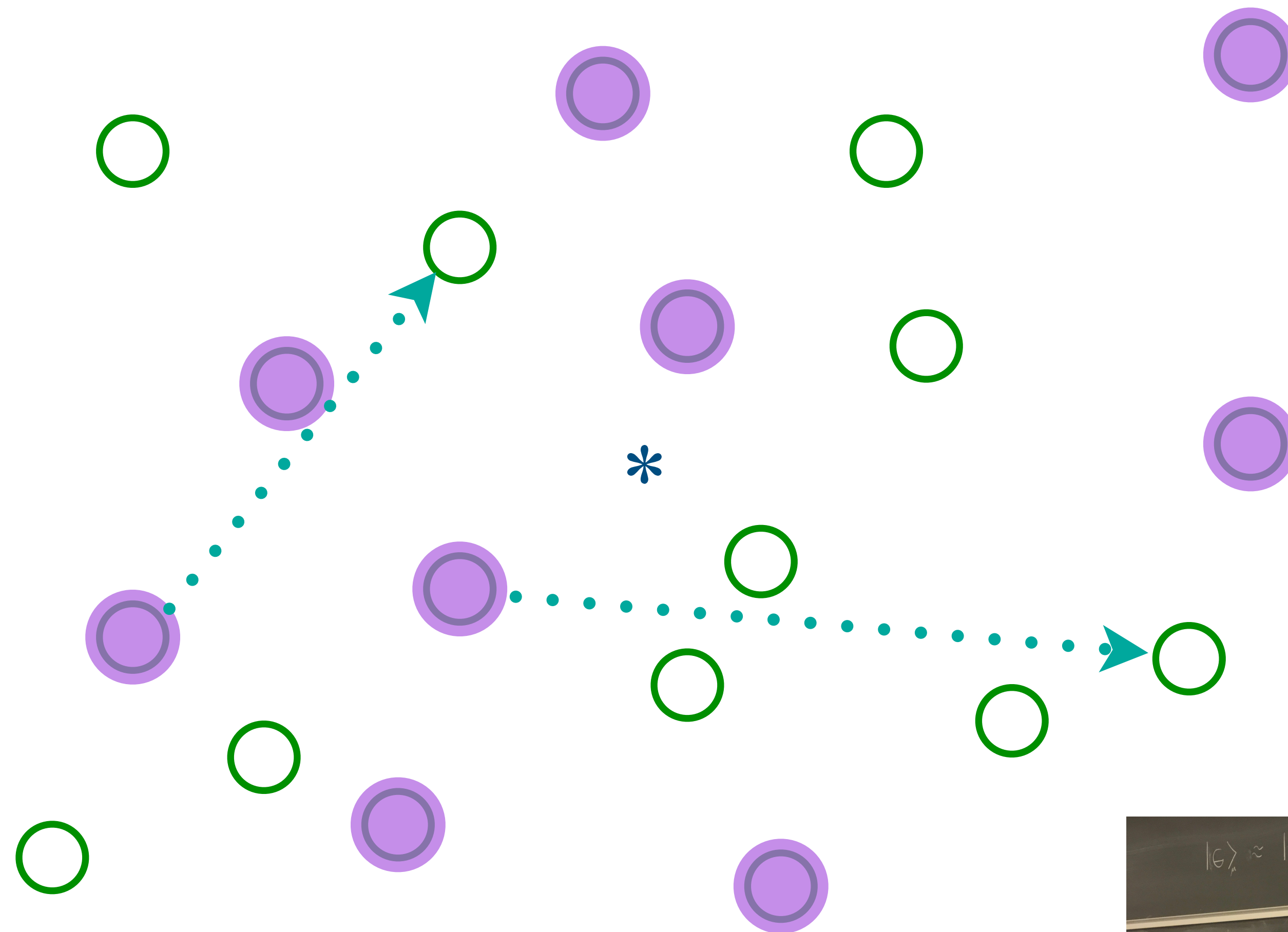


Place electrons randomly on some sites

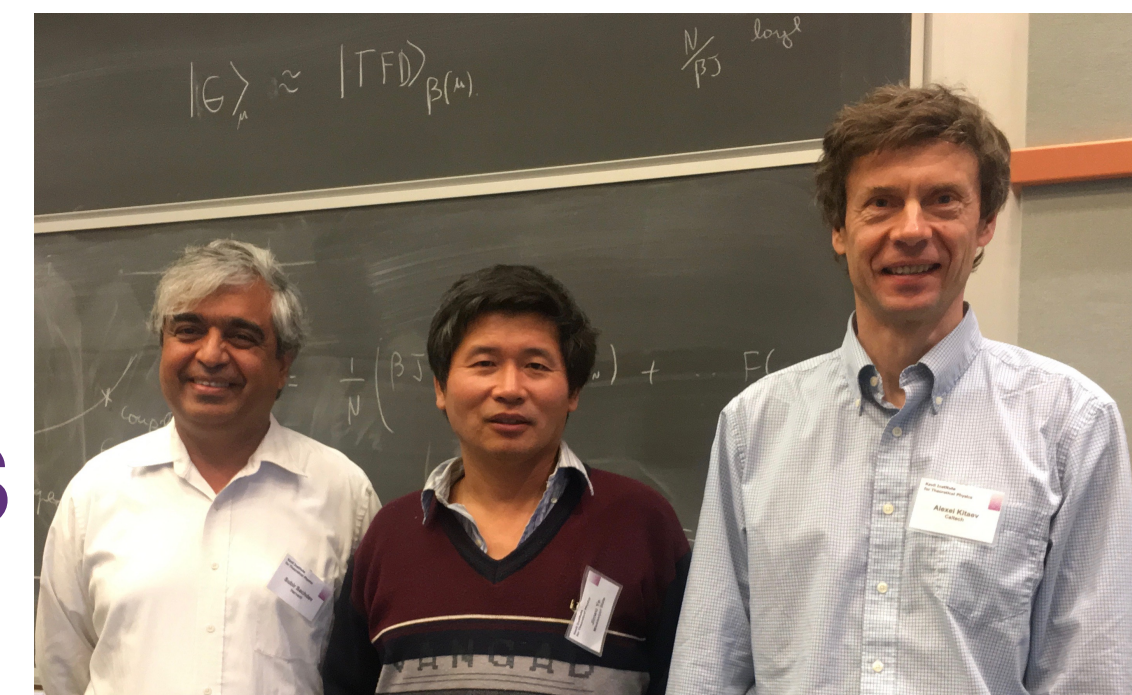


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

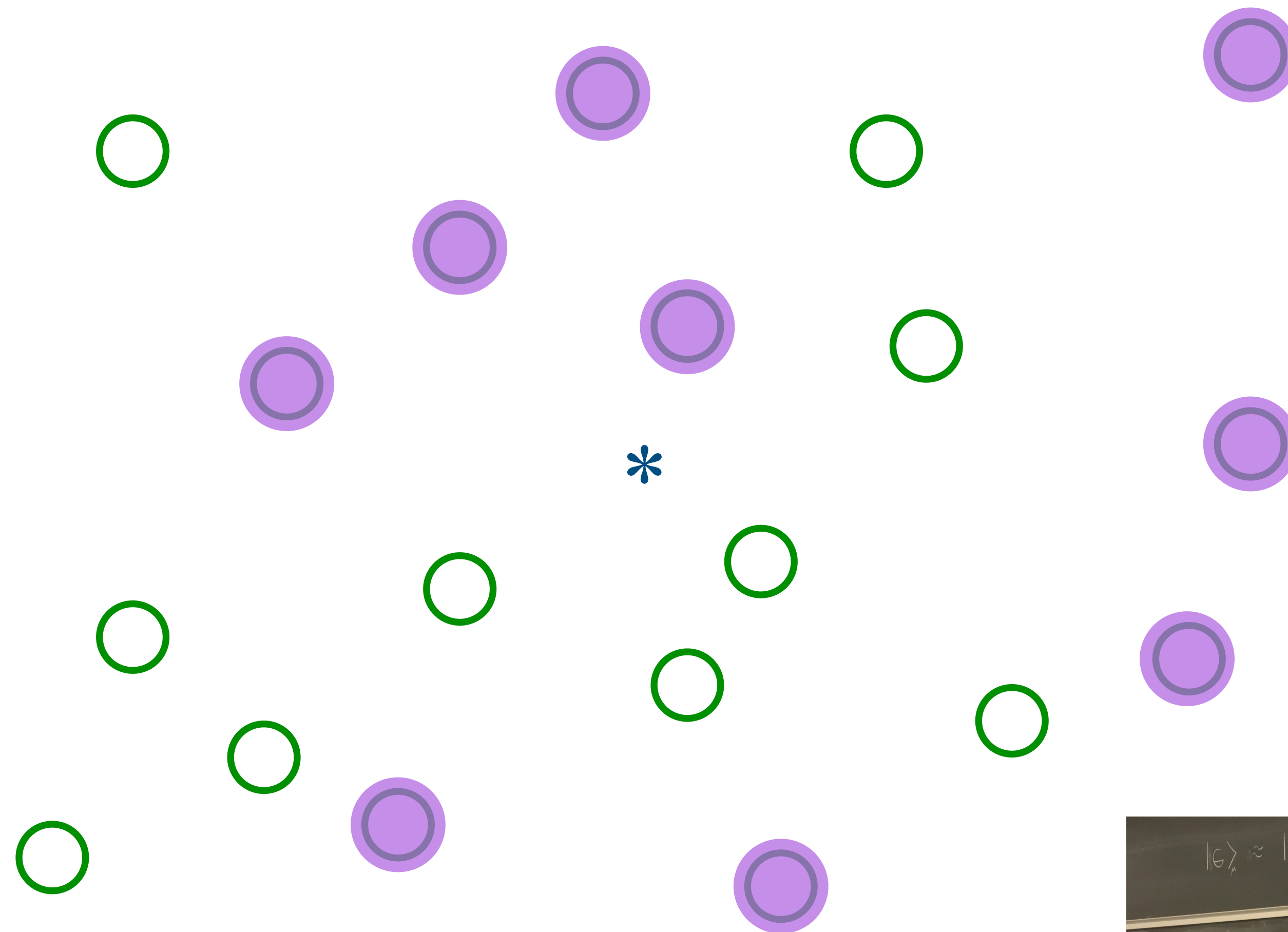


Place electrons randomly on some sites

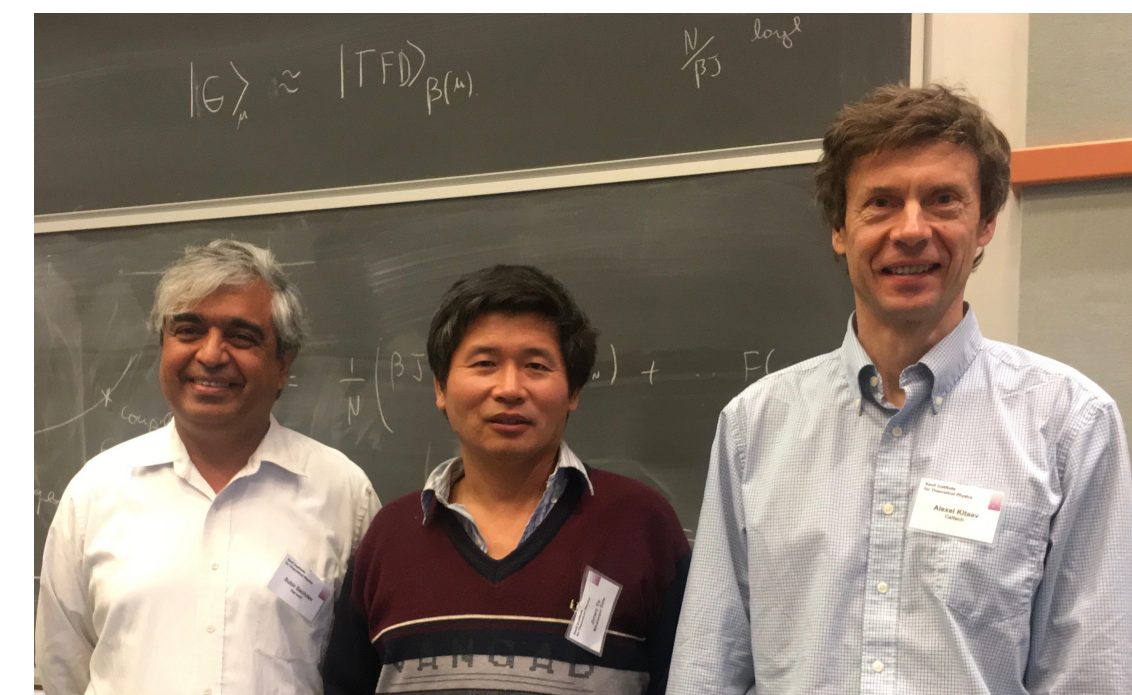


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

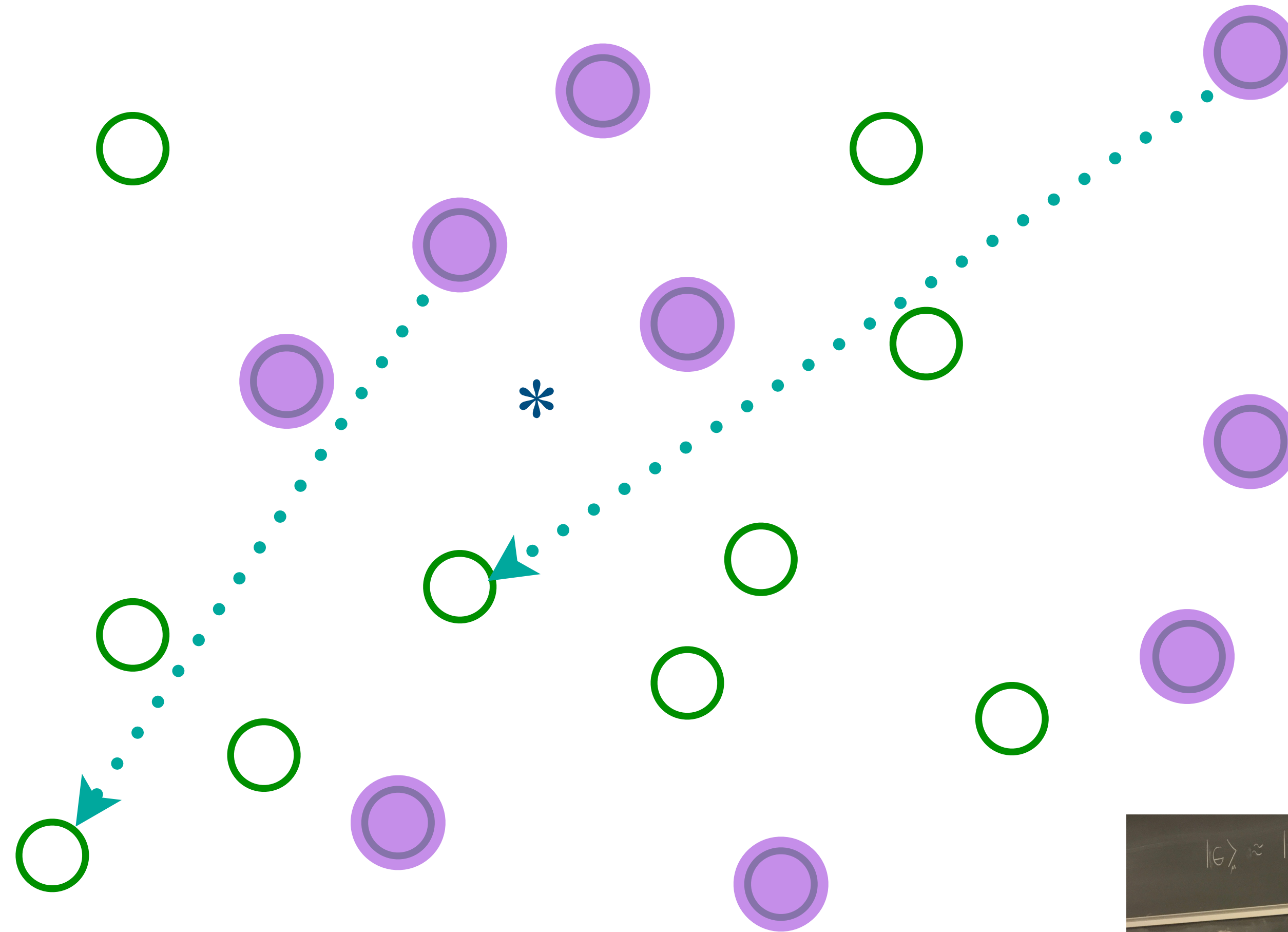


Entangle electrons pairwise randomly

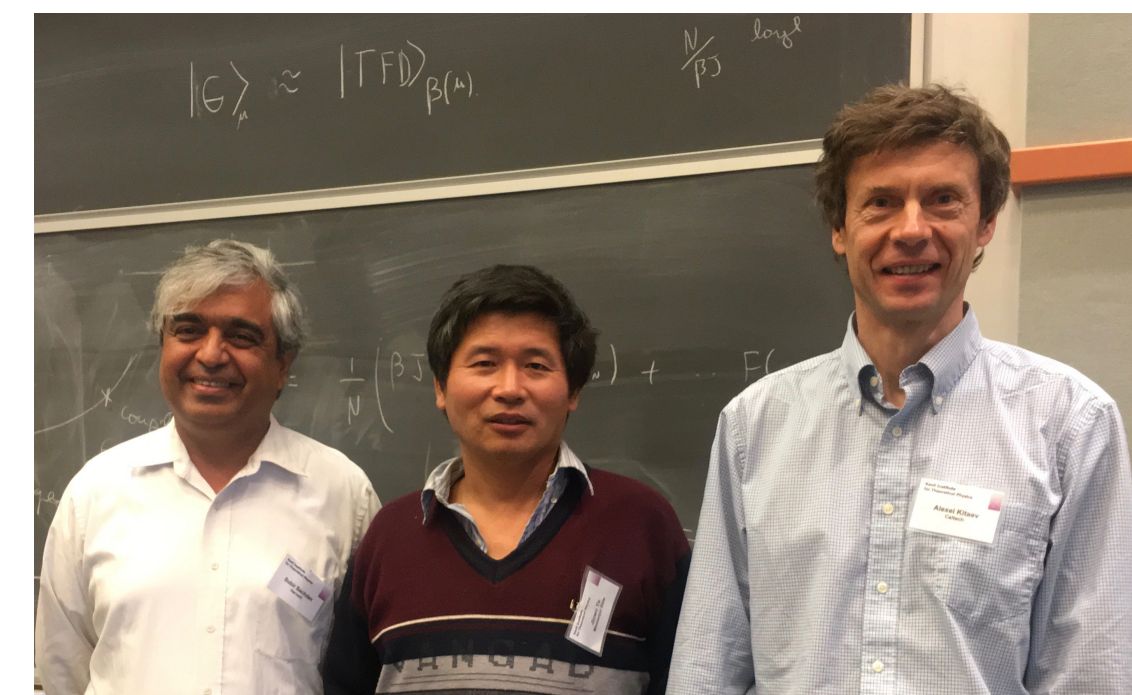


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

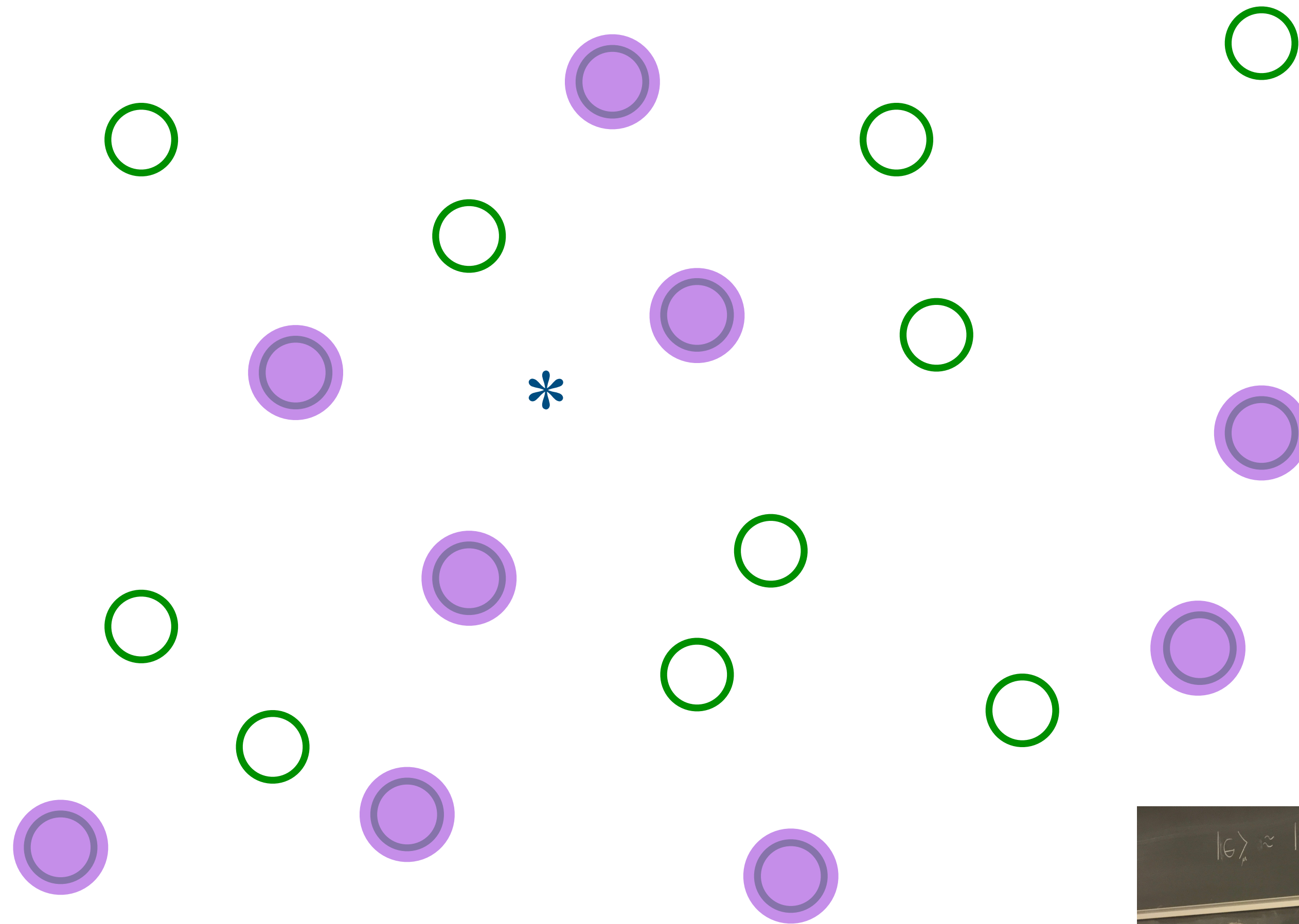


Entangle electrons pairwise randomly

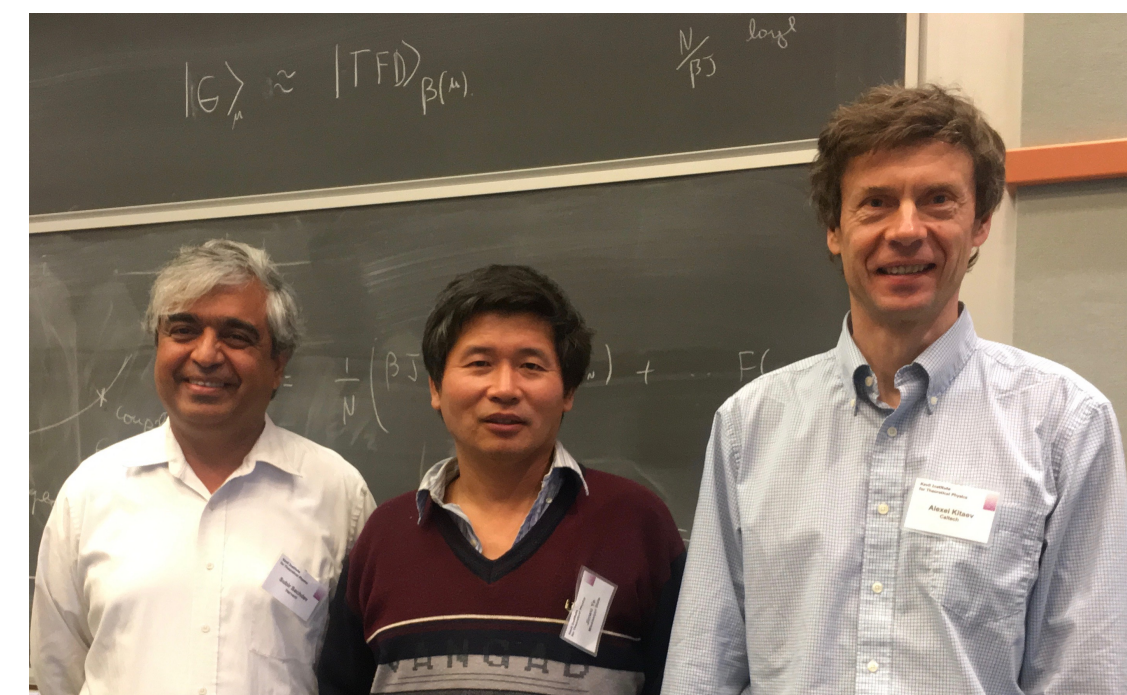


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

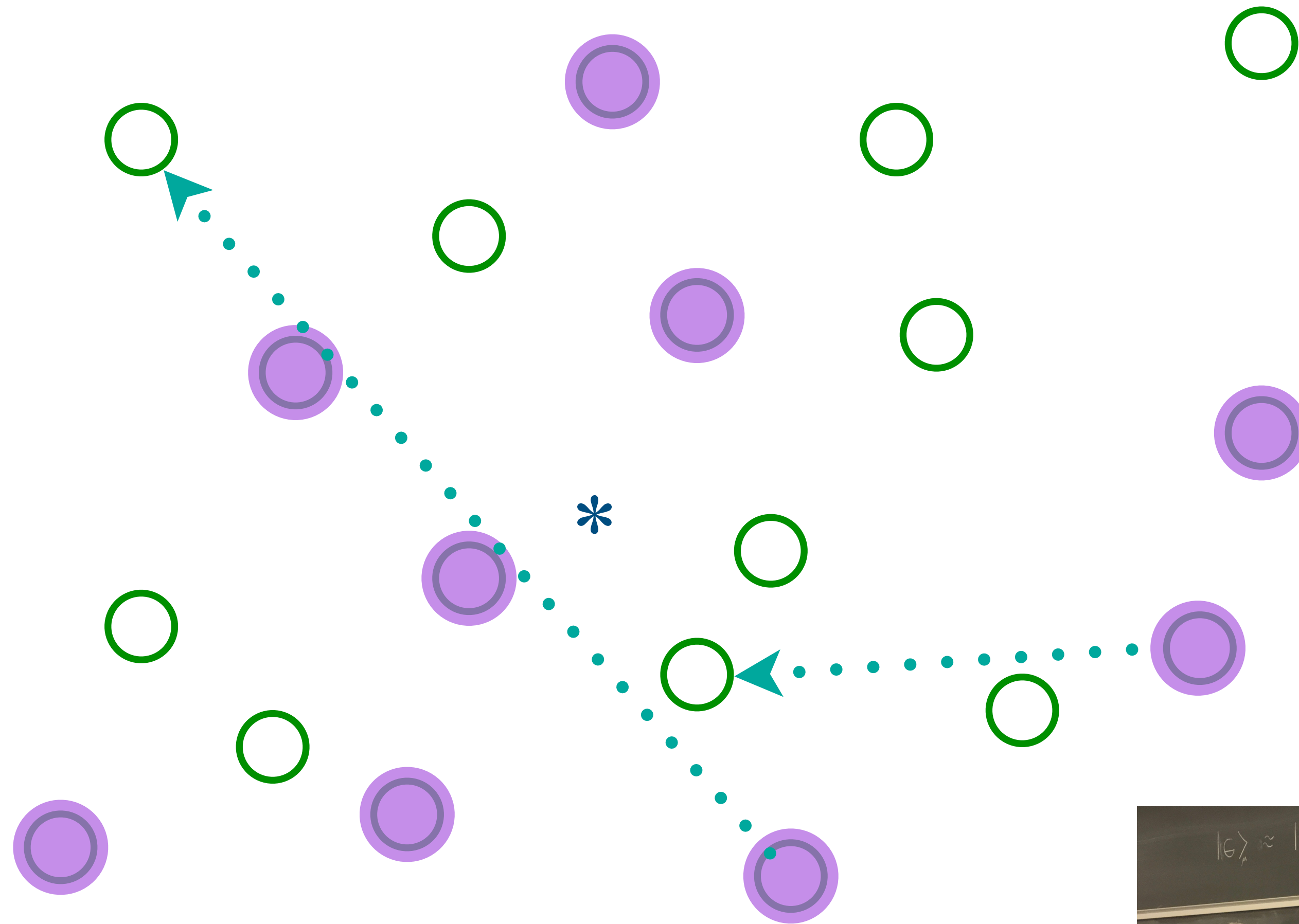


Entangle electrons pairwise randomly

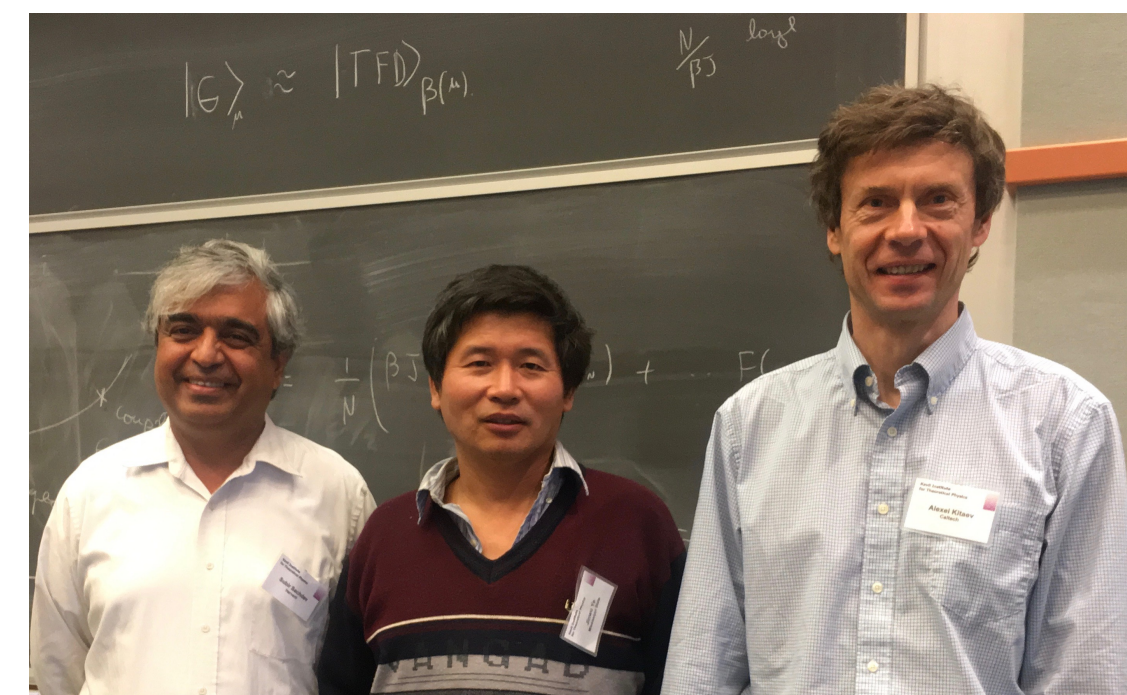


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

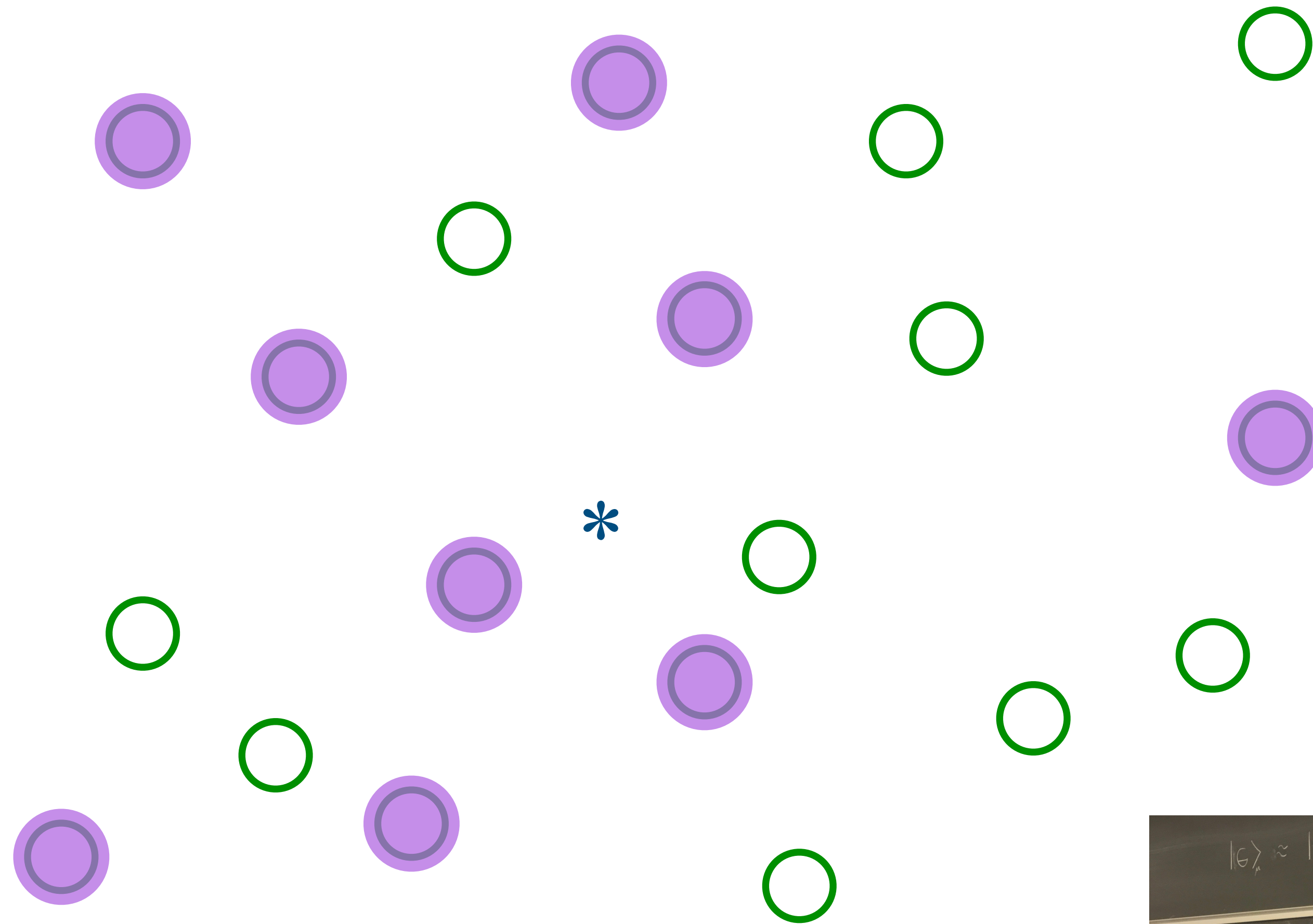


Entangle electrons pairwise randomly

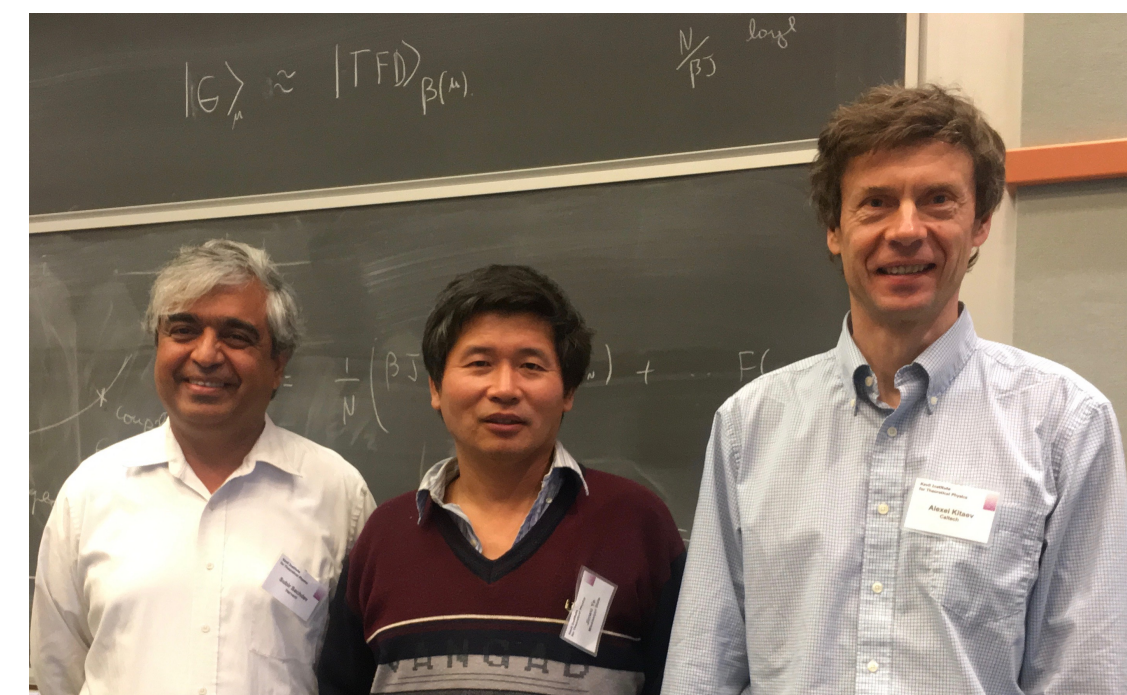


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

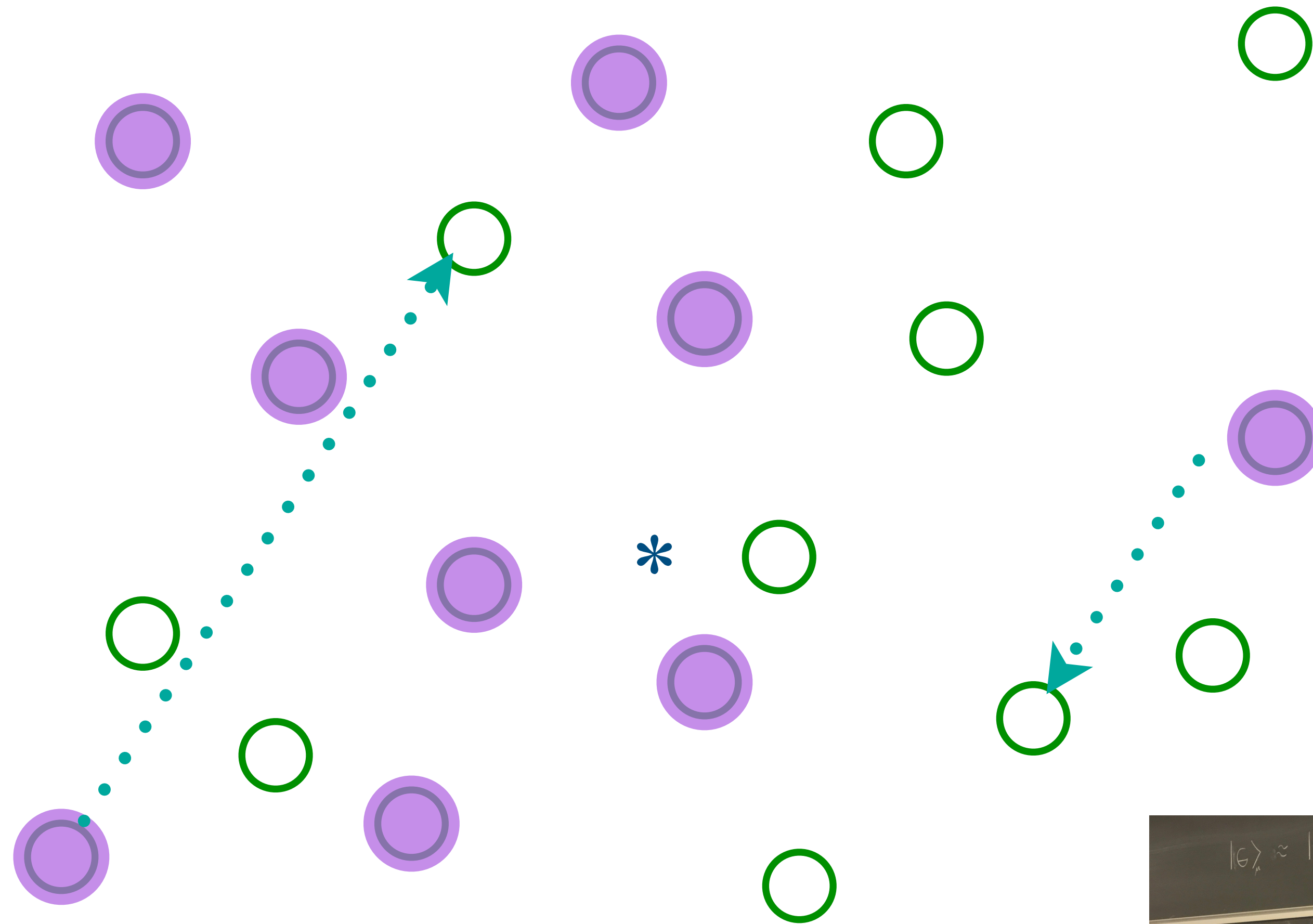


Entangle electrons pairwise randomly

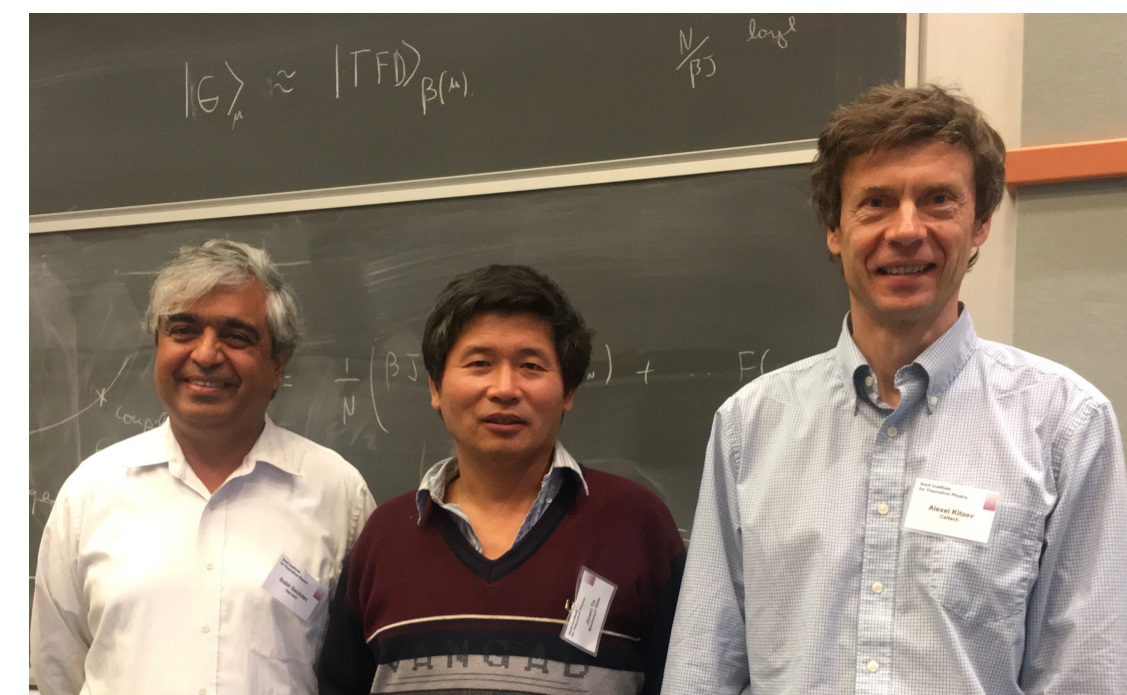


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

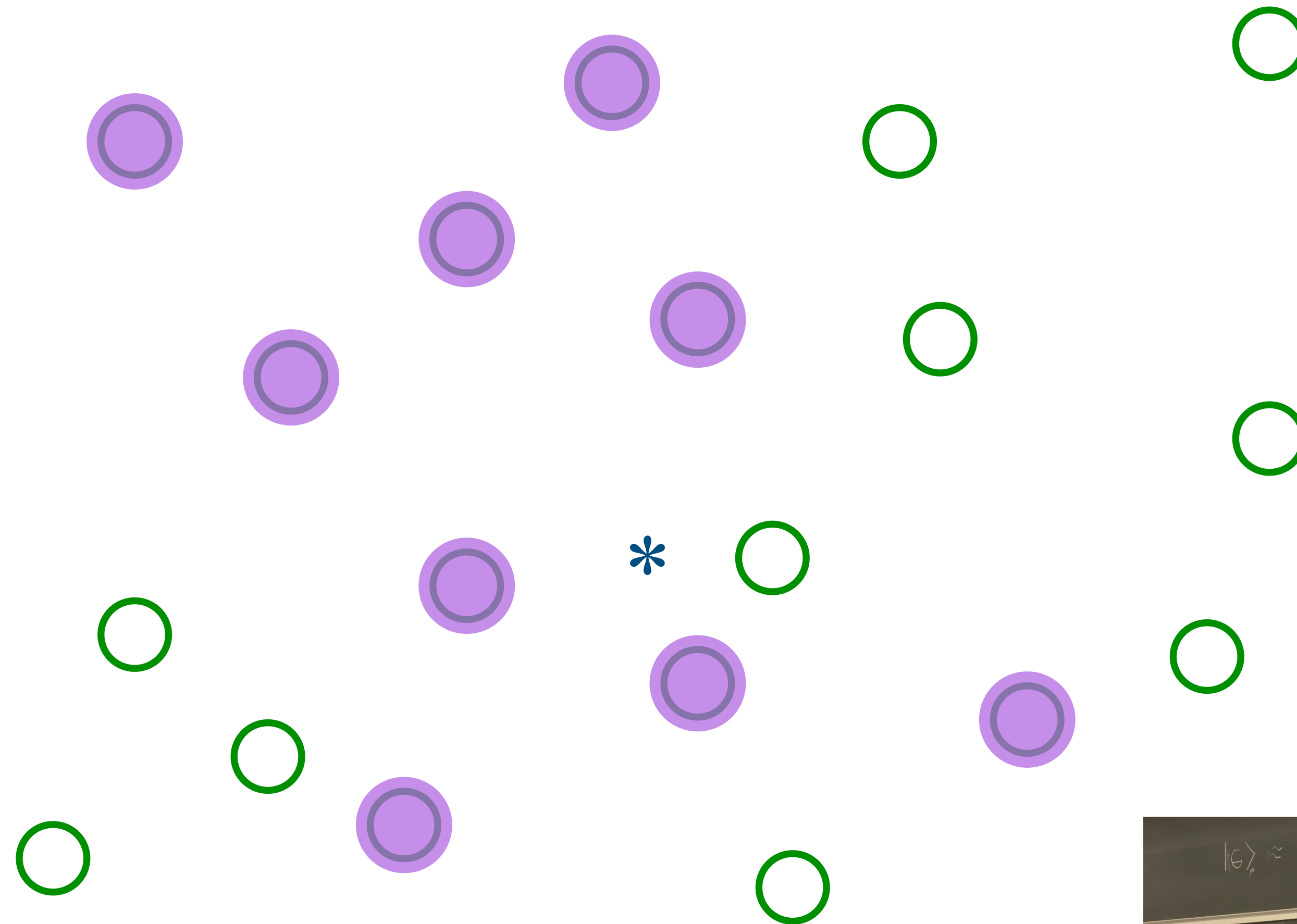


Entangle electrons pairwise randomly

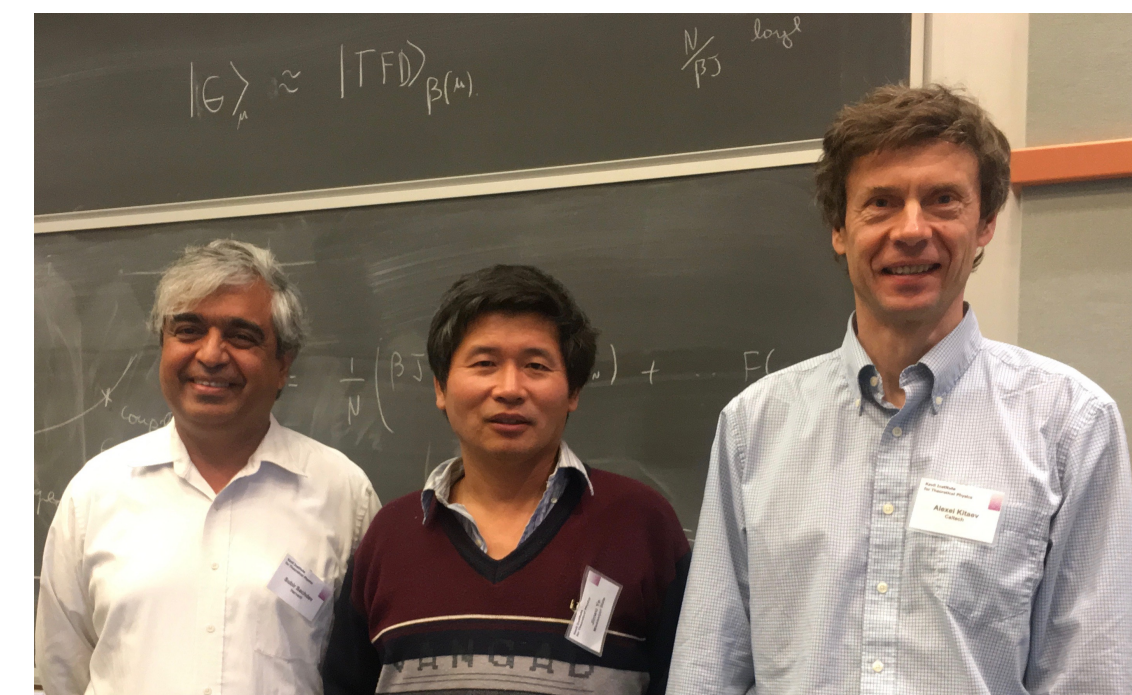


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

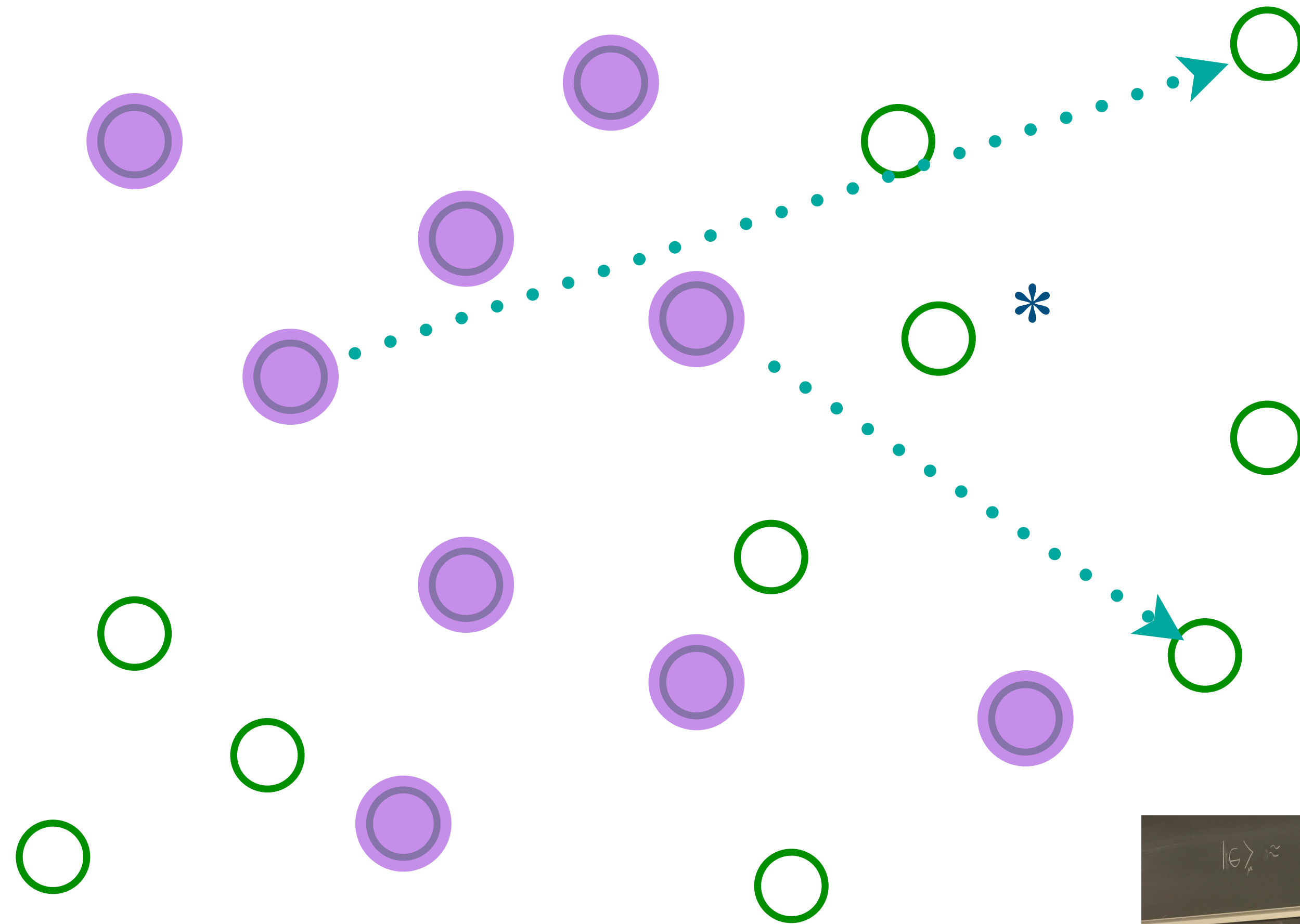


Entangle electrons pairwise randomly

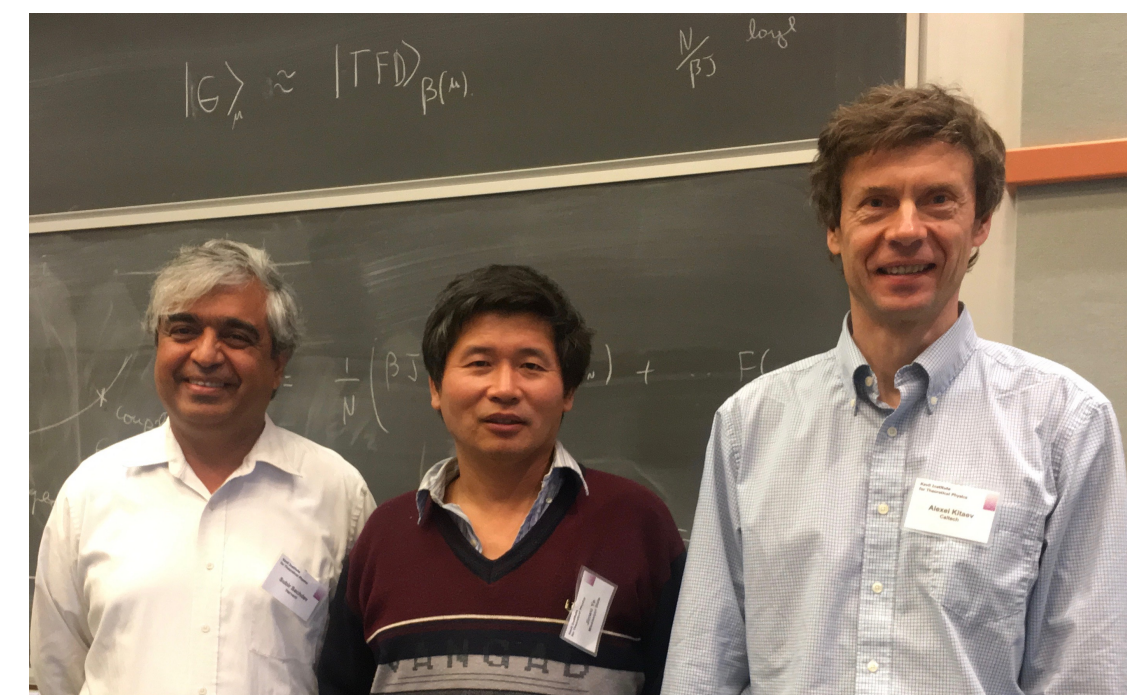


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

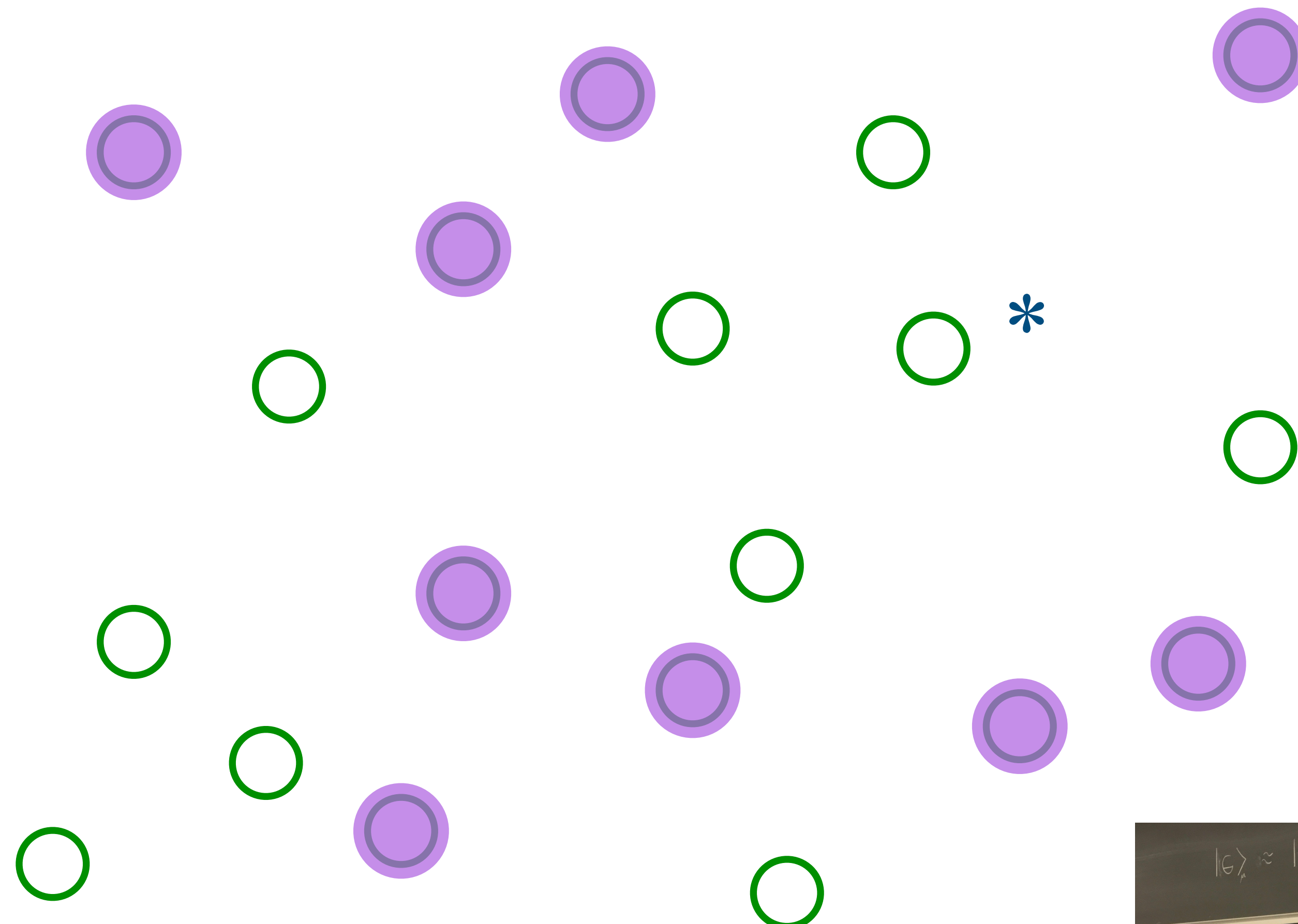


Entangle electrons pairwise randomly

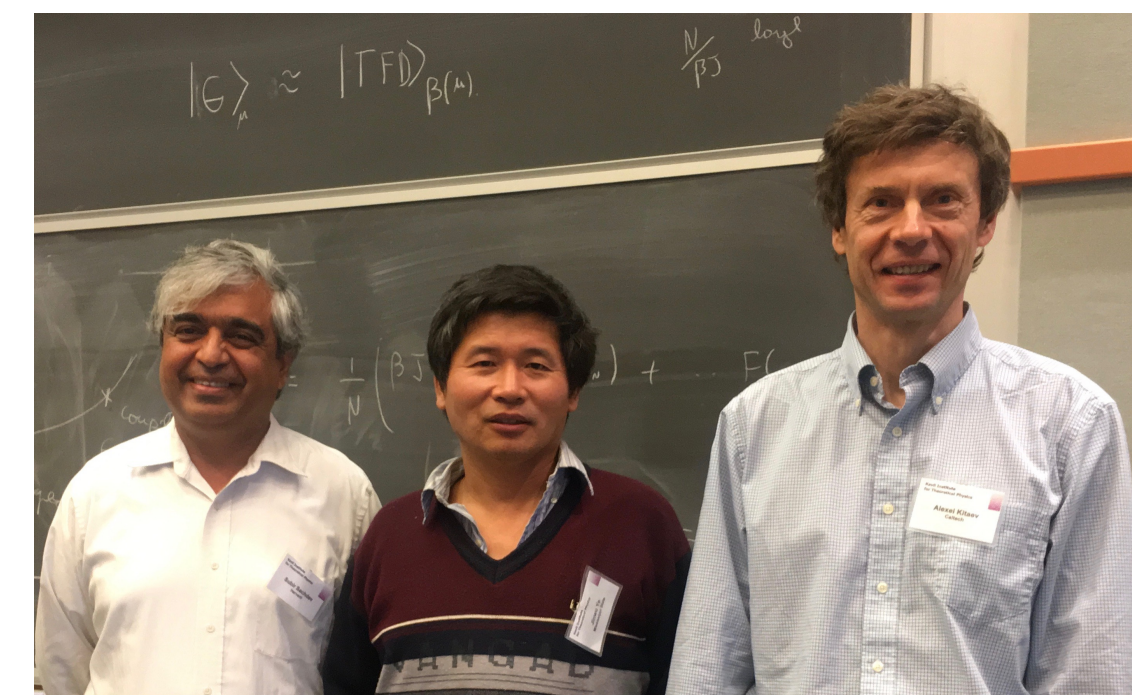


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

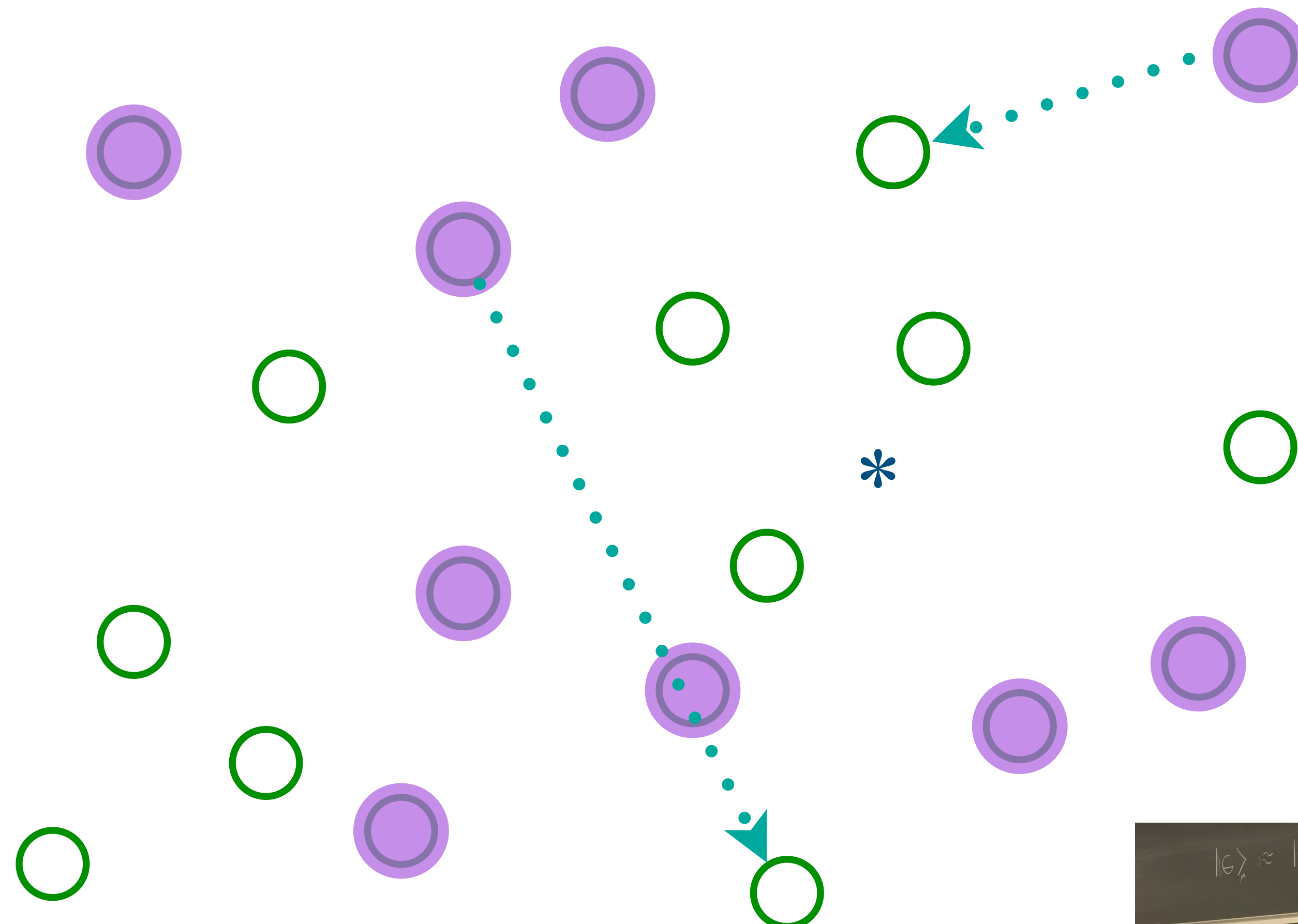


Entangle electrons pairwise randomly

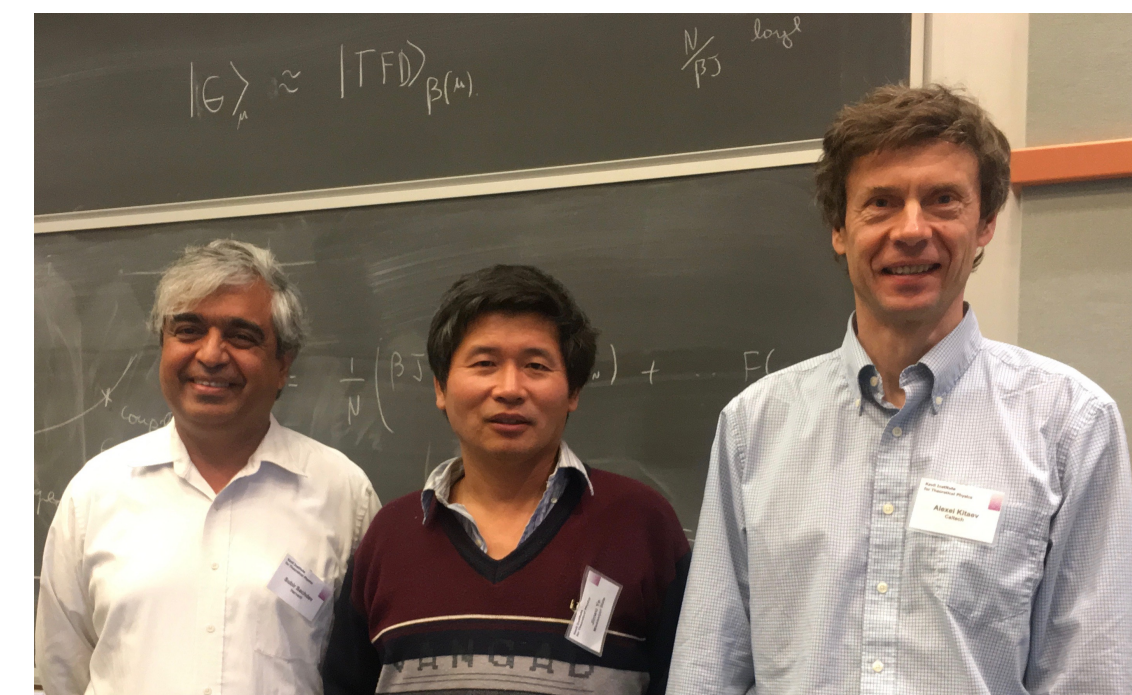


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)

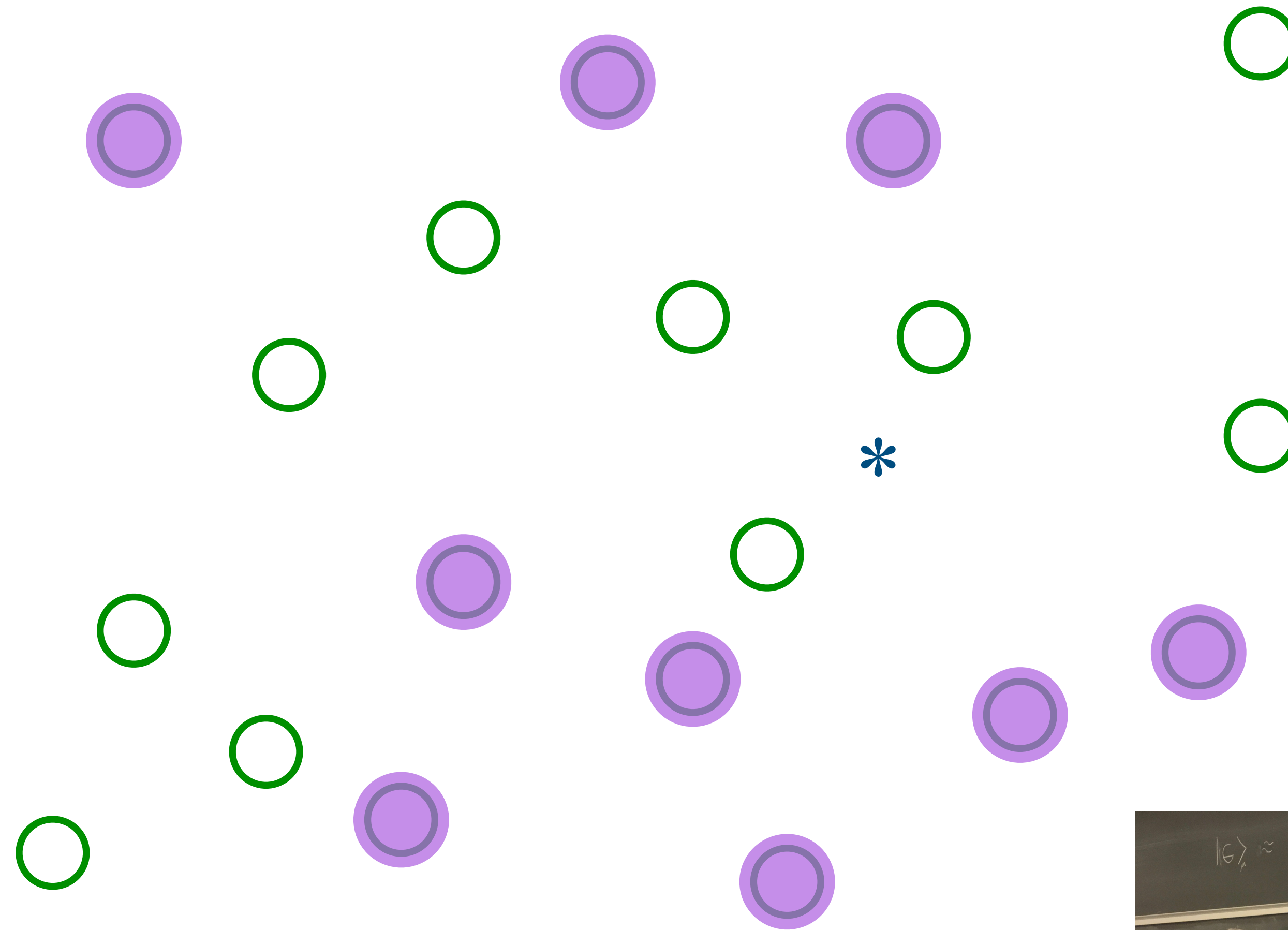


Entangle electrons pairwise randomly

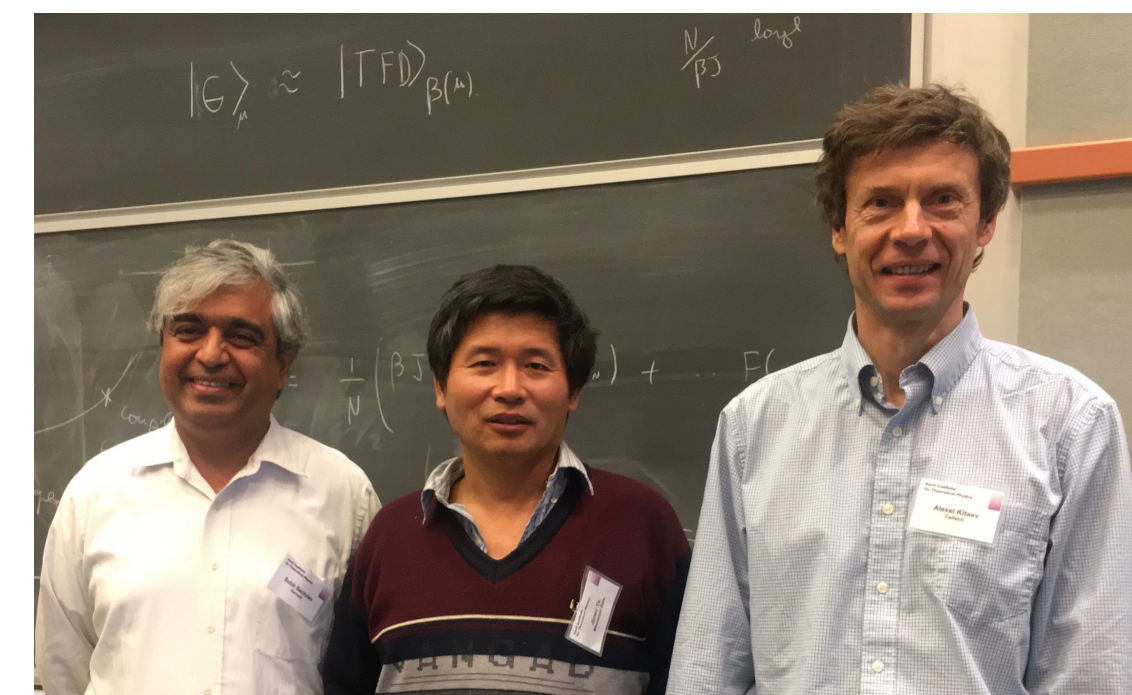


# The SYK model

Sachdev, Ye (1993); Kitaev (2015)



Entangle electrons pairwise randomly



# The Sachdev-Ye-Kitaev (SYK) model

(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large  $N$  limit;  
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} - \mu \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

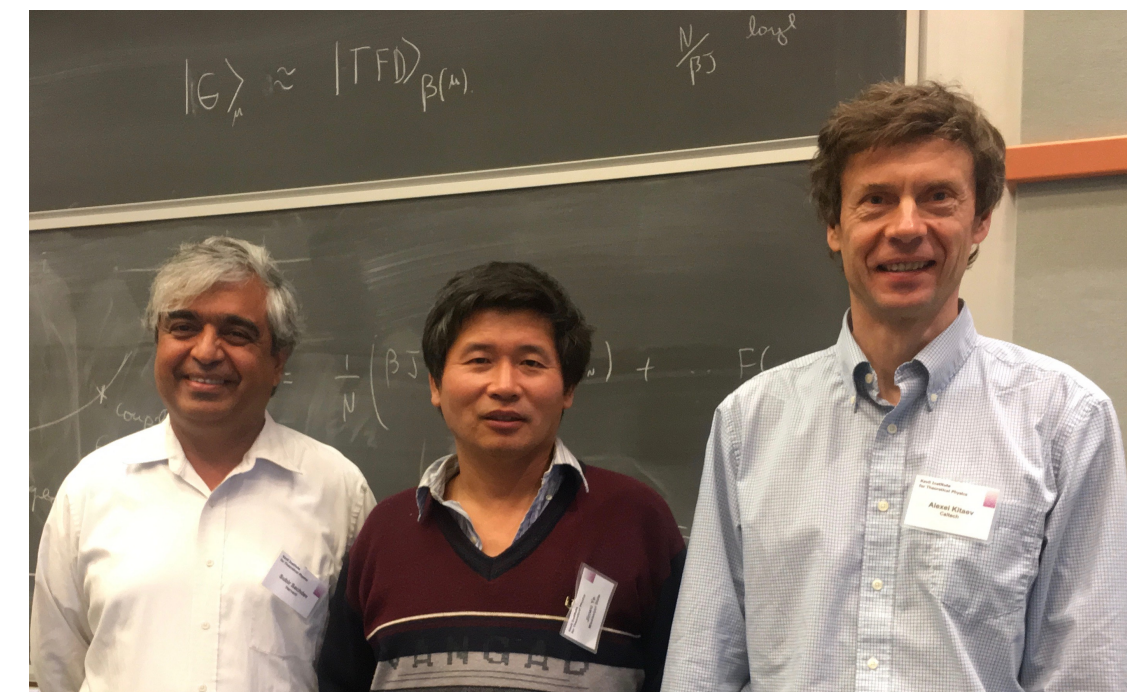
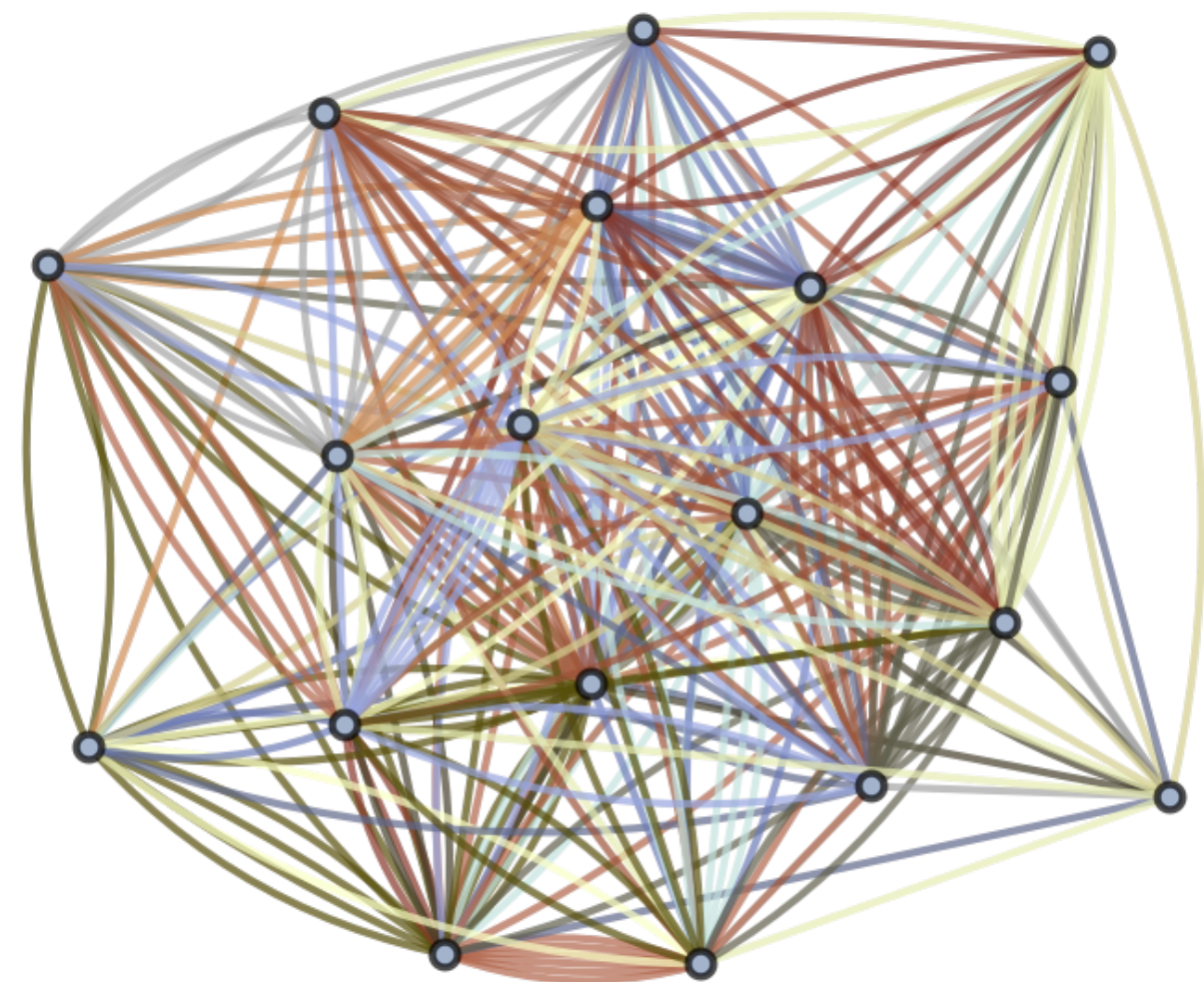
$$c_{\alpha} c_{\beta} + c_{\beta} c_{\alpha} = 0 \quad , \quad c_{\alpha} c_{\beta}^{\dagger} + c_{\beta}^{\dagger} c_{\alpha} = \delta_{\alpha\beta}$$

$$Q = \frac{1}{N} \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$U_{\alpha\beta;\gamma\delta}$  are independent random variables with  $\overline{U_{\alpha\beta;\gamma\delta}} = 0$  and  $\overline{|U_{\alpha\beta;\gamma\delta}|^2} = U^2$   
 $N \rightarrow \infty$  yields critical strange metal.

S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)



# A simple model of a metal with quasiparticles

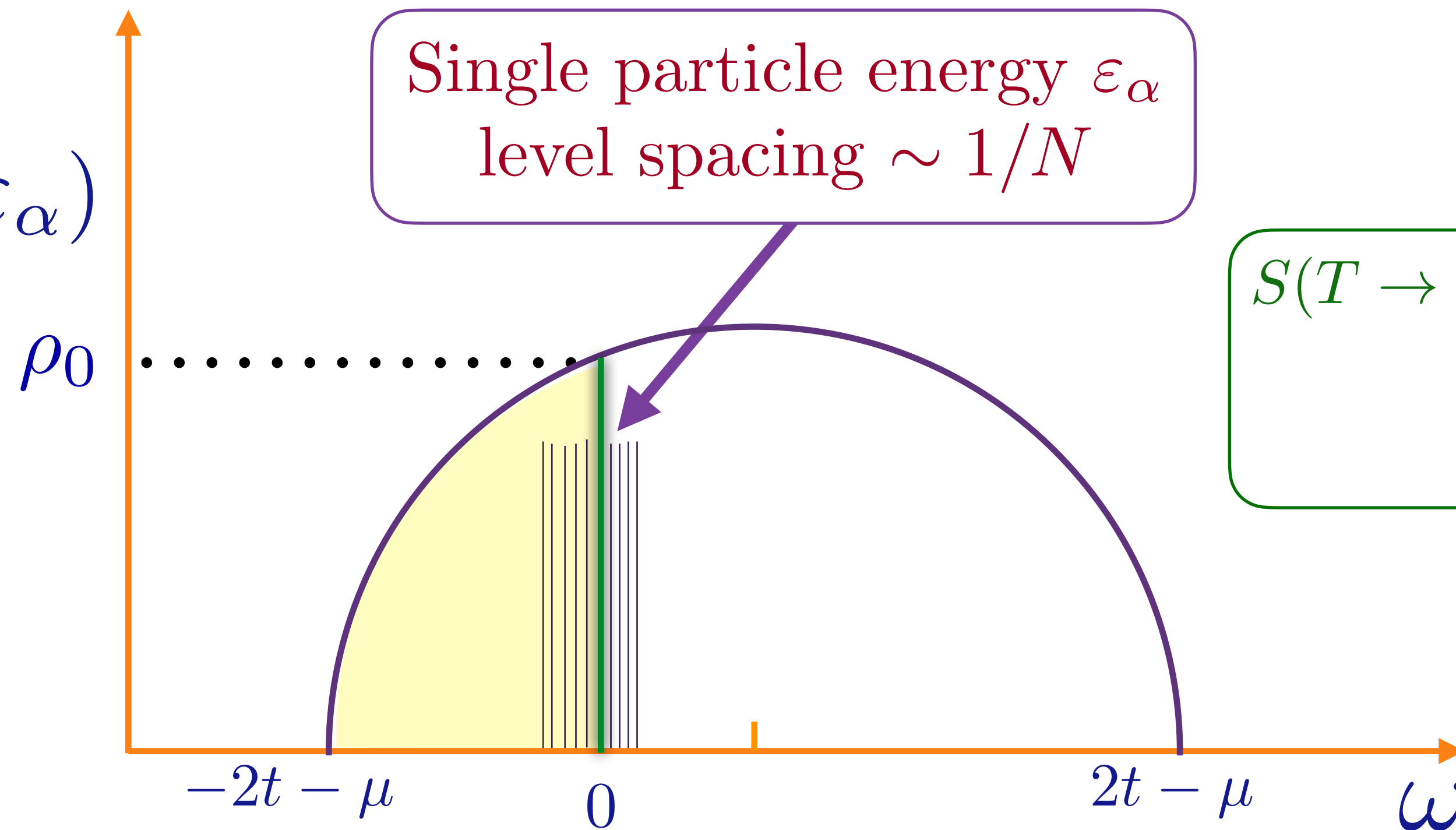
$$H = \frac{1}{(N)^{1/2}} \sum_{i,j=1}^N t_{ij} c_i^\dagger c_j - \mu \sum_i c_i^\dagger c_i$$

$$c_i c_j + c_j c_i = 0 \quad , \quad c_i c_j^\dagger + c_j^\dagger c_i = \delta_{ij}$$

$$\frac{1}{N} \sum_i c_i^\dagger c_i = Q$$

$t_{ij}$  are independent random variables with  $\overline{t_{ij}} = 0$  and  $\overline{|t_{ij}|^2} = t^2$

$$\rho(\omega) = \frac{1}{N} \sum_{\alpha} \delta(\omega - \varepsilon_{\alpha})$$

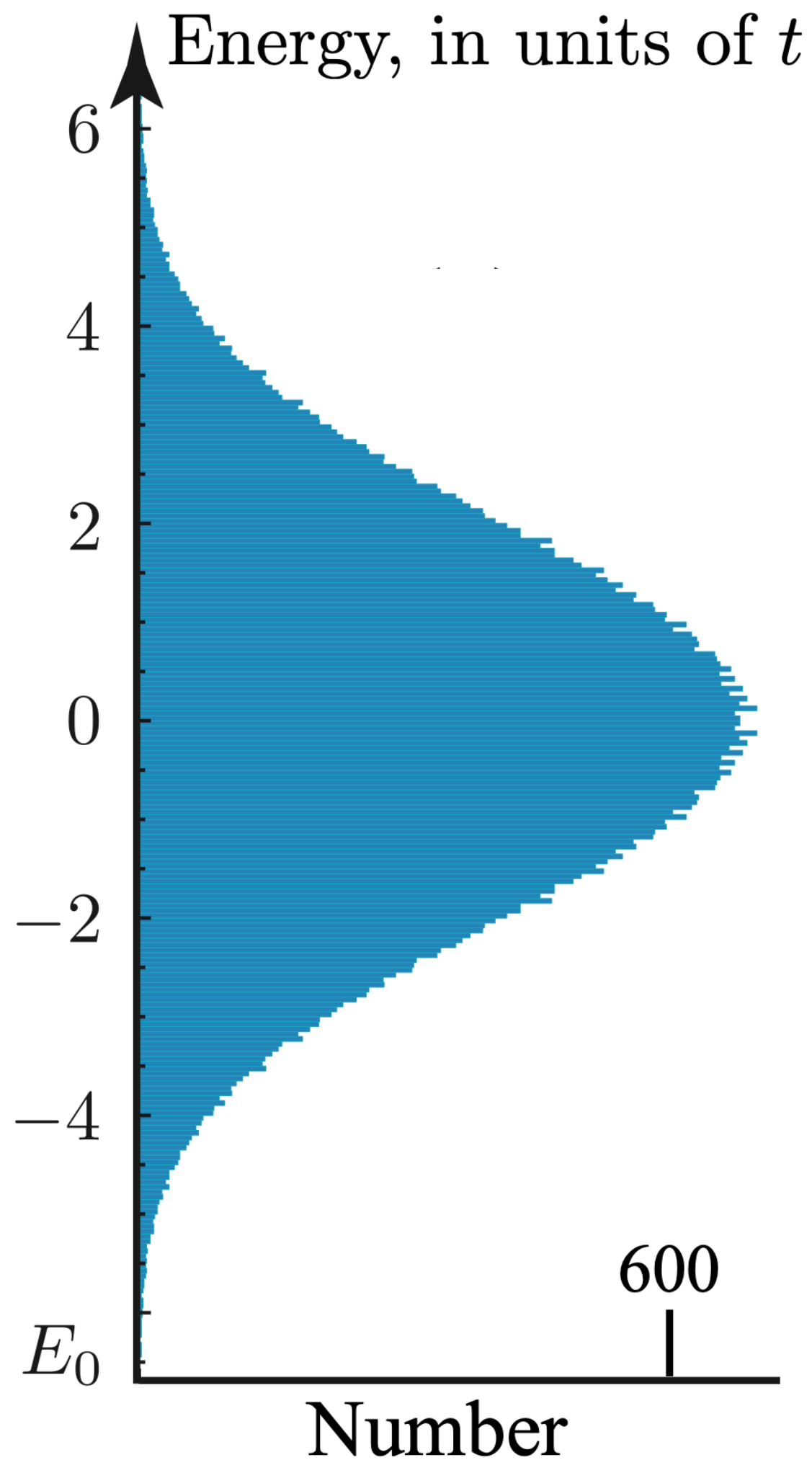


$$S(T \rightarrow 0) = N\gamma T$$

$$\gamma = \frac{\pi^2}{3} \rho_0$$

# Random matrix model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



For random matrix model:

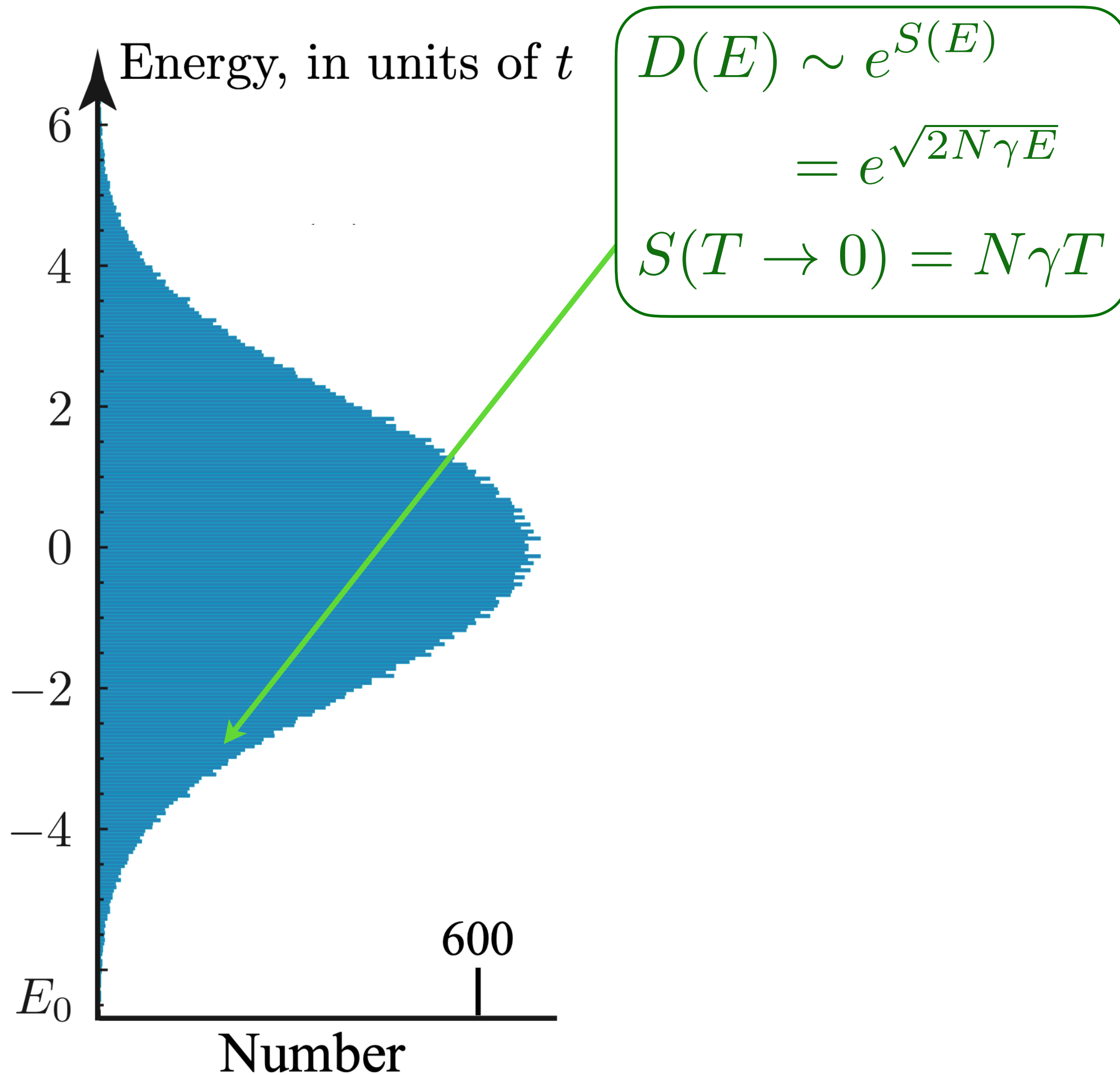
$$E_0 + E_i = \sum_{\alpha} n_{\alpha} \varepsilon_{\alpha}$$

$n_{\alpha} = 0, 1,$   
occupation number

Many-body density of states

# Random matrix model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



For random matrix model:

$$E_0 + E_i = \sum_{\alpha} n_{\alpha} \varepsilon_{\alpha}$$

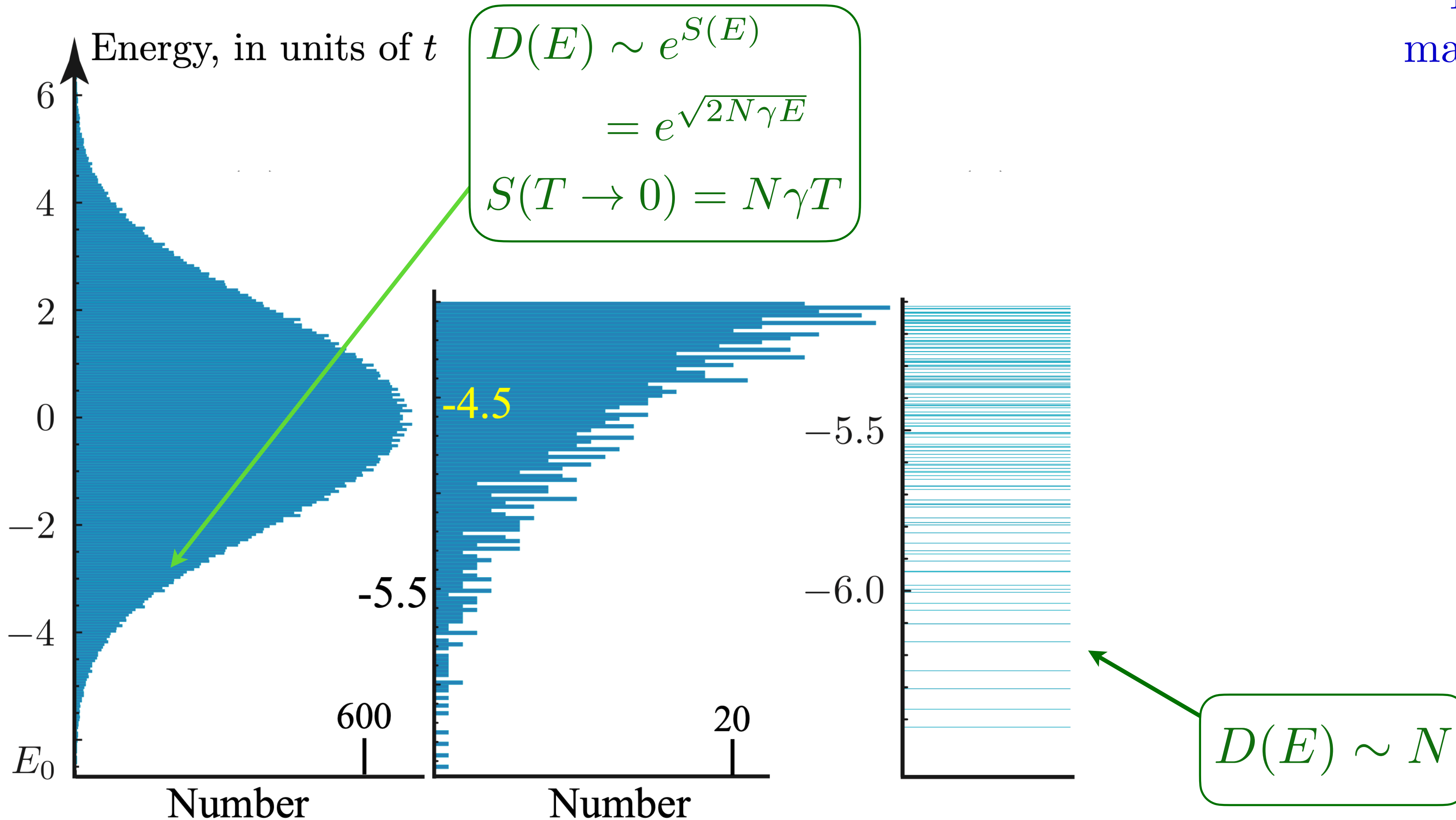
$n_{\alpha} = 0, 1,$   
occupation number

Many-body density of states

# Random matrix model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$

For random matrix model:  
 $E_0 + E_i = \sum_{\alpha} n_{\alpha} \epsilon_{\alpha}$   
 $n_{\alpha} = 0, 1,$   
occupation number



Many-body density of states

# The Sachdev-Ye-Kitaev (SYK) model

(See also: the “2-Body Random Ensemble” in nuclear physics; did not obtain the large  $N$  limit;  
T.A. Brody, J. Flores, J.B. French, P.A. Mello, A. Pandey, and S.S.M. Wong, Rev. Mod. Phys. **53**, 385 (1981))

$$H = \frac{1}{(2N)^{3/2}} \sum_{\alpha, \beta, \gamma, \delta=1}^N U_{\alpha\beta;\gamma\delta} c_{\alpha}^{\dagger} c_{\beta}^{\dagger} c_{\gamma} c_{\delta} - \mu \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

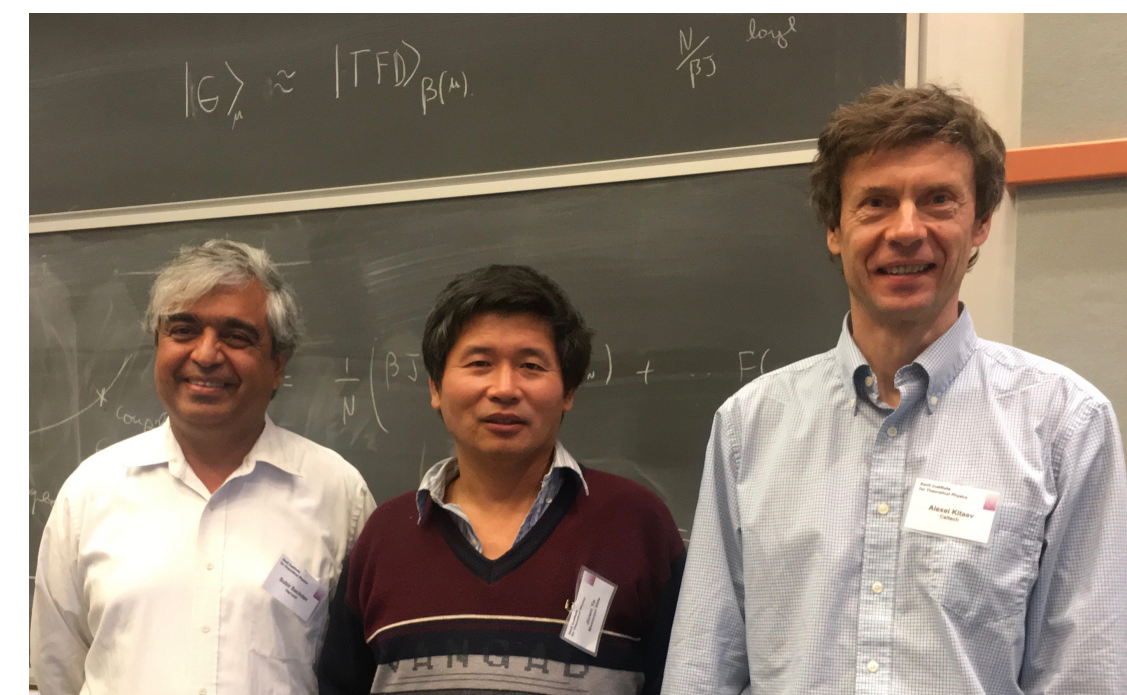
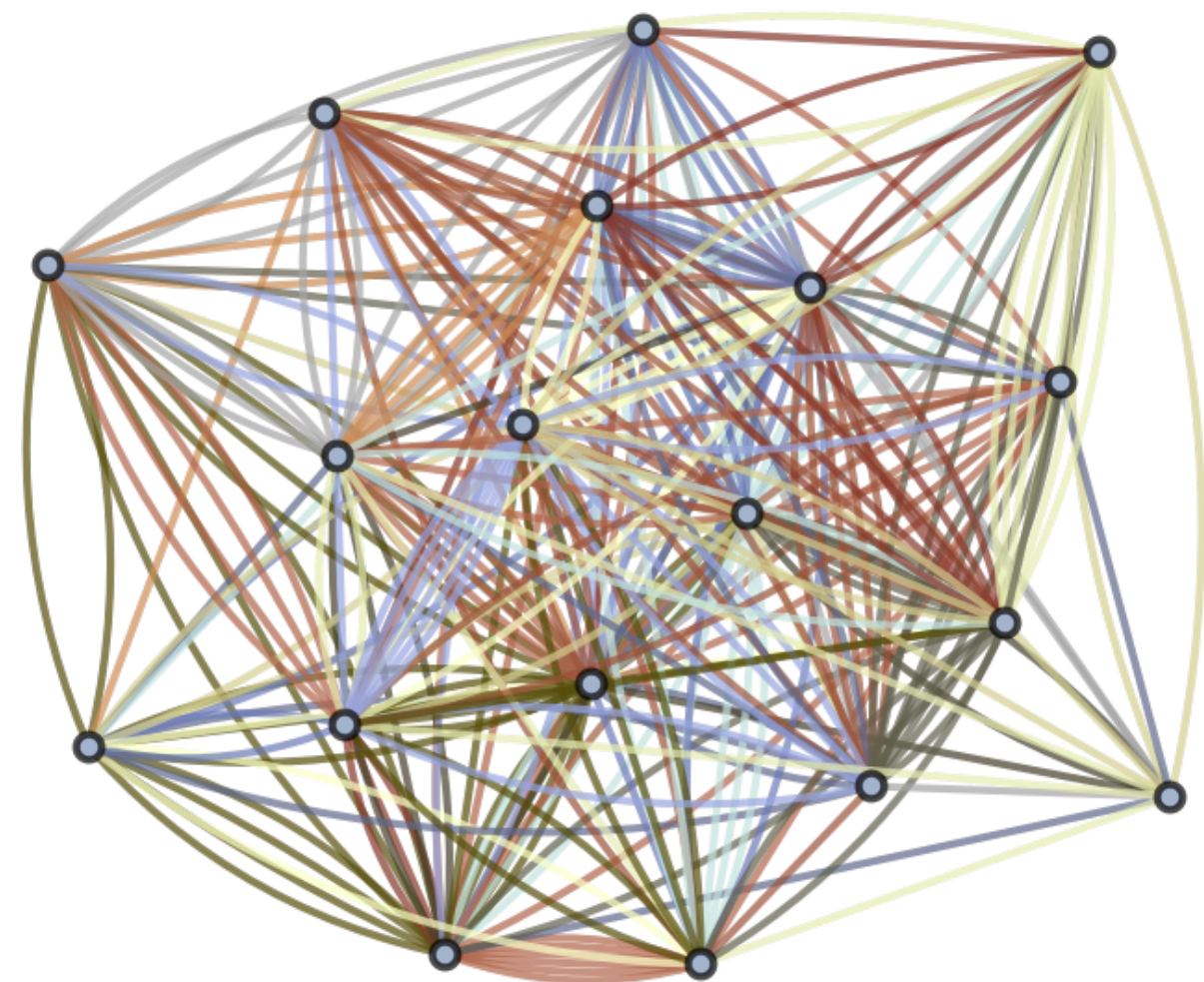
$$c_{\alpha} c_{\beta} + c_{\beta} c_{\alpha} = 0 \quad , \quad c_{\alpha} c_{\beta}^{\dagger} + c_{\beta}^{\dagger} c_{\alpha} = \delta_{\alpha\beta}$$

$$Q = \frac{1}{N} \sum_{\alpha} c_{\alpha}^{\dagger} c_{\alpha}$$

$U_{\alpha\beta;\gamma\delta}$  are independent random variables with  $\overline{U_{\alpha\beta;\gamma\delta}} = 0$  and  $\overline{|U_{\alpha\beta;\gamma\delta}|^2} = U^2$   
 $N \rightarrow \infty$  yields critical strange metal.

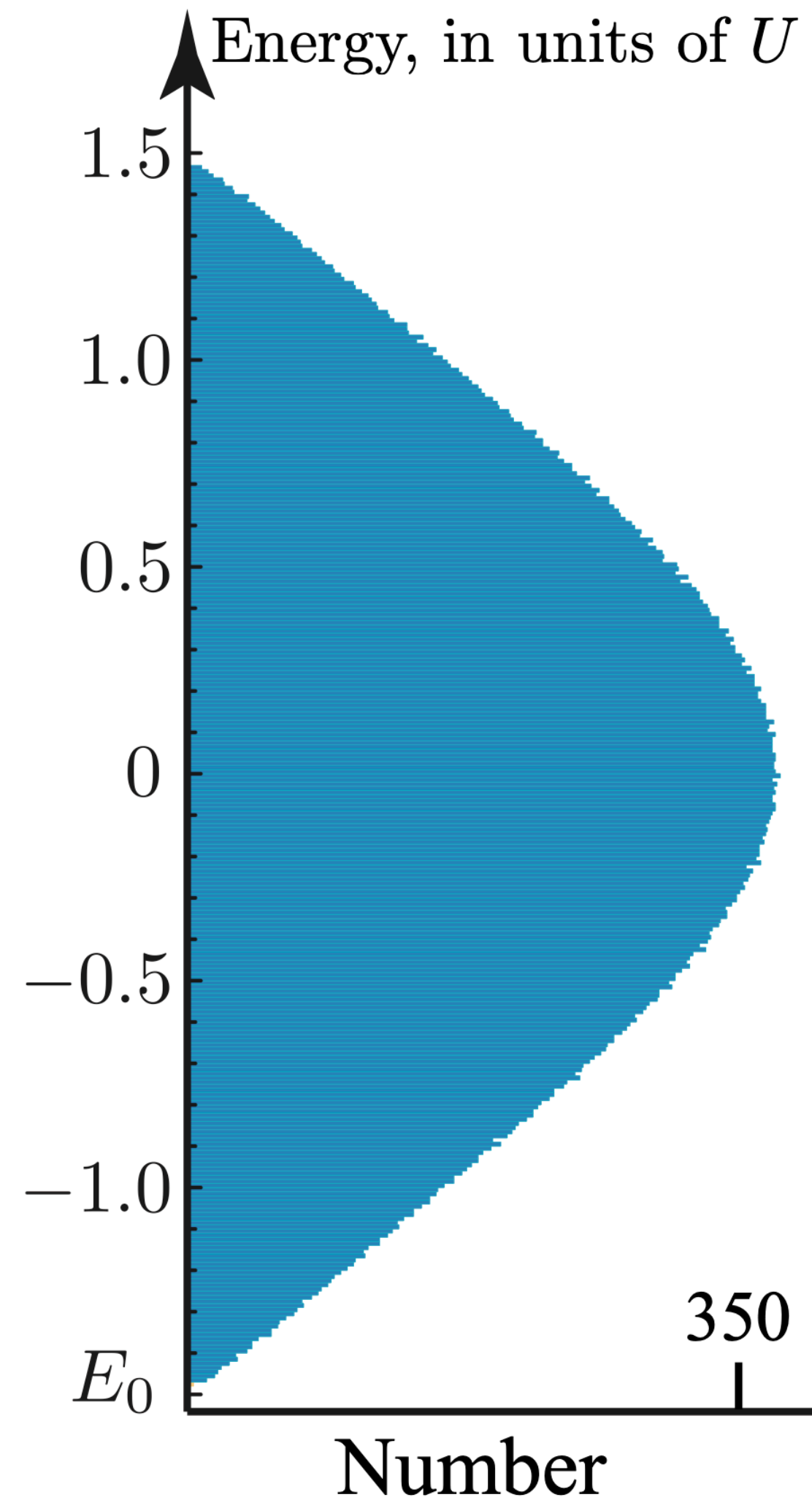
S. Sachdev and J. Ye, PRL **70**, 3339 (1993)

A. Kitaev, unpublished; S. Sachdev, PRX **5**, 041025 (2015)



# Complex SYK model

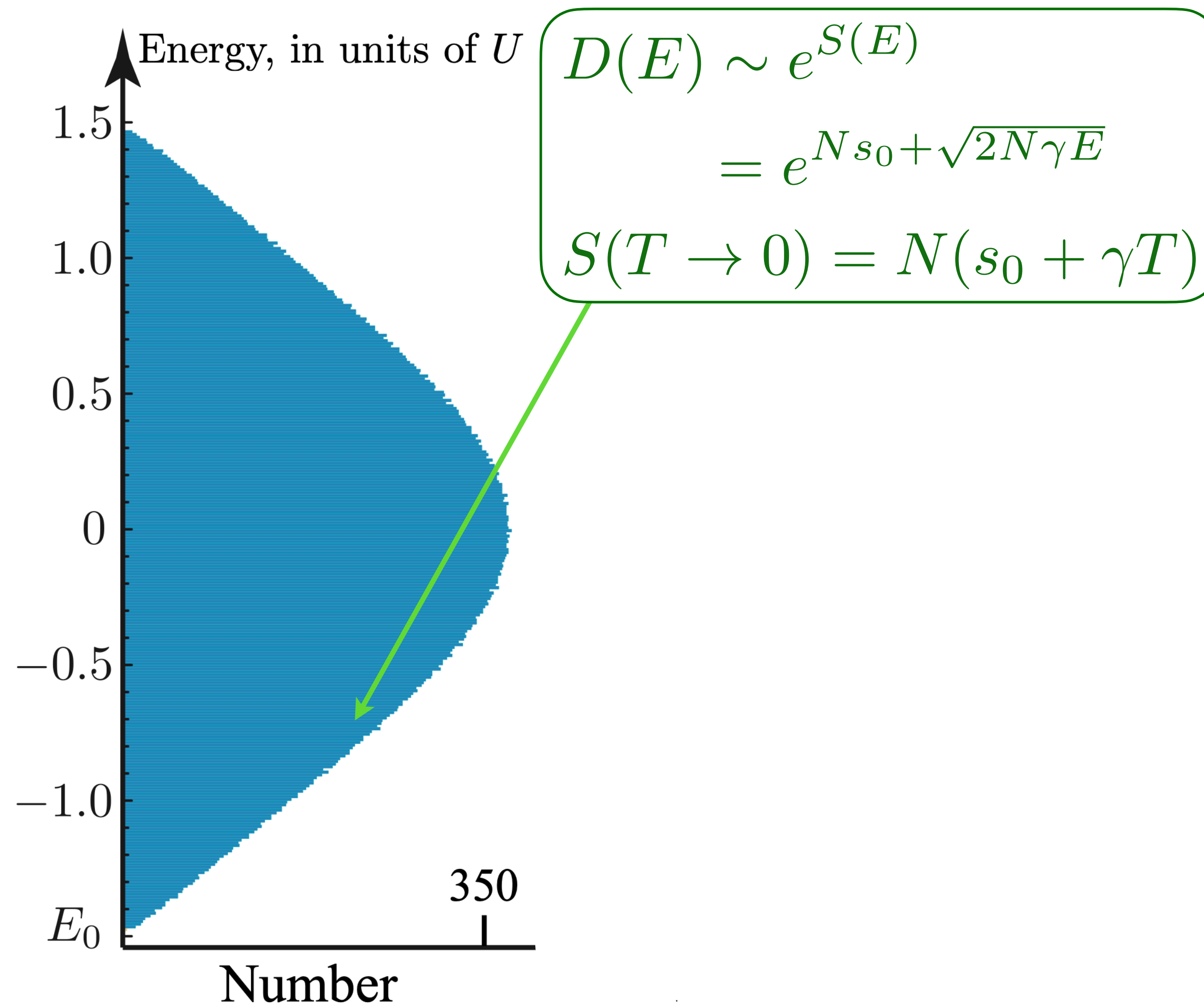
$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



Many-body density of states

# Complex SYK model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



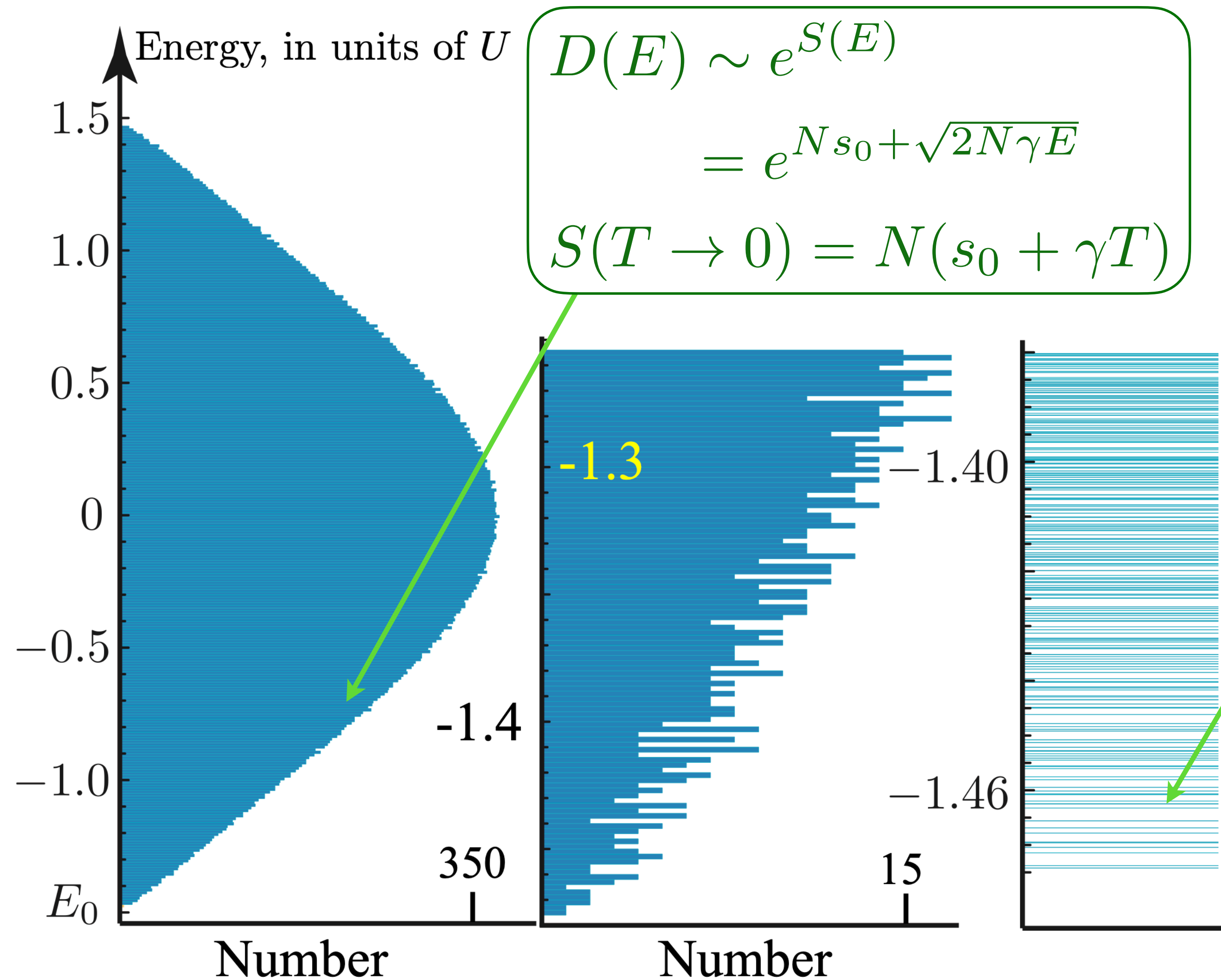
$$s_0 = 0.464848 \dots$$

A. Georges, O. Parcollet, and  
S. Sachdev,  
PRB **63**, 134406 (2001)

Many-body density of states

# Complex SYK model

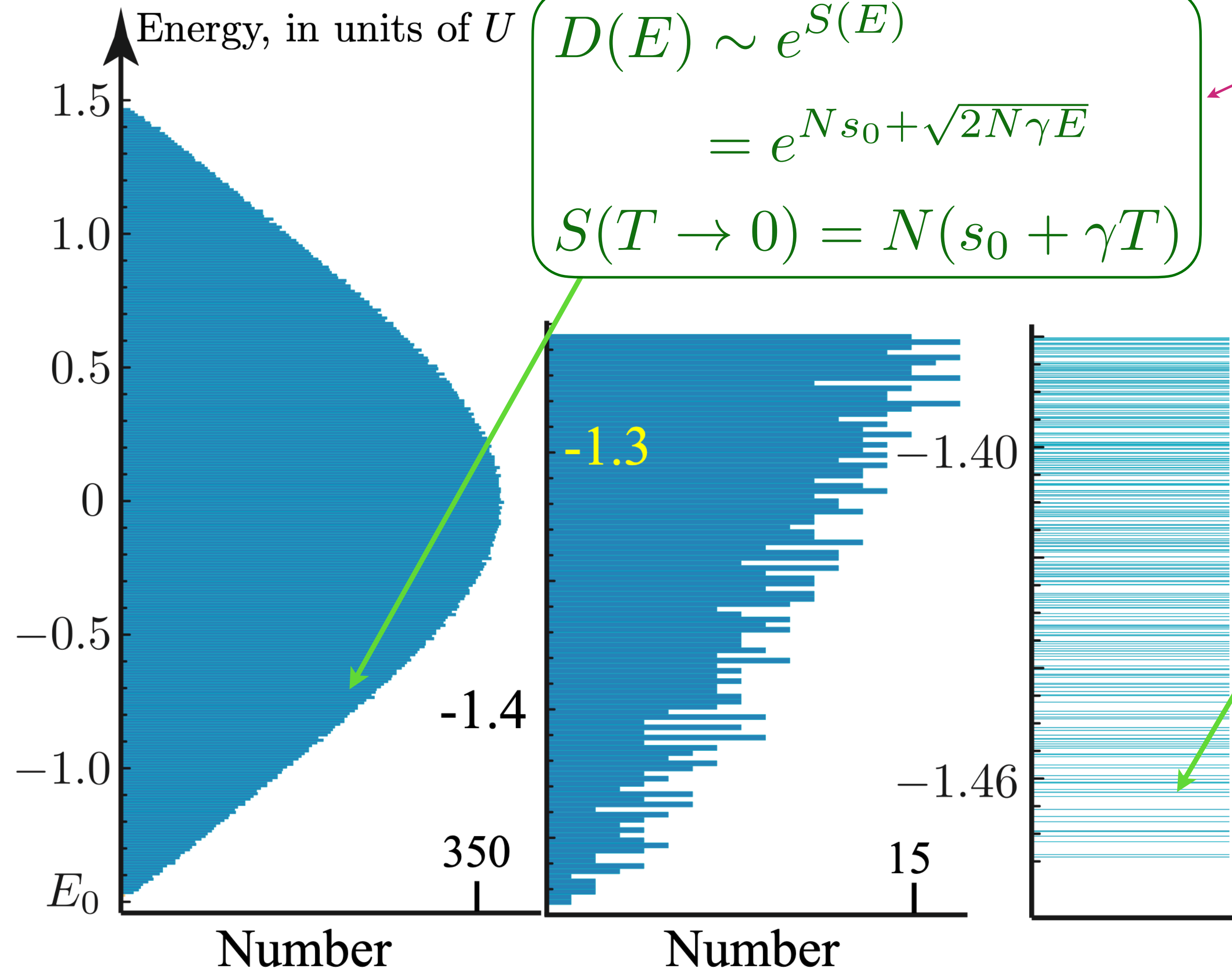
$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



Many-body density of states

# Complex SYK model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim e^{S(E)}$$

$$= e^{N s_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{N s_0} \sinh(\sqrt{2N\gamma E})$$

$$D(E) \sim 2 e^{N s_0} \sqrt{2N\gamma E}$$

No quasiparticle decomposition:  
wavefunctions change chaotically  
from one state to the next.

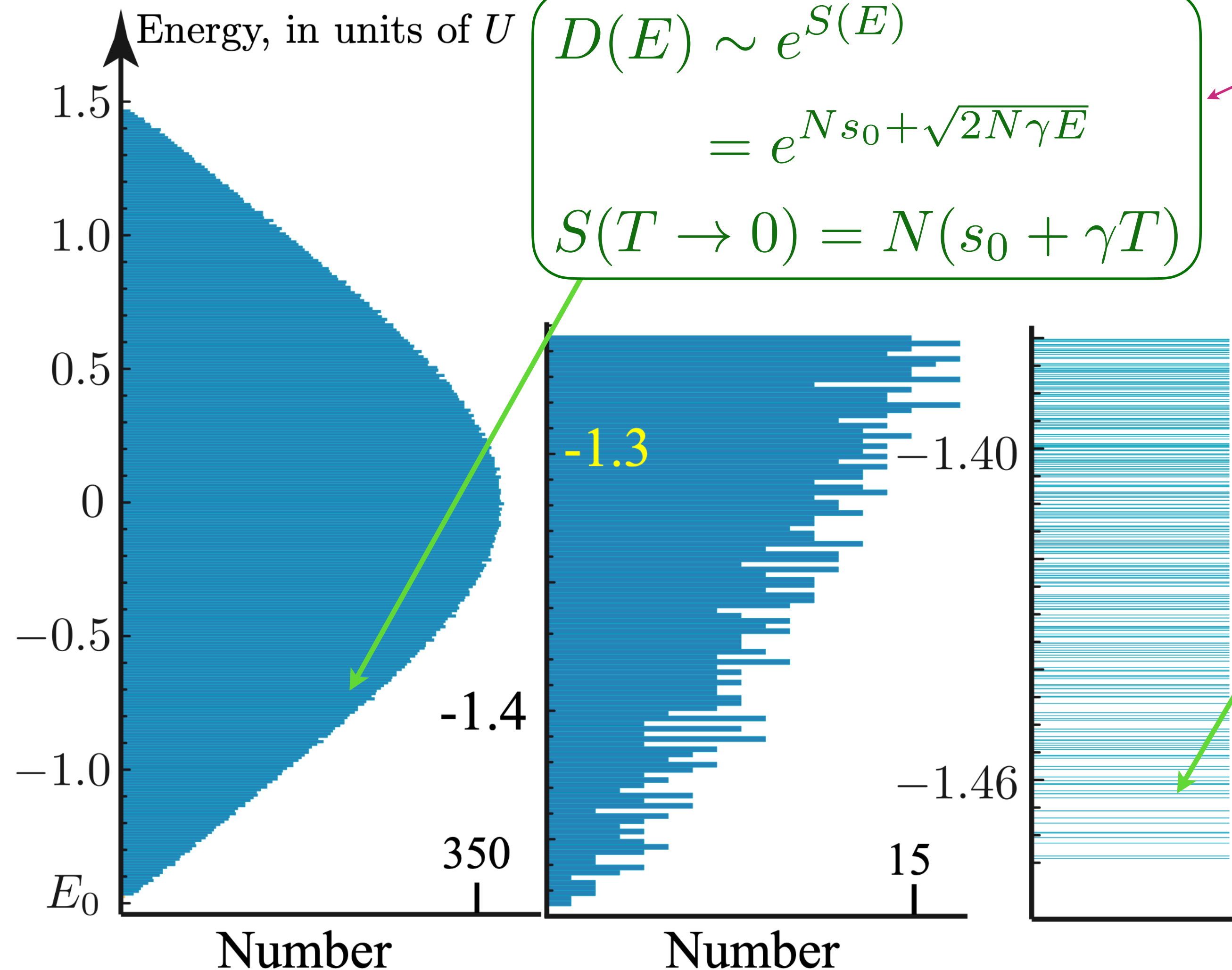
$$s_0 = 0.464848 \dots$$

A. Georges, O. Parcollet, and  
S. Sachdev,  
PRB **63**, 134406 (2001)

Many-body density of states

# Complex SYK model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim e^{S(E)}$$

$$= e^{N s_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{N s_0} \sinh(\sqrt{2N\gamma E})$$

$$e^{-F(T)/T} = \int_0^\infty dE D(E) e^{-E/T}$$

$$S(T) = -\partial F / \partial T$$

$$D(E) \sim$$

$$2 e^{N s_0} \sqrt{2N\gamma E}$$

No quasiparticle decomposition:  
wavefunctions change chaotically  
from one state to the next.

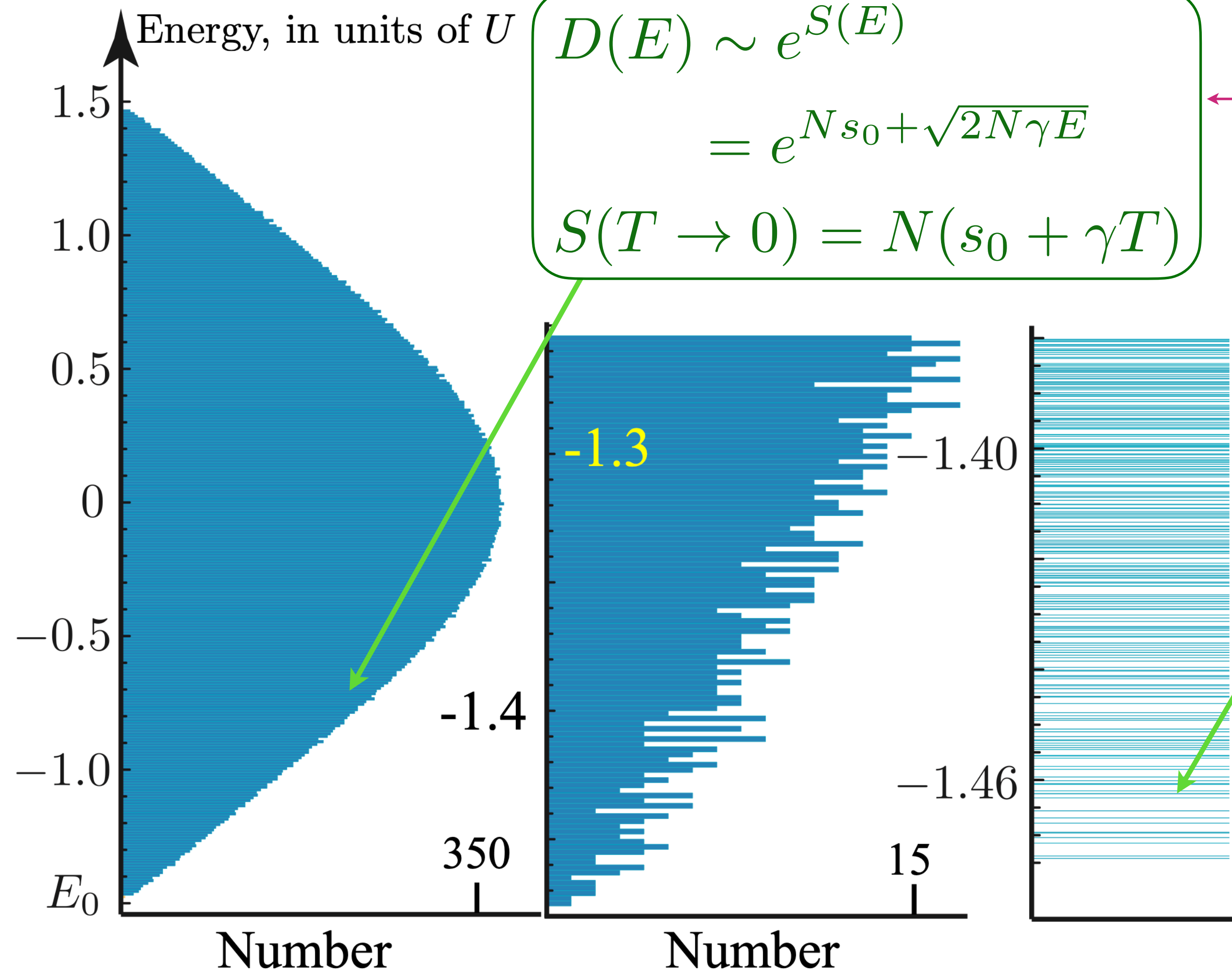
$$s_0 = 0.464848 \dots$$

A. Georges, O. Parcollet, and  
S. Sachdev,  
PRB **63**, 134406 (2001)

Many-body density of states

# Complex SYK model

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim e^{S(E)}$$

$$= e^{N s_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{N s_0} \sinh(\sqrt{2N\gamma E})$$

$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln \left( \frac{U}{T} \right)$$

$$D(E) \sim 2 e^{N s_0} \sqrt{2N\gamma E}$$

No quasiparticle decomposition:  
wavefunctions change chaotically  
from one state to the next.

$$s_0 = 0.464848 \dots$$

A. Georges, O. Parcollet, and  
S. Sachdev,  
PRB **63**, 134406 (2001)

Many-body density of states

# The Sachdev-Ye-Kitaev (SYK) model

## Universal Planckian time dynamics

- Green's function has Planckian time scaling  
 $G(\omega, T) \sim \omega^{-1/2} F(\hbar\omega/k_B T)$ .
- Leading (dangerously) irrelevant operator is a time reparameterization soft mode  $\tau \rightarrow f(\tau)$ .
- Time reparameterization mode leads to many-body quantum chaos in the out-of-time-order correlator (OTOC) with maximal Lyapunov exponent  $\lambda_L = 2\pi k_B T/\hbar$ . Kitaev (2015) Maldacena, Shenker, Stanford (2015) Maldacena Stanford (2016)

# The Sachdev-Ye-Kitaev (SYK) model

## Universal Planckian time dynamics

- Green's function has Planckian time scaling  
 $G(\omega, T) \sim \omega^{-1/2} F(\hbar\omega/k_B T)$ .
- Leading (dangerously) irrelevant operator is a time reparameterization soft mode  $\tau \rightarrow f(\tau)$ .
- Time reparameterization mode leads to many-body quantum chaos in the out-of-time-order correlator (OTOC) with maximal Lyapunov exponent  $\lambda_L = 2\pi k_B T/\hbar$ . Kitaev (2015) Maldacena, Shenker, Stanford (2015) Maldacena Stanford (2016)
- The  $T$ -dependence of the entropy also arises from the time reparameterization soft mode:  $S = N(s_0 + \gamma T) - (3/2) \ln(U/T)$ .

1. Introduction to Planckian metals

2. Introduction to black holes

3. The SYK model

4. Progress on the theory of black holes

5. Progress on the theory of Planckian metals

*A. Random  $t$ - $J$  model*

*B. Fermi surface coupled to a critical boson*

# Thermodynamics of quantum black holes with charge $Q$ :



$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right)$$
$$= \exp(S_{BH}) \times \left( \dots????\dots \right)$$

Gibbons, Hawking (1977)

Chambin, Emparan, Johnson, Myers (1999)

$$S_{BH}(T \rightarrow 0, Q) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left( 1 + \frac{2(\pi A_0)^{1/2}T}{\hbar c} \right)$$

$A_0$  is the area of the charged black hole horizon at  $T = 0$ .

$Q$  is the black hole charge.

$A_0$  is a function of  $Q$ .

Thermodynamics of quantum black holes with charge  $Q$ :



$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right)$$
$$= \exp(S_{BH}) \times \left( \dots????\dots \right)$$

Gibbons, Hawking (1977)

Chambin, Emparan, Johnson, Myers (1999)

$$S_{BH}(T \rightarrow 0, Q) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left( 1 + \frac{2(\pi A_0)^{1/2}T}{\hbar c} \right)$$

$A_0$  is the area of the charged black hole horizon at  $T = 0$ .

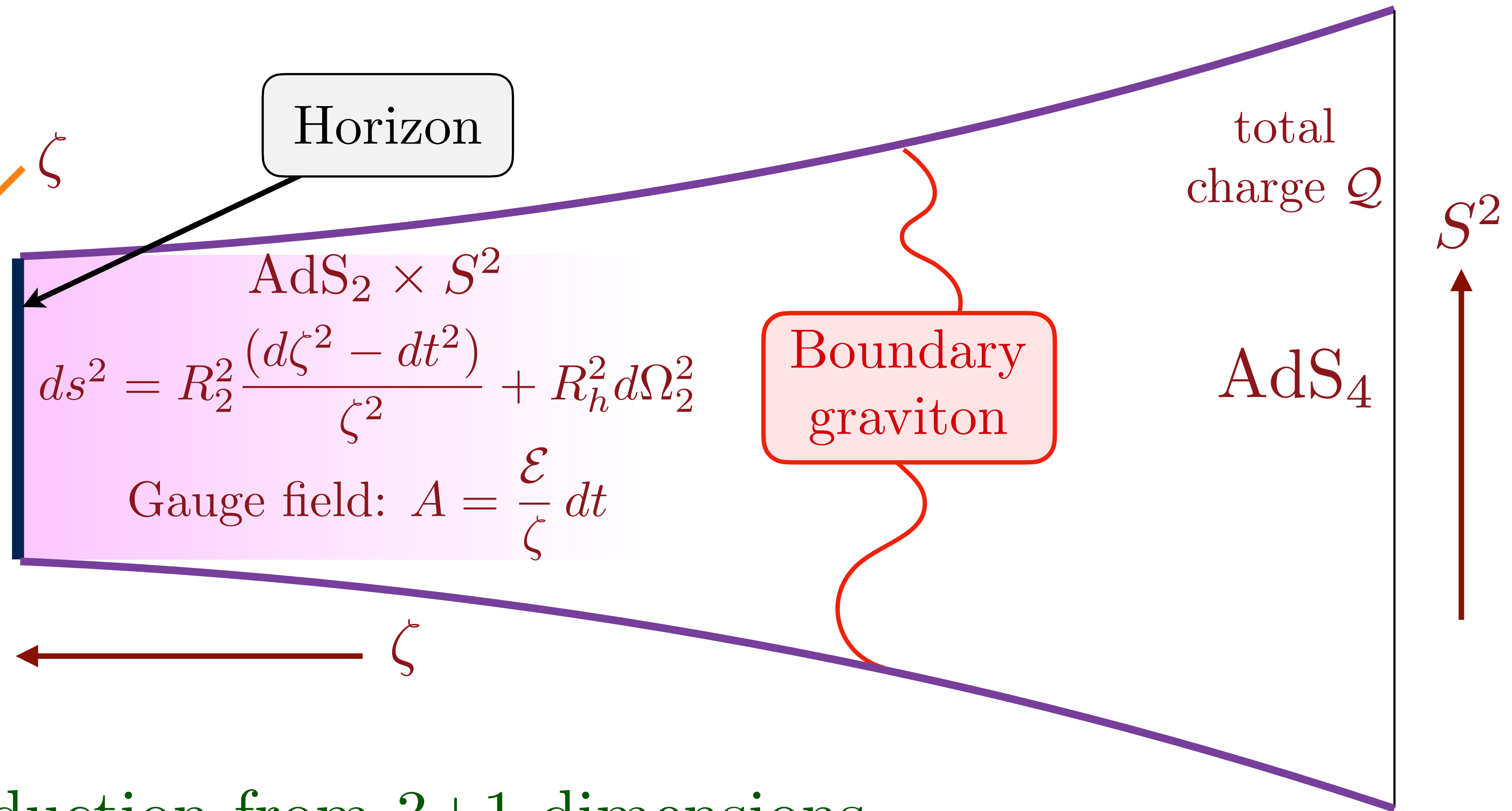
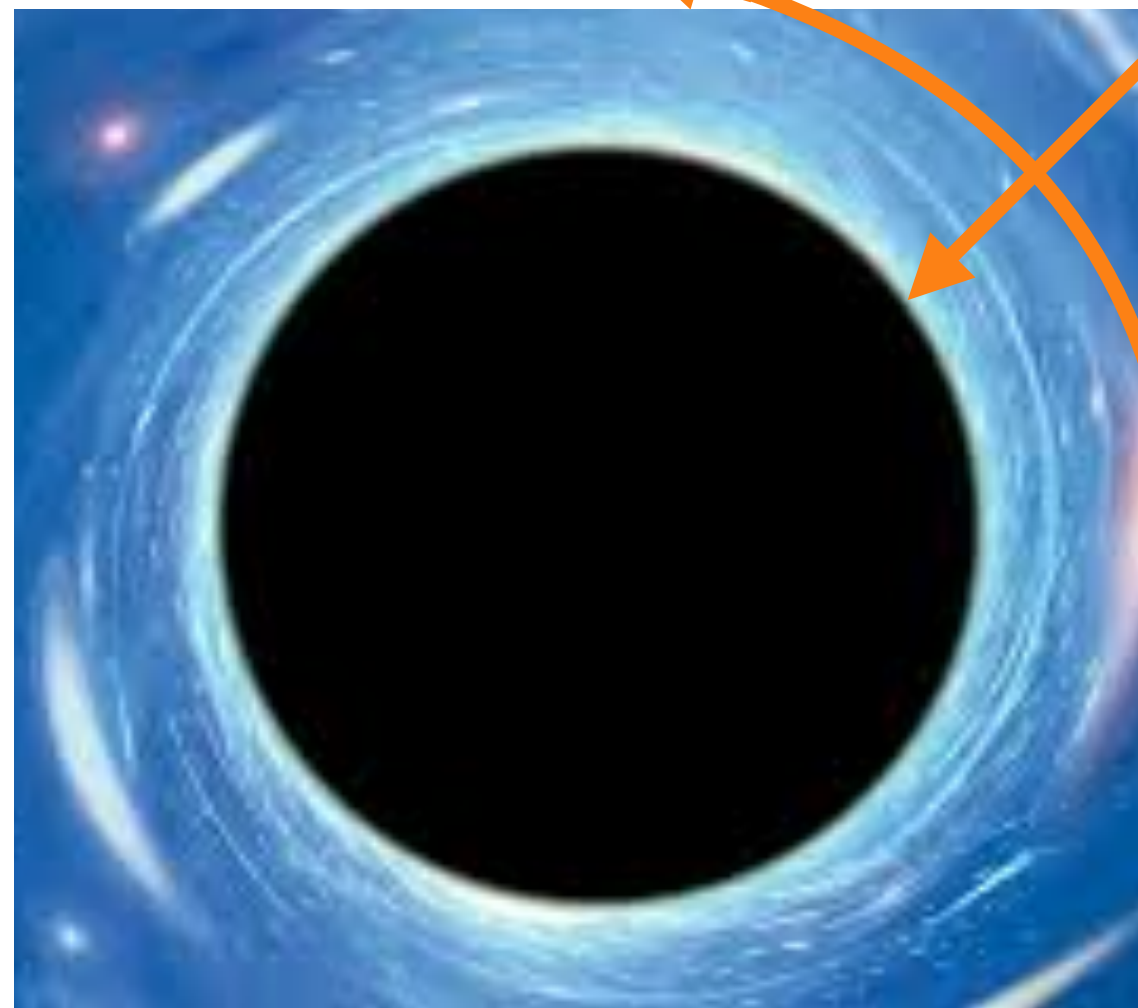
$Q$  is the black hole charge.

$A_0$  is a function of  $Q$ .

Note the similarity to the large  $N$  entropy of the SYK model !

(along with other similarities) Sachdev PRL 2010

# Reissner-Nordstrom black hole of Einstein-Maxwell theory



Dimensional reduction from 3+1 dimensions to 1+1 dimensions ( $AdS_2$ ) at low energies!

Thermodynamics of quantum black holes with charge  $Q$ :



$$\int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right) \\ \approx \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{JT gravity of AdS}_2+\text{boundary graviton}}^{(1+1)}[g_{\mu\nu}, A_{\mu}] \right)$$

$$S_{BH}(T \rightarrow 0, Q) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left( 1 + \frac{2(\pi A_0)^{1/2}T}{\hbar c} \right)$$

$A_0$  is the area of the charged black hole horizon at  $T = 0$ .

$Q$  is the black hole charge.

$A_0$  is a function of  $Q$ .

Thermodynamics of quantum black holes with charge  $Q$ :



$$\begin{aligned}
 & \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right) \\
 & \approx \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{JT gravity of AdS}_2+\text{boundary graviton}}^{(1+1)}[g_{\mu\nu}, A_{\mu}] \right) \\
 & = \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{SYK}}[\text{time reparameterizations } f(\tau), \text{ phase rotations } \phi(\tau)] \right)
 \end{aligned}$$

$$S_{BH}(T \rightarrow 0, Q) = \frac{A(T)c^3}{4G\hbar} = \frac{A_0c^3}{4G\hbar} \left( 1 + \frac{2(\pi A_0)^{1/2}T}{\hbar c} \right)$$

$A_0$  is the area of the charged black hole horizon at  $T = 0$ .

$Q$  is the black hole charge.

$A_0$  is a function of  $Q$ .

# Thermodynamics of quantum black holes with charge $Q$ :



$$\begin{aligned}
 & \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{Einstein gravity+Maxwell EM}}^{(3+1)}[g_{\mu\nu}, A_{\mu}] \right) \\
 & \approx \int \mathcal{D}g_{\mu\nu} \mathcal{D}A_{\mu} \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{JT gravity of AdS}_2+\text{boundary graviton}}^{(1+1)}[g_{\mu\nu}, A_{\mu}] \right) \\
 & = \int \mathcal{D}f(\tau) \mathcal{D}\phi(\tau) \exp \left( -\frac{1}{\hbar} \mathcal{S}_{\text{SYK}}[\text{time reparameterizations } f(\tau), \text{ phase rotations } \phi(\tau)] \right)
 \end{aligned}$$

$$S(T \rightarrow 0, Q) = S_{BH} - \frac{3}{4} \ln \left( \frac{\hbar c^5}{GT^2} \right)$$

$$S_{BH} = \frac{A(T)c^3}{4G\hbar} = \frac{A_0 c^3}{4G\hbar} \left( 1 + \frac{2(\pi A_0)^{1/2} T}{\hbar c} \right)$$

$A_0$  is the area of the charged black hole horizon at  $T = 0$ ,  $Q$  is the black hole charge. The  $\ln T$  term is the contribution of the boundary graviton.

(There is also a  $-(241/45) \ln(A_0/G)$  correction at  $T = 0$   
A. Sen 2011)

# Many-body density of states

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$

Same entropy and (coarse-grained) density of states in a model of interacting (fermionic) qubits with a discrete spectrum!

Energy, in units of  $U$

$$D(E) \sim e^{S(E)}$$

$$= e^{Ns_0 + \sqrt{2N\gamma E}}$$

$$S(T \rightarrow 0) = N(s_0 + \gamma T)$$

$$D(E) \sim 2 e^{Ns_0} \sinh(\sqrt{2N\gamma E})$$

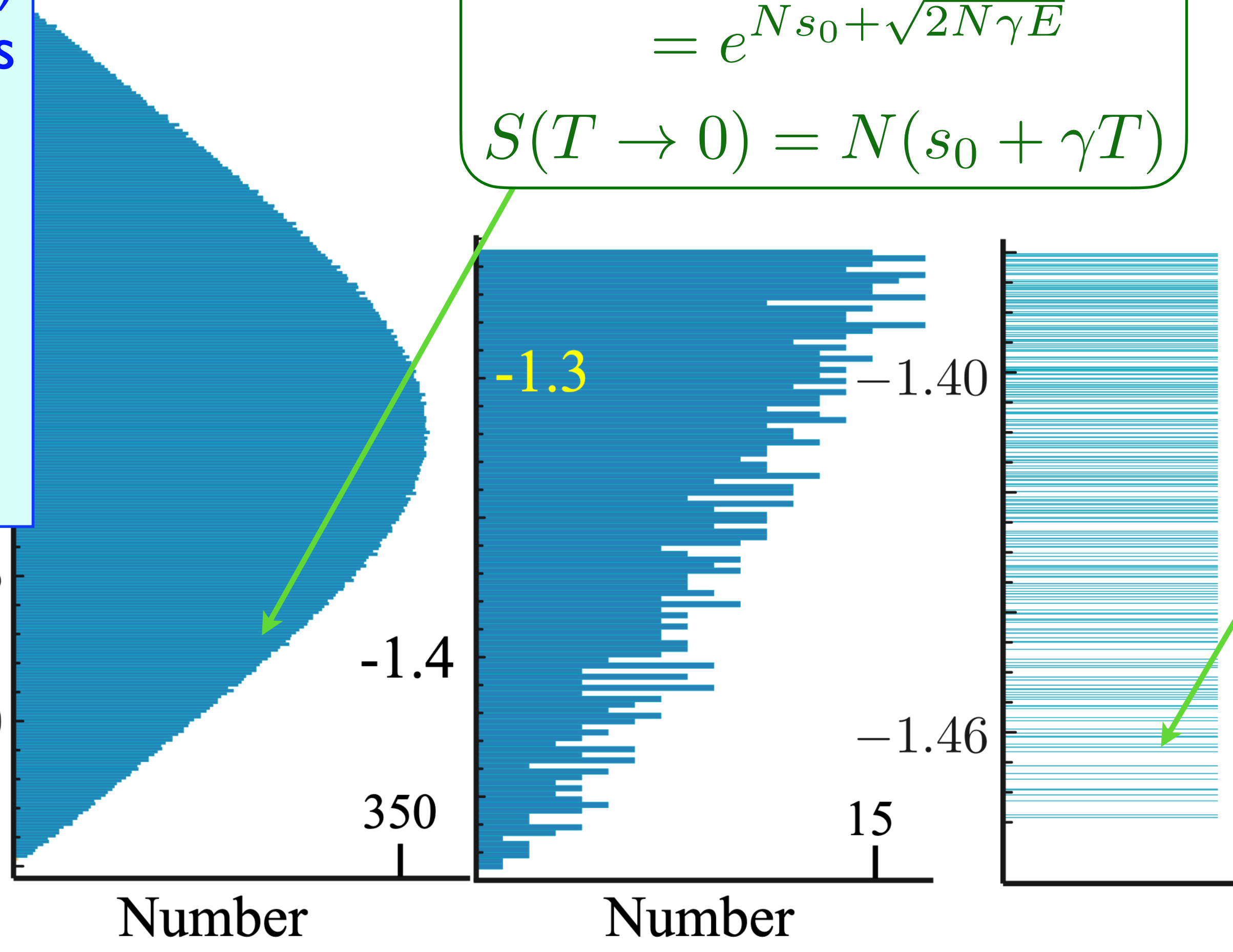
$$S(T) = N(s_0 + \gamma T) - \frac{3}{2} \ln \left( \frac{U}{T} \right)$$

$$D(E) \sim 2 e^{Ns_0} \sqrt{2N\gamma E}$$

No quasiparticle decomposition: wavefunctions change chaotically from one state to the next.

$$s_0 = 0.464848 \dots$$

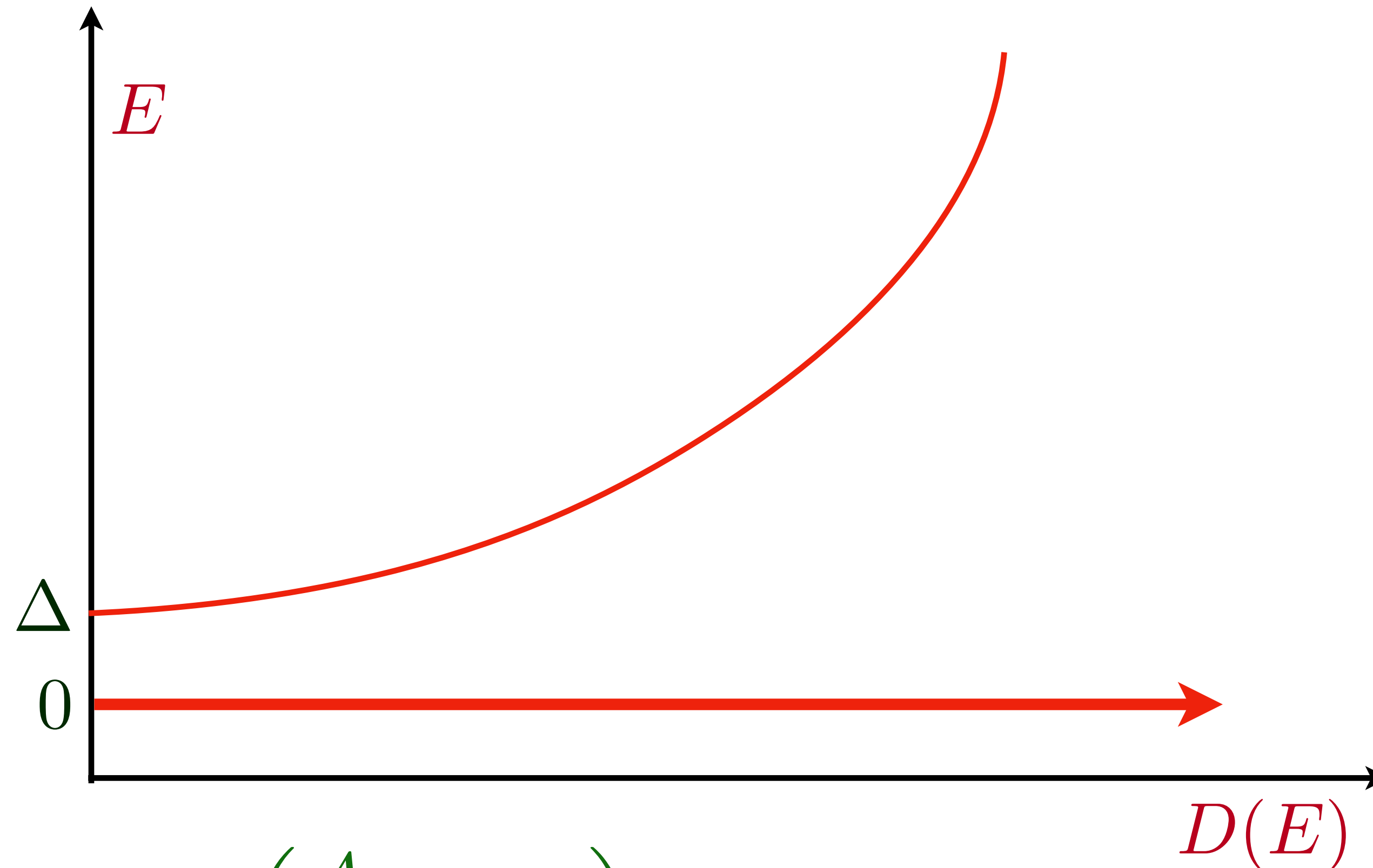
A. Georges, O. Parcollet, and S. Sachdev,  
PRB **63**, 134406 (2001)



## Complex SYK model

## Many-body density of states

$$D(E) = \sum_i \delta(E - E_i); \quad E_0 + E_i \Rightarrow \text{Many body eigenvalue}$$



$$D(E) \sim \exp\left(\frac{A_0}{4G} + \dots\right) \delta(E) + f_{\text{reg}}(E - \Delta), \quad \Delta \sim R_h^{-1}$$

## Supersymmetric black holes and SYK models

1. Introduction to Planckian metals
2. Introduction to black holes
3. The SYK model
4. Progress on the theory of black holes
5. Progress on the theory of Planckian metals
  - A. Random  $t$ - $J$  model*
  - B. Fermi surface coupled to a critical boson*

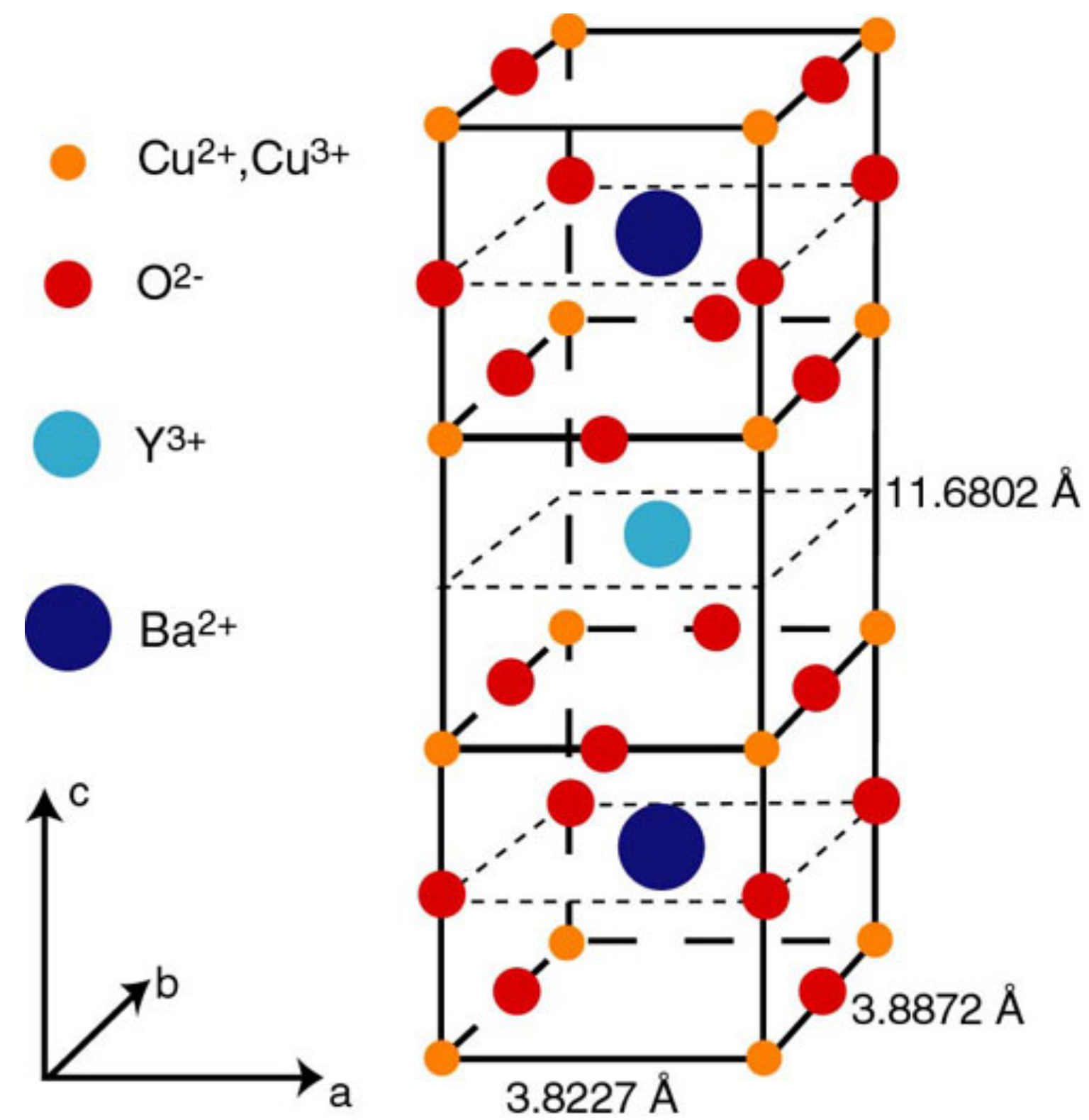
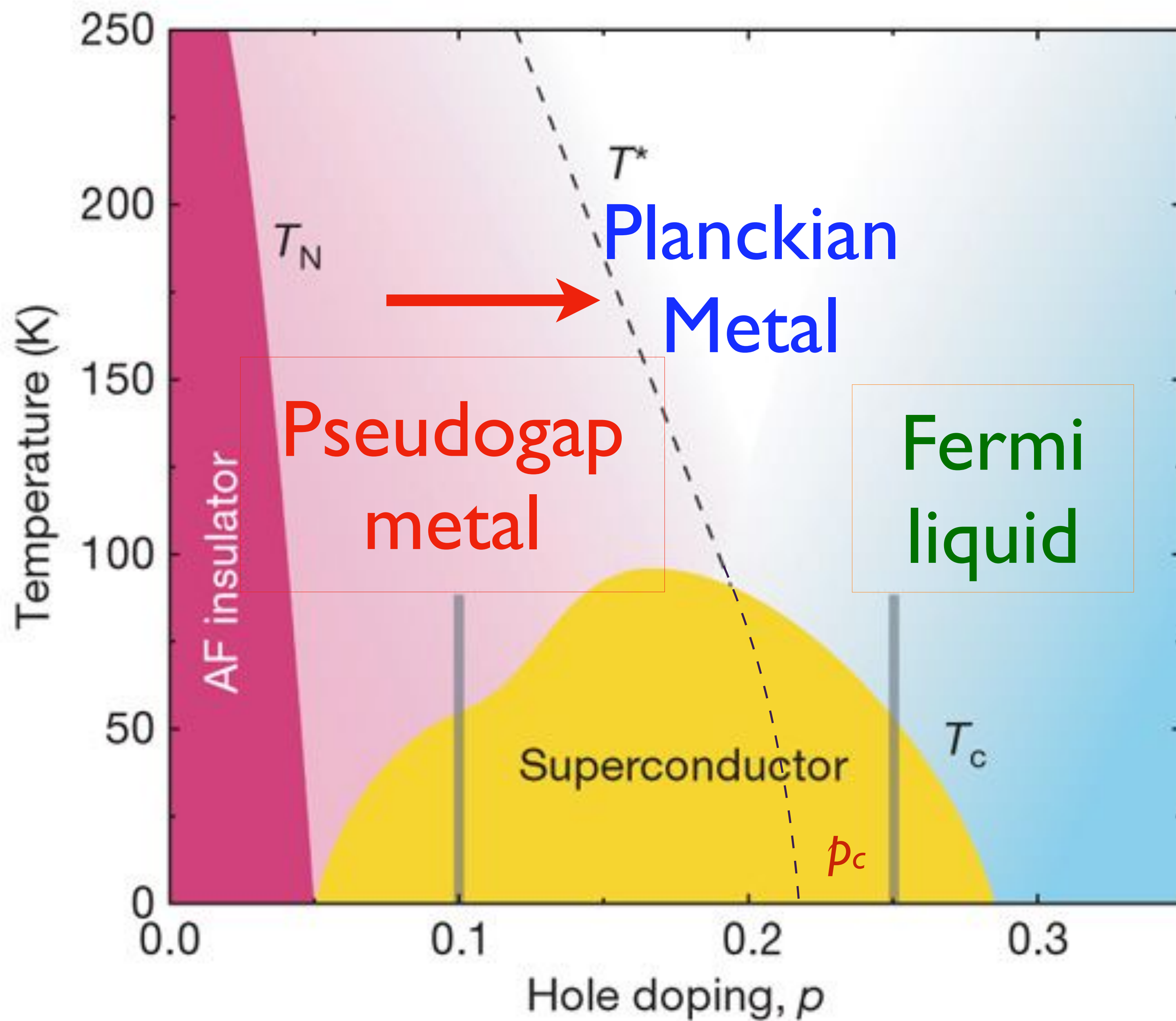
1. Introduction to Planckian metals

2. Introduction to black holes

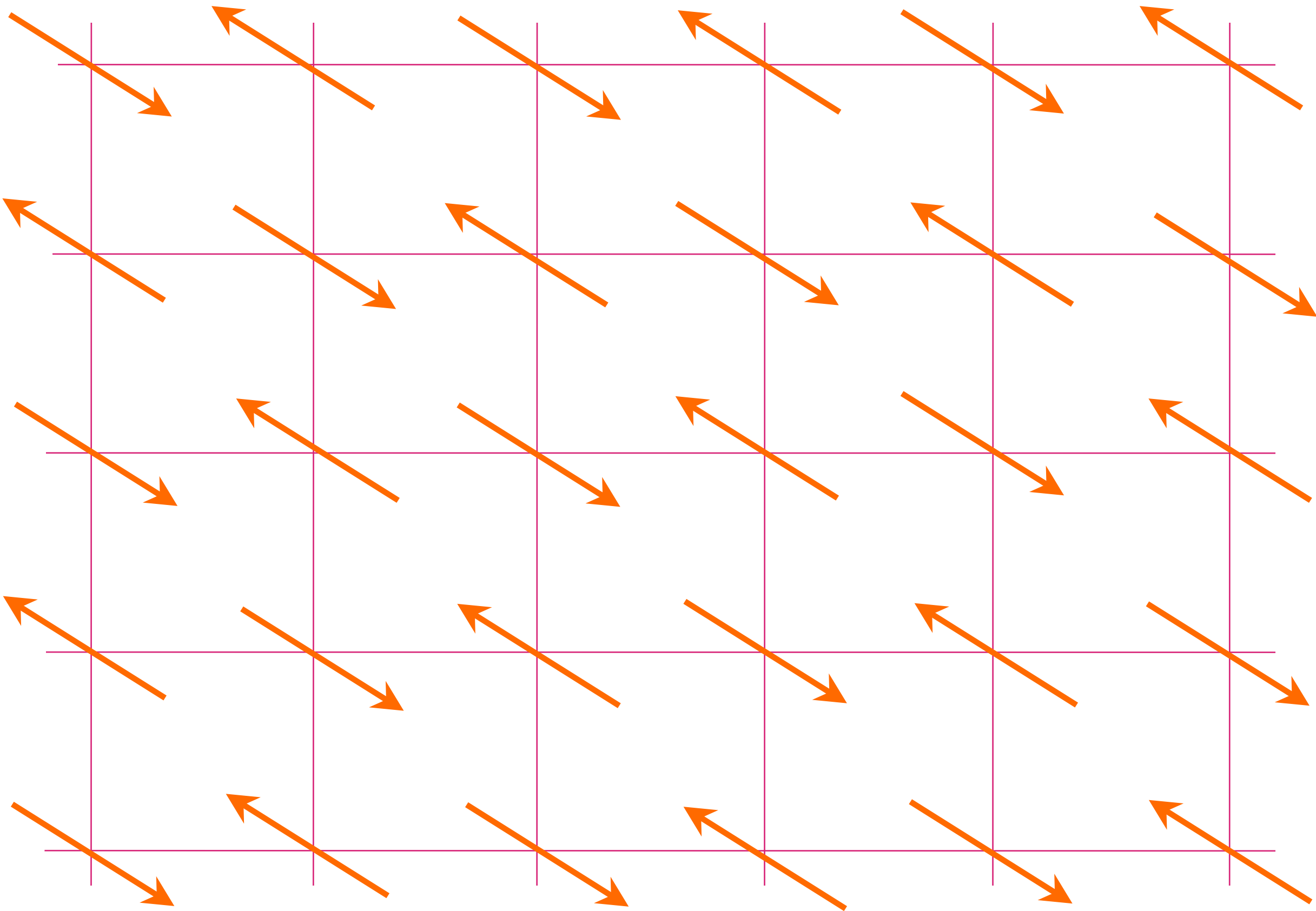
3. The SYK model

4. Progress on the theory of black holes

5. Progress on the theory of Planckian metals

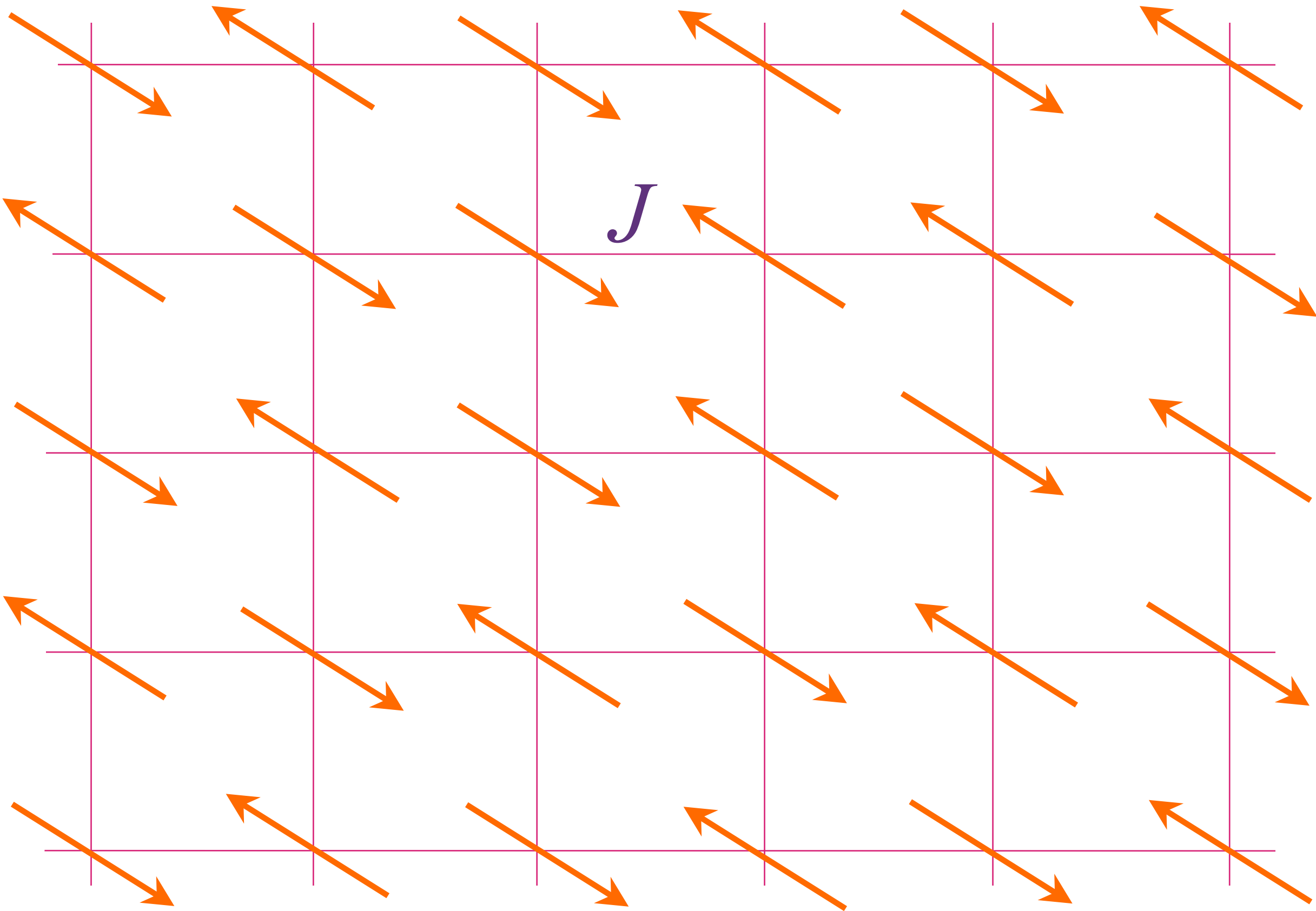


# Insulating antiferromagnet



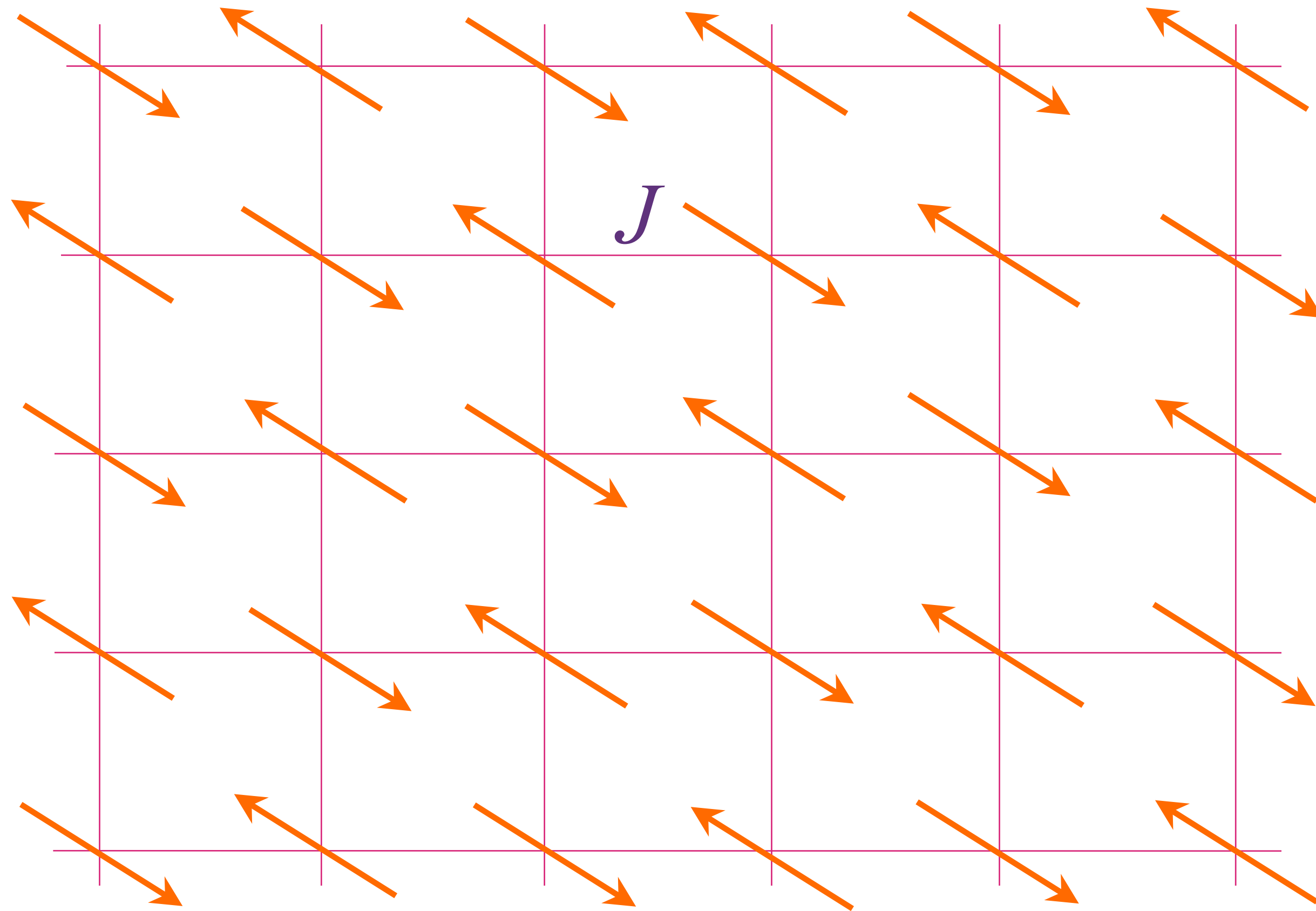
$p=0$

# Insulating antiferromagnet



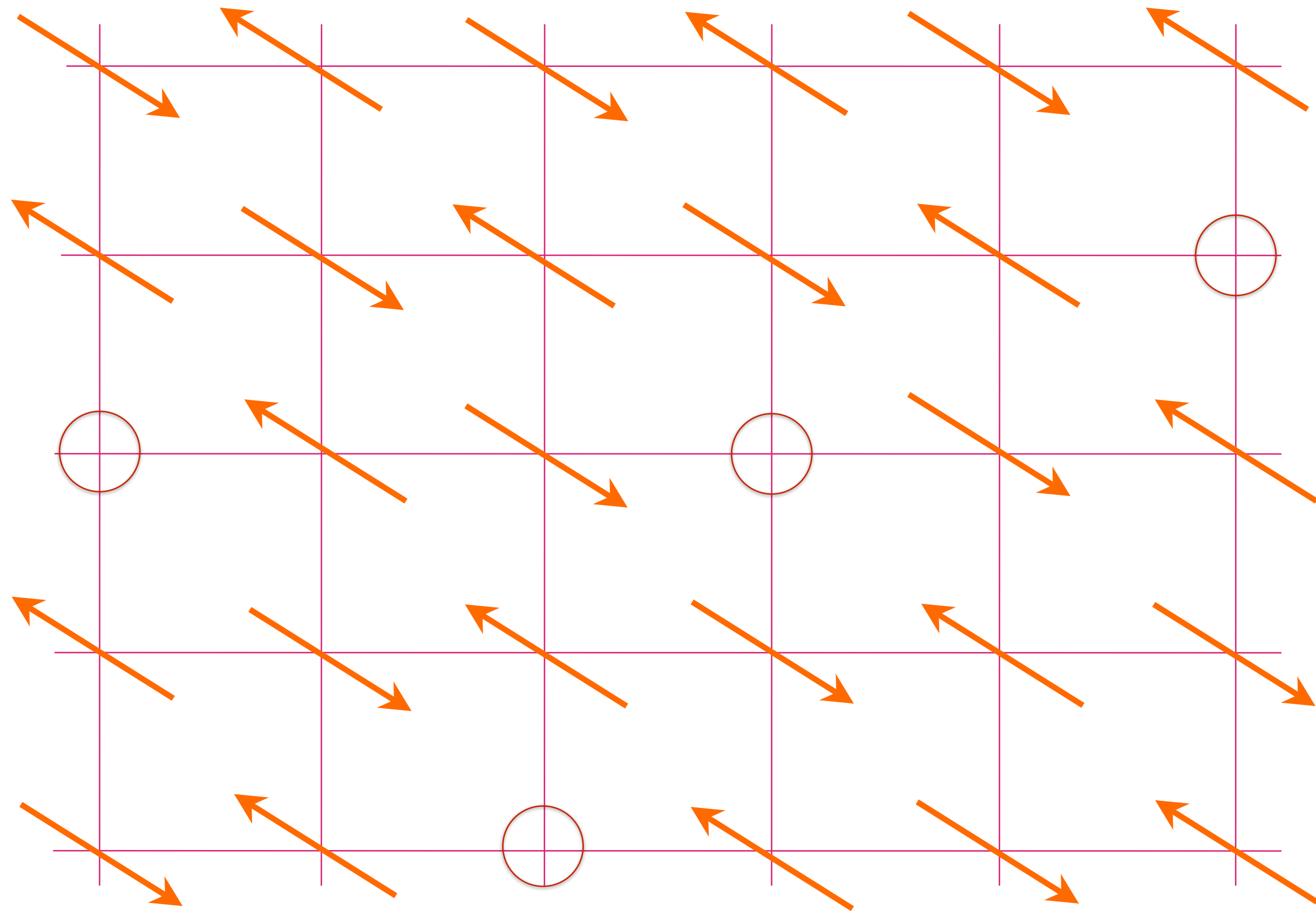
$$p=0$$

# Insulating antiferromagnet

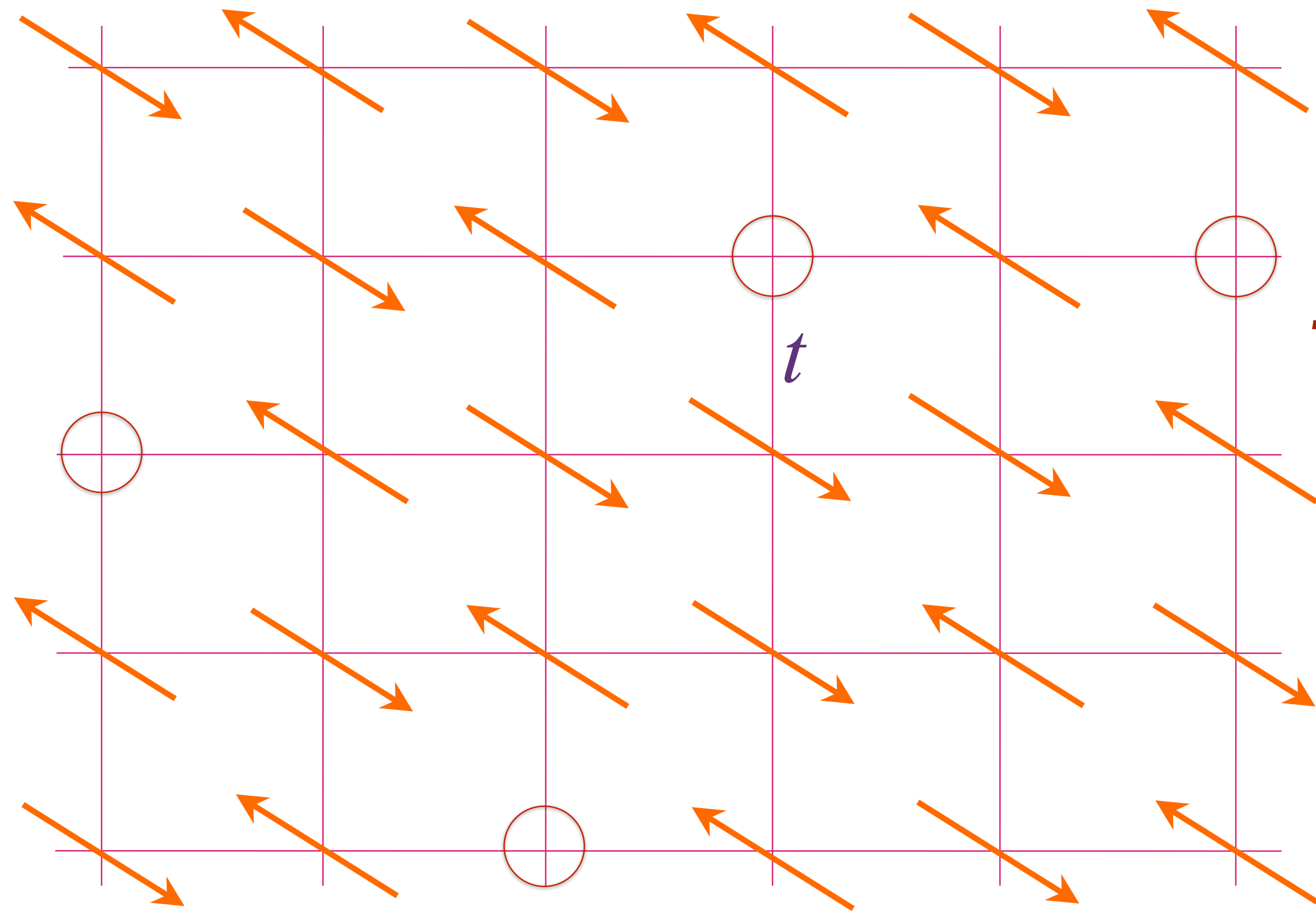


$$p=0$$

# Antiferromagnet doped with hole density $p$

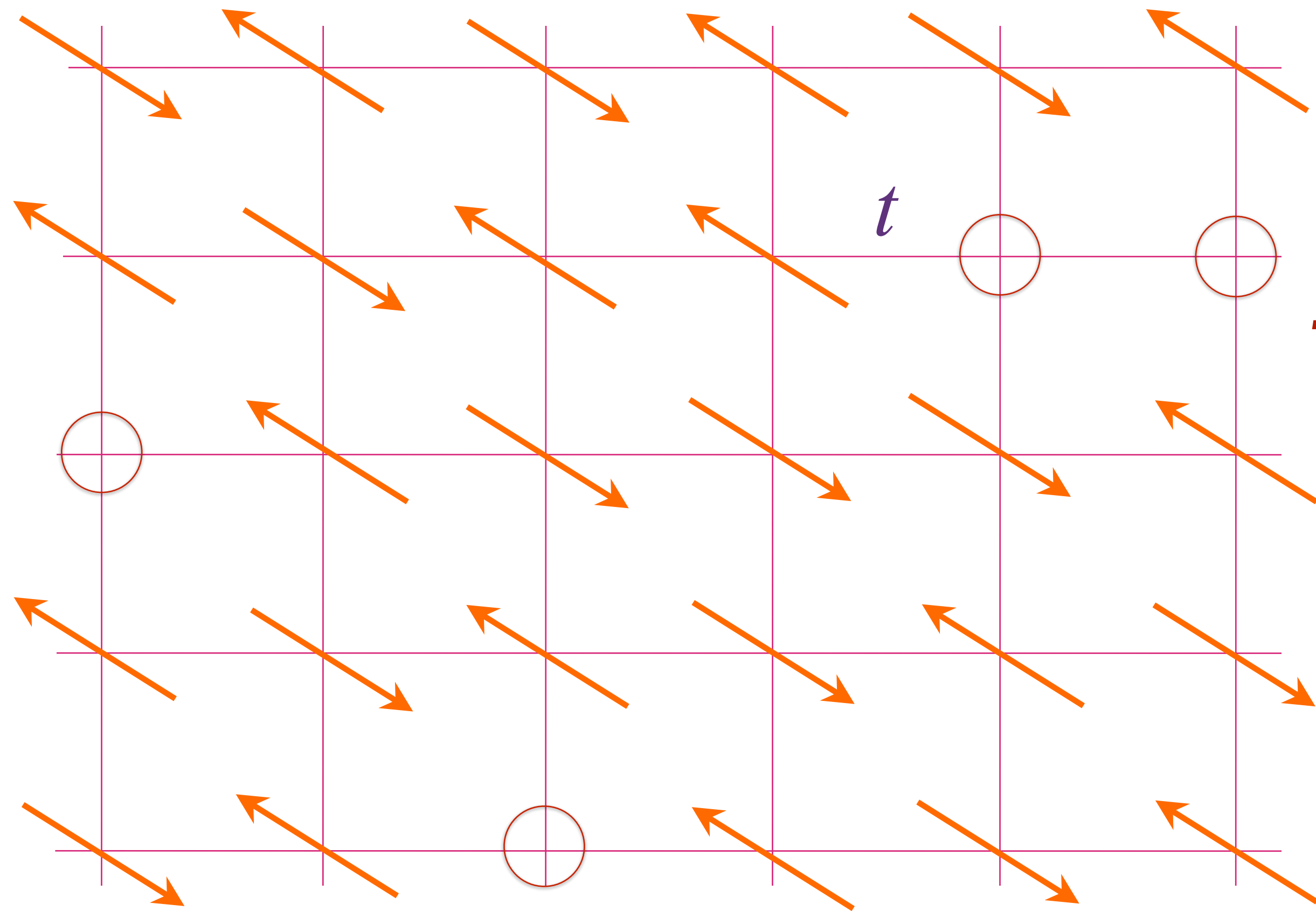


# Antiferromagnet doped with hole density $p$



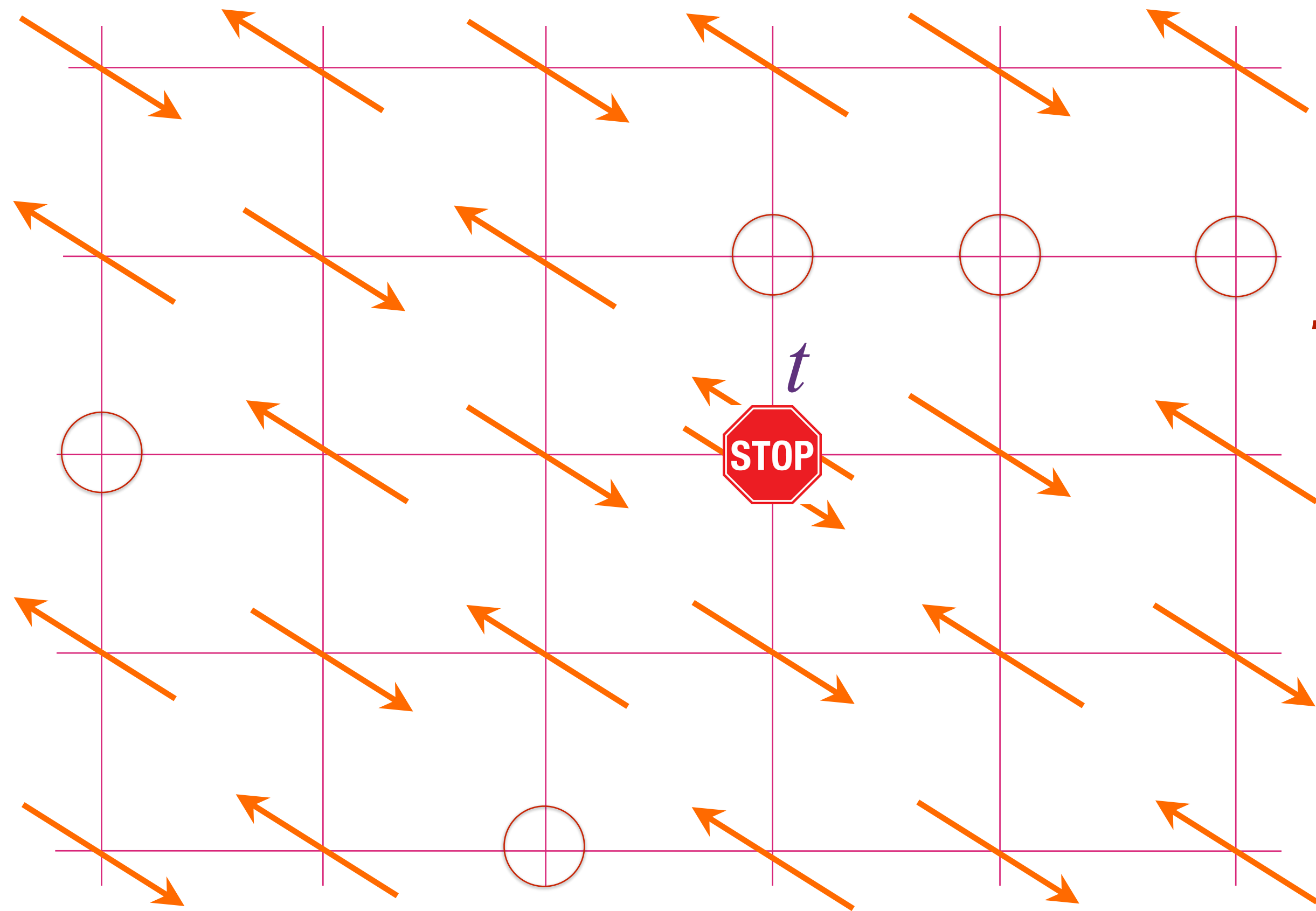
**$t$ - $J$  model**

# Antiferromagnet doped with hole density $p$



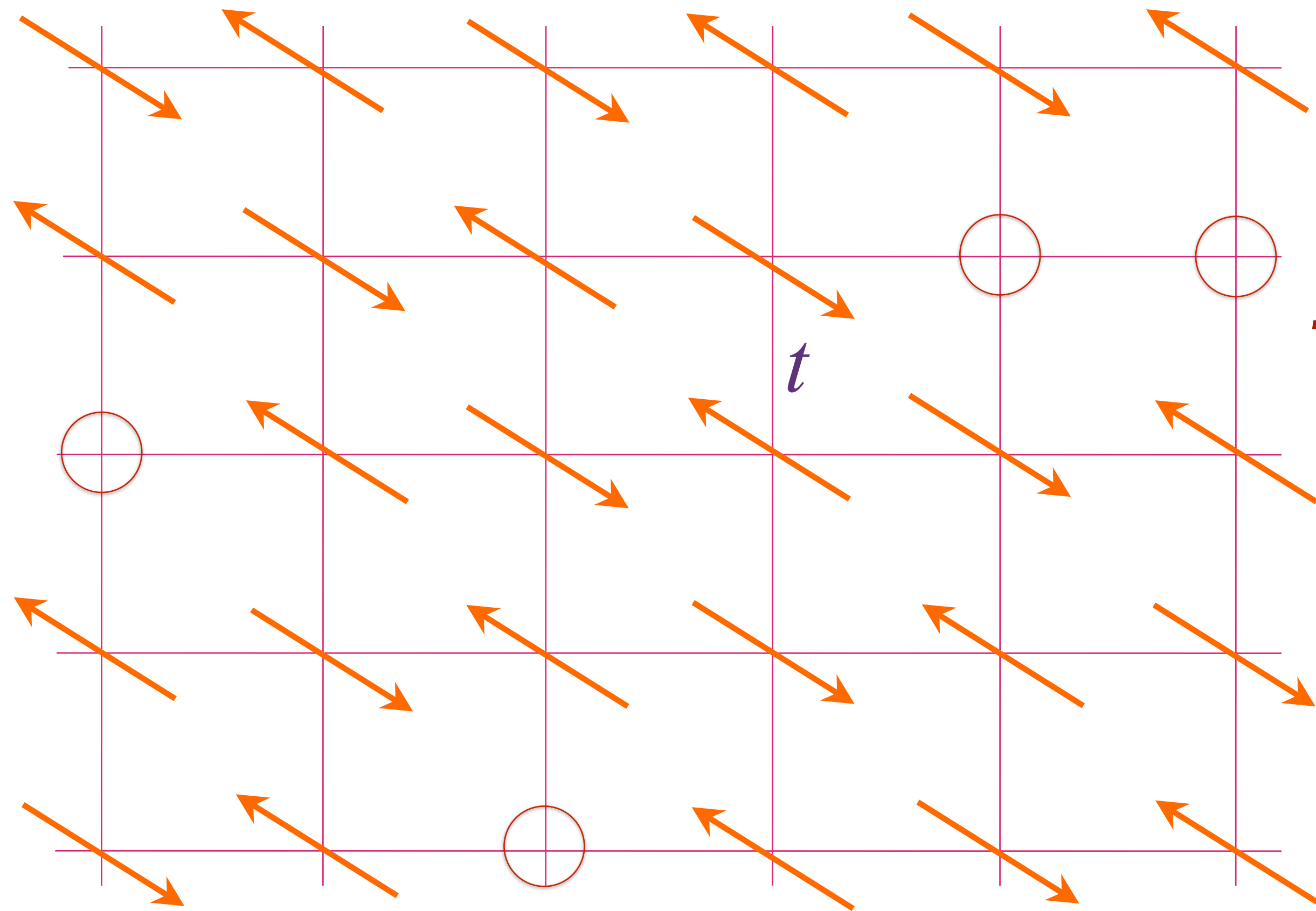
**$t$ - $J$  model**

# Antiferromagnet doped with hole density $p$



**$t$ - $J$  model**

# Antiferromagnet doped with hole density $p$



**$t$ - $J$  model**

## Random $t$ - $J$ model doped with hole density $p$

$$H = -\frac{1}{\sqrt{N}} \sum_{i,j=1}^N t_{ij} \mathcal{P}_d c_{i\alpha}^\dagger c_{j\alpha} \mathcal{P}_d + \frac{1}{\sqrt{N}} \sum_{i<j=1}^N J_{ij} \vec{S}_i \cdot \vec{S}_j$$

$$\vec{S}_i = \frac{1}{2} c_{i\alpha}^\dagger \vec{\sigma} c_{i\alpha}$$

$\mathcal{P}_d$  projects out doubly-occupied sites.

$$J_{ij} \text{ random, } \overline{J_{ij}} = 0, \overline{J_{ij}^2} = J^2$$

$J \Rightarrow$  two-particle interaction, similar to that in SYK  
 $t \Rightarrow$  one-particle hopping, can be regular or random

## Random $t$ - $J$ model doped with hole density $p$

$$H = -\frac{1}{\sqrt{N}} \sum_{i,j=1}^N t_{ij} \mathcal{P}_d c_{i\alpha}^\dagger c_{j\alpha} \mathcal{P}_d + \frac{1}{\sqrt{N}} \sum_{i<j=1}^N J_{ij} \vec{S}_i \cdot \vec{S}_j$$

$$\vec{S}_i = \frac{1}{2} c_{i\alpha}^\dagger \vec{\sigma} c_{i\alpha}$$

$\mathcal{P}_d$  projects out doubly-occupied sites.

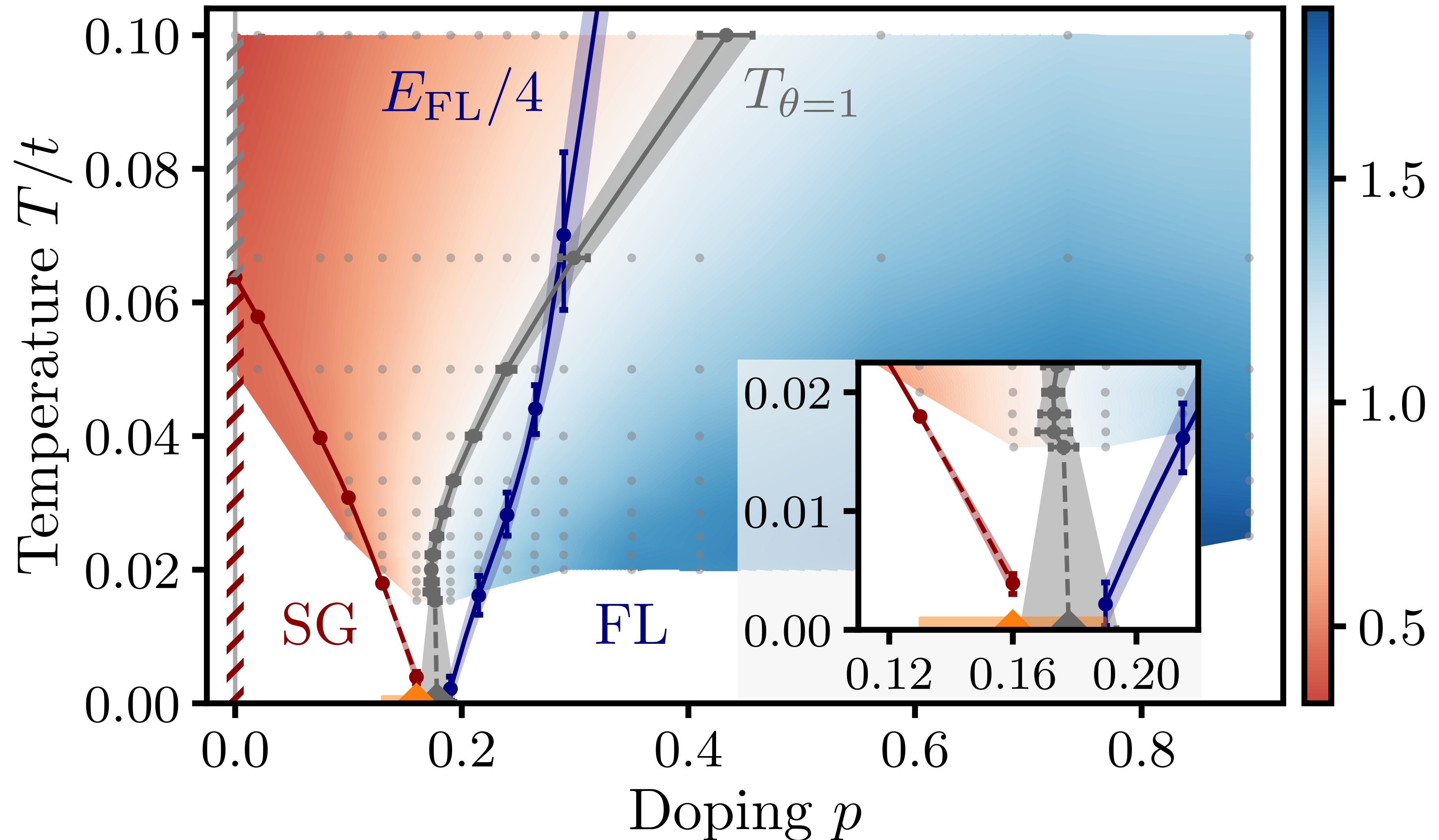
$$J_{ij} \text{ random, } \overline{J_{ij}} = 0, \overline{J_{ij}^2} = J^2$$

Parisi solved the model with this term only with  $\vec{S}_i$  replaced by classical Ising spins  $\sigma_i = \pm 1$

$J \Rightarrow$  two-particle interaction, similar to that in SYK

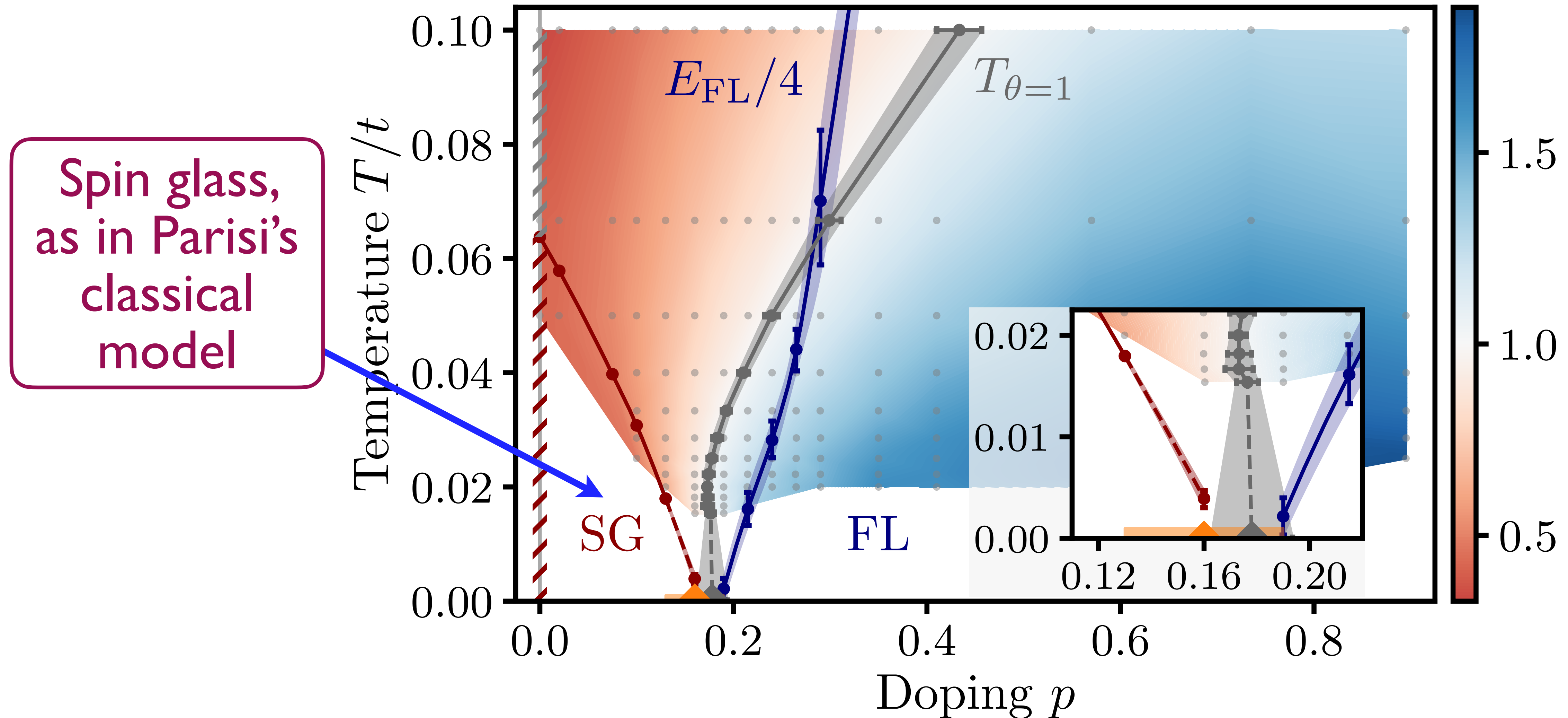
$t \Rightarrow$  one-particle hopping, can be regular or random

Numerical solution of  $t$ - $J$  model on a fully-connected cluster  
with all-to-all and random  $t_{ij}$  and  $J_{ij}$

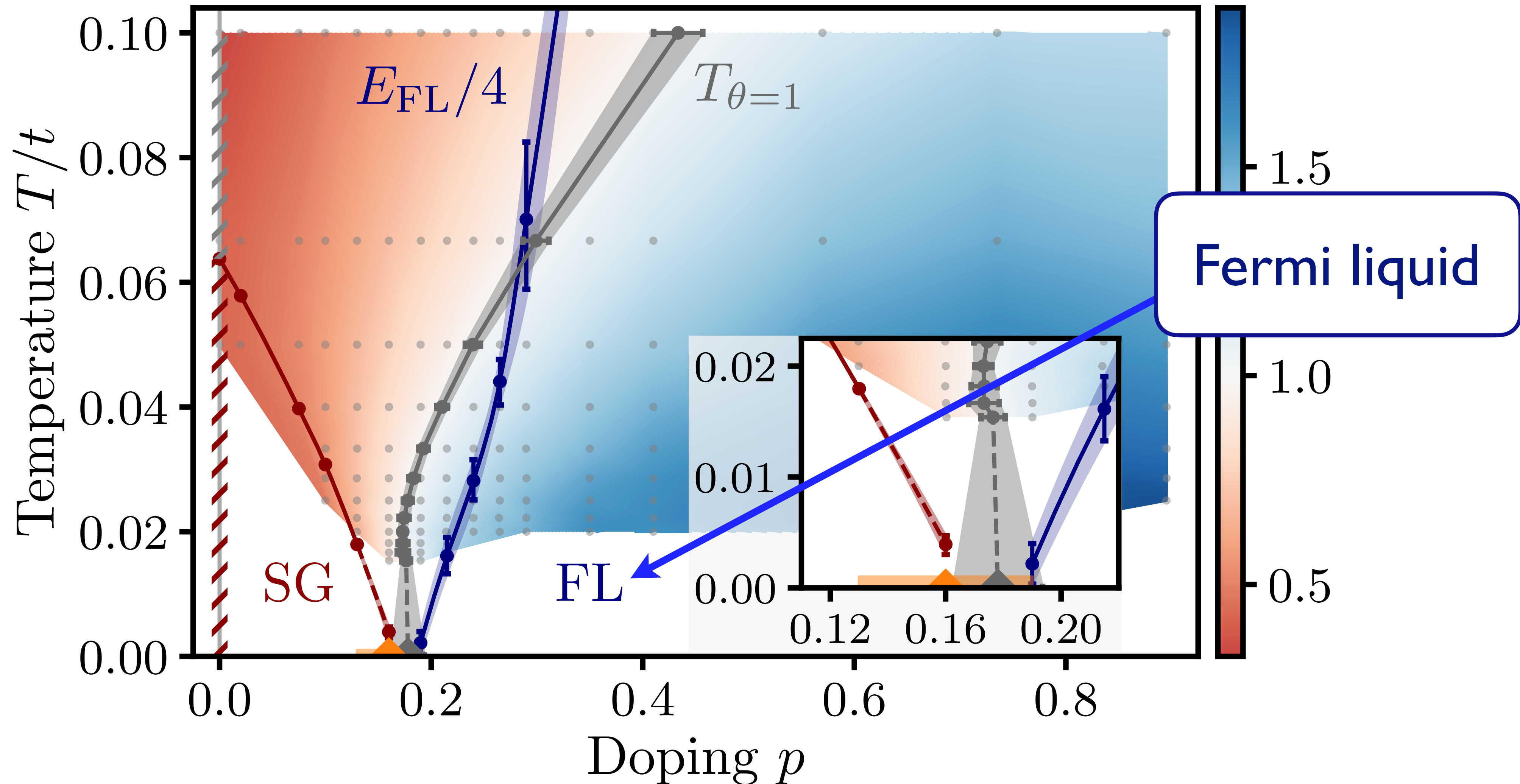


P. T. Dumitrescu, N. Wentzell, A. Georges, O. Parcollet, arXiv:2103.08607  
H. Shackleton, A. Wietek, A. Georges, and S. Sachdev, PRL **126**, 136602 (2021)

Numerical solution of  $t$ - $J$  model on a fully-connected cluster  
with all-to-all and random  $t_{ij}$  and  $J_{ij}$

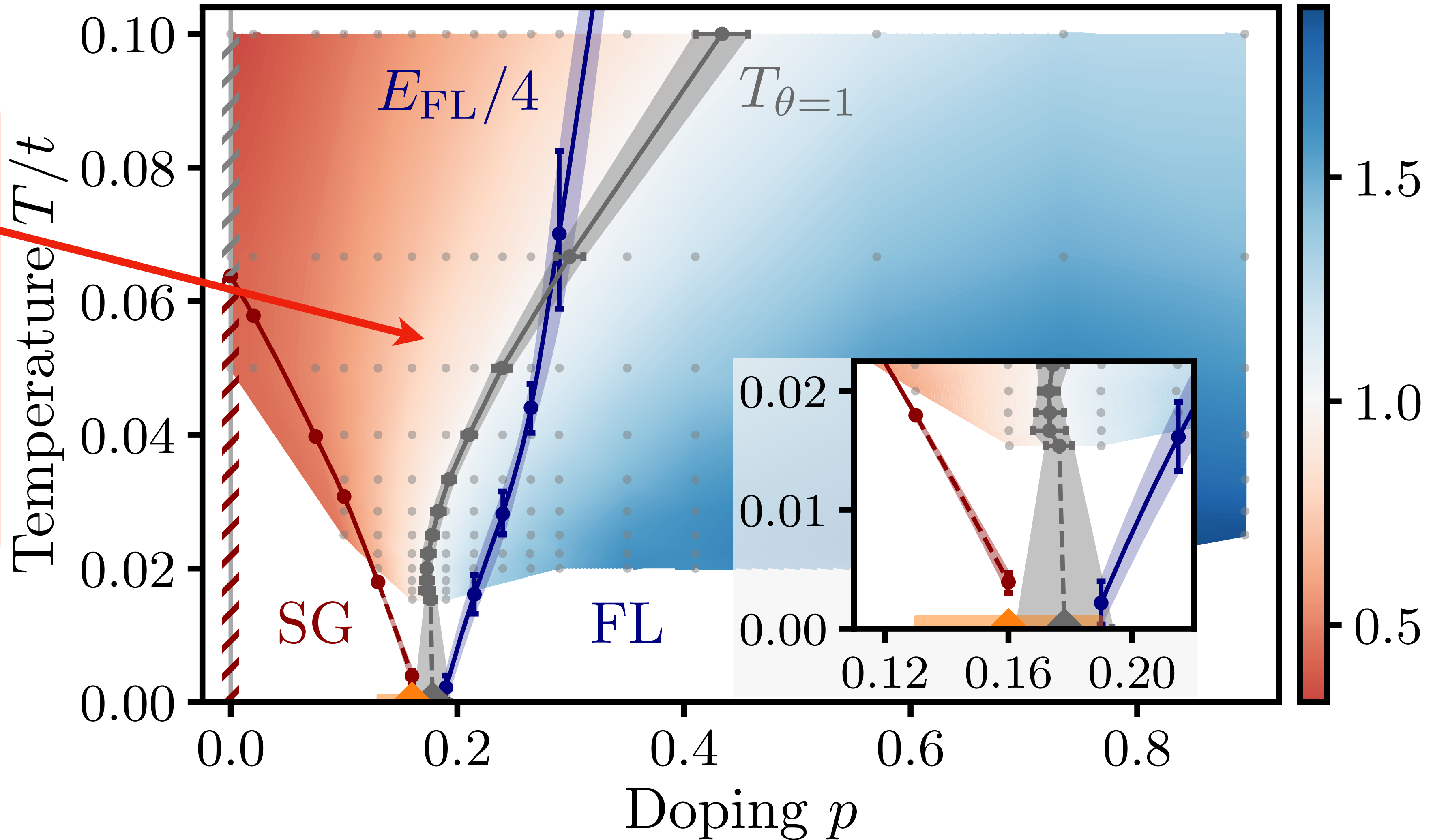


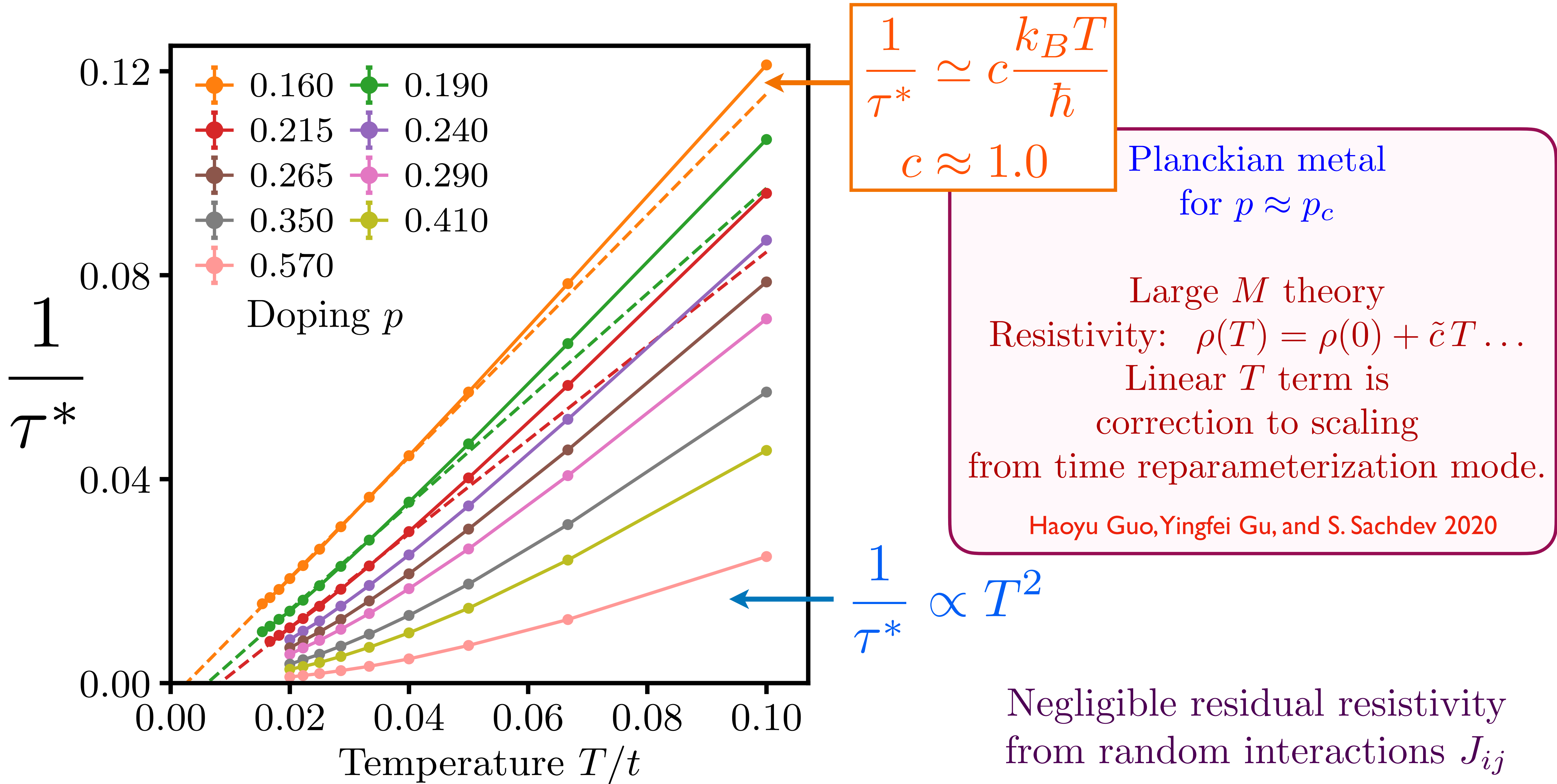
Numerical solution of  $t$ - $J$  model on a fully-connected cluster  
with all-to-all and random  $t_{ij}$  and  $J_{ij}$



Numerical solution of  $t$ - $J$  model on a fully-connected cluster  
with all-to-all and random  $t_{ij}$  and  $J_{ij}$

Planckian  
metal with  
SYK dynamics  
of  
fractionalized  
spinon  
excitations

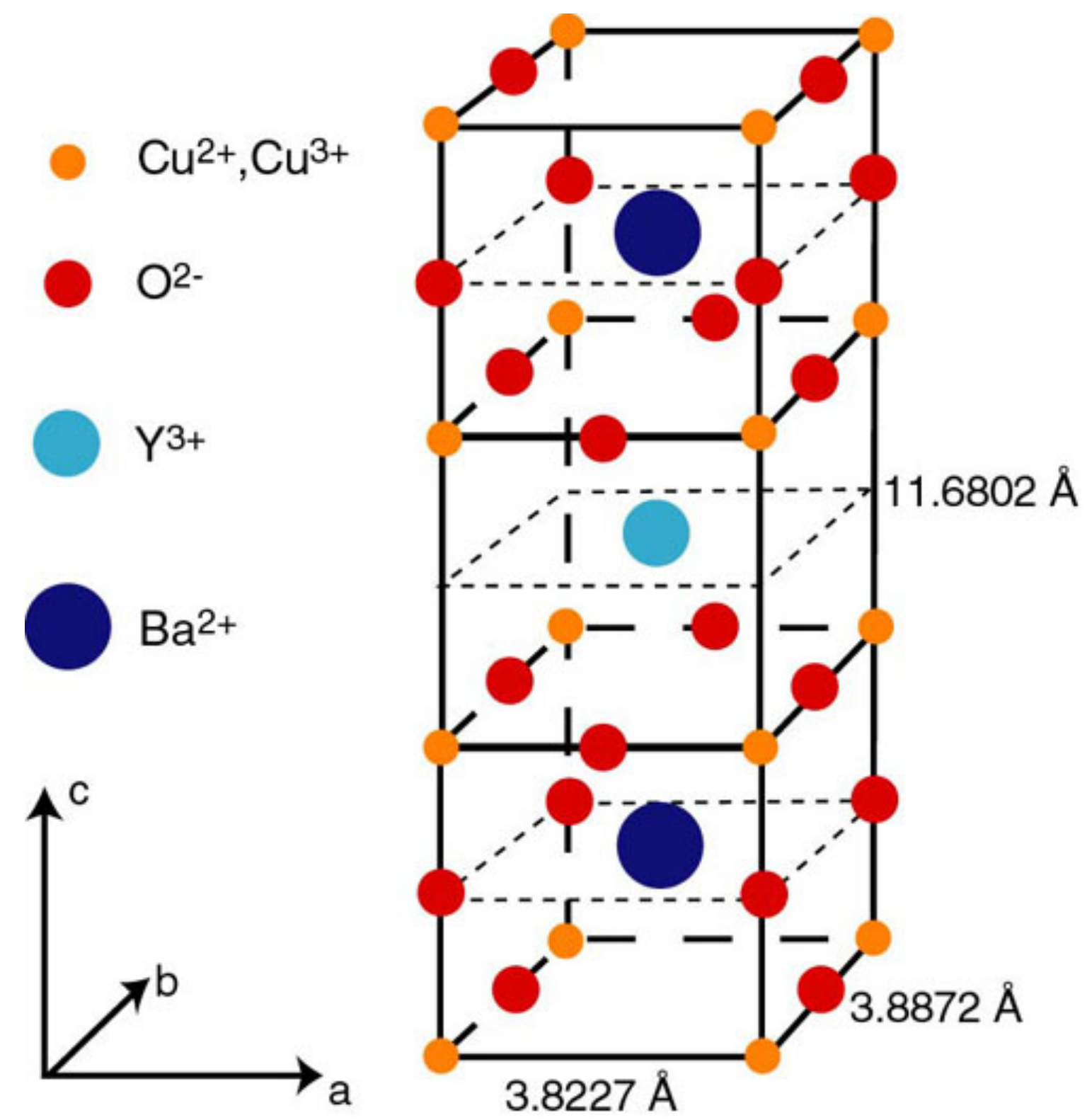
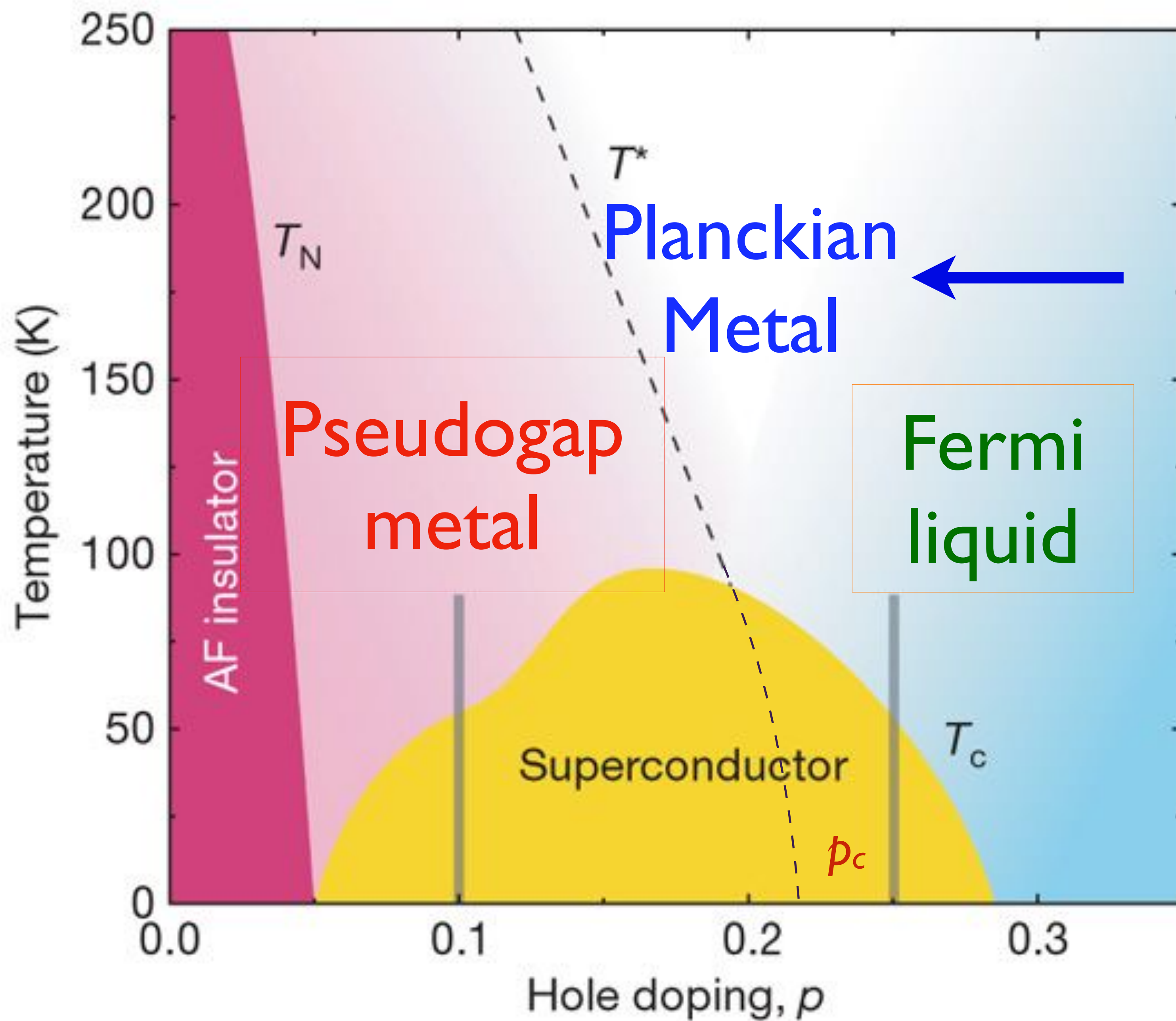




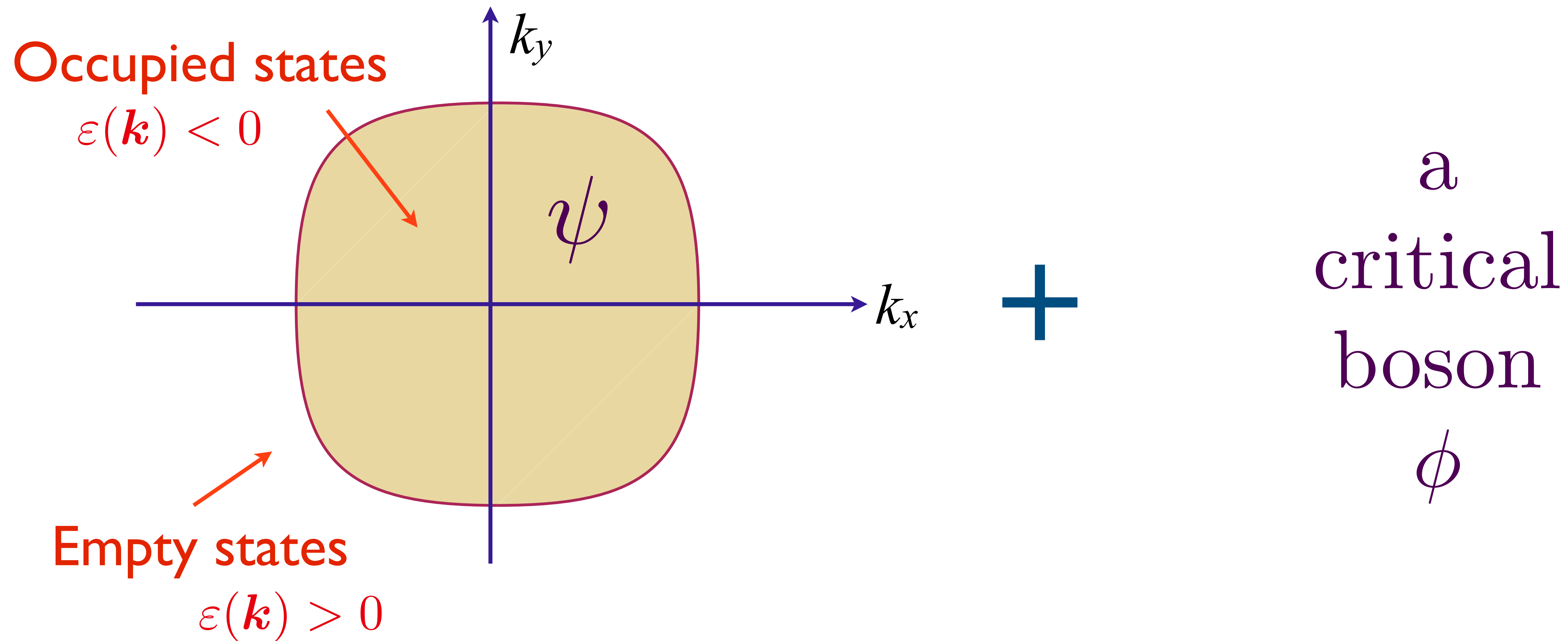
1. Introduction to Planckian metals
2. Introduction to black holes
3. The SYK model
4. Progress on the theory of black holes
5. Progress on the theory of Planckian metals

*A. Random  $t$ - $J$  model*

*B. Fermi surface coupled to a critical boson*



# Fermi surface coupled to a critical boson



## Fermi surface coupled to a critical boson

“Yukawa” coupling:  $g \int d^2r d\tau \psi^\dagger(r, \tau) \psi(r, \tau) \phi(r, \tau)$

Yields a state without quasiparticle excitations, but the theory is not systematic at large  $N$

Sung-Sik Lee (2009)

# Fermi surface coupled to a critical boson

“Yukawa” coupling:  $\frac{g_{ijl}}{N} \int d^2r d\tau \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$

$$\overline{g_{ijl}} = 0 \quad , \quad \overline{|g_{ijl}|^2} = g^2$$

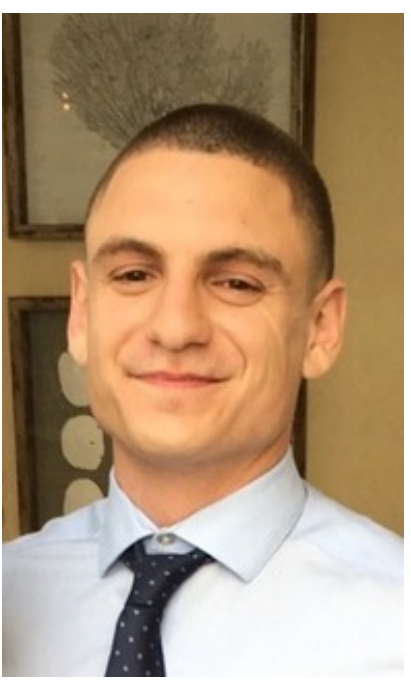
# Fermi surface coupled to a critical boson

“Yukawa” coupling: 
$$\frac{g_{ijl}}{N} \int d^2r d\tau \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$$
$$\overline{g_{ijl}} = 0 \quad , \quad \overline{|g_{ijl}|^2} = g^2$$

- Yields a systematic large  $N$  theory of a “non-Fermi liquid”: a compressible state without quasiparticle excitations. The Fermi surface is sharp in momentum space, but the spectral functions are diffuse in frequency space.
- There is Planckian dynamics at the Fermi surface:

$$G(k = k_F, \omega, T) \sim \omega^{-2/3} F(\hbar\omega/k_B T)$$

- There is many-body quantum chaos in the out-of-time-order correlator (OTOC) with maximal Lyapunov exponent  $\lambda_L = 2\pi k_B T/\hbar$ .



# Fermi surface coupled to a critical boson

“Yukawa” coupling: 
$$\frac{g_{ijl}}{N} \int d^2r d\tau \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$$

$$\overline{g_{ijl}} = 0 \quad , \quad \overline{|g_{ijl}|^2} = g^2$$

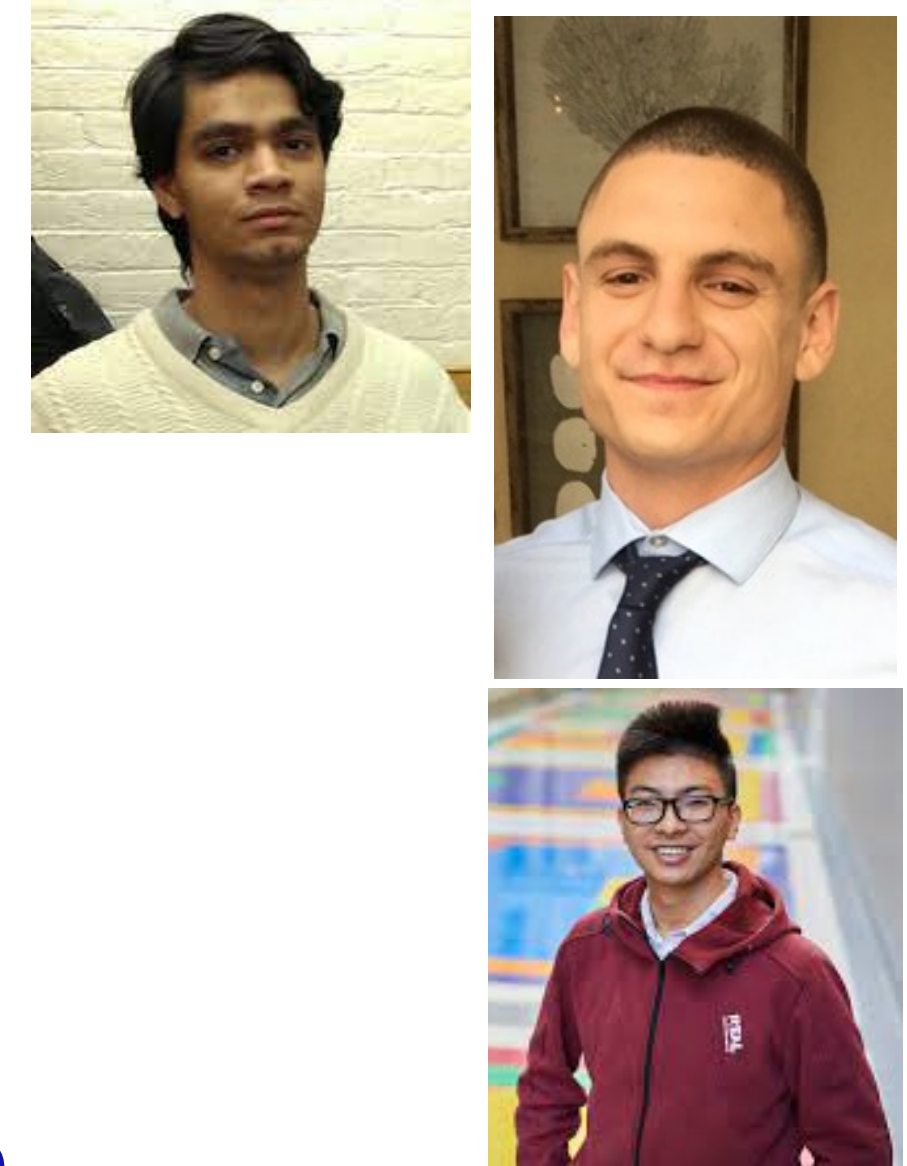
Conservation of momentum implies the d.c. conductivity is infinite

A. Eberlein, I. Mandal, and S. S., PRB **94**, 045133 (2016)

$$\text{Re } \sigma(\omega) = D\delta(\omega) + \text{Re } \sigma_{\text{reg}}(\omega)$$

$$\text{Re } \sigma_{\text{reg}}(\omega, T = 0) \sim \frac{1}{\omega^{2/3}}$$

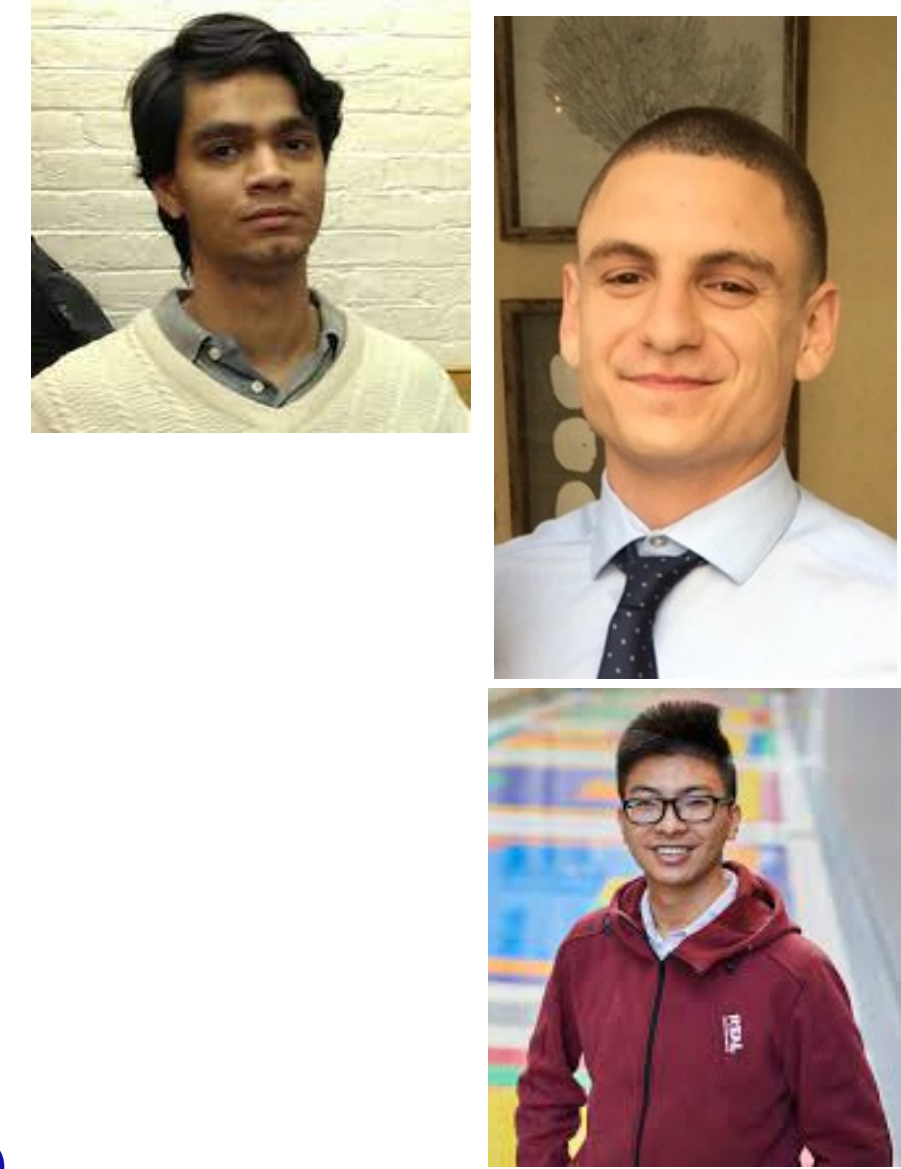
Yong Baek Kim, A. Furusaki, Xiao-Gang Wen, P. A. Lee, PRB **50**, 17917 (1994)



# Fermi surface coupled to a critical boson

“Yukawa” coupling: 
$$\frac{g_{ijl}}{N} \int d^2r d\tau \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$$

$$\overline{g_{ijl}} = 0 \quad , \quad \overline{|g_{ijl}|^2} = g^2$$



Conservation of momentum implies the d.c. conductivity is infinite

A. Eberlein, I. Mandal, and S. S., PRB **94**, 045133 (2016)

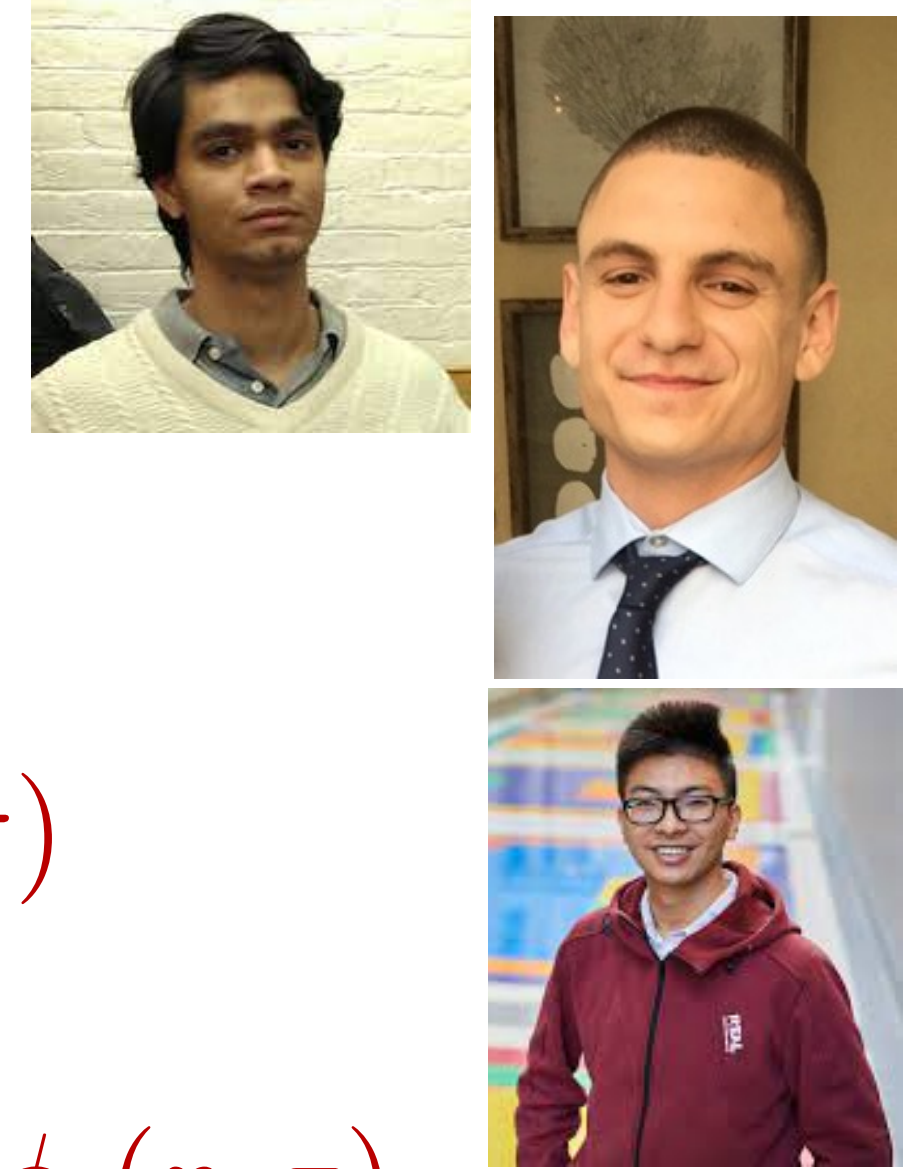
$$\text{Re } \sigma(\omega) = D\delta(\omega) + \text{Re } \sigma_{\text{reg}}(\omega)$$
$$\text{Re } \sigma_{\text{reg}}(\omega, T = 0) \sim \frac{1}{\omega^{2/3}}$$

Yong Baek Kim, A. Furusaki, Xiao-Gang Wen, P. A. Lee, PRB **50**, 17917 (1994)

Have to include the effects of spatial disorder or umklapp

Ilya Esterlis, Haoyu Guo, Aavishkar Patel, S.S. to appear

# Fermi surface coupled to a critical boson



“Yukawa” coupling:  $\frac{g_{ijl}}{N} \int d^2r d\tau \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$

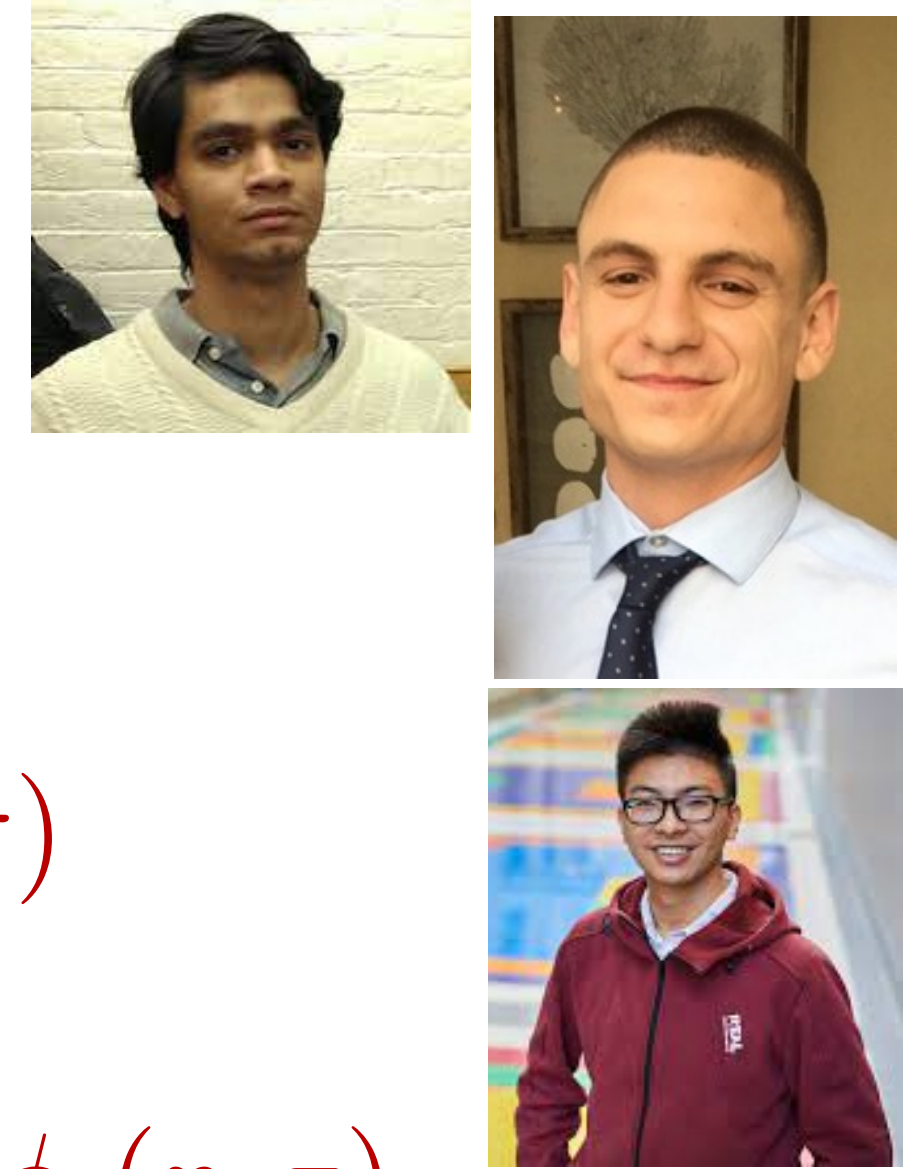
Random potential:  $+\frac{1}{\sqrt{N}} \int d^2r d\tau v_{ij}(r) \psi_i^\dagger(r, \tau) \psi_j(r, \tau)$

Random interactions:  $+\frac{1}{N} \int d^2r d\tau g'_{ijl}(r) \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$

$$\overline{g_{ijl}} = 0 \quad , \quad \overline{g_{ijl}^* g_{abc}} = g^2 \delta_{ia} \delta_{jb} \delta_{lc} \quad , \quad \overline{v_{ij}(r)} = 0 \quad , \quad \overline{v_{ij}^*(r) v_{lm}(r')} = v^2 \delta(r - r') \delta_{il} \delta_{jm}$$

$$\overline{g'_{ijl}(r)} = 0 \quad , \quad \overline{g'_{ijl}^*(r) g'_{abc}(r')} = g'^2 \delta(r - r') \delta_{ia} \delta_{jb} \delta_{lc}$$

# Fermi surface coupled to a critical boson



“Yukawa” coupling:  $\frac{g_{ijl}}{N} \int d^2r d\tau \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$

Random potential:  $+\frac{1}{\sqrt{N}} \int d^2r d\tau v_{ij}(r) \psi_i^\dagger(r, \tau) \psi_j(r, \tau)$

Random interactions:  $+\frac{1}{N} \int d^2r d\tau g'_{ijl}(r) \psi_i^\dagger(r, \tau) \psi_j(r, \tau) \phi_l(r, \tau)$

$$\overline{g_{ijl}} = 0 \quad , \quad \overline{g_{ijl}^* g_{abc}} = g^2 \delta_{ia} \delta_{jb} \delta_{lc} \quad , \quad \overline{v_{ij}(r)} = 0 \quad , \quad \overline{v_{ij}^*(r) v_{lm}(r')} = v^2 \delta(r - r') \delta_{il} \delta_{jm}$$

$$\overline{g'_{ijl}(r)} = 0 \quad , \quad \overline{g'_{ijl}^*(r) g'_{abc}(r')} = g'^2 \delta(r - r') \delta_{ia} \delta_{jb} \delta_{lc}$$

With  $g, v, g'$  all non-zero, resistivity  $\rho(T) = \rho(0) + \tilde{c}T \dots$   
 $\rho(0)$  is determined by  $v$ ,  
 while  $\tilde{c}$  is determined by a subleading random interaction  $g'$ .  
 These features are just as in the random  $t$ - $J$  model.

# Summary

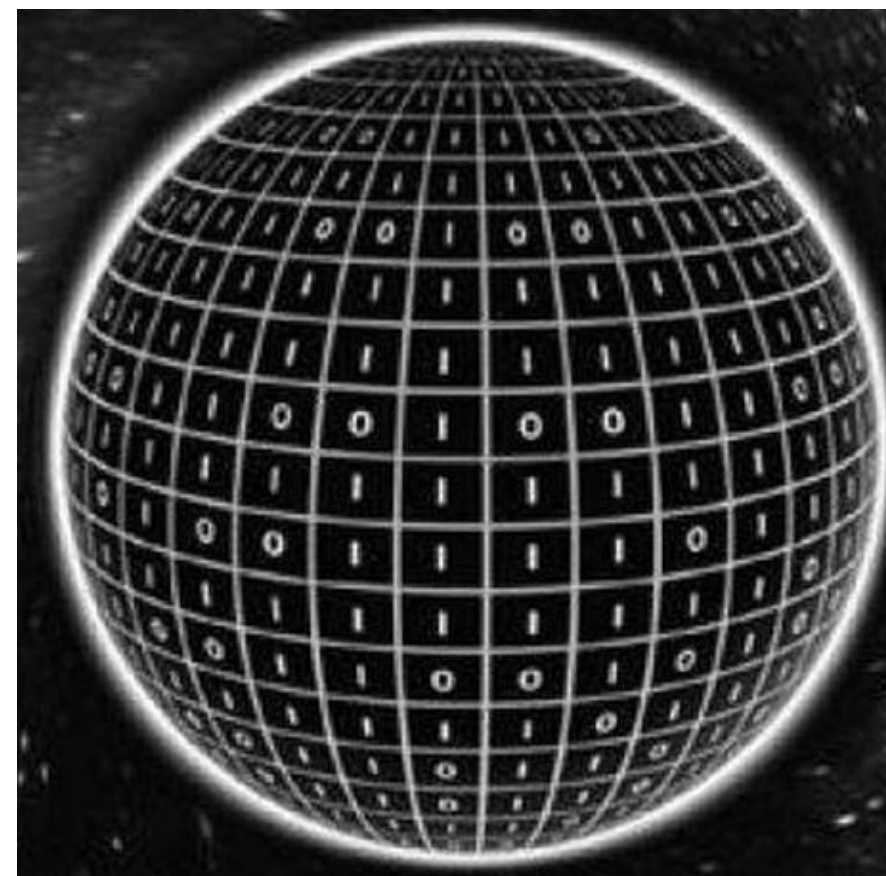
- SYK: a solvable model without quasiparticle excitations, with Planckian time dynamics, and maximal chaos with Lyapunov exponent  $2\pi k_B T / \hbar$ .

# Summary

- SYK: a solvable model without quasiparticle excitations, with Planckian time dynamics, and maximal chaos with Lyapunov exponent  $2\pi k_B T / \hbar$ .
- Low energy theory of time reparameterizations is the theory of the boundary graviton in 2D quantum gravity on  $\text{AdS}_2$ .

# Summary

- Boundary graviton leads to universal  $-3/2 \ln(1/T)$  correction to Bekenstein-Hawking entropy of low  $T$  charged black holes in Einstein gravity, and to the SYK model. So the semiclassical entropy of Einstein gravity is reproduced by a unitary quantum system with a discrete spectrum. Further work along these lines has led to progress on the Page curve describing the time evolution of the entropy of an evaporating black hole.



# Summary

- Random  $t$ - $J$  model captures many aspects of the cuprates over a wide intermediate temperature range, including the Planckian metal behavior. The linear- $T$  resistivity arises from the first subleading operator of random interactions.

# Summary

- Random  $t$ - $J$  model captures many aspects of the cuprates over a wide intermediate temperature range, including the Planckian metal behavior. The linear- $T$  resistivity arises from the first subleading operator of random interactions.
- Two-dimensional Fermi surface coupled to a critical boson has no quasiparticle excitations, and exhibits Planckian time dynamics and maximal chaos with Lyapunov exponent  $2\pi k_B T/\hbar$ . The linear- $T$  resistivity arises from the first subleading operator of random interactions.

# Summary

- Random  $t$ - $J$  model captures many aspects of the cuprates over a wide intermediate temperature range, including the Planckian metal behavior. The linear- $T$  resistivity arises from the first subleading operator of random interactions.
- Two-dimensional Fermi surface coupled to a critical boson has no quasiparticle excitations, and exhibits Planckian time dynamics and maximal chaos with Lyapunov exponent  $2\pi k_B T/\hbar$ . The linear- $T$  resistivity arises from the first subleading operator of random interactions.